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1 **Mineral inclusions in diamonds from Karowe Mine, Botswana:**
2 **super-deep sources for super-sized diamonds?**

3
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14
15 **Abstract**

16 Mineral inclusions in diamonds play a critical role in constraining the relationship between diamonds
17 and mantle lithologies. Here we report the first major and trace element study of mineral inclusions in
18 diamonds from the Karowe Mine in north-east Botswana, along the western edge of the Zimbabwe
19 Craton. From a total of 107 diamonds, 134 silicate, 15 oxide, and 22 sulphide inclusions were
20 recovered. The results reveal that 53 % of Karowe inclusion-bearing diamonds derived from eclogitic
21 sources, 44 % are peridotitic, 2 % have a sublithospheric origin, and 1 % are websteritic. The
22 dominant eclogitic diamond substrates sampled at Karowe are compositionally heterogeneous, as
23 reflected in wide ranges in the CaO contents (4-16 wt%) of garnets and the Mg# (69-92) and jadeite
24 contents (14-48 mol%) of clinopyroxenes. Calculated bulk rock REE_N patterns indicate that both
25 shallow and deep levels of the subducted slab(s) were sampled, including cumulate-like protoliths.
26 Peridotitic garnet compositions largely derive from harzburgite/dunite substrates (~90 %), with almost

27 half the garnets having CaO contents <1.8 wt %, consistent with pyroxene-free (dunitic) sources. The
28 highly depleted character of the peridotitic diamond substrates is further documented by the high mean
29 and median Mg# (93.1) of olivine inclusions. One low-Ca garnet records a very high Cr₂O₃ content
30 (14.7 wt%), implying that highly depleted cratonic lithosphere at the time of diamond formation
31 extended to at least 220 km depth. Inclusion geothermobarometry indicates that the formation of
32 peridotitic diamonds occurred along a 39-40 mW/m² model geotherm. A sublithospheric inclusion
33 suite is established by three eclogitic garnets containing a majorite component, a feature so far unique
34 within the Orapa cluster. These low- and high-Ca majoritic garnets follow pyroxenitic and eclogitic
35 trends of majoritic substitution, respectively. The origin of the majorite-bearing diamonds is estimated
36 to be between 330 to 420 km depth, straddling the asthenosphere–transition zone boundary. This new
37 observation of superdeep mineral inclusions in Karowe diamonds is consistent with a sublithospheric
38 origin for the exceptionally large diamonds from this mine.

39

40 Key words:

41 Zimbabwe Craton

42 Orapa kimberlite cluster

43 Sublithospheric

44 Majorite

45

46

47 **Introduction**

48

49 The recent recovery of exceptionally large diamonds (including the 1,109 carats Lesedi la
50 Rona) at the Karowe mine in north-eastern Botswana raises the question: what controls such
51 an exceptionally coarse size-frequency distribution? To address this question we studied the
52 mineralogy and chemistry of inclusions in a representative suite of diamonds from this mine.

53 Based on recent conflicting views on the origin of very large (typically nitrogen-free) gem
54 diamonds, invoking either a megacryst-like origin within the lithosphere (Moore 2014) or
55 crystallization in the sublithospheric mantle (Smith et al. 2016), particular emphasis was
56 placed on evaluating the presence or absence of inclusions indicative of a superdeep origin. A
57 second goal of the study was to employ lithospheric inclusions in diamonds from Karowe to
58 establish the compositional characteristics, the thermal conditions, and the thickness of the
59 lithospheric mantle at the time of diamond formation.

60

61 **Geological framework**

62

63 The geology of Botswana is mainly defined by the domains of the Kalahari Super-Craton,
64 which include the Archean basement of the Kaapvaal and Zimbabwe cratons sutured by the
65 Archean to Paleoproterozoic Limpopo Mobile Belt (Fig. 1). In the west, the Kalahari Super-
66 Craton is bounded by the Paleoproterozoic Kheis and Magondi orogenic belts (Fig. 1). The
67 Kaapvaal Craton nucleated first and stabilised between 3.7 Ga to 2.6 Ga (De Wit et al. 1992).
68 The formation of the core of the Zimbabwe Craton and its crustal growth occurred between
69 3.5 Ga to 2.6 Ga (Horstwood et al. 1999; Rollinson and Whitehouse 2011). The collision of
70 the Kaapvaal and Zimbabwe cratons at about 2.7 to 2.6 Ga resulted in the formation of the
71 Limpopo Mobile Belt (Van Reenen et al. 1987). The western Kheis and Magondi mobile belts
72 range between 1.8 to 2.0 Ga in age (Majaule et al. 2001; Treloar 1988).

