Coexisting serpentine and quartz from carbonate-bearing serpentinized peridotite in the Samail Ophiolite, Oman

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Abstract

Tectonically exposed mantle peridotite in the Oman Ophiolite is variably serpentinized and carbonated. Networks of young carbonate veins are prevalent in highly serpentinized peridotite, particularly near low-temperature alkaline springs emanating from the peridotite. An unusual feature in some samples is the coexistence of serpentine and quartz, which is not commonly observed in serpentinites. This assemblage is unstable with respect to serpentine + talc or talc + quartz under most conditions. Serpentine in the carbonated serpentinites in this study is more iron-rich than in most serpentinites reported in previous studies, and samples with co-existing quartz contain the most iron-rich serpentines. Calculations of thermodynamic equilibria in the MgO-SiO₂-H₂O-CO₂ system suggest that serpentine + quartz may be a stable assemblage at low temperatures (e.g., less than ~15-50°C), and is stabilized to higher temperatures by preferential cation substitutions in serpentine over talc. Based on these calculations, serpentine + quartz assemblages could result from serpentinization at near-surface temperatures. Clumped isotope thermometry of carbonate veins yields temperatures within error of the observed temperatures in Oman groundwater for all samples analyzed, while the $\delta^{18}O$ of water calculated to be in equilibrium with carbonate precipitated at those temperatures is within error of the observed isotopic composition of Oman groundwater for the majority of samples analyzed. As groundwater geochemistry suggests that carbonate precipitation and serpentinization occur concomitantly, this indicates that both hydration and carbonation of peridotite are able to produce extensive alteration at the relatively low temperatures of the near-surface weathering environment.

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1. Introduction

Mantle peridotite is far from thermodynamic equilibrium with the atmosphere and surface waters. Peridotite exposed at Earth's surface reacts readily with aqueous fluids to form products including serpentine and carbonate minerals. Carbonation and serpentinization of exposed mantle peridotite play an important role in numerous natural processes and may also provide effective means of permanent carbon storage (e.g., Seifritz 1990). Natural CO₂ uptake via carbonation of peridotite on the seafloor is a significant carbon sink and an important element of the carbon cycle (Früh-Green et al. 2004; Kelemen et al. 2011). It has even been proposed that the emergence of large areas of ultramafic seafloor in the New Caledonia region was responsible for rapid CO₂ draw-down leading to profound climate change (Reusch 2011).

The fate of carbon in these environments is closely linked with hydration of peridotite. For example, serpentinization produces very alkaline and reduced fluids, which affect other geochemical reactions in those environments (e.g., Barnes and O'Neil 1969; Bruni et al. 2002). H₂-rich fluids resulting from serpentinization may provide chemical energy for microbial communities and were perhaps vital in the emergence of early life (Sleep et al. 2004; Schulte et al. 2006; McCollom and Bach 2009), as well as acting as a reductant for abiotic synthesis of hydrocarbons in the presence of FeNi metal and reduced carbon species (Sleep et al. 2004; McCollom et al. 2010). Serpentinization and carbonation of peridotite are exothermic, and heat released by serpentinization of seafloor peridotite may contribute to hydrothermal circulation (Schuiling 1964; Fyfe 1974; Kelley et al. 2001; Emmanuel and Berkowitz 2006). Exothermic self-heating could also lead to increased reaction rates during peridotite carbonation (Kelemen and Matter 2008). In situ carbonation of peridotite has the potential to store large amounts of CO₂ as stable carbonate minerals (Kelemen and Matter 2008). The efficiency of peridotite carbonation depends on temperature, pressure, water composition, and reactions involving the magnesium silicates that are dissolved and precipitated during alteration of peridotite (e.g., O'Connor et al. 2004; Hansen et al. 2005; Chizmeshya et al. 2007). Thus, understanding the circumstances of hydration and carbonation of peridotite during natural alteration can constrain the optimal conditions and limitations of mineral carbon sequestration in peridotite.

High degrees of carbonation and serpentinization are observed in carbonate-bearing serpentinized peridotite in the Samail Ophiolite of Oman. The extent of serpentinization in the tectonically exposed mantle section of ophiolites is typically 30-60%, reaching 100% in some cases (Boudier et al. 2010). In peridotite that hosts abundant carbonate veins in the Samail Ophiolite, we commonly observe 100% serpentinization, and relict olivine is rare. Most of the sub-surface carbonate veins now exposed by erosion are relatively young, with 95% of 40 samples yielding ¹⁴C ages ranging from 0 to 50,000 years (Clark and Fontes 1990; Kelemen and Matter 2008; Kelemen et al. 2010). Low-temperature alkaline springs emanating from the peridotite, as well as the spectacular travertine terraces that precipitate around these springs, are also linked to ongoing carbonation and serpentinization. Ca-OH-rich, carbon-poor waters are formed as groundwater reacts with peridotite isolated from the atmosphere and precipitate calcite and dolomite as they emerge at the surface to react with atmospheric CO2 or mix with Mg-HCO3 waters (Barnes and O'Neil 1969; Neal and Stanger 1985; Bruni et al. 2002; Kelemen and Matter 2008). Spring water temperatures are ~30°C, similar to the mean annual air temperature in Oman. Kelemen et al. (2011) report d¹⁸O values in peridotite-hosted carbonate veins and travertines that are consistent with equilibrium with Oman groundwater and spring-water at near-surface temperatures. Together, carbonate ages, geochemical trends in groundwater, and preliminary stable isotope thermometry suggest that lowtemperature serpentinization and carbonation in these rocks has been ongoing for more than 50,000 years and continues to the present day.

The products of ongoing alteration of peridotite at near-surface temperatures in Oman are primarily serpentine + magnesite, dolomite, or calcite, with calcite more prevalent near alkaline springs and associated travertine deposits. Talc is very rare in our carbonate-veined, altered peridotite samples, found only as scattered grains in one sample (OM07-28A). However, we observe the surprising occurrence of abundant quartz + serpentine without talc in two of our samples (OM08-206A and OM08-206D). Here we discuss the mineral parageneses and hermometry in low temperature carbonated and serpentinized peridotites from the Samail Ophiolite in Oman and consider the implications for conditions of alteration.

2. Analytical Methods

Samples were studied by petrographic microscope, powder X-ray diffraction (XRD), X-ray fluorescence (XRF), and electron microprobe. Powdered samples were analyzed by X-ray diffraction using a Philips Model 1830 X-ray diffractometer located at Woods Hole Oceanographic Institution and peaks for constituent minerals were identified using MacDiff and Match! Software. Whole rock major and trace element composition were determined by X-ray fluorescence at the Washington State University GeoAnalytical Lab, where loss on ignition (LOI), primarily H₂O + CO₂, was also measured (http://www.sees.wsu.edu/Geolab/equipment/xrf.html). Polished thin sections were quantitatively analyzed with a 5-spectrometer Cameca SX-100 electron microprobe located at the American Museum of Natural History in New York, using a 10µm beam diameter, with 15kV accelerating voltage, 10nA current, and 20-30 second peak time, to analyze the major element composition of phases in thin section in terms of Si, Ti, Al, Cr, Fe, Mn, Mg, Ca, and Ni, using natural and synthetic standards. To correct for matrix effects in serpentine. Si and Mg concentrations in serpentine analyses were corrected using a calibration based on analyses of reference antigorite standard, BM66586. The reported analytical precision is based on replicate analyses of this reference material. Calibration information for each measured oxide, including spectrometer crystal, calibration standard, peak time, wt% in antigorite sample, standard deviation, and post-measurement corrections, are shown in the supplementary table in Online Resource 1.

Clumped isotope and conventional δ¹⁸O carbonate-water exchange thermometry were employed to determine the temperature of carbonate precipitation during alteration of peridotite. Clumped isotope thermometry is a relatively new carbonate paleothermometer based on temperature-dependent "clumping" of ¹³C and ¹⁸O relative to a stochastic distribution in CO₂ released by acid digestion of carbonate, and does not depend on the bulk isotopic composition of the fluid from which the carbonate minerals precipitated (Ghosh et al. 2006; Eiler 2007; Huntington et al. 2009). CO₂ was extracted from carbonate minerals in our samples by phosphoric acid digestion, with samples containing only calcite veins reacting at 25°C for 12-16 hours in sealed vessels (OM08-01 and OM09-130) or at 90°C for 20 minutes with evolved CO₂ continuously trapped cryogenically (samples OM08-206D and OM10-34B), samples containing only dolomite veins reacting at 90°C for 20 minutes in sealed vessels, and samples containing magnesite veins

reacting at 80°C for 2-3 days in sealed vessels. Samples containing more than one carbonate mineral were subjected to a stepped phosphoric acid digestion procedure following the method of Guo (2008). The extracted gas was purified by cryogenic separation and gas chromatography, and analyzed by gas-source isotope ratio mass spectrometry using a modified Thermo-Finnegan 253 in the laboratories for stable isotope geochemistry at Caltech to simultaneously make measurements of clumped isotope ratios, reported as Δ 47 values (Ghosh et al. 2006), δ ¹³C, and δ ¹⁸O.

