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# Multiple point principle of the Standard Model with scalar singlet dark matter and right-handed neutrinos 

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#### Abstract

We consider the multiple point principle of the Standard Model with scalar singlet dark matter (DM) and three heavy right-handed neutrinos at the scale where the beta function $\beta_{\lambda}$ of the effective Higgs self-coupling $\lambda_{\text {eff }}$ becomes zero. We do a two-loop analysis and find that the top quark mass $M_{t}$ and the Higgs portal coupling $\kappa$ are strongly related to each other. One of the good points of this model is that a larger $M_{t}(\gtrsim 171 \mathrm{GeV})$ is allowed. This fact is consistent with the recent experimental value $M_{t}=173.34 \pm 0.76 \mathrm{GeV}$ [ATLAS, CDF, CMS, and D0 Collaborations, arXiv: 1403.4427 [hep-ex]], which corresponds to the DM mass $769 \mathrm{GeV} \leq m_{\mathrm{DM}} \leq 1053 \mathrm{GeV}$.


Subject Index B32, B37

## 1. Introduction

The discovery of the Higgs-like particle and its mass [1,2] is a very meaningful result for the Standard Model (SM). It suggests that the Higgs potential can be stable up to the Planck scale $M_{\mathrm{pl}}$ and also that both the Higgs self-coupling $\lambda$ and its beta function $\beta_{\lambda}$ become very small around the Planck scale. This fact has attracted much attention, and there are many works trying to find its physical meaning [3-20].
One interesting and meaningful study is to consider how the physics beyond the SM affects such a criticality. For example, recently there has been a two-loop analysis of the Higgs portal $Z_{2}$ scalar model [21]. In this model, the SM singlet scalar is a dark matter (DM) candidate, and it is found that its mass can be predicted to be $400 \mathrm{GeV}<m_{\mathrm{DM}}<470 \mathrm{GeV}$ from the requirement that $\lambda$ and $\beta_{\lambda}$ simultaneously become zero at $10^{17} \mathrm{GeV}$; this is usually called the multiple point principle (MPP) [3-5].
In this paper, we study the MPP of the next minimal extension of the SM, namely, besides the Higgs portal $Z_{2}$ scalar, we include SM singlet heavy right-handed neutrinos [19,22,23]. The MPP of this model at the (reduced) Planck scale $M_{\mathrm{pl}}$ has already been investigated in Ref. [19]. There, by using the two-loop beta functions and the tree-level Higgs potential, they concluded that $m_{\mathrm{DM}}$ and the heavy Majorana mass $M_{\mathrm{R}}$ of the right-handed neutrino should be

$$
\begin{align*}
& 8.5(8.0) \times 10^{2} \mathrm{GeV} \leq m_{\mathrm{DM}} \leq 1.4(1.2) \times 10^{3} \mathrm{GeV},  \tag{1}\\
& 6.3(5.5) \times 10^{13} \mathrm{GeV} \leq M_{\mathrm{R}} \leq 1.6(1.2) \times 10^{14} \mathrm{GeV}, \tag{2}
\end{align*}
$$

within $172.6 \mathrm{GeV} \leq M_{t} \leq 174.1 \mathrm{GeV}$. The different points in this paper are as follows:
(1) We consider the MPP at the scale where $\beta_{\lambda}$ becomes zero. Namely, we do not fix the MPP scale at $M_{\mathrm{pl}}$. As a result, the condition $\beta_{\lambda}=0$ does not reduce the degrees of freedom of the parameters.
(2) In addition to the two-loop beta functions, we also calculate the one-loop effective potential.
(3) We fix $M_{\mathrm{R}}$ to $10^{13} \mathrm{GeV}$, and include the Yukawa coupling $Y_{\mathrm{R}}$ between the $Z_{2}$ scalar and the right-handed neutrinos.

Although, within the renormalizable Lagrangian, there are also two scalar couplings in this model (see Eq. (12)), we focus on $\lambda$ (and $\beta_{\lambda}$ ) in this paper. ${ }^{1}$ The existence of heavy right-handed neutrinos is naturally needed if we try to explain the light neutrino masses by the seesaw mechanism. Thus, this model is phenomenologically interesting because it can explain both DM and the light neutrino masses.
This paper is organized as follows. In Sect. 2, we review the MPP of the pure SM for the later discussion. In Sect. 3, we give a two-loop analysis of the SM with the scalar singlet DM and three right-handed neutrinos. In Sect. 4, a summary is given.

## 2. Preliminary: Multiple point principle of SM

In the SM, the one-loop effective potential in the Landau gauge is given by

$$
\begin{equation*}
V_{\text {eff }}(\phi, \mu)=V_{\text {tree }}(\phi, \mu)+V_{1 \text { loop }}^{\mathrm{SM}}(\phi, \mu), \tag{3}
\end{equation*}
$$

where

$$
\begin{align*}
V_{\text {tree }}(\phi, \mu):= & e^{4 \Gamma(\phi)} \frac{\lambda(\mu)}{4} \phi^{4},  \tag{4}\\
V_{1 \text { loop }}(\phi):= & e^{4 \Gamma(\phi)}\left\{-6 \cdot \frac{M_{t}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{t}^{2}(\phi)}{\mu^{2}}\right)-\frac{3}{2}+2 \Gamma(\phi)\right]\right. \\
& +3 \cdot \frac{M_{W}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{W}^{2}(\phi)}{\mu^{2}}\right)-\frac{5}{6}+2 \Gamma(\phi)\right] \\
& \left.+3 \cdot \frac{M_{Z}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{Z}^{2}(\phi)}{\mu^{2}}\right)-\frac{5}{6}+2 \Gamma(\phi)\right]\right\},  \tag{5}\\
M_{t}(\phi)= & \frac{y_{t}(\mu)}{\sqrt{2}} \phi, \quad M_{W}(\phi)=\frac{g_{2}(\mu)}{2} \phi, \quad M_{Z}=\frac{\sqrt{g_{2}^{2}(\mu)+g_{\mathrm{Y}}^{2}(\mu)}}{2} \phi . \tag{6}
\end{align*}
$$