73 The Orapa, Letlhakane, Damtshaa, and Karowe mines form part of the Orapa
74 kimberlite cluster, located in north-east Botswana, along the western edge of the Zimbabwe
75 Craton (Fig. 1). The lithospheric mantle beneath this region has been studied for the past three
76 decades both through diamonds and their mineral inclusions and through mantle xenoliths
77 collected from the Orapa, Letlhakane, and Damtshaa kimberlites (Shee and Gurney 1979;
78 Gurney et al. 1984). Inclusion studies revealed that the diamond sources beneath Orapa,

79 Letlhakane, and Damtshaa are compositionally distinct (Deines and Harris 2004; Deines et al.
80 2009). At Letlhakane and Damtshaa, peridotitic inclusions are more common compared to a
81 predominately eclogitic diamond association at Orapa. Peridotitic diamond substrates in the
82 lithospheric mantle beneath Orapa were investigated by Stachel et al. (2004b), who found that
83 inclusions in diamond reflect mild metasomatic overprint through CHO fluids compared to
84 the extensive metasomatic modification observed in peridotite xenoliths (Griffin et al. 2003).
85 The authors attributed this difference to peridotitic diamond formation that predated major
86 modification of the lithosphere during Proterozoic rifting and compression. Stiefenhofer et al.
87 (1997) conducted a detailed study on mantle xenoliths from Letlhakane, located
88 approximately 40km southeast of Orapa. With the Orapa and Letlhakane kimberlites having
89 been emplaced into the Proterozoic Magondi Belt (Fig. 1), this study, together with the Re-Os
90 dating presented by Carlson et al. (1999), has established that both Letlhakane and Orapa are
91 underlain by lithospheric mantle that is chemically depleted, cold, and old. This indicates that
92 in the Orapa area, the Magondi Belt is thrust over the western edge of the Zimbabwe Craton
93 (Stiefenhofer et al. 1997).

94 Outside the Orapa kimberlite cluster, information about the lithospheric mantle of the
95 Zimbabwe Craton is principally based on mantle xenoliths and inclusion-bearing diamonds
96 from the Cambrian kimberlites at Murowa and Sese, located in the southern part of the craton.
97 The present evidence indicates that the lithospheric mantle beneath the southern Zimbabwe
98 Craton is exceptionally depleted with diamond substrates of harzburgitic-dunitic paragenesis
99 (Smith et al. 2009). Similar strongly depleted harzburgitic diamond sources were also
100 established for the River Ranch kimberlite, located in the Central Zone of the Limpopo
101 Mobile Belt (Kopylova et al. 1997). The Zimbabwe Craton and parts of the Limpopo Mobile
102 Belt are underlain by a thick (~225 to 250 km) mantle root, with lithosphere thickness
103 increasing even further towards the south beneath the Kaapvaal Craton (~250 to 300 km;
104 Fouch et al. 2004).

105

106 **Samples and analytical techniques**

107

108 A total of 120 inclusion-bearing diamonds was selected from run-of-mine production at the
109 sorting office of Karowe Diamond Mine in Gaborone, Botswana. The collected diamonds
110 range in size from about 2 to 3 mm (the equivalent to -7+5, -9+7, and -11+9 Diamond
111 Trading Company sieve classes). The diamonds were visually inspected under a binocular
112 microscope and fully documented for shape, colour and surface features, before being
113 photographed. Subsequently, inclusions were extracted by crushing of the diamonds in a
114 purpose-built steel cracker, mounted individually in epoxy and polished to a 0.25 μm finish.
115 The inclusions measured between 60 μm and 500 μm in longest dimension. From the initial
116 120 diamonds, 171 mineral inclusions were successfully recovered and prepared from 107
117 diamonds. For the remaining 13 diamonds, inclusions were either not recovered or lost during
118 preparation.

119 Major and minor element compositions of mineral inclusions were analysed using a
120 JEOL 8900R electron probe micro-analyser (EPMA). All elements were measured with an
121 accelerating voltage of 20 kV, beam current of 20 nA and ≤ 2 μm beam diameter. Peak count
122 times were 15-30 seconds and combined background times were 30-60 seconds. Standards for
123 the two light major elements Si ($\text{K}\alpha$) and Mg ($\text{K}\alpha$) were pyrope garnet, diopside or forsterite
124 (Fo90.5), depending on the mineral analyzed. For the remaining elements, standards were
125 albite (Na- $\text{K}\alpha$), pyrope garnet (Al- $\text{K}\alpha$), apatite (P- $\text{K}\alpha$), sanidine (K- $\text{K}\alpha$), labradorite (Ca- $\text{K}\alpha$),
126 rutile (Ti- $\text{K}\alpha$), chromium oxide (Cr- $\text{K}\alpha$), spessartine (Mn- $\text{K}\alpha$), fayalite (Fe- $\text{K}\alpha$), nickel wire
127 (Ni- $\text{K}\alpha$). The standards are described in more detail in Czap et al. (this volume; their Table
128 S1), with exception of Fo90.5, which is from a spinel peridotite mantle xenoliths collected at
129 Harrat al Kishb, Saudi Arabia (McGuire et al. 1992). The CITZAF correction (Armstrong

130 1995) was utilized for data reduction. Three to five spots were analysed on each sample
131 depending on the size of the mineral grain. After assessing homogeneity, the oxide
132 concentrations measured were then averaged for individual grains. The resulting detection
133 limits typically are ≤ 0.02 wt% oxide.