Clumped isotope analytical methods and data processing were carried out according to the protocols described in Huntington et al. (2009). Δ47 values were normalized based on analysis of heated CO₂ gases with a stochastic distribution of isotopologues, and were corrected for fractionation during phosphoric acid digestion at temperatures above that of the original calibration (25°C) (Ghosh et al. 2006) via linear interpolation between that and the experimentally-determined correction factor of 0.081‰ at 90°C (Passey et al. 2010). This agrees relatively well with the theoretically predicted correction of 0.0078‰ for phosphoric digestion of calcite at 90°C (Guo et al. 2009). Although Guo et al. (2009) predict different (slightly non-linear) hot acid fractionation corrections for calcite, dolomite, and magnesite, the difference between those corrections and the experimental value does not result in significantly different clumped isotope temperatures (~3°C higher for magnesite digested at 90°C, ~2°C higher for dolomite digested at 90°C, and ~1°C lower for calcite and dolomite digested at 50°C for low-temperature carbonates).

Clumped isotope temperatures were calculated using a polynomial fit to the experimental calcite calibration data between 1 and 50°C and at 1100°C from Ghosh et al. (2006) and experimental dolomite calibration data between 25 and 350°C (Bonifacie et al. 2011; Bristow et al. 2011; M. Bonifacie, manuscript in preparation). The synthetic dolomite data are consistent with inorganic calcite calibrations as well as the theoretical dolomite calibration of Guo et al. (2009). Although experimental calibration data are not available for magnesite, Guo et al. (2009) calculate a temperature dependence of Δ 47 in magnesite that is indistinguishable from that in dolomite. Varying the choice of equation used for Δ 47 temperature-dependence (single polynomial fit to all available calibration data (Bristow et al. 2011), separate fits to calibration data for calcite and dolomite (Ghosh et al. 2006; M. Bonifacie et al., manuscript in preparation),

or separate theoretical calibrations for each mineral (Guo et al. 2009)) results in less than 4°C variation in the calculated temperatures of carbonate precipitation at low temperatures, which is similar in magnitude to the analytical error. As such, a single polynomial fit is used for all three minerals.

Oxygen isotope ratios of water in equilibrium with the carbonates at these clumped isotope temperatures were calculated using the calcite-water fractionation factors from Kim and O'Neil(1997), dolomite-water fractionation factors from Vasconcelos et al. (2005), and magnesite-water fractionation factors calculated ratioing the reduced partition function ratios for magnesite and water from, Schauble et al. (2006) and Rosenbaum (1997), respectively. Since experimental data are not available for magnesite-water oxygen isotope fractionation, and different theoretical calculations of fractionation factors can yield significantly different results (Schauble et al. 2006; Chacko and Deines 2008; Zheng 2011), less confidence can be given to the $\delta^{18}O_{water}$ values derived from magnesite.

3. Results

3.1 Mineralogy

Most samples have mesh textures typical of serpentine replacing olivine (Wicks and Whittaker 1977; O'Hanley 1996). Rare olivine relicts are present in only a few carbonate-veined serpentinite samples, with sparse pyroxene the most common relict primary silicate phase. X-ray diffraction (XRD) analysis was employed to identify the major minerals in most of the samples. XRD, combined with optical petrography, reveals that serpentine is composed of lizardite and chrysotile, and the vein carbonates are primarily dolomite and magnesite. However, in some veins near active alkaline springs, calcite is the most abundant carbonate mineral. Minerals other than serpentine and carbonate commonly reported in carbonate-bearing serpentinized peridotites during low-temperature alteration and weathering of peridotite, such as brucite, talc, and hydrated carbonates including hydromagnesite, nesquehonite, dypingite, and huntite (e.g., Stanger and Neal 1994; Wilson et al. 2006; Klein et al. 2009; Bach and Klein 2009)), are not present in sufficient quantities (more than 5 to 10%) in our samples to be identified in XRD, nor were they found by electron microprobe in most samples. Our samples are chemically heterogeneous, as magnesium and calcium are

mobilized in the fluid during serpentinization and precipitated elsewhere in millimeter-to-decimeter-scale carbonate veins. Bulk-rock chemistry and mineralogy is presented in Table 1.

The fine-grained nature of most of our samples limits the usefulness of the petrographic microscope for identifying phases and observing textures. Electron microprobe analysis, including both point measurements and elemental mapping, and back-scattered electron (BSE) imaging are therefore valuable for distinguishing different types of carbonates in situ, observing fine-scale textural relationships between minerals, identifying accessory minerals, and determining the composition of the minerals in our samples.

Accessory minerals in carbonate-veined serpentinites identified using the electron microprobe include chromian spinel, magnetite, hematite, goethite, olivine, diopside, enstatite, talc, quartz, ankerite, barite, gypsum, pyrite, and silver sulfide, as summarized in Table 1. Hydrated carbonates and brucite can be distinguished from their anhydrous counterparts (including magnesite) using the microprobe totals. Hydrated carbonate phases are not observed in any of our samples, and brucite is found at only a single point, in keeping with the likelihood of brucite breakdown to form magnesium carbonates via reactions such as $Mg(OH)_2 + CO_2 = MgCO_3 + H_2O$. Minor phases show signs of disequilibrium. For example, in isolated, small veins and vugs up to ~50 µm in size, some apparently late-forming calcite is found in contact with magnesite, which under equilibrium conditions should lead to the formation of dolomite. Many studies have emphasized the role of magnetite in serpentinization (e.g., Hansen et al. 2005; Klein et al. 2009), since serpentines generally accommodate less iron than their precursor silicate minerals (Trommsdorff and Evans 1972; Vance and Dungan 1977), making it possible in some cases to use magnetic susceptibility as an indicator of the extent of serpentinization (Toft et al. 1990; Hansen et al. 2005; Bach et al. 2006). In some of our samples, scattered magnetite grains are found with serpentine, but in most samples magnetite occurs only as rims on relict chromian spinels, if at all, and hematite is more common than magnetite, while other samples have essentially no iron oxides. Our data are consistent with the observation that hematite, rather than magnetite, is common in carbonate-altered peridotites worldwide (Eckstrand 1975; Frost 1985).

The presence of minerals such as hematite and sulfates in many of our samples suggests that oxygen fugacity was relatively high, at least locally, despite the reducing nature of serpentinization reactions that commonly reduce H₂O while oxidizing Fe²⁺ from olivine (Frost 1985). In some samples hematite and magnetite are present in the same thin section, implying buffering along redox reactions, localized heterogeneities in oxygen fugacity, or disequilibrium.

A surprising assemblage observed in a few samples is coexisting serpentine + quartz. In two samples, OM08-206A and OM08-206D, quartz makes up a large enough proportion of the rock to be identifiable by XRD (Figure 1). Small amounts of a pure silica phase were also detected by electron microprobe in a third serpentinite sample, OM08-01. In OM08-206A and OM08-206D, microcrystalline quartz is found in the cores of mesh-texture serpentine in 5-10 mm wide, formerly pyroxene-rich bands that still contain relatively abundant relict pyroxene. ~30 volume % of these bands is composed of quartz. The bands are also evident in outcrop morphology and comprise ~10 volume % of the outcrop.

Although the silica appears isotropic in thin section, XRD spectra of OM08-206A and OM08-206D show the silica phase to be quartz, and not opal or amorphous silica. Unlike some silica reported in weathered and silicified serpentinites elsewhere (O'Hanley 1996; Tsikouras et al. 2006; Boschi et al. 2009), the texture of coexisting serpentine and quartz in our samples is similar to serpentine mesh texture. Boschi et al. (2009) report chalcedony rimming carbonates, which in turn surround serpentine relicts, without direct contact between chalcedony and serpentine. Tsikouras et al. (2006) report microcrystalline quartz surrounding relict serpentine. In our quartz-bearing serpentinite samples, serpentine forms a meshwork that surrounds cores of quartz (Figure 2), whereas in a typical mesh serpentine the mesh centers are composed of relict olivine or late-forming serpentine that statically replaces olivine (O'Hanley 1996). Beinlich et al. (2010) report a similar occurrence of quartz filling mesh cells defined by serpentine veins, formed during low temperature, near-surface weathering of ultramafic clasts in a sedimentary basin.

Carbonate-free, partially serpentinized peridotite found at the same location as our quartz-bearing serpentinite has a similar mesh texture with serpentine surrounding olivine, as well as surrounding epoxy-filled holes in the thin section. However, the serpentine meshwork surrounding quartz commonly includes a layer of more iron-rich serpentine rims surrounding the quartz cores, whereas the serpentine mesh

surrounding olivine or epoxy-filled holes does not. The cryptocrystalline quartz is commonly very finely inter-grown with sub-micron-sized iron compounds, which results in up to 25 wt% FeO* in some mixed analyses. Aside from this variable iron component, microprobe analyses show that quartz regions are nearly pure SiO₂ and cannot be confused with the soda lime glass slides on which the thin sections are mounted (Table 2). Although serpentine + quartz is unstable under typical metamorphic conditions with respect to serpentine + talc or talc + quartz (Evans and Guggenheim 1988), the low temperatures of ongoing serpentinization in Oman could allow this assemblage to coexist stably, as discussed below.