Here, $\mu$ is the renormalization scale, $\Gamma(\phi)$ is the wave function renormalization, and $\lambda(\mu), y_{t}(\mu)$, $g_{2}(\mu)$, and $g_{Y}(\mu)$ are the renormalized couplings. ${ }^{2}$ By using these results, the effective Higgs selfcoupling $\lambda_{\text {eff }}(\phi, \mu)$ can be defined as

$$
\begin{equation*}
V_{\mathrm{eff}}(\phi, \mu):=\frac{\lambda_{\mathrm{eff}}(\phi, \mu)}{4} \phi^{4} . \tag{7}
\end{equation*}
$$

To minimize the contribution of $V_{1 \text { loop }}^{\mathrm{SM}}(\phi, \mu)$, we set $\phi=\mu$ in the following discussion.

[^0]

Fig. 1. Left: the running effective Higgs self-coupling $\lambda_{\text {eff }}$ as a function of the Higgs field $\phi$. The blue band corresponds to $95 \%$ confidence level (CL) deviation of the top quark pole mass $M_{t}$. Right: the scale $\Lambda_{\beta}$ where $\beta_{\lambda_{\text {eff }}}$ becomes zero as a function of $M_{t}$.

The left panel of Fig. 1 shows $\lambda_{\text {eff }}(\phi)$ as a function of $\phi$. For the initial values, we have used the numerical results of Ref. [24], and the Higgs mass is fixed at

$$
\begin{equation*}
M_{h}=125.15 \mathrm{GeV} \tag{8}
\end{equation*}
$$

We use Eq. (8) as a typical value in the following discussion. The band corresponds to $95 \%$ CL deviation of the top quark pole mass $M_{t}$. For the $1 \sigma$ level, this is given by [25]

$$
\begin{equation*}
M_{t}=171.2 \pm 2.4 \mathrm{GeV} \tag{9}
\end{equation*}
$$

If we assume that all the other parameters of the SM except for $M_{t}$ are fixed, we can find the scale $\Lambda_{\beta}$ where $\beta_{\lambda_{\text {eff }}}$ becomes zero as a function of $M_{t}$. Here, $\beta_{\lambda_{\text {eff }}}$ means

$$
\begin{equation*}
\beta_{\lambda_{\mathrm{eff}}}(\phi):=\frac{d \lambda_{\mathrm{eff}}(\phi)}{d \log \phi} . \tag{10}
\end{equation*}
$$

The right panel of Fig. 1 shows $\Lambda_{\beta}$ as a function of $M_{t}$. The MPP requires that $\lambda_{\text {eff }}\left(\Lambda_{\beta}\right)$ should become zero, and predicts

$$
\begin{equation*}
M_{t}=170.9 \mathrm{GeV} \tag{11}
\end{equation*}
$$

This is the MPP of the pure SM. In the next section, we discuss the MPP of the SM with the scalar singlet DM and three right-handed neutrinos.

## 3. MPP of the SM with scalar singlet dark matter and right-handed neutrinos

We consider the following renormalizable Lagrangian:

$$
\begin{align*}
\mathcal{L}= & \mathcal{L}_{\mathrm{SM}}+\frac{1}{2} \partial_{\mu} S \partial^{\mu} S-\frac{m_{\mathrm{DM}}^{2}}{2} S^{2}-\frac{\kappa}{2} S^{2} H^{\dagger} H-\frac{\lambda_{\mathrm{DM}}}{4!} S^{4}+\sum_{j=1}^{3} \bar{\nu}_{\mathrm{R} j} i \gamma^{\mu} \partial_{\mu} \nu_{\mathrm{R} j} \\
& -\sum_{i, j}\left(y_{v i j} \bar{L}_{i} H^{\dagger} \nu_{\mathrm{R} j}+\mathrm{h} . \mathrm{c}\right)-\sum_{i, j}\left(M_{\mathrm{R} i j}+\frac{Y_{\mathrm{R} i j}}{\sqrt{2}} S\right) \bar{v}_{\mathrm{R} i}^{c} \nu_{\mathrm{R} j} . \tag{12}
\end{align*}
$$

Here, $H$ is the Higgs field, $S$ is the SM singlet real scalar field, $m_{\mathrm{DM}}$ is its mass, $\nu_{\mathrm{R} i}$ are right-handed neutrinos, $M_{\mathrm{R} i j}$ are their Majorana masses, and ( $Y_{\mathrm{R} i j}, y_{v i j}$ ) are the Yukawa couplings. For simplicity,
we assume that $M_{\mathrm{R} i j}, Y_{\mathrm{R} i j}$, and $y_{v i j}$ are diagonalized, and also that they are respectively equal for the three generations. In this case, Eq. (12) becomes

$$
\begin{align*}
\mathcal{L}= & \mathcal{L}_{\mathrm{SM}}+\frac{1}{2} \partial_{\mu} S \partial^{\mu} S-\frac{m_{\mathrm{DM}}^{2}}{2} S^{2}-\frac{\kappa}{2} S^{2} H^{\dagger} H-\frac{\lambda_{\mathrm{DM}}}{4!} S^{4}+\sum_{i=1}^{3} \bar{\nu}_{\mathrm{R} i} i \gamma^{\mu} \partial_{\mu} \nu_{\mathrm{R} i} \\
& -y_{v} \sum_{i=1}^{3}\left(\bar{L}_{i} H^{\dagger} \nu_{\mathrm{R} i}+\text { h.c }\right)-\sum_{i=1}^{3}\left(M_{\mathrm{R}}+\frac{Y_{\mathrm{R}}}{\sqrt{2}} S\right) \bar{v}_{\mathrm{R} i}^{c} \nu_{\mathrm{R} i} . \tag{13}
\end{align*}
$$