134 Rare earth element (REE) and trace element concentrations for selected garnet and
135 clinopyroxene inclusions were measured using a Resonetics M-50 193 nm excimer laser
136 coupled with a Thermo Element XR 2 sector-field inductively coupled plasma mass
137 spectrometer (LA-ICP-MS). Grains were ablated with a spot size of 50 or 75 μm at a laser
138 frequency of 10 Hz and energy density of ~ 4 J/cm². For each sample two or three spots were
139 analyzed and an average of 100 sweeps through the mass spectrum were made for each
140 analysis. Measurement time comprised of 40 seconds background collection followed by 60
141 seconds sample ablation. The ICP-MS was operated at low mass resolution mode ($M/\Delta M = \text{ca.}$
142 300). The ThO/Th signal was monitored to ensure that oxide production remained below 0.5
143 %. Calibration of relative element sensitivities was performed using the NIST SRM 612 glass
144 reference material and ⁴³Ca was employed as an internal standard to normalize the signal
145 intensities. Data processing was performed offline using Iolite v3 (Paton et al. 2011).
146 Detection limits are ≤ 30 ppb for REE, Nb, Zr, Y, Sr, and ~ 1 ppm for Ti and Ni.

147

148

149 **Results**

150

151 Most of the 107 diamonds contained a single mineral inclusion, commonly olivine (n=23;
152 Table S1 in Electronic Supplementary Material A [ESM A]). Twenty-three diamonds hosted
153 two-phase assemblages and one diamond a three-phase assemblage (Table S1 in ESM A).
154 Inclusions were subdivided into the peridotitic (lherzolititic and harzburgitic), eclogitic, and

155 websteritic suites based on mineralogy and major element composition (Grütter et al. 2004;
156 Gurney et al. 1984; Meyer 1987.

157

158 **Major element compositions of inclusions and paragenetic associations**

159

160 **Eclogitic suite**

161

162 The 33 recovered eclogitic garnets can be divided into two groups on the basis of their CaO
163 contents (Grütter et al. 2004): low-Ca (CaO <6 wt%; n=11) and high-Ca (CaO ≥6 wt%;
164 n=22). As observed by Stachel and Harris (2008), there is a crude positive correlation
165 between Na and Ca content (Table S1 in Electronic Supplementary Material B [ESM B]),
166 both generally increasing from low- to high-Ca garnets and with the most Na₂O-rich garnet
167 (0.54 wt%, diamond KW79) also having the highest CaO concentration (16.1 wt%). A similar
168 crude positive correlation with Ca content exists for Ti, with the exception of unusually high
169 TiO₂ (1.7 wt%) being observed in two garnets with intermediate CaO content (10.1 wt%)
170 from diamond KW56. The two Ti-rich garnets show unusually low Al₂O₃ contents (average
171 19.9 wt%), with the typically six-fold coordinated cations Ti, Al, and Cr summing to 3.77
172 ([O]=24). As there is no excess of Si over the available tetrahedral sites in these garnets, this
173 apparent deficiency can only be compensated through the presence of ferric iron (Canil and
174 O'Neill 1996). In diamond KW36, a single garnet was associated with a deep blue kyanite
175 (Table S1 in ESM B) and contained 7.0 wt% CaO, which is too low for derivation from
176 typical grosspydrite (Sobolev et al. 1968). The three (low-Ca) garnets with the highest Mg#
177 [~70; molar 100×Mg/(Mg+Fe)] also have elevated Cr₂O₃ contents (ranging from 0.52-0.80
178 wt%). These characteristics indicate an affinity to the websteritic inclusion suite (Deines et al.

179 1993) but fall in a compositional range where a clear distinction between low-Ca eclogitic and
180 websteritic garnets is not possible (Stachel and Harris 2008).

181 The 40 analysed eclogitic clinopyroxene inclusions range in composition from
182 augites (n=16) to omphacites [n=24; jadeite component calculated as
183 $100 \times 2\text{Na} / (2\text{Na} + \text{Ca} + \text{Mg} + \text{Fe})$; Morimoto 1988]. The augites are unusual among eclogitic
184 clinopyroxene inclusions worldwide with Cr_2O_3 contents extending to high values (average
185 0.15 wt%, max 0.27 wt%), high NiO (average 0.10 wt%), high MgO (average 19.8 wt%), and
186 low jadeite contents (14.2-16.3 mol%) plus very low CaO (average 10.3, minimum 9.3 wt%).
187 Their extremely low molar Ca# [$100 \times \text{Ca} / (\text{Ca} + \text{Mg} + \text{Fe})$] of 23.3 (range 21.4-24.9) is unique
188 among eclogitic inclusions worldwide, with the exception of inclusions from Letlhakane in
189 the same kimberlite cluster (Deines and Harris 2004). For omphacites (25.8-48.2 mol%
190 jadeite) there is a crude negative correlation between Mg# and Ti content. The two
191 omphacites with the highest jadeite content (from diamond KW66) have an unusually high
192 Mg# (91.4-91.7). With increasing Al content, Karowe omphacites deviate from the near 1:1
193 correlation between Al and Na cation content shown by eclogitic clinopyroxene inclusions
194 worldwide towards excess Al (Fig. S1 in ESM A), suggestive of an additional, higher than
195 normal Tschermaks component (Stachel and Harris 2008). Seven omphacites coexist with
196 either low-Ca garnets (n=3 from one diamond) or high-Ca garnets (n=4 from two diamonds)
197 garnets. For this sample set, clinopyroxenes associated with low-Ca garnet, compared to those
198 occurring with high-Ca garnet, have a higher Mg# (average of 77.8 versus 71.7), Cr_2O_3
199 content (0.10 versus 0.04 wt%) and K_2O abundance (0.16 versus 0.06 wt%) at decreased CaO
200 contents (12.2 versus 14.7).