3.2 Iron in serpentine

Average serpentine compositions for each sample are reported in Table 3, and a complete table of all electron microprobe analyses of serpentine can be found in Online Resource 2. Serpentine in our carbonated serpentinite samples from Oman is, on average, slightly more iron-rich than lizardite from serpentinized mantle peridotites compiled from the literature by Evans et al. (2009), and the serpentine in our quartz-bearing samples is significantly more iron-rich, as shown in Figure 3. Lizardite analyses from the literature have an average $X_{Fe} = \text{molar Fe/[molar Fe + Mg] of 5.7\%}$, serpentines in our quartz-free serpentinite samples have an average X_{Fe} of 6.3%, and serpentines in our samples containing coexisting serpentine + quartz have an average X_{Fe} of 10.1%. The iron content of much of the serpentine in our samples, particularly serpentine coexisting with quartz, exceeds the iron content of the relict olivine ($X_{Fe} \sim 9.0\%$) and pyroxene ($X_{Fe} \sim 8.1\%$) found in these samples and elsewhere in the ophiolite (Hanghoj et al. 2010). Variation in serpentine composition across the boundaries between serpentine and quartz is evident in WDS elemental maps of Fe (Figure 4), with serpentine forming Fe-rich rims around quartz containing fine-grained Fe-compounds.

Projection of our serpentine microprobe analyses onto a molar MgO-SiO₂-FeO ternary diagram (Figure 5), reveals that, in addition to higher than normal Fe-content, several of our serpentine analyses are also anomalously silica-rich. This may be due to the presence of talc inter-layers within the serpentine structure (Veblen and Buseck 1979). Fe substitution and/or the presence of talc interlayers may stabilize serpentine relative to talc, and may help explain why end-member talc is so rarely observed in our samples.

Many studies model iron substitution in serpentine as if all iron in serpentine is Fe(II) on M sites to produce $Fe_3^{2+}Si_2O_5(OH)_4$ (e.g., Frost and Beard 2007; McCollom and Bach 2009), but significant Fe(III) substitution is also possible (Wicks and Plant 1979; O'Hanley and Dyar 1993; Klein et al. 2009) and perhaps is dominant in many lizardites (Evans 2008, 2010; Evans et al. 2009). Evans (2008) emphasized the importance of ferric iron substitution in serpentine, primarily in the form of a ferri-Tschermak substitution, as proposed by Wicks and Plant (1979), where Fe^{3+} substitutes on both M sites and T sites to form a cronstedtite component of serpentine, producing an end-member serpentine with the stoichiometry $(Mg_2Fe^{3+})(Fe^{3+}Si)O_5(OH)_4$. In some cases ferric iron may substitute only on the M site, charge balanced by M site vacancies, yielding an end-member ferrian serpentine with the formula $(\Box Fe_2^{3+})Si_2O_5(OH)_4$ (Wicks and Plant 1979; Evans 2008; Evans et al. 2009). Comparison of the compositions of serpentine from our samples with these end-members (Figure 5), suggests that all three substitution mechanisms may be present in our serpentines.

Regardless of the particular iron substitution mechanism in serpentine, the reaction of magnetite with water and aqueous silica to produce an Fe component of serpentine will be favored at lower temperatures because the reaction involves a net consumption of volatiles (given nearly constant pressure in the near-surface alteration environment). For example,

(R1)
$$Fe^{2+}Fe_2^{3+}O_4 + 2 SiO_{2 aq} + H_2O + H_{2 aq} = Fe_3^{2+}Si_2O_5(OH)_4$$

magnetite ferroan lizardite

(R2)
$$2 \text{ Fe}^{2+} \text{Fe}_2^{3+} \text{O}_4 + 6 \text{ SiO}_{2 \text{ aq}} + 7 \text{ H}_2 \text{O} = 3 \left(\Box \text{Fe}_2^{3+} \right) \text{Si}_2 \text{O}_5 (\text{OH})_4 + \text{H}_{2 \text{ aq}}$$
 magnetite vacancy-coupled Fe³⁺lizardite

(R3)
$$4 \operatorname{Fe^{2+}Fe_2^{3+}O_4} + 3 \operatorname{SiO_{2 \, aq}} + 5 \operatorname{H_2O} + \operatorname{H_{2 \, aq}} = 3 (\operatorname{Fe_2^{2+}Fe^{3+}}) (\operatorname{Fe^{3+}Si}) \operatorname{O_5(OH)_4}$$
 magnetite cronstedtite

The source of aqueous SiO_2 in these reactions can be understood as the result of incongruent carbonation of silicates, e.g., olivine, serpentine, pyroxene, or talc + CO_2 = magnesite + SiO_{2aq} . The observation that the formation of iron-rich serpentine should be favored at lower temperatures is borne out in geochemical reaction path models by Klein et al. (2009), which calculate decreasing magnetite and increasing iron in serpentine with decreasing temperature.

The samples containing abundant quartz + serpentine also have the lowest bulk rock Mg#'s of our samples (87.2 for OM08-206A and 82.7 for OM08-206D), slightly lower than in typical Oman mantle peridotites, while relict clinopyroxenes in these samples have typical mantle Mg#'s of 91-93. The low bulk rock Mg# is probably due to small amounts of Mg-extraction during alteration (e.g., Snow and Dick 1995). Even so, the average Fe-content of serpentine in these samples ($X_{Fe} \sim 11\%$) is still higher than would be predicted ($X_{Fe} \sim 6\%$) for antigorite serpentine in Fe/Mg exchange equilibrium with Mg# 82-87 olivine (Trommsdorff and Evans 1972). We infer that the high Fe content of serpentine in these samples is due mainly to low temperature reactions consuming Fe-oxides (e.g., reactions R1-R3), and not to the slightly Fe-rich bulk compositions.

3.3 Stability of coexisting serpentine + quartz at low temperatures

T-X diagrams for end-member reactions in the system MgO-SiO₂-H₂O-CO₂ were calculated at constant pressure using THERMOCALC (Holland and Powell 1998). Thermodynamic data are not available for lizardite, and chrysotile and lizardite have very similar free energy (Evans 2004), so the stability of both phases is modeled using thermodynamic data for chrysotile. Calculations were done under fluid saturated conditions to produce T-X diagrams for Mg-end-member magnesite (mgs), chrysotile (chr), talc (ta), and quartz (q), at a variety of near-surface pressures and mineral activities.

For pure Mg-end-members (unit activities of Mg-components in all minerals), the isobaric invariant point involving magnesite, chrysotile, talc, quartz, and fluid is predicted to be at 15°C, which is slightly lower than observed ground- and surface-water temperatures of 17-40°C in Oman (Neal and Stanger 1985; Matter 2001, 2005; Dewandel et al. 2005; our unpublished data). Increasing pressure does not have a significant effect on the temperature of the invariant point, but shifts it to lower values of X_{CO2} , from $log(X_{CO2})$ of -5.06 at 3.5 bars to -6.6 at 1kbar. T- X_{CO2} phase relations for magnesite, chrysotile, talc, and quartz with unit activities are plotted at 5 bars pressure in Figure 6.

The temperature of the invariant point on the T- X_{CO2} plot is controlled mainly by the reaction chrysotile + 2 quartz = talc + H_2O , which is not sensitive to X_{CO2} over the range considered. THERMOCALC output reports the uncertainty of the temperature of this reaction as \pm 4°C (one sigma).

However, the THERMOCALC database also lists uncertainties for the standard enthalpies of each mineral. If the standard enthalpies are allowed to vary by two standard deviations, the temperature of this reaction for pure end-members at 1 bar can be as low as -33°C or as high as 65°C. Note that these extreme bounds arise via taking the 2 sigma high values of enthalpy for reactants (4% probability) and the 2 sigma low values for products (also 4% probability), or vice versa. Thus, the resulting temperature extremes lie well outside 2 sigma bounds for the reaction temperature. Using other internally-consistent thermodynamic databases further illustrates the uncertainty associated with these calculations; using the thermodynamic data of Gottschalk (1997) yields a temperature of 46°C, while using that of Berman (1988) yields a temperature of 37°C, not only illustrating the uncertainty associated with choosing a thermodynamic database, but also suggesting that the temperatures calculated by THERMOCALC may be low.