Thus, including the top mass $M_{t}$, there are seven unknown parameters,

$$
\begin{equation*}
M_{t}, \quad m_{\mathrm{DM}}, \quad \kappa, \quad \lambda_{\mathrm{DM}}, \quad y_{v}, \quad M_{\mathrm{R}}, \quad Y_{\mathrm{R}} \tag{14}
\end{equation*}
$$

in this model. In the following discussion, to distinguish the initial values of these parameters at $\mu=M_{t}$ from their running couplings, we use the subscript 0 for their initial values, like $\kappa_{0}$, except for $M_{t}$. Because $S$ is the candidate for the $\mathrm{DM}, m_{\mathrm{DM}}$ and $\kappa$ must satisfy some relation such that they can explain the observed energy density [27]:

$$
\begin{equation*}
\Omega_{\mathrm{DM}} h^{2}:=\frac{\rho_{\mathrm{DM}} h^{2}}{\rho_{\mathrm{tot}}}=0.1196 \pm 0.0031(68 \% \mathrm{CL}) \tag{15}
\end{equation*}
$$

For $m_{\mathrm{DM}} \gtrsim M_{h}$, this relation is approximately given by [28]

$$
\begin{equation*}
\log _{10} \kappa \simeq-3.63+1.04 \log _{10}\left(\frac{m_{\mathrm{DM}}}{\mathrm{GeV}}\right) \tag{16}
\end{equation*}
$$

Moreover, if we assume that the neutrino mass is $0.1 \mathrm{eV}, y_{\nu}$ and $M_{\mathrm{R}}$ must satisfy

$$
\begin{equation*}
-\frac{M_{\mathrm{R}}}{2}\left(1-\sqrt{1+\frac{2 y_{v}^{2} v_{h}^{2}}{M_{\mathrm{R}}^{2}}}\right) \simeq \frac{y_{v}^{2} v_{h}^{2}}{2 M_{\mathrm{R}}}=0.1 \mathrm{eV} \tag{17}
\end{equation*}
$$

where $v_{h}$ is the Higgs expectation value. This is the usual relation of the seesaw mechanism. In the following discussion, we choose $M_{\mathrm{R}}=10^{13} \mathrm{GeV}$, so $y_{v}$ is fixed by Eq. (17). As a result, four of the seven parameters remain free; they are

$$
\begin{equation*}
M_{t}, \quad \kappa, \quad \lambda_{\mathrm{DM}}, \quad \text { and } \quad Y_{\mathrm{R}} . \tag{18}
\end{equation*}
$$

To discuss how the effective couplings behave at the high-energy scale, we must know the renormalization group equations (RGEs) of this model. Their results are presented in Appendix A. Here, note that the contributions from the heavy right-handed neutrinos should be taken into account at the scale where $\mu \geq M_{\mathrm{R}}$. The one-loop effective potential of the Higgs field is given by

$$
V_{1 \text { loop }}(\phi, \mu):=\left\{\begin{array}{l}
V_{1 \text { loop }}^{\mathrm{SM}}(\phi)+\frac{M_{\mathrm{DM}}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{\mathrm{DM}}(\phi)^{2}}{\mu^{2}}\right)-\frac{3}{2}\right]-6 \cdot \frac{M_{v}^{-}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{v}^{-}(\phi)^{2}}{\mu^{2}}\right)-\frac{3}{2}\right]  \tag{19}\\
\left(\text { for } \phi<M_{\mathrm{R}}\right), \\
V_{1 \text { loop }}^{\mathrm{SM}}(\phi)+\frac{M_{\mathrm{DM}}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{\mathrm{DM}}(\phi)^{2}}{\mu^{2}}\right)-\frac{3}{2}\right]-6 \cdot \frac{M_{v}^{-}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{v}^{-}(\phi)^{2}}{\mu^{2}}\right)-\frac{3}{2}\right] \\
-6 \cdot \frac{M_{v}^{+}(\phi)^{4}}{64 \pi^{2}}\left[\log \left(\frac{M_{v}^{+}(\phi)^{2}}{\mu^{2}}\right)-\frac{3}{2}\right] \quad\left(\text { for } \phi>M_{\mathrm{R}}\right),
\end{array}\right.
$$

where

$$
\begin{equation*}
M_{\mathrm{DM}}(\phi):=\sqrt{e^{2 \Gamma(\phi)} \frac{\kappa \phi^{2}}{2}+m_{\mathrm{DM}}^{2}}, \quad M_{\nu}^{ \pm}(\phi):=\frac{M_{\mathrm{R}}}{2}\left(1 \pm \sqrt{1+\frac{2 y_{\nu}^{2} e^{2 \Gamma(\phi)} \phi^{2}}{M_{\mathrm{R}}^{2}}}\right) \tag{20}
\end{equation*}
$$



Fig. 2. The running effective Higgs self-coupling $\lambda_{\text {eff }}$ as a function of $\phi$. The upper left (right) panel shows the $M_{t}\left(\kappa_{0}\right)$ dependence. For $M_{t}$, the blue band corresponds to $95 \% \mathrm{CL}$ deviation from 171.2 GeV . The lower left (right) panel shows the $\lambda_{\mathrm{DM} 0}\left(Y_{\mathrm{R} 0}\right)$ dependence.