201 Four SiO_2 (99.5-100.4 wt%) inclusions were recovered from four diamonds and are
202 assumed to represent primary coesite. Of these, three co-exist with (1) a pair of low-Ca
203 eclogitic garnets, (2) an omphacitic clinopyroxene, and (3) a low-Ni sulphide (monosulphide
204 solid solution) inclusion, consistent with an eclogitic association. The fourth SiO_2 inclusion

205 coexists with olivine (KW32e). Coesite may occur in peridotitic assemblages as a
206 consequence of intense source carbonation (Wyllie and Huang 1976). As previously observed
207 for inclusions in diamonds from the Renard Mine (Superior Craton; Hunt et al. 2012), the
208 disequilibrium assemblage olivine + coesite requires continued diamond growth during
209 progressive source carbonation, with the olivine inclusion representing an earlier growth
210 stage. Two colourless and one deep blue kyanites inclusions were recovered. The latter
211 coexisted with an orange garnet (KW36a; Table S1 in ESM B) and has elevated levels of
212 TiO_2 (0.31 wt%), Cr_2O_3 (0.17 wt%), FeO (0.34 wt%), and MgO (0.18 wt%) compared to the
213 two colourless kyanites that are ~99.8 % pure (Table S1 in ESM B).

214

215 Sulphide

216

217 Eighteen sulphide inclusions (12 single crystals and 6 polyphase grains; Table S2 in ESM B)
218 are considered eclogitic. From the polyphase inclusions, five consist of two-phase and one of
219 a three-phase assemblage. Inclusions were assigned an eclogitic paragenesis on the basis of
220 co-existing eclogitic clinopyroxenes ($n=2$) and coesite ($n=1$) or <12 wt% Ni (Bulanova et al.
221 1996). A total of 25 sulphide phases were identified and plotted in a Fe-Ni- S quadrilateral
222 diagram (Fig. S2 in ESM A). These include 12 occurrences of monosulphide solid solution
223 (mss), of which 10 are Ni-poor (<10 wt%) and have a molar Ni/Fe ratio <0.2 . Two are
224 associated with separate chalcopyrite phases. Two remaining mss are Ni-rich (13.9-16.1
225 wt%), exceeding the cut-off value for typical eclogitic sulphides. Both are, however, still
226 considered eclogitic on the basis of coexisting clinopyroxene inclusions (omphacite and low-
227 Cr augite). The mss occurring with low-Cr augite additionally coexists with pentlandite
228 containing 16.1 wt% Ni. Furthermore, the mss and pentlandite inclusions have detectable Cr
229 contents (0.33 wt% and 0.29 wt%, respectively). Without the coexisting low-Cr augite
230 inclusion, these two sulphides would have been assigned a peridotitic paragenesis (Stachel

231 and Harris 2008), their compositions clearly reflect a chemically more depleted substrate than
232 commonly observed for eclogitic inclusions. From the six low-Ni (<0.70 wt%) pyrrhotites,
233 one coexisted with pyrite, one with a pyrite-chalcopyrite assemblage and one with
234 pentlandite. This pentlandite is distinct with an elevated Co content (1.48 wt%) compared to
235 all other eclogitic sulphide inclusions (<0.9 wt%). The remaining three inclusions mainly
236 consist of pyrrhotite, with homogenous compositions. The Cu concentrations in chalcopyrite
237 (n=3) range from 26.5 to 27.3 wt%.

238

239 **Peridotitic suite**

240

241 Of the nine peridotitic garnets recovered, eight are harzburgitic and one is lherzolitic (Fig. 2).
242 Cr₂O₃ contents range from intermediate (5.7 wt%) to very high (14.7 wt%) and exceed the
243 highest Cr₂O₃ concentrations previously recorded for a garnet inclusion from the Orapa
244 kimberlite cluster (13.8 wt %; Deines and Harris 2004). Four garnets classify as low-Ca (CaO
245 <1.8 wt%), indicative of derivation from extremely depleted, potentially dunitic sources
246 (Grütter et al. 1999). Two of the harzburgitic and the one lherzolitic garnet record unusually
247 high TiO₂ contents (0.19-0.24 wt%).

248 Thirty eight olivine inclusions have a narrow range in Mg# from 92.3 to 94.2 (Fig.
249 3), with a mean and median of 93.1. On the basis of coexisting garnets, six olivines can be
250 assigned to the harzburgitic paragenesis with a Mg# mean (93.3) and median (93.1), similar to
251 the remaining unassigned olivines, suggesting that all olivines may be harzburgitic. Karowe
252 olivines fall into the normal ranges established for inclusions worldwide with the exception of
253 three with unusual compositions: (1) two olivines have unusually low NiO (0.25 wt%; Fig. 3).
254 One (KW31d) coexists with harzburgitic garnet and orthopyroxene and the other (KW32b)
255 with coesite. (2) A single olivine in diamond KW111 has very high Cr₂O₃ (0.19 wt%), a

256 feature that has been related to low $\text{Cr}^{3+}/\text{Cr}^{2+}$ in the growth environment associated with
257 unusually reducing conditions (Li et al. 1995; Bell et al. 2014).