Additionally, the single parameter activity model used by THERMOCALC for H_2O - CO_2 activity is calibrated over the PT range of 0.5–20 kbar, 400–1,000 °C, and is likely to be imprecise near the solvus involving immiscible H_2O - and CO_2 -rich fluids (Holland and Powell 2003). Applying THERMOCALC to our low-pressure, low-temperature system is thus an approximation that will become less accurate near the solvus. For calculations at pressures greater than 3.5 bars, serpentine + quartz are predicted to be stable when X_{CO2} in fluid is more than one order of magnitude lower than the solubility limit of CO_2 in aqueous fluid (Duan and Sun 2003). Even so, because the non-ideal behavior of H_2O - CO_2 fluids at these low temperatures and pressures may not be well fit by the activity model used in THERMOCALC, the results of our thermodynamic calculations in this section should be viewed as illustrative, rather than fully quantitative.

Serpentine accommodates more iron than does talc (Trommsdorff and Evans 1972), and thus adding FeO to MgO-SiO₂-H₂O-CO₂ stabilizes serpentine with respect to talc. A T-X diagram for magnesite-chrysotile-talc-quartz was calculated with reduced activities for chrysotile and talc based on an ideal mixing-on-sites model (Spear 1993). Magnesite in our serpentinite samples contains very little iron (less than 0.5 wt% on average), so magnesite was modeled with unit activity. The average Mg# of our (limited number of) of talc analyses was about 95.5, which is approximately consistent with Fe partitioning between talc and antigorite observed by Trommsdorff and Evans (1972). A fit to their data predicts an Mg#

of 96.4 for talc when antigorite has an Mg# of 92.8, the average Mg# of serpentine in our single talc-bearing sample, OM07-28A. We therefore used the fit of Trommsdorff and Evans (1972) to predict the Mg# of talc that could coexist with our higher-Fe serpentine. While iron is the most abundant element substituting into serpentine and talc, the concentrations of aluminum and chromium present in these minerals are also considered in the activity model. Serpentine in quartz-bearing samples has an average Mg# of 89.9 and an average chrysotile activity of 0.710. Based on the partitioning data of Trommsdorff and Evans (1972), talc in Fe/Mg exchange equilibrium with this serpentine should have an Mg# of 94.9, giving an Mg-end-member talc activity of 0.856. When the T-X phase diagram is recalculated with these reduced activities, the invariant point is predicted by THERMOCALC to be at 19°C. While this temperature is no doubt imprecise, it does reflect the fact that preferential Fe, Al and Cr substitution into serpentine will raise the temperature of the invariant point, compared to the MgO-SiO₂-H₂O-CO₂ system. An average composition of the three most iron-rich serpentine analyses in our samples yields a chrysotile activity of 0.236 and a theoretical talc activity of 0.474, which places the calculated invariant point at a temperature of 28°C. The location of the calculated invariant point for the range of chrysotile activities observed in our samples is plotted for pressures between 3.5 bar and 1 kbar in Figure 7.

Again, because the H₂O-CO₂ fluid model used in THERMOCALC is uncertain at low temperature and pressure, and because different methods of fitting and extrapolating the available thermodynamic data result in invariant point temperatures that vary by more than 30°C, the calculated temperatures reported in this section should be viewed as approximate. However, they demonstrate that serpentine + quartz could be stable during near-surface weathering of Oman peridotites. Also, even if talc were stable with respect to serpentine + quartz at, e.g., 30°C and 50 bars, these calculations demonstrate that the thermodynamic driving force for talc crystallization would be very low under these conditions, and this could enhance kinetic inhibition of talc formation.

3.4 Clumped isotope thermometry in carbonate veins

Results of clumped isotope analysis of carbonate veins in serpentinized peridotite from the Oman ophiolite are presented in Table 4, including: the CO2 clumped isotope index Δ 47 (a measure of the enrichment in mass 47 isotopologues relative to a stochastic composition); temperature corresponding to

the $\Delta 47$ value (based on calibration of the relationship between carbonate growth temperature and the $\Delta 47$ value of CO_2 extracted from carbonate by phosphoric acid dissolution); bulk isotopic composition of extracted CO_2 in terms of $\delta^{13}C$ and $\delta^{18}O$; the calculated bulk $\delta^{18}O$ of the carbonate mineral (i.e., correcting for the acid dissolution fractionation); and the calculated $\delta^{18}O$ of water in equilibrium with the carbonate at the calculated clumped isotope temperature. Clumped isotope temperatures range from 23°C to 43°C, placing all samples within error of the observed temperatures of shallow groundwater and surface water in Oman, ~17-40°C (Neal and Stanger 1985; Matter 2001, 2005; Dewandel et al. 2005; our unpublished data). $\delta^{13}C$ ranges from -11.4% to 0.3% (V-PDB), and $\delta^{18}O$ of the carbonates range from 26.4% to 38.8% (VSMOW). We believe the $\delta^{18}O$ data for magnesite presented here are more accurate than those presented in Kelemen et al. (2011), as those were derived from magnesite reacted for only 10 minutes at 70°C, so that much of the CO_2 could have been produced by dissolution of dolomite rather than magnesite. $\delta^{18}O$ values of water calculated to be in equilibrium with the carbonate at the measured clumped isotope temperature range from -7.7% to 8.0% (VSMOW), with the majority of samples falling within error of the -2.1% to 3.6% (VSMOW) range of $\delta^{18}O$ observed in Oman groundwater (Neal and Stanger 1985; Clark et al. 1992; Matter 2005; J. Matter, pers. comm. 2012), as shown in Figure 8.

4. Discussion

The observation of high Fe content in serpentines and the coexistence of serpentine and quartz are consistent with low-temperature serpentinization. This supports the hypothesis that serpentine in our samples formed at lower temperatures than most lizardite and chrysotile whose compositions are published. Oxygen isotope fractionation between serpentine and magnetite in drilled and dredged samples of seafloor outcrops suggests serpentinization occurred at temperatures of ~130-400°C (Wenner and Taylor 1971; Früh-Green et al. 1996; Barnes et al. 2009). Continental lizardites and chrysotiles yield serpentine-magnetite temperatures above 85°C (Wenner and Taylor 1971).

Johannes (1969) placed the reaction serpentine + 2 quartz = talc $+ H_2O$ at $\sim 300^{\circ}C$ at 2 kbar for Mg-end-member minerals. However, the experiments to determine the equilibrium conditions for this reaction were unsuccessful in that study, and serpentinite is observed to have altered to talc during formation of listwanite at temperatures less than 300°C (e.g., Hansen et al. 2005). Tsikouras et al. (2006)

attributed the coexistence of serpentine + quartz in their samples to Mg loss to an extremely CO_2 -rich fluid at temperatures below 500°C, via the reaction chrysotile = $2 \text{ SiO}_2 + 3 \text{ MgO}(aq) + 2 \text{ H}_2\text{O}$, but our calculations suggest that serpentine + quartz is unstable at all X_{CO2} above ~ 50 °C. This low-temperature interpretation is also consistent with the formation of mesh-texture serpentine + quartz during surface and near-surface weathering of ultramafic clasts, as reported by Beinlich et al. (2010). Our calculations show that serpentine coexists stably with quartz at low temperatures, from a calculated temperature of ~ 15 to 46°C for Mg end-members, extending to slightly higher temperatures for bulk compositions containing iron, in which Fe partitions preferentially into serpentine rather than talc. Thus the presence of serpentine + quartz supports the hypothesis that serpentinization is ongoing in the Samail ophiolite in Oman, in a low-temperature weathering horizon (Neal and Stanger 1985; Barnes and O'Neil 1969; Bruni et al. 2002; Kelemen and Matter 2008; Kelemen et al. 2011).

The low temperatures required for serpentine + quartz stability are within the range of carbonate crystallization temperatures derived from conventional carbonate-water oxygen isotope thermometry and clumped stable isotope thermometry, and of observed groundwater temperatures in Oman. Clumped isotope thermometry yields carbonate crystallization temperatures of 23-43°C (Fig. 8), consistent with the observed temperatures of shallow groundwater and surface water in Oman, ~17-40°C (Neal and Stanger 1985; Matter 2001, 2005; Dewandel et al. 2005; our unpublished data). Similarly, Kelemen et al. (2011) used observed δ^{18} O in monomineralic magnesite and calcite samples, together with the observed range of δ¹⁸O in Oman ground water and spring water and the water-carbonate oxygen isotope thermometer of Chacko and Deines (2008) to determine crystallization temperatures of 23-60°C. This is consistent with our observation that calculated values of δ^{18} O for water in equilibrium with carbonate are mostly consistent with the observed modern range of -2.1 to 3.6% (VSMOW) in groundwater, surface waters, and alkaline springs in peridotite catchments in Oman (Neal and Stanger 1985; Clark et al. 1992; Matter 2005; J. Matter, pers. comm. 2012). The agreement of carbonate stable isotope data with temperatures and δ^{18} O values in modern groundwater demonstrates that alteration in the form of carbonate vein precipitation in peridotite occurred under conditions similar to those in the modern environment. Together with geochemical trends in groundwater (Barnes and O'Neil 1969; Neal and Stanger 1985; Bruni et al. 2002; Kelemen and Matter 2008; Paukert et al. 2012), this suggests that carbonation and serpentinization may be occurring simultaneously via reactions between modern groundwater and peridotite, as proposed by Kelemen and Matter (2008) and Barnes and O'Neil (1969).