In these expressions, we have put $S=0$ because we now focus on the MPP of the Higgs sector. ${ }^{3}$ Furthermore, we can neglect $m_{\text {DM }}$ in Eq. (20) because its effect is very small when $\phi \gg m_{\text {DM }}$. As in Sect. 2, we set $\phi=\mu$, and define the effective Higgs self-coupling $\lambda_{\text {eff }}$ as

$$
\begin{equation*}
\lambda_{\text {eff }}(\phi):=\frac{4}{\phi^{4}} V(\phi)=\frac{4}{\phi^{4}}\left(V_{\text {tree }}(\phi)+V_{1 \text { loop }}(\phi)\right) . \tag{21}
\end{equation*}
$$

Figure 2 shows $\lambda_{\text {eff }}(\phi)$ for various values of the parameters. Here, the typical values are chosen to be

$$
\begin{equation*}
\lambda_{\mathrm{DM} 0}=0.2, \quad \kappa_{0}=0.2, \quad Y_{\mathrm{R} 0}=0.2 . \tag{22}
\end{equation*}
$$

One can see that $\lambda_{\text {eff }}$ depends mainly on $M_{t}$ and $\kappa_{0}$, and hardly on $\lambda_{\text {DM } 0}$ and $Y_{\mathrm{R} 0}$. This is because $\lambda_{\mathrm{DM}}$ does not appear in $\beta_{\lambda}$ and $Y_{\mathrm{R}}$ appears at the two-loop level (see Eq. (A7) in Appendix A). Therefore, by fixing $\lambda_{\mathrm{DM}}$ and $Y_{\mathrm{R}}$, we can relate $M_{t}$ and $\kappa_{0}$ from the MPP.
By the same procedure of Sect. 2, we can calculate the scale $\Lambda_{\beta}$ where $\beta_{\lambda_{\text {eff }}}$ becomes zero, and obtain $\lambda_{\text {eff }}\left(\Lambda_{\beta}\right)$ as a function of $M_{t}$ and $\kappa_{0}$. Figure 3 shows the results. In the upper (lower) panels, $Y_{\mathrm{R} 0}$ is fixed to $0.2(0.7)$. The difference between the left and right panels is whether the tree- or one-loop-level potential is used. The parameter regions where $\lambda_{\text {eff }}\left(\Lambda_{\beta}\right)<0$ and $\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)<0$ are filled, respectively, with blue and red. Both of them are excluded from the stability of the potentials. The MPP predicts that $M_{t}$ and $\kappa_{0}$ should exist on the green contour. One of the good points of this

[^1]

Fig. 3. The parameter dependences of $\lambda_{\text {eff }}\left(\Lambda_{\beta}\right)$. Here, $\lambda_{\mathrm{DM} 0}$ is fixed at 0.2 , and $Y_{\mathrm{R} 0}$ is fixed at 0.2 (0.7) in the upper (lower) panels. The left (right) panels show the calculations by using the tree- (one-loop-) level potential. The green lines show the prediction by the MPP. The contours that represent $\Lambda_{\beta}=10^{16} \mathrm{GeV}, 10^{17} \mathrm{GeV}$, and $10^{18} \mathrm{GeV}$ are also shown by red, blue, and orange, respectively.
model is that a larger $M_{t}$ is allowed, unlike in the SM. This is consistent with the recent experimental value [26],

$$
\begin{equation*}
M_{t}=173.34 \pm 0.76 \mathrm{GeV} \tag{23}
\end{equation*}
$$

which corresponds to the DM mass (see Eq. (16)),

$$
\begin{equation*}
769 \mathrm{GeV} \leq m_{\mathrm{DM}} \leq 1053 \mathrm{GeV} \tag{24}
\end{equation*}
$$

Two comments are needed.

1. The contours that represent $\Lambda_{\beta}=10^{16} \mathrm{GeV}, 10^{17} \mathrm{GeV}$, and $10^{18} \mathrm{GeV}$ are also shown in Fig. 3 by red, blue, and orange, respectively. Thus, a larger $M_{t}$ (such as Eq. (23)) means that, in this model, the MPP of the Higgs potential occurs at the relatively low-energy scale ( $\lesssim 10^{16} \mathrm{GeV}$ ).
2. As is seen from the lower panels of Fig. 3, we can also require $\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)=0$ in addition to $\lambda_{\text {eff }}\left(\lambda_{\beta}\right)=0$. Because $\kappa_{0}$ and $Y_{\mathrm{R} 0}$ appear in the one-loop part of $\beta_{\lambda_{\mathrm{DM}}}$, we can obtain a further relation between them by $\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)=0$. Although one might think that the one remaining parameter can be determined by $\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)=0$, we have checked that it is difficult to satisfy $\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)=\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)=0$ simultaneously. See Appendix B for more details.