258 Six orthopyroxene inclusions have a narrow range in Mg# (93.6-94.5; mean and
259 median of 94.0) and show high CaO contents (0.31-0.62 wt% CaO) relative to the more
260 common Ca-depleted harzburgitic orthopyroxenes ($\text{CaO} \leq 0.16$ wt%; Stachel and Harris
261 2008). They fall into a compositional range where harzburgitic and lherzolitic enstatites
262 overlap; based on coexisting garnets, one inclusion is harzburgitic (KW31) and one lherzolitic
263 (KW49). The lone recovered diopside inclusion recovered has a Mg# of 93.1, similar to the
264 median value (92.9) for lherzolitic clinopyroxene inclusions worldwide but is unusually low
265 in Al_2O_3 (0.67 wt%), Cr_2O_3 (0.55 wt%) and Na_2O (0.29 wt%) content (Fig. S1 in ESM A).
266 Eleven Mg-chromites recovered show compositions typical for inclusions in diamond, with
267 high Cr# [$85.6-93.3$; $100 \times \text{Cr}/(\text{Cr}+\text{Al})$] and generally low TiO_2 contents (0.06-0.40 wt%), with
268 a single exception (KW5; 1.1 wt% TiO_2).

269

270 Sulphide

271

272 Four sulphide inclusions were classified as peridotitic. Of these, two are from a single
273 diamond (KW101) where they co-existed with olivine, the other two are lone inclusions
274 (Fig.S2 in ESM A). The two sulphides occurring with olivine are mss and have the elevated
275 Cr contents (0.43 wt%) typically associated with a peridotitic paragenesis (Stachel and Harris
276 2008). The mss with the lower Ni content (16.3 wt%) shows exsolution of pentlandite around
277 the edges and includes a platinum nugget (1 μm diameter; detected by EDS). Of the two
278 remaining sulphides, one is mainly composed of pentlandite (36.1 wt% Ni) and the other
279 consists of pentlandite with a contact phase of Ni-Fe alloy (68.0 wt % Ni, 27.6 wt% Fe and
280 1.50 wt% Co) approximating the chemical composition of awaruite (Ni_2Fe to Ni_3Fe). The
281 whole inclusion measured 120 μm in diameter.

282

283 **Websteritic suite**

284

285 A single orthopyroxene inclusion (KW109b, Table S1 in ESM B) shows an exceptionally low
286 Mg# (56.7), very high CaO (1.43 wt%), and elevated P₂O₅ (0.05 wt%; detection limit =0.02
287 wt%) and TiO₂ (0.21 wt%) concentrations. The inclusion falls far below the lower cut-off for
288 peridotitic orthopyroxene inclusions at Mg# ~86 (Stachel and Harris 2008) and consequently
289 is assigned to the websteritic suite.

290

291 **Sublithospheric inclusions**

292

293 Three “eclogitic” garnets recovered from two diamonds (KW50 and 57, Table S1 in ESM B)
294 exhibit a variable majorite component. Based on 24 oxygens per formula unit, the single low-
295 Ca eclogitic garnet contains 6.46 Si cations and has the high concentrations of TiO₂ (1.4 wt%)
296 and Na₂O (0.77 wt%) often seen among majoritic garnets. The two high-Ca (9.3 wt% CaO)
297 eclogitic garnets have a Si cation content of 6.12, which is only marginally elevated, but these
298 garnets also contain high TiO₂ (1.0 wt%) and Na₂O (0.51 and 0.55 wt%). Their Mg# (42.6-
299 42.8) is unusually low for garnets with a majorite component.

300

301 **REE and trace element compositions of the inclusions**

302

303 **Eclogitic suite**

304

305 For the eclogitic inclusion suite, trace element concentrations were determined on 13 garnets
306 (12 diamonds; Fig. 4) and 13 clinopyroxenes (nine diamonds; Fig. S3). Three low-Ca garnets
307 exhibit steep positive slopes within the LREE_N that become less steep within the MREE_N-

308 HREE_N, rising from ~0.1-0.2 times chondritic abundance for La to about 20-40 times
309 chondritic level for Lu (Fig. 4a). The REE_N patterns for six high-Ca garnets also show steep
310 positive slopes within the LREE but become flat from MREE to HREE at about 10 to 40
311 times chondritic abundance (Fig. 4a). The same distinction in MREE_N-HREE_N slopes for low
312 and high-Ca garnets was previously made by Beard et al. (1996) for eclogite xenoliths from
313 the Mir kimberlite (Siberia). The remaining four high-Ca garnets (with the two garnets from
314 KW81 being indistinguishable; Fig. 4b) have very unusual, flat REE_N patterns at about 2 to 20
315 times chondritic abundance. Mild positive Eu anomalies [geometric Eu/Eu* of 1.21-1.26;
316 defined as $Eu_N / \sqrt{Sm_N \times Gd_N}$] are seen in high-Ca garnets KW79 (positive LREE_N slope
317 and flat MREE_N-HREE_N), KW81a and b (flat REE_N). Similar REE_N patterns are documented
318 for some eclogitic garnet inclusions from the Premier (Cullinan) Mine (Viljoen et al. 2010).
319 From the 13 REE_N patterns for clinopyroxenes, six omphacites and three low-Cr augites are
320 characterized by a mild positive slope within the LREE_N and a steady decline within MREE_N-
321 HREE_N from Nd at 2-10 and Lu 0.2-3 chondritic abundance (Fig. S3a in ESM A). Of the
322 remaining omphacitic clinopyroxenes, three have humped patterns, i.e. they show a more
323 pronounced peak in the LREE_N at about 20-90 times chondritic abundances followed by steep
324 negative slopes in MREE_N-HREE_N (Fig. S3b in ESM A). The last omphacite (KW68b) shows
325 the mild positive slope within the LREE_N, as characteristic for the first group, followed by
326 steeply declining MREE_N-HREE_N like the second group (Fig. S3b). It is the only
327 clinopyroxene to show a strong positive Eu anomaly (Eu/Eu*=1.51). Mild positive Eu
328 anomalies (Eu/Eu*=1.21) are seen for the two omphacite inclusions from diamond KW65
329 (showing the mildly sloping REE_N patterns of the first group and coexisting with high-Ca
330 garnet with Eu/Eu* of 1.15).