The range of fluid X_{CO2} within the serpentine + quartz stability field is also consistent with observed groundwater compositions in the Samail Ophiolite, which range from a mole fraction of $\sim 10^{-3.5}$ in neutral to slightly alkaline groundwater to below detection limits (less than 10^{-5}) in alkaline springs (Neal and Stanger 1985; Matter 2001, 2005; Dewandel et al. 2005; E. Shock, pers. comm., 2011). Groundwater at the high X_{CO2} end of this range would be too CO_2 -rich for serpentine + quartz to coexist stably according to our calculations. However, the presence of calcite in our quartz-bearing serpentinites, together with the observation that calcite veins in serpentine are abundant in the vicinity of alkaline springs, suggests that alteration occurred in the presence of Ca-OH type waters, which have exceptionally low dissolved CO_2 contents in peridotite catchments worldwide (Barnes and O'Neil 1969; Barnes et al. 1978; Bruni et al. 2002; Kelemen and Matter 2008). Thus, these samples were likely subject to alteration under conditions of low fluid CO_2 concentrations, favorable for serpentine + quartz stability.

It is likely that serpentinization of peridotite in Oman began during hydrothermal alteration of shallow suboceanic mantle prior to obduction (Boudier et al. 2010). Some of this may have occurred at higher temperatures, in the ~130-400°C range inferred from δ^{18} O in ocean-floor serpentinites (e.g., Wenner and Taylor 1971; Früh-Green et al. 1996; Barnes et al. 2009). This initial, oceanic serpentinization may have yielded Fe-poor serpentine and magnetite, as is observed in a partially serpentinized peridotite sample (OM08-206I), collected at the same location as samples with serpentine + quartz. During later alteration at lower temperature, early magnetite may have been consumed to form more Fe-rich serpentine as in reactions R1-R3 in Section 3.2. Much of the Fe-rich serpentine we observe could represent re-equilibration of the products of previous serpentinization rather than direct serpentinization of olivine, particularly where serpentine is locally more Fe-rich than the initial olivine in the peridotite.

Elevated silica activity may favor the formation of iron-rich serpentine (Frost and Beard 2007), as in reactions R1-3, thus indirectly stabilizing the serpentine + quartz assemblage. If the initial peridotite is harzburgite or lherzolite rather than dunite, the presence of pyroxene will shift the bulk composition towards silica. We infer that clinopyroxene dissolution is occurring during present-day serpentinization of

peridotite due to the generation of Ca-OH fluid (Barnes and O'Neil 1969; Neal and Stanger 1985; Bruni et al. 2002). In the samples where we see coexisting serpentine + quartz, calcite is the most abundant carbonate phase, and relict clinopyroxenes are observed. Centimeter-scale pyroxene-rich bands comprise ~10% of the outcrop where serpentine + quartz assemblage and found, and quartz is restricted to these bands. Small-scale heterogeneities in fluid composition due to pyroxene dissolution may shift the silica activity to higher values, locally stabilizing Fe-rich serpentine and thereby expanding the stability of serpentine + quartz.

Precipitation of magnesite and dolomite veins could also introduce heterogeneities that would favor the local formation of Fe-rich serpentine in some areas. Serpentine is more silica-rich than olivine, so during alteration of olivine alone without SiO₂ mass transfer, it must coexist with a less silica-rich Mg phase, such as magnesite, dolomite, or brucite e.g.,

(R4)
$$2 \text{ Mg}_2 \text{SiO}_4 + 2 \text{ H}_2 \text{O} + \text{CO}_2 = \text{Mg}_3 \text{Si}_2 \text{O}_5 (\text{OH})_4 + \text{MgCO}_3$$

forsterite serpentine magnesite

Higher concentrations of CO₂ favor additional magnesium carbonate precipitation, thereby raising the silica activity of the fluid. In Oman, dense networks of magnesite veins constitute large volume proportions in serpentinized peridotite at meter scales, indicating that silica activity was elevated, at least locally. Additionally, magnesium dissolution and removal during alteration (e.g., Snow & Dick, 1995) may have produced slightly lower bulk Mg# in the quartz-bearing samples, compared to typical Oman mantle peridotite. Both lower bulk Mg# and elevated silica could favor the formation of Fe-rich serpentine, and as a result, stabilize the serpentine + quartz assemblage. However, as discussed in Section 3.3 above, the main factor leading to the absence of talc in our samples is low temperature, stabilizing serpentine + quartz even in MgO-SiO₂-H₂O-CO₂, and leading to reactions producing Fe-rich serpentine from magnetite + Fe-poor serpentine.

5. Conclusions

Some samples of serpentinized peridotite in the Oman ophiolite contain the unusual assemblage, serpentine + quartz, instead of talc. Phase equilibrium calculations suggest that this assemblage is stable at

low temperature, below approximately 15 to 46°C in an Fe-free system depending on the thermodynamic data used. Consistent with this low inferred temperature, clumped isotope analyses of carbonate veins minerals in these and other, similar samples yield temperature estimates of 15 to 50°C. Serpentine in serpentinized peridotite with carbonate veins in the Oman Ophiolite is relatively Fe-rich in comparison with other serpentinites. The formation of high-Fe serpentine rather than magnetite is favored at low temperatures, at a given pressure, and at high silica activities. Thus the presence of Fe-rich serpentine is another indicator of low-temperature alteration. Also, because Fe partitions into serpentine relative to tale, relatively high Fe content can help to stabilize serpentine + quartz. Temperatures and oxygen isotope compositions of fluids, inferred from phase equilibria and clumped isotope data, are consistent with the range measured in groundwater and alkaline spring water in peridotite catchments in Oman. These data all support the hypothesis that serpentinization and carbonation of peridotite are occurring at ambient temperatures in a shallow weathering horizon in the Oman Ophiolite.

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Figure Captions

Fig. 1 XRD spectrum of quartz-bearing carbonate-veined serpentinized peridotite sample OM08-206D compared to XRD spectra for lizardite (RRUFFID=R060006), calcite (RRUFFID=R040070), and quartz (RRUFFID=R040031) from the RRUFF database, (Downs 2006), Cu-Kα X-ray radiation. Axes are intensity (counts) vs. degrees 2-theta

Fig. 2 Back-scattered electron image showing serpentine (dark grey), coexisting with quartz (light grey "islands") in (a) sample OM08-206A and (b) OM08-206D. The lighter grey veins in both images are calcite. Calcite is also found along with quartz in some of the cores of the mesh texture in OM08-206A

Fig. 3 Frequency histogram of total mol% Fe/[Fe + Mg] contents in serpentine from our quartz-bearing serpentinite (red, n= 96 analyses), our quartz-free carbonated serpentinite (blue, n=486 analyses), and lizardites from the literature (black, n=503 analyses) (Evans et al. 2009)

Fig. 4 WDS element map of Fe content in coexisting serpentine + quartz in sample OM08-206D. Strands of Fe-rich material (very light grey and white) are distributed within the quartz regions (black), concentrated at the edge of the quartz "mesh centers." Although the serpentine (grey) throughout these samples containing coexisting serpentine + quartz is already more Fe-rich than serpentine in other samples of carbonated serpentinites from Oman, the mesh serpentine surrounding these quartz regions forms even more Fe-rich rims (light grey) around the quartz mesh centers

Fig. 5 Projection of all serpentine data and compositions of forsterite, talc, and serpentine end-members onto a ternary MgO-SiO₂-FeO diagram (molar proportions). Red crosses represent serpentine from samples that contain coexisting serpentine and quartz (OM08-206A, OM08-206D, OM08-01). The size of the circles representing these data approximates the uncertainty in these measurements, given as the 2σ of repeated measurements of antigorite standard BM66586 analyzed as unknown using the same calibrations used to analyze serpentine in our samples. See Online Resource 1 for data

Fig. 6 T-X diagram of the phase relations between pure (activity =1) Mg end-member magnesite (Mgs), chrysotile (Chr), talc (Ta), and quartz (Q) at 5 bar

Fig. 7 Location of the invariant point for magnesite-chrysotile-talc-quartz calculated in THERMOCALC at pressures of 3.5, 10, 100 and 1000 bars, for chrysotile and talc activities corresponding to Fe-contents ranging from pure Mg end-members to our most Fe-rich serpentines (A through D). Talc activity was calculated from the predicted Mg# of talc coexisting with serpentine of a given Mg# based on the data of Trommsdorff and Evans (1972). (A) unit activities for all mineral phases (pure Mg end-members); (B) chrysotile activity = 0.71, talc activity = 0.86, corresponds to the average composition of serpentine coexisting with quartz; (C) chrysotile activity = 0.51, talc activity = 0.72, corresponds to the twenty most Fe-rich analyses of serpentine; D) chrysotile activity = 0.24, talc activity = 0.47, corresponds to the three most Fe-rich analysis of serpentine. The water-rich limb of the CO_2 - H_2O solvus, as calculated by Duan and Sun (2003) is plotted as dashed lines, showing that even at the low pressures considered (>3.5 bar), unmixing should not occur in the H_2O -rich fluids in equilibrium with serpentine + quartz