## 4. Summary

We have discussed the MPP of the SM with scalar singlet DM and right-handed neutrinos. We have found that $\lambda_{\text {eff }}$ and $\beta_{\lambda_{\text {eff }}}$ can simultaneously become zero within a reasonable parameter region. The MPP predicts a strong relation between the portal coupling $\kappa$ and the top mass $M_{t}$. Unlike the pure SM, a larger $M_{t}$ is allowed in this model, which is favorable for the recent experimental values [25,26]:

$$
\begin{equation*}
M_{t}=173.34 \pm 0.76 \mathrm{GeV} \tag{25}
\end{equation*}
$$

Although we have found that the MPP can be satisfied for the Higgs potential, it is difficult to realize the exact flatness of the scalar potential at some high-energy scale $\Lambda$ :

$$
\begin{equation*}
\lambda(\Lambda)=\beta_{\lambda}(\Lambda)=\lambda_{\mathrm{DM}}(\Lambda)=\beta_{\lambda_{\mathrm{DM}}}(\Lambda)=\kappa(\Lambda)=\beta_{\kappa}(\Lambda)=0 ; \tag{26}
\end{equation*}
$$

see Appendix B for details. It would be interesting to consider a generalization of this model in such a way that the MPP can be realized for whole scalar fields.

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## Appendix A. Two-loop renormalization group equations

The two-loop RGEs where the Lagrangian is given by Eq. (13) are as follows: ${ }^{4}$

$$
\begin{align*}
\frac{d g_{Y}}{d t} & =\frac{1}{(4 \pi)^{2}} \frac{41}{6} g_{Y}^{3}+\frac{g_{Y}^{3}}{(4 \pi)^{4}}\left(\frac{199}{18} g_{Y}^{2}+\frac{9}{2} g_{2}^{2}+\frac{44}{3} g_{3}^{2}-\frac{17}{6} y_{t}^{2}-\frac{3}{2} y_{v}^{2}\right),  \tag{A1}\\
\frac{d g_{2}}{d t} & =-\frac{1}{(4 \pi)^{2}} \frac{19}{6} g_{2}^{3}+\frac{g_{2}^{3}}{(4 \pi)^{4}}\left(\frac{3}{2} g_{Y}^{2}+\frac{35}{6} g_{2}^{2}+12 g_{3}^{2}-\frac{3}{2}\left(y_{t}^{2}+y_{v}^{2}\right)\right),  \tag{A2}\\
\frac{d g_{3}}{d t} & =-\frac{7}{(4 \pi)^{2}} g_{3}^{3}+\frac{g_{3}^{3}}{(4 \pi)^{4}}\left(\frac{11}{6} g_{Y}^{2}+\frac{9}{2} g_{2}^{2}-26 g_{3}^{2}-2 y_{t}^{2}\right), \tag{A3}
\end{align*}
$$