331 Eclogitic bulk rock REE_N patterns are calculated for five non-touching pairs of
332 garnet and clinopyroxene (assuming a 1:1 modal abundance ratio; Table S3 in ESM B). Of
333 the three bulk rocks calculated involving high-Ca garnets, two have REE_N patterns sub-

334 parallel to N-MORB with overall lower REE concentrations and more pronounced depletions
335 in LREE (Fig. S4a in ESM A). The third sample (KW56) shows similar HREE concentrations
336 but is distinctly enriched in LREE (to concentrations greater than N-MORB; Fig.S4a). The
337 two pairs involving low-Ca garnets exhibit a gradual increase from $LREE_N$ to $HREE_N$ with N-
338 MORB like HREE (Fig S4b in ESM A). The overall REE depletion and, in two cases, low
339 LREE/HREE for bulk rocks involving high-Ca garnets are all similar to the observations of
340 Beard et al. (1996) for high-Ca eclogite xenoliths. The negative $LREE_N$ - $HREE_N$ slope for the
341 KW56 bulk rock suggests that at Karowe the eclogitic substrates for the high-Ca group were
342 locally affected by cumulate enriched protoliths with metasomatically added LREE.

343

344 **Peridotitic Suite**

345

346 Six harzburgitic and one lherzolitic garnet were analysed for trace element contents. Three
347 harzburgitic garnets show typical sinusoidal REE_N patterns peaking in the LREE (Ce-Nd), a
348 low at Dy-Er and steep positive slopes towards Lu (Fig. S5a in ESM A). A fourth sinusoidal
349 pattern is unusual in peaking only in the MREE (at ~ 10 chondritic abundance for Eu) and
350 having overall super-chondritic MREE and HREE (Fig. S5a). The remaining two harzburgitic
351 garnets have comparatively flat, slightly U-shaped REE_N patterns in the range of 0.3-4 times
352 chondritic abundance (Fig. S5b in ESM A). The single lherzolitic garnet has a normal REE_N
353 pattern with LREE depleted and flat $MREE_N$ - $HREE_N$ at about 4 times chondritic abundance
354 (Fig. S5b).

355

356 **Sublithospheric inclusions**

357

358 The two analyzed majoritic garnets show opposite REE_N patterns (Fig. 5): the low-Ca
359 majorite displays a steep negative slope from $LREE_N$ at ~ 100 times to $HREE_N$ at ~ 1 times

360 chondritic abundance. The high-Ca majoritic garnet rises from chondritic LREE_N abundance
361 to 300 times chondritic HREE_N.

362

363 **Geothermobarometry**

364

365 In ten diamonds two or more mineral inclusion species coexisted that are suitable for the
366 calculation of equilibrium temperatures and/or pressures (Table S2 in ESM A). Five non-
367 touching pairs of eclogitic garnet and clinopyroxene yield equilibrium temperatures of 1260-
368 1480 °C, using the Fe-Mg exchange thermometer of Krogh (1988) at a pre-set pressure of 5
369 GPa. Two pairs involving low-Ca garnets give lower temperatures (1260-1270 °C), compared
370 to three involving high-Ca garnets (1420-1480 °C; Table S2 in ESM A). A similar dichotomy
371 is not observed for Orapa, where pairs of clinopyroxene and low-Ca (n=3) or high-Ca garnets
372 (n=6) yield average equilibration temperatures of 1150 and 1160 °C, respectively (calculated
373 from the dataset of Deines et al. 1993 and Gurney et al. 1984). For Letlhakane, only one pair
374 of clinopyroxene and low-Ca garnet was analysed (yielding 1320 °C), the eight pairs
375 involving high-Ca garnets yield an average of 1210 °C (calculated from the data set of Deines
376 and Harris, 2004). A separation into cool (shallow) low-Ca and hot (deep) high-Ca eclogites,
377 therefore, is not a general feature of the lithospheric mantle beneath the Orapa kimberlite
378 cluster.