Fig. 8 (a) Calculated δ^{18} O of water from which carbonates precipitated, versus clumped isotope temperatures of carbonate veins; blue circles = calcite, yellow diamonds = dolomite, vermillion triangles = magnesite. Dashed and dotted lines connect different minerals from the same sample. Error bars represent one standard error of the mean. The grey box showing the range of temperatures and δ^{18} O measured in modern groundwater in Oman (Neal and Stanger 1985; Clark et al. 1992; Matter 2005; our unpublished data), is consistent with most clumped isotope temperatures and calculated fluid oxygen isotope compositions. (b) Temperatures calculated by conventional oxygen isotope thermometry, representing a range of δ^{18} O_{water} from -2.1 to +3.6, versus clumped isotope temperatures; blue boxes = calcite, yellow boxes = dolomite, vermillion boxes = magnesite. Box width represents two standard errors of the mean of clumped isotope temperature

Electronic Supplementary Materials

Online Resource 1 Electron microprobe calibration setup for serpentine analysis

Online Resource 2 All electron microprobe analyses of serpentine in our Oman serpentinites and antigorite standard BM66586. Analyses have been corrected based on repeat analyses of the antigorite standard, as in Online Resource 1. Points close to the edge of serpentine grains that appear to include mixed analyses with Fe-oxides or carbonates, points with totals <73%, and points with >0.3 Al atoms per formula unit (most likely chlorite) have been excluded.

Table 1 Bulk rock mineralogy and major element chemistry of Oman serpentinite samples

Sample	OM07-05	OM07-17	OM07-27	OM07-28A	OM07-34A	OM07-39	OM07-53	OM07-54	OM08-01
UTM ^a E	451254	586007	598089	598021	608365	634006	640914	646240	608957
UTM N	2652948	2523245	2555805	2555719	2526870	2571609	2533993	2537136	2529931
Mineralogy ^b									
serp	X	X	x	X	tr	x	X	X	Χ
mgs				X		X			
dol	X	X	X	X	X	x	X	X	tr
cc	х	X	x		X	x	Х	tr	Χ
other trace minerals:	ank	Cr-sp, Fe- ox, AgS, hal	Cr-sp, chl, di, hem, bar	Cr-sp, tc, di	bar, gyp	Cr-sp	chl, di	hem, chl, mgt, Cr-sp, en, AgS	cr-sp, qtz, chl
wt%									
SiO2	19.78	10.22	7.82	-	2.25	1.09	20.30	21.46	28.10
TiO2	0.01	0.03	0.04	-	0.02	0.00	0.20	0.13	0.02
A12O3	0.34	0.48	1.12	-	0.39	0.01	3.22	2.27	0.80
FeO*	3.37	1.79	2.12	-	0.16	0.27	4.67	1.91	4.10
MnO	0.08	0.02	0.04	-	0.00	0.01	0.09	0.08	0.07
MgO	28.25	18.49	21.78	-	3.11	44.94	23.00	29.30	26.94
CaO	17.54	31.09	26.45	-	52.28	2.97	20.23	15.00	16.45
Na2O	0.02	0.07	0.01	-	0.02	0.01	0.04	0.02	0.04
K2O	0.00	0.06	0.07	-	0.01	0.00	0.05	0.00	0.02
P2O5	0.01	0.01	0.01	-	0.01	0.00	0.01	0.01	0.01
LOI (%)	29.92	37.14	39.38	-	41.71	48.68	27.51	29.02	22.84
Total	99.32	99.41	98.84	-	99.96	97.99	99.34	99.19	99.38
Mg#	0.94	0.95	0.95		0.97	1.00	0.90	0.96	0.92
Table 1 (C	Continued)								
Sample	OM08-201	OM08-206A	OM08-206D	OM08-206I	OM09-13	OM09-121	OM09-122	OM09-129	OM09-130
UTM E	600090	596370	596370	596370	621536	487520	487520	487520	487520
UTM N	2528421	2531270	2531270	2531270	2585653	2576000	2576000	2576000	2576000
Mineralogy									
	X	X	X	X	X	X	X	X	X
serp		^	^					^	
serp mgs	Α	Λ	^				Х	X	
serp mgs dol		Χ	^			X			
mgs dol	×				-	X	X X	x	
mgs dol cc	x	X	X				х		X
mgs dol cc qtz	x					X	X X	x	
mgs dol cc	x	X	X	X mgt, Cr-sp, chl	Cr-sp, goeth, chl, bar	X	X X	x	
mgs dol cc qtz ol other trace	X X Cr-sp, hem,	X x di, Fe-ox,	X x di, Fe-ox,	X mgt, Cr-sp,	Cr-sp, goeth, chl,	X x mgt, Cr-sp,	X X tr	x X	X Cr-sp, hem,
mgs dol cc qtz ol other trace minerals:	X X Cr-sp, hem,	X x di, Fe-ox,	X x di, Fe-ox,	X mgt, Cr-sp,	Cr-sp, goeth, chl,	X x mgt, Cr-sp,	X X tr	x X	X Cr-sp, hem,
mgs dol cc qtz ol other trace minerals:	X X Cr-sp, hem, pyr	X x di, Fe-ox, Cr-sp	X x di, Fe-ox, Cr-sp, en	X mgt, Cr-sp, chl	Cr-sp, goeth, chl, bar	X x mgt, Cr-sp, di	x X tr	X X mgt	X Cr-sp, hem, mgt
mgs dol cc qtz ol other trace minerals: wt% SiO2	X X Cr-sp, hem, pyr	X x di, Fe-ox, Cr-sp	X x di, Fe-ox, Cr-sp, en	X mgt, Cr-sp, chl	Cr-sp, goeth, chl, bar	x mgt, Cr-sp, di	x X tr	x X mgt	X Cr-sp, hem, mgt
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2	X X Cr-sp, hem, pyr	x di, Fe-ox, Cr-sp	x di, Fe-ox, Cr-sp, en	X mgt, Cr-sp, chl 37.77 0.01	Cr-sp, goeth, chl, bar 40.04 0.03	x mgt, Cr-sp, di 22.93 0.01	x x tr	x X mgt	X Cr-sp, hem, mgt 23.59 0.01
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3	X X Cr-sp, hem, pyr 19.36 0.01 0.95	X x di, Fe-ox, Cr-sp 27.80 0.02 0.54	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75	X mgt, Cr-sp, chl 37.77 0.01 1.65	Cr-sp, goeth, chl, bar 40.04 0.03 1.47	x mgt, Cr-sp, di 22.93 0.01 0.41	x x tr hem	x X mgt 27.49 0.01 0.36	X Cr-sp, hem, mgt 23.59 0.01 0.36
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO*	X X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12	X x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62	x x tr hem 18.12 0.01 0.33 1.68	x X mgt 27.49 0.01 0.36 5.90	23.59 0.01 0.36 5.13
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO* MnO	X X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12 0.06	X x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29 0.09	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47 0.11	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09 0.10	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80 0.08	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62 0.05	x x tr hem 18.12 0.01 0.33 1.68 0.04	x x mgt 27.49 0.01 0.36 5.90 0.10	X Cr-sp, hem, mgt 23.59 0.01 0.36 5.13 0.08
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO* MnO MgO	X X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12 0.06 18.95	X x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29 0.09 23.96	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47 0.11 17.37	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09 0.10 36.10	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80 0.08 38.12	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62 0.05 28.97	x x tr hem 18.12 0.01 0.33 1.68 0.04 27.02	x mgt 27.49 0.01 0.36 5.90 0.10 30.17	X Cr-sp, hem, mgt 23.59 0.01 0.36 5.13 0.08 23.67
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO* MnO MgO CaO	X X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12 0.06 18.95 28.35	x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29 0.09 23.96 17.13	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47 0.11 17.37 24.17	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09 0.10 36.10 0.21	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80 0.08 38.12 0.38	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62 0.05 28.97 15.62	x x tr hem 18.12 0.01 0.33 1.68 0.04 27.02 20.12	x mgt 27.49 0.01 0.36 5.90 0.10 30.17 12.51	X Cr-sp, hem, mgt 23.59 0.01 0.36 5.13 0.08 23.67 21.06
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO* MnO MgO CaO Na2O K2O	x X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12 0.06 18.95 28.35 0.15 0.01	x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29 0.09 23.96 17.13 0.06 0.03	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47 0.11 17.37 24.17 0.04 0.02	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09 0.10 36.10 0.21 0.06 0.00	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80 0.08 38.12 0.38 0.03 0.00	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62 0.05 28.97 15.62 0.05 0.01	x x tr hem 18.12 0.01 0.33 1.68 0.04 27.02 20.12 0.33 0.04	x mgt 27.49 0.01 0.36 5.90 0.10 30.17 12.51 0.07 0.00	X Cr-sp, hem, mgt 23.59 0.01 0.36 5.13 0.08 23.67 21.06 0.14 0.01
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO* MnO MgO CaO Na2O K2O P2O5	x X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12 0.06 18.95 28.35 0.15 0.01	x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29 0.09 23.96 17.13 0.06 0.03 0.01	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47 0.11 17.37 24.17 0.04 0.02 0.02	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09 0.10 36.10 0.21 0.06 0.00 0.01	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80 0.08 38.12 0.38 0.03 0.00 0.01	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62 0.05 28.97 15.62 0.05 0.01	x x tr hem 18.12 0.01 0.33 1.68 0.04 27.02 20.12 0.33 0.04 0.01	x mgt 27.49 0.01 0.36 5.90 0.10 30.17 12.51 0.07 0.00 0.00	X Cr-sp, hem, mgt 23.59 0.01 0.36 5.13 0.08 23.67 21.06 0.14 0.01 0.01
mgs dol cc qtz ol other trace minerals: wt% SiO2 TiO2 Al2O3 FeO* MnO MgO CaO Na2O K2O	x X Cr-sp, hem, pyr 19.36 0.01 0.95 3.12 0.06 18.95 28.35 0.15 0.01	x di, Fe-ox, Cr-sp 27.80 0.02 0.54 6.29 0.09 23.96 17.13 0.06 0.03	x di, Fe-ox, Cr-sp, en 25.29 0.01 0.75 6.47 0.11 17.37 24.17 0.04 0.02	X mgt, Cr-sp, chl 37.77 0.01 1.65 8.09 0.10 36.10 0.21 0.06 0.00	Cr-sp, goeth, chl, bar 40.04 0.03 1.47 6.80 0.08 38.12 0.38 0.03 0.00	x mgt, Cr-sp, di 22.93 0.01 0.41 4.62 0.05 28.97 15.62 0.05 0.01	x x tr hem 18.12 0.01 0.33 1.68 0.04 27.02 20.12 0.33 0.04	x mgt 27.49 0.01 0.36 5.90 0.10 30.17 12.51 0.07 0.00	X Cr-sp, hem, mgt 23.59 0.01 0.36 5.13 0.08 23.67 21.06 0.14 0.01