[^2]\[

$$
\begin{align*}
\frac{d y_{t}}{d t}= & \frac{y_{t}}{(4 \pi)^{2}}\left(\frac{9}{2} y_{t}^{2}+3 y_{v}^{2}-\frac{17}{12} g_{Y}^{2}-\frac{9}{4} g_{2}^{2}-8 g_{3}^{2}\right) \\
& +\frac{y_{t}}{(4 \pi)^{4}}\left\{-12 y_{t}^{4}-\frac{27}{4} y_{v}^{4}-\frac{27}{4} y_{t}^{2} y_{v}^{2}-\frac{9}{8} Y_{\mathrm{R}}^{2} y_{v}^{2}+6 \lambda^{2}+\frac{1}{4} \kappa^{2}-12 \lambda y_{t}^{2}\right. \\
& +g_{Y}^{2}\left(\frac{131}{16} y_{t}^{2}+\frac{15}{8} y_{v}^{2}\right)+g_{2}^{2}\left(\frac{225}{16} y_{t}^{2}+\frac{45}{8} y_{v}^{2}\right)+36 g_{3}^{2} y_{t}^{2}+\frac{1187}{216} g_{Y}^{4} \\
& \left.-\frac{23}{4} g_{2}^{4}-108 g_{3}^{4}-\frac{3}{4} g_{Y}^{2} g_{2}^{2}+9 g_{2}^{2} g_{3}^{2}+\frac{19}{9} g_{3}^{2} g_{Y}^{2}\right\},  \tag{A4}\\
\frac{d y_{v}}{d t}= & \frac{y_{v}}{(4 \pi)^{2}}\left(\frac{9}{2} y_{v}^{2}+3 y_{t}^{2}+\frac{1}{4} Y_{\mathrm{R}}^{2}-\frac{3}{4} g_{Y}^{2}-\frac{9}{4} g_{2}^{2}\right) \\
& +\frac{y_{v}}{(4 \pi)^{4}}\left\{-12 y_{v}^{4}-\frac{27}{4} y_{t}^{4}-\frac{19}{32} Y_{\mathrm{R}}^{4}-y_{v}^{2}\left(\frac{27}{4} y_{t}^{2}+\frac{21}{16} Y_{\mathrm{R}}^{2}\right)+6 \lambda^{2}+\frac{1}{4} \kappa^{2}\right. \\
& -12 \lambda y_{v}^{2}-\kappa Y_{\mathrm{R}}^{2}+g_{Y}^{2}\left(\frac{123}{16} y_{v}^{2}+\frac{85}{24} y_{t}^{2}+\frac{9}{16} Y_{\mathrm{R}}^{2}\right)+g_{2}^{2}\left(\frac{225}{16} y_{v}^{2}+\frac{45}{8} y_{t}^{2}+\frac{27}{16} Y_{\mathrm{R}}^{2}\right) \\
& \left.+20 g_{3}^{2} y_{t}^{2}+\frac{35}{24} g_{Y}^{4}-\frac{23}{4} g_{2}^{4}-\frac{9}{4} g_{Y}^{2} g_{2}^{2}\right\},  \tag{A5}\\
\frac{d Y_{\mathrm{R}}}{d t}= & \frac{Y_{\mathrm{R}}}{(4 \pi)^{2}}\left(3 Y_{\mathrm{R}}^{2}+2 y_{v}^{2}\right)+\frac{Y_{\mathrm{R}}}{(4 \pi)^{4}}\left\{-\frac{81}{16} Y_{\mathrm{R}}^{4}-\frac{27}{4} Y_{\mathrm{R}}^{2} y_{v}^{2}-9 y_{t}^{2} y_{v}^{2}-\frac{27}{2} y_{v}^{4}\right. \\
& \left.-\lambda \mathrm{DM} Y_{\mathrm{R}}^{2}-8 \kappa y_{v}^{2}-\frac{1}{4} g_{Y}^{2} y_{v}^{2}-\frac{3}{4} g_{2}^{2} y_{v}^{2}\right\},  \tag{A6}\\
\frac{d \lambda}{d t}= & \frac{1}{(4 \pi)^{2}}\left(\lambda\left(24 \lambda-9 g_{2}^{2}-3 g_{Y}^{2}+12 y_{v}^{2}+12 y_{t}^{2}\right)+\frac{3}{4} g_{2}^{2} g_{Y}^{2}+\frac{9 g_{2}^{4}}{8}\right. \\
& \left.+\frac{3 g_{Y}^{4}}{8}+\frac{\kappa^{2}}{2}-6 y_{v}^{4}-6 y_{t}^{4}\right)+\frac{1}{(4 \pi)^{4}}\left\{-2 \kappa^{3}-5 \kappa^{2} \lambda-312 \lambda^{3}+36 \lambda^{2}\left(g_{Y}^{2}+3 g_{2}^{2}\right)\right. \\
& +\lambda\left(\frac{629}{24} g_{Y}^{4}+\frac{39}{4} g_{2}^{2} g_{Y}^{2}-\frac{73}{8} g_{2}^{4}\right)+\frac{305}{16} g_{2}^{6}-\frac{289}{48} g_{Y}^{2} g_{2}^{4}-\frac{559}{48} g_{Y}^{4} g_{2}^{2}+\frac{379}{48} g_{Y}^{6} \\
& -32 g_{3}^{2} y_{t}^{4}-\frac{8}{3} g_{Y}^{2} y_{t}^{4}-\frac{9}{4} g_{2}^{4}\left(y_{t}^{2}+y_{v}^{2}\right)+\lambda y_{t}^{2}\left(\frac{85}{6} g_{Y}^{2}+\frac{45}{2} g_{2}^{2}+80 g_{3}^{2}\right) \\
& -144 y_{v}^{2}\left(\frac{15}{2} g_{Y}^{2}+\frac{45}{2} g_{2}^{2}\right)+g_{Y}^{2} y_{t}^{2}\left(-\frac{19}{4} g_{Y}^{2}+\frac{21}{2} g_{2}^{2}\right)-g_{V}^{2} y_{v}^{2}\left(\frac{3}{4} g_{Y}^{2}+\frac{3}{2} g_{2}^{2}\right) \\
& \left.3 \lambda\left(y_{t}^{4}+y_{v}^{4}+\frac{3}{2} Y_{\mathrm{R}}^{2} y_{v}^{2}\right)+30\left(y_{t}^{6}+y_{v}^{6}+\frac{1}{5} Y_{\mathrm{R}}^{2} y_{v}^{4}\right)-\frac{3}{2} Y_{\mathrm{R}}^{2} \kappa^{2}\right\} \tag{A7}
\end{align*}
$$,
\]

$$
\begin{align*}
\frac{d \lambda_{\mathrm{DM}}}{d t}= & \frac{1}{(4 \pi)^{2}}\left(3 \lambda_{\mathrm{DM}}^{2}+12 \kappa^{2}+6 \lambda_{\mathrm{DM}} Y_{\mathrm{R}}^{2}-18 Y_{\mathrm{R}}^{4}\right) \\
& +\frac{1}{(4 \pi)^{4}}\left\{-\frac{17}{3} \lambda_{\mathrm{DM}}^{3}-20 \kappa^{2} \lambda_{\mathrm{DM}}-48 \kappa^{3}-72\left(y_{t}^{2}+y_{v}^{2}\right) \kappa^{2}+24\left(g_{Y}^{2}+3 g_{2}^{2}\right) \kappa^{2}\right. \\
& \left.+72 Y_{\mathrm{R}}^{4}\left(Y_{\mathrm{R}}^{2}+y_{v}^{2}\right)+\lambda_{\mathrm{DM}} Y_{\mathrm{R}}^{2}\left(\frac{21}{2} Y_{\mathrm{R}}^{2}-18 y_{v}^{2}\right)-9 Y_{\mathrm{R}}^{2} \lambda_{\mathrm{DM}}^{2}\right\}, \tag{A8}
\end{align*}
$$