379 For the peridotitic suite, one touching (KW86) and three non-touching garnet-olivine
380 pairs (KW31, 69 and 74) give a temperature range of 900-1245 °C (Table S2 in ESM A),
381 based on the garnet-olivine thermometer of O'Neill and Wood (1979; calculated at 5 GPa).
382 Diamond KW31, in addition to garnet and olivine, also hosted a non-touching orthopyroxene;
383 application of the garnet-orthopyroxene Fe-Mg exchange thermometer of Harley (1984)
384 yielded a temperature of 1180 °C at pre-set pressure of 5 GPa (Table S2 in ESM A), which is
385 within error of the 1245 °C estimate obtained with the olivine-garnet thermometer (O'Neill

386 and Wood 1979). This suggests that olivine, orthopyroxene and garnet in this diamond likely
387 are in equilibrium. Simultaneous estimation of pressure and temperature using garnet-
388 orthopyroxene geothermobarometry (Brey and Köhler 1990; Harley 1984) yields 5.2 GPa and
389 1200 °C as a condition of inclusion entrapment. Application of the same geothermobarometer
390 combination to a second non-touching garnet-orthopyroxene pair in diamond KW49 gives 4.1
391 GPa at 950 °C (Table S2 in ESM A), resulting in the two samples plotting along a 39-40
392 mW/m² (Hasterok and Chapman 2011) model geotherm (Fig. S6 in ESM A). This result is in
393 very good agreement with the 40 mW/m² paleogeotherm derived by Stiefenhofer et al. (1997)
394 for peridotite xenoliths from the nearby Cretaceous Letlhakane kimberlite. The single
395 lherzolitic clinopyroxene inclusion (KW 93) could not be used for geothermobarometric
396 calculations as it fails the compositional filters of Grütter (2009).

397 Based on the Ca-in-opx thermometer (Brey and Köhler 1990), the single websteritic
398 inclusions last equilibrated at 1380 °C. Such a high temperature is not typically associated
399 with diamond formation along steady state cratonic geotherms (Stachel and Luth 2015) but
400 suggests either association with a thermal perturbation (magmatic intrusion) within the
401 lithosphere or derivation from below the lithospheric mantle. Given that this temperature falls
402 within the range observed for eclogitic (non-majoritic) garnet and clinopyroxene pairs from
403 Karowe (see above), a lithospheric origin appears plausible. Pressure determinations for the
404 three majoritic garnet inclusions from Karowe using the barometer of Beyer and Frost (2017)
405 range between 11 and 14 GPa, clearly implying a sublithospheric origin (Fig. 6).

406

407

408 **Discussion**

409

410 Of the 107 inclusion-bearing diamonds studied here, 53 % are eclogitic, 44 % peridotitic, 2 %
411 sublithospheric, and 1 % websteritic. This predominantly eclogitic diamond production at
412 Karowe is similar to the Orapa mine but unlike the predominantly peridotitic productions of
413 the adjacent Letlhakane and Damtshaa mines.

414

415 **Origin and evolution of eclogitic and websteritic diamond substrates**

416

417 Eclogitic garnet and clinopyroxene inclusions from Karowe are compositionally variable but
418 for the most part, compare well to previous studies on inclusions in diamonds from the Orapa
419 kimberlite cluster. A comparison of Karowe garnet compositions with other garnets inclusions
420 from the Orapa cluster is shown in Figure S7 in ESM A. A websteritic inclusion suite was
421 first recognized in diamonds from the Orapa kimberlite (Gurney et al. 1984) and documents
422 the presence of an unusually wide spectrum of “mafic” diamond substrates in the local
423 lithospheric mantle. This is also evident from the compositions of a large suite of eclogite and
424 pyroxenite xenoliths from the Orapa (Aulbach et al. 2017). In Karowe, the presence of
425 orthopyroxene-bearing (websteritic) substrates is evidenced by a single enstatite inclusion
426 with low Mg# (57). Possible saturation in orthopyroxene may also be indicated by Ca# as low
427 as 21.5 in augite inclusions with elevated Cr, Ni, and Mg contents. Mineralogical variability
428 in the eclogitic substrates is also highlighted by the occasional presence of kyanite and coesite
429 inclusions. Kyanite occurs in Group I diamondiferous eclogites from Orapa (Shee and Gurney
430 1979) and was recently described in Group II diamond-free eclogites by Aulbach et al.
431 (2017). It is a characteristic phase of aluminous eclogites in general (Spetsius 2004; Shu et
432 al.2016).

433 The REE characteristics (bulk rock $LREE_N/HREE_N$ generally <1 ; presence of
434 positive Eu anomalies in some high-Ca garnets and coexisting clinopyroxenes) are consistent

435 with a derivation of Karowe low- and high-Ca eclogitic diamond substrates from subducted
436 protoliths. Moderate positive Eu anomalies in two of the calculated bulk rocks involving high-
437 Ca garnets (Fig. S8 in ESM A) suggest cumulate enriched protoliths for these samples (e.g.,
438 Aulbach and Viljoen 2015). The absence of distinct positive Sr anomalies for the same
439 calculated bulk rocks (Fig. S8), however, indicates decoupling of Sr and Eu/Eu*. A
440 decoupling of Sr content and Eu/Eu* is also observed for garnet inclusions in diamonds
441 world-wide and was related to Sr depletion during secondary partial melting events (Stachel et
442 al. 2015). Positive Sr anomalies without paired Eu anomalies in the two calculated low-Ca
443 group bulk rocks (Fig. S8) again show decoupling of Sr content and Eu/Eu*. Positive slopes
444 from MREE to HREE ($Yb_N/Gd_N=2.2-3.3$) indicate that the protoliths of the low-Ca eclogitic
445 diamond substrates likely experienced partial melting and associated depletion in LREE (Fig.
446 S-4), followed by metasomatic re-enrichment in Sr (Aulbach et al. 2017). Estimated bulk rock
447 Mg# of 61-72 are on the high side of average present-day N-MORB (~60; Floyd 1991) and
448 could be inherited from gabbroic protoliths (Aulbach and Jacob 2016) or again reflect a
449 secondary melting extraction event.