a UTM coordinates are in zone 40Q.
b serp=serpentine, mgs=magnesite, dol=dolomite, cc=calcite, qtz=quartz, ol=olivine, Cr-sp=chromian spinel, chl=chlorite, tc=talc, hem=hematite, mgt=magnetite, goeth=goethite, Fe-ox=undetermined Fe-oxides, di=diopside, en=enstatite, ank=ankerite, hal=halite, bar=barite, gyp=gypsum. Large X indicates major phase, smaller x indicates minor phase, and 'tr' indicates only trace amounts of phase are present.

Table 2 Select electron microprobe analyses of quartz in Oman serpentinites compared to thin section slide glass

Sample	Slide glass	OM08-206A		OM08-2	2σ	
wt%						
SiO2	72.2	96.6	99.7	97.9	97.4	0.4
TiO2		0.00	0.00	0.00	0.01	0.06
A12O3	1.20	0.02	0.00	0.02	0.00	0.02
FeO*	0.03	3.1	0.3	1.6	2.1	0.1
MgO	4.3	0.1	0.0	0.1	0.1	0.2
MnO		0.00	0.03	0.02	0.00	0.03
CaO	6.4	0.0	0.0	0.0	0.0	0.4
Na2O	14.30	0.02	0.01	0.04	0.01	0.09
K2O	1.20	0.03	0.02	0.01	0.01	0.01
Cr2O3		0.13	0.00	0.00	0.02	0.04
NiO		0.01	0.02	0.02	0.01	0.06
SO3	0.03					
Total	99.7	100.0	100.2	99.8	99.7	

Table 3 Serpentine analyses, averaged by sample

Sample #	OM07-05	OM07-17	OM07-27	OM07-28A	OM07-34A	OM07-39	OM07-53	OM07-54	OM08-01
n	6	111	19	2	3	6	2	59	49
wt%									
SiO2	42.0	42.0	42.1	44.9	47.7	42.4	40.3	42.0	40.7
TiO2	0.00	0.01	0.01	0.01	0.06	0.01	0.01	0.05	0.01
A12O3	0.42	0.20	0.39	0.07	2.96	0.07	4.51	1.62	0.62
FeO	3.94	6.51	4.47	5.32	0.95	2.26	5.07	2.72	6.04
MnO	0.06	0.08	0.05	0.02	0.04	0.04	0.10	0.05	0.10
MgO	41.0	39.4	41.0	38.6	36.7	40.0	37.1	40.6	39.7
CaO	0.04	0.11	0.09	0.12	0.71	0.16	0.14	0.11	0.19
Cr2O3	0.05	0.02	0.01	0.01	0.04	0.01	0.03	0.09	0.13
NiO	0.09	0.03	0.15	0.53	0.03	0.08	0.22	0.13	0.26
Total	84.6	85.5	85.1	86.7	86.5	82.0	84.6	84.3	84.7
Mg#	0.95	0.92	0.94	0.93	0.99	0.97	0.93	0.96	0.92
Sample #	OM08-201	OM08-206A	OM08-206D	OM08-206I	OM09-13	OM09-121	OM09-122	OM09-129	OM09-130
n	101	27	69	25	4	29	17	13	40
wt%									
SiO2	40.3	40.8	42.1	41.2	42.2	43.2	43.1	41.1	42.1
TiO2	0.01	0.01	0.01	0.01	0.03	0.01	0.00	0.01	0.01
A12O3	0.71	0.13	0.47	0.99	1.44	0.34	0.35	0.01	0.07
FeO*	5.85	7.65	7.42	4.77	4.11	2.19	2.40	3.69	3.89
MnO	0.08	0.08	0.05	0.05	0.05	N.A.	0.07	N.A.	0.05
MgO	39.8	38.1	37.1	40.7	40.4	42.1	42.1	42.2	41.7
CaO	0.06	0.20	0.26	0.05	0.04	0.12	0.03	0.06	0.05
Cr2O3	0.03	0.04	0.12	0.02	0.25	0.15	0.03	0.01	0.02
Cr2O3 NiO	0.03 0.27	0.04 0.33		0.02 0.34	0.25 0.23	0.15 0.17	0.03 0.18	0.01 0.18	0.02 0.23
			0.12						

Table 4 Clumped isotope measurements of carbonate veins from Oman serpentinites

Sample	Min.	Digestion T, °C	Corrected $\Delta 47^a$	T, °C ^b	$\delta^{13}C$	$\delta^{18} O CO_2$ VSMOW	δ ¹⁸ O carbonate VSMOW ^c	δ ¹⁸ O water VSMOW ^d
OM08-206D	cc	90	0.600 ± 0.005	38 ± 2	-9.003 ± 0.003	37.286 ± 0.005	28.874 ± 0.002	2.8 ± 0.3
OM08-206D	cc	90	0.635 ± 0.014	28 ± 4	-8.986 ± 0.002	37.285 ± 0.004	28.874 ± 0.001	0.8 ± 0.8
OM10-34B	cc	90	0.598 ± 0.010	39 ± 3	-9.090 ± 0.002	36.313 ± 0.005	27.909 ± 0.002	2.0 ± 0.6
OM10-34B	cc	90	0.604 ± 0.007	37 ± 2	-8.926 ± 0.004	37.840 ± 0.011	27.898 ± 0.002	1.6 ± 0.4
OM09-130	cc	25	0.594 ± 0.014	40 ± 5	$-11.409 \pm .001$	36.897 ± 0.002	26.376 ± 0.002	0.7 ± 0.8
OM08-01	cc	50	0.625 ± 0.005	31 ± 2	$-10.230 \pm .001$	37.840 ± 0.004	28.274 ± 0.004	0.7 ± 0.3
OM07-54	dol	90	0.630 ± 0.008	29 ± 2	$-2.019 \pm .001$	48.166 ± 0.002	38.831 ± 0.002	8.0 ± 0.4
OM07-17	dol	50	0.587 ± 0.006	43 ± 2	$0.343 \pm .001$	43.704 ± 0.002	33.203 ± 0.002	5.1 ± 0.3
OM08-01	dol	50	0.588 ± 0.010	42 ± 3	$-10.173 \pm .001$	37.789 ± 0.002	27.348 ± 0.002	-0.7 ± 0.6
OM07-61a	dol	50	0.638 ± 0.009	27 ± 3	$-4.933 \pm .001$	45.923 ± 0.005	35.400 ± 0.005	4.3 ± 0.5
OM07-61a	mgs	80	0.643 ± 0.008	26 ± 2	$-8.666 \pm .001$	43.225 ± 0.001	33.146 ± 0.001	-4.9 ± 0.6
OM07-07	mgs	80	0.652 ± 0.007	23 ± 2	$-9.944 \pm .001$	40.832 ± 0.001	30.776 ± 0.003	-7.7 ± 0.6

 $[\]pm$ uncertainties represent one standard error of the mean of all acquisitions for a sample

Calcite: $10^3 \ln \alpha_T = 5.608 * 10^5/T^2 + 3.89$, Dolomite: $10^3 \ln \alpha_T = 5.858 * 10^5/T^2 + 5.51$, Magnesite: $10^3 \ln \alpha_T = 6.845 * 10^5/T^2 + 4.22$,

where $10^3 \ln \alpha_T$ refers to fractionation at digestion temperatures T in Kelvin.