$$
\begin{align*}
\frac{d \kappa}{d t}= & \frac{1}{(4 \pi)^{2}}\left(4 \kappa^{2}+12 \kappa \lambda+\kappa \lambda_{\mathrm{DM}}+3 \kappa\left(2 y_{t}^{2}+2 y_{v}^{2}+Y_{\mathrm{R}}^{2}\right)-\frac{3}{2} \kappa\left(g_{Y}^{2}+3 g_{2}^{2}\right)-12 Y_{\mathrm{R}}^{2} y_{v}^{2}\right) \\
& +\frac{\kappa}{(4 \pi)^{4}}\left\{-\frac{21}{2} \kappa^{2}-72 \kappa \lambda-60 \lambda^{2}-6 \kappa \lambda_{\mathrm{DM}}-\frac{5}{6} \lambda_{\mathrm{DM}}^{2}-\left(y_{t}^{2}+y_{v}^{2}\right)(12 \kappa+72 \lambda)\right. \\
& -3 Y_{\mathrm{R}}^{2}\left(2 \kappa+\lambda_{\mathrm{DM}}\right)-\frac{27}{2} y_{t}^{4}-\frac{27}{2} y_{v}^{4}-\frac{3}{4} Y_{\mathrm{R}}^{4}+\frac{51}{4} Y_{\mathrm{R}}^{2} y_{v}^{2}+g_{Y}^{2}(\kappa+24 \lambda)+3 g_{2}^{2}(\kappa+24 \lambda) \\
& \left.+y_{t}^{2}\left(\frac{85}{12} g_{Y}^{2}+\frac{45}{4} g_{2}^{2}+40 g_{3}^{2}\right)+y_{v}^{2}\left(\frac{15}{4} g_{Y}^{2}+\frac{45}{4} g_{2}^{2}\right)+\frac{557}{48} g_{Y}^{4}-\frac{145}{16} g_{2}^{4}+\frac{15}{8} g_{Y}^{2} g_{2}^{2}\right\} \\
& +\frac{Y_{\mathrm{R}}^{2} y_{v}^{2}}{(4 \pi)^{4}}\left\{\frac{3}{2}\left(g_{Y}^{2}+3 g_{2}^{2}\right)+27 Y_{\mathrm{R}}^{2}+66 y_{v}^{2}\right\}  \tag{A9}\\
\frac{d \Gamma}{d t}= & \frac{1}{(4 \pi)^{2}}\left(\frac{9}{4} g_{2}^{2}+\frac{3}{4} g_{Y}^{2}-3 y_{t}^{2}-3 y_{v}^{2}\right) . \tag{A10}
\end{align*}
$$

## Appendix B. Is an exact flat potential possible?

One question is whether the MPP can be realized exactly, namely, if

$$
\begin{equation*}
\lambda\left(\Lambda_{\beta}\right)=\beta_{\lambda}\left(\Lambda_{\beta}\right)=\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)=\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)=\kappa\left(\Lambda_{\beta}\right)=\beta_{\kappa}\left(\Lambda_{\beta}\right)=0 \tag{B1}
\end{equation*}
$$

is possible or not. Here, for simplicity, we also define $\Lambda_{\beta}$ as the scale where $\beta_{\lambda}$ becomes zero. To discuss this possibility, it is qualitatively enough to consider one-loop RGEs. One can easily understand that it is impossible to realize Eq. (B1) as follows: even if $\lambda\left(\Lambda_{\beta}\right), \beta_{\lambda}\left(\Lambda_{\beta}\right), \lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)$, and $\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)$ become simultaneously zero, we cannot make $\kappa\left(\Lambda_{\beta}\right)$ zero because the one-loop part of $\beta_{\lambda_{\mathrm{DM}}}$ at $\Lambda_{\beta}$ becomes

$$
\begin{equation*}
\left.\beta_{\lambda_{\mathrm{DM}}}\right|_{\Lambda_{\beta}}=\frac{1}{(4 \pi)^{2}}\left(12 \kappa^{2}-18 Y_{\mathrm{R}}^{4}\right), \tag{B2}
\end{equation*}
$$




Fig. B1. The blue (red) lines show the contours where $\beta_{\lambda_{\mathrm{DM}}}\left(\lambda_{\mathrm{DM}}\right)\left(\Lambda_{\beta}\right)=0$. The left (right) panel is the $M_{t}=170(176) \mathrm{GeV}$ case. Here, note that, if $\kappa_{0} \gtrsim 0.3, \Lambda_{\beta}$ becomes less than $M_{\mathrm{R}}=10^{13} \mathrm{GeV}$, and there is no solution of $\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)=0$ because the one-loop part of $\beta_{\lambda_{\mathrm{DM}}}$ is always positive when $\mu \leq M_{\mathrm{R}}$.
and we need $\kappa\left(\Lambda_{\beta}\right) \neq 0$ to satisfy $\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)=0 .{ }^{5}$ Furthermore, it is also difficult even to satisfy $\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)=\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)=0$ simultaneously; see Fig. B1. This shows the contours such that $\lambda_{\mathrm{DM}}\left(\Lambda_{\beta}\right)$ and $\beta_{\lambda_{\mathrm{DM}}}\left(\Lambda_{\beta}\right)$ become zero, respectively. Here, we have used two-loop RGEs. One can see that the two contours do not intersect.