^a Δ 47 corrected for mass spectrometer nonlinearity by normalization based on heated gas measurements (Huntington et al. 2009), and for fractionation during acid digestion at temperatures above 25°C (Passey et al. 2010). For these measurements, the linearity correction was given by: Δ 47(corrected for linearity) = -0.8453/ -0.779*(Δ 47_{measured}-(0.0076* Δ 47+ -0.779)) for the first four rows and Δ 47(corrected for linearity) = -0.8453/ -0.815*(Δ 47_{measured}-(0.0077* Δ 47+ -0.815)) for the remaining rows.

^b Calculated from a polynomial fit to inorganic calcite and dolomite experimental calibration data (Ghosh et al. 2006; M. Bonifacie, manuscript in preparation)

 $^{^{\}rm c}$ CO₂-carbonate δ^{18} O fractionation factors for phosphoric digestion from Das Sharma et al. (2002):

^d Carbonate-water δ^{18} O fractionation factors for calcite from Kim and O'Neil (1997), for dolomite from Vasconcelos et al. (2005), and for magnesite from reduced partition coefficients from Schauble et al. (2006) and Rosenbaum (1997)

Figure 1

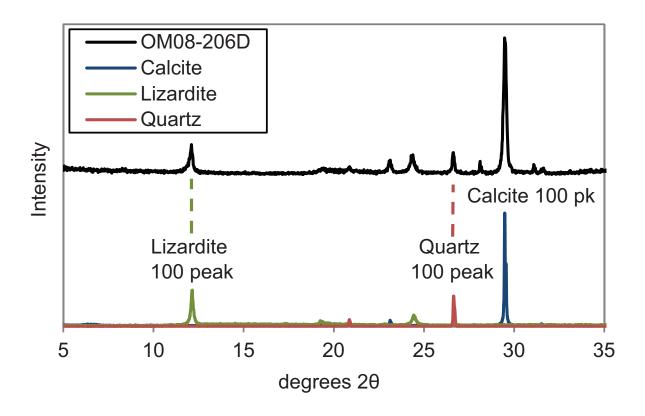


Figure 2

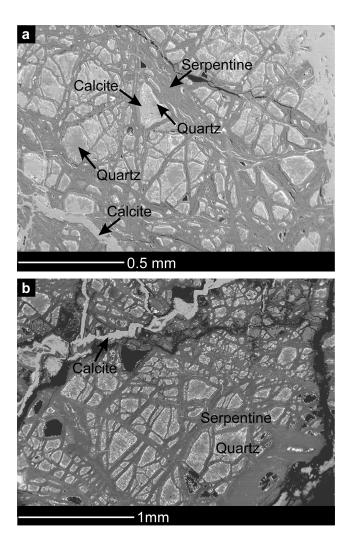


Figure 3

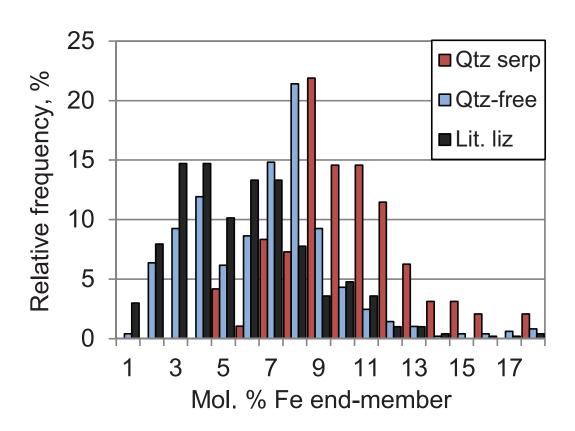


Figure 4

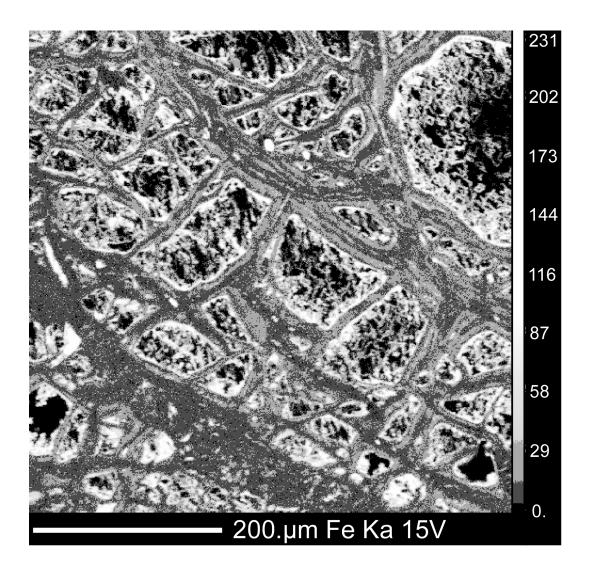


Figure 5

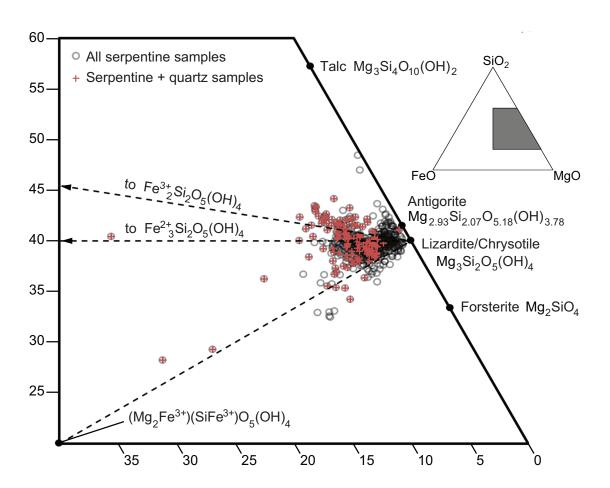


Figure 6

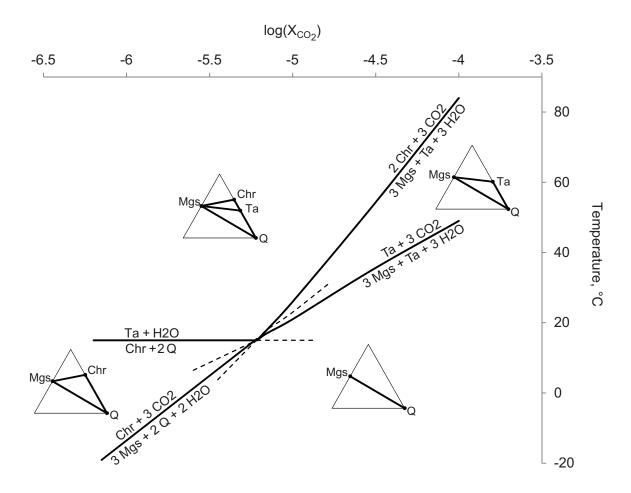


Figure 7

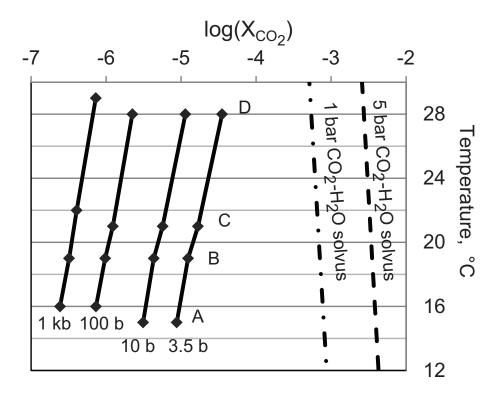
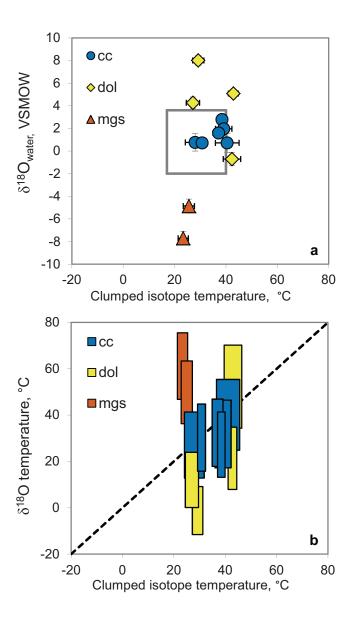


Figure 8



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