## References

[1] G. Aad et al. [ATLAS Collaboration], Phys. Lett. B 716, 1 (2012) [arXiv:1207.7214 [hep-ex]] [Search InSPIRE].
[2] S. Chatrchyan et al. [CMS Collaboration], Phys. Lett. B 716, 30 (2012) [arXiv:1207.7235 [hep-ex]] [Search inSPIRE].
[3] C. D. Froggatt and H. B. Nielsen, Phys. Lett. B 368, 96 (1996) [arXiv:9511371 [hep-ph]] [Search InSPIRE].
[4] C. D. Froggatt, H. B. Nielsen, and Y. Takanishi, Phys. Rev. D 64, 113014 (2001) [arXiv:0104161 [hep-ph]] [Search InSPIRE].
[5] H. B. Nielsen, [arXiv:1212.5716 [hep-ph]] [Search InSPIRE].
[6] M. Shaposhnikov and C. Wetterich, Phys. Lett. B 683, 196 (2010) [arXiv:0912.0208 [hep-th]] [Search inSPIRE].
[7] K. A. Meissner and H. Nicolai, Phys. Lett. B 660, 260 (2008) [arXiv:0710.2840 [hep-th]] [Search inSPIRE].
[8] V. V. Khoze, C. McCabe, and G. Ro, J. High Energy Phys. 1408, 026 (2014) [arXiv:1403.4953 [hep-ph]] [Search inSPIRE].
[9] H. Kawai and T. Okada, Prog. Theor. Phys. 127, 689 (2012) [arXiv:1110.2303 [hep-th]] [Search InSPIRE].
[10] H. Kawai, Int. J. Mod. Phys. A 28, 1340001 (2013).
[11] Y. Hamada, H. Kawai, and K. Kawana, Int. J. Mod. Phys. A 29, 1450099 (2014) [arXiv:1405.1310 [hep-ph]] [Search InSPIRE].
[12] Y. Hamada, H. Kawai, and K. Kawana, [arXiv:1409.6508 [hep-ph]] [Search inSPIRE].
[13] Y. Hamada, H. Kawai, and K. y. Oda, Prog. Theor. Exp. Phys. 2014, 023B02 (2014) [arXiv:1308.6651 [hep-ph]] [Search inSPIRE].
[14] Y. Hamada, H. Kawai, K. y. Oda, and S. C. Park, Phys. Rev. Lett. 112, 241301 (2014) [arXiv:1403.5043 [hep-ph]] [Search inSPIRE].
[15] Y. Hamada, H. Kawai, K. y. Oda, and S. C. Park, [arXiv:1408.4864 [hep-ph]] [Search inSPIRE].
[16] Y. Hamada, K. y. Oda, and F. Takahashi, [arXiv:1408.5556 [hep-ph]] [Search InSPIRE].
[17] Y. Kawamura, Prog. Theor. Exp. Phys. 2013, 113B04 (2013) [arXiv:1308.5069 [hep-ph]] [Search InSPIRE].
[18] K. A. Meissner and H. Nicolai, Phys. Lett. B 648, 312 (2007) [arXiv:0612165 [hep-th]] [Search InSPIRE].
[19] N. Haba, H. Ishida, K. Kaneta, and R. Takahashi, Phys. Rev. D 90, 036006 (2014) [arXiv:1406.0158 [hep-ph]] [Search InSPIRE].
[20] S. Iso and Y. Orikasa, Prog. Theor. Exp. Phys. 2013, 023B08 (2013) [arXiv:1210.2848 [hep-ph]] [Search InSPIRE].
[21] Y. Hamada, H. Kawai, and K. y. Oda, J. High Energy Phys. 1407, 026 (2014)
[arXiv:1404.6141 [hep-ph]] [Search inSPIRE].
[22] N. Haba and R. Takahashi, Phys. Rev. D 89, 115009 (2014) [arXiv:1404.4737 [hep-ph]] [Search inSPIRE].
[23] N. Haba, H. Ishida, and R. Takahashi, [arXiv:1405.5738 [hep-ph]] [Search inSPIRE].
[24] D. Buttazzo, G. Degrassi, P. P. Giardino, G. F. Giudice, F. Sala, A. Salvio, and A. Strumia, J. High Energy Phys. 1312, 089 (2013) [arXiv:1307.3536 [hep-ph]] [Search INSPIRE].
[25] S. Moch et al., [arXiv:1405.4781 [hep-ph]] [Search inSPIRE].
[26] ATLAS, CDF, CMS, and D0 Collaborations, [arXiv:1403.4427 [hep-ex]] [Search inSPIRE].

[^3][27] P. A. R. Ade et al. [Planck Collaboration], Astron. Astrophys. (2014) [arXiv:1303.5076 [astro-ph.CO]] [Search inSPIRE].
[28] J. M. Cline, K. Kainulainen, P. Scott, and C. Weniger, Phys. Rev. D 88, 055025 (2013) [arXiv:1306.4710 [hep-ph]] [Search inSPIRE].
[29] Y. Hamada, H. Kawai, and K. y. Oda, Phys. Rev. D 87, 053009 (2013) [arXiv:1210.2538 [hep-ph]] [Search inSPIRE].
[30] M. E. Machacek and M. T. Vaughn, Nucl. Phys. B 222, 83 (1983).
[31] M. E. Machacek and M. T. Vaughn, Nucl. Phys. B 236, 221 (1984).
[32] M. E. Machacek and M. T. Vaughn, Nucl. Phys. B 249, 70 (1985).


[^0]:    ${ }^{1}$ It is difficult to realize the MPP of the other scalar couplings simultaneously in addition to $\lambda$. This is discussed in Appendix B.
    ${ }^{2}$ For the beta functions of the SM, see, e.g., Refs. [21,24,29]. Alternatively, we can reproduce them by using the results in Appendix A.

[^1]:    ${ }^{3}$ Of course, we can consider the MPP of the DM sector. We study such a situation in Appendix B.

[^2]:    ${ }^{4}$ The calculations in this appendix are based on Refs. [30-32], and our results are in agreement with the recent result [23] when there is only one right-handed neutrino and $Y_{\mathrm{R}}=0$.

[^3]:    ${ }^{5}$ Typically, a non-zero positive $Y_{\mathrm{R} 0}$ is needed to make $\beta_{\lambda_{\mathrm{DM}}}$ negative. As a result, $Y_{\mathrm{R}}\left(\Lambda_{\beta}\right)$ is non-zero because the one-loop part of $\beta_{Y_{\mathrm{R}}}$ is always positive.