450 Combining the observed variability in the major element compositions of garnet and
451 clinopyroxene inclusions with the range of bulk rock REE_N patterns indicates that a large
452 cross section of oceanic crust (not necessarily related to a single subduction event) was
453 sampled. Beard et al. (1996) suggested an origin of high-Ca eclogites as cumulate-enriched
454 rocks of the deeper oceanic crust (causing overall REE depletion increasing towards the
455 LREE) and low-Ca eclogites as upper oceanic crust (with relatively unfractionated REE). At
456 Karowe, the presence of positive Eu anomalies only in high-Ca garnets and associated
457 clinopyroxenes is consistent with this model. Subsequent modification during partial melting
458 in the garnet stability field (LREE loss; Foley et al. 2002; Jacob 2004), metasomatic
459 modification (LREE addition; Ireland et al. 1994) and interaction with mantle peridotite
460 (adding Cr and Mg; Smart et al. 2009) may all have acted to increase the compositional

461 variance further and, in combination, cause compositionally diverse eclogitic to pyroxenitic
462 diamond substrates beneath Karowe.

463

464 **Peridotitic diamond substrates in the lithospheric mantle beneath Karowe**

465

466 Peridotitic olivine inclusions in Karowe diamonds have high Mg# with a mean and median of
467 93.1, comparable to olivine inclusions from elsewhere in the Orapa cluster (e.g., Deines et al.
468 2009) and harzburgitic olivine inclusions worldwide (mean: 93.2; Stachel and Harris 2008).
469 Such high forsterite contents imply the presence of peridotitic mantle sources that experienced
470 primary melt extraction approaching and exceeding the exhaustion of orthopyroxene
471 (Bernstein et al. 2007). Two Karowe olivines with normal Mg# have unusually low NiO
472 contents (mean of 0.26 wt% as opposed to 0.36 wt% worldwide; Fig. 3). Unusually low Ni
473 contents were shown to occur in olivine inclusions that originated as ringwoodite in
474 equilibrium with ferropericlase in the lower transition zone (Brey et al. 2004). However, the
475 two low-Ni olivines from Karowe have low P₂O₅, Cr₂O₃ and Na₂O, inconsistent with a
476 superdeep origin, and in one case (KW31) occur in apparent equilibrium with normal
477 lithospheric inclusions (harzburgitic garnet and orthopyroxene).

478 Peridotitic garnet inclusions from Karowe are dominantly (~90 %) harzburgitic and
479 almost half have CaO contents < 1.8 wt%, i.e. likely originating from primary residues that
480 were molten beyond the point of orthopyroxene exhaustion (Grütter et al. 1999). One of the
481 low-Ca garnets has a very high Cr₂O₃ content (14.7 wt%) that implies that at the time of
482 diamond formation highly depleted cratonic lithosphere extended down to at least 220 km
483 depth (equivalent to 7 GPa pressure; Fig. 2; Grütter et al. 2006) beneath Karowe. These
484 observations establish that the lithospheric mantle beneath Karowe, at least at the time of
485 peridotitic diamond formation, was strongly depleted in easily fusible components and very
486 thick, similar to the southern part of the Zimbabwe Craton (Smith et al. 2009). REE_N patterns

487 for harzburgitic garnets (sinusoidal to slightly U-shaped) and the one lherzolitic garnet
488 (normal REE_N) are very distinct and imply different styles of metasomatic re-enrichment:
489 fluid metasomatism for the harzburgitic and melt metasomatism for the lherzolitic substrates
490 (Stachel et al. 2004a). The Y and Zr concentrations of these garnets (Fig. S9 in ESM A) are
491 consistent with this conclusion: all but one harzburgitic garnet plot in the depleted field but
492 clearly define a trend that leads into the field for low temperature, fluid-style metasomatism.
493 Whereas the single lherzolitic garnet documents a much higher Y/Zr ratio, consistent with
494 melt metasomatism. Very similar garnet REE_N patterns to those observed at Karowe
495 (sinusoidal and slightly U-shaped for harzburgitic and normal for the one lherzolitic garnet)
496 and metasomatic styles (fluid metasomatism for harzburgitic and melt metasomatism for
497 lherzolitic garnets) were also documented for peridotitic garnet inclusions in diamonds from
498 Orapa (Fig. S9; Stachel et al. 2004b).

499 The presence of a Ni-Fe alloy inclusion (intergrown with pentlandite) indicates that,
500 at least locally, fO_2 conditions as reducing as the iron-wüstite buffer occurred in the peridotitic
501 diamond substrates. Generally, such reducing conditions are not observed within lithospheric
502 mantle, even for cratons of similar depth extent (Stagno et al. 2013), and hence are interpreted
503 as a localized environment rather than widespread metal saturation in the deep mantle
504 lithosphere below the Zimbabwe Craton.

505

506 **Sublithospheric substrates**

507

508 Pressure determinations for the three majoritic garnet inclusions from two Karowe diamonds
509 place their origin in the deep asthenosphere and uppermost transition zone (between 330 and
510 420 km depth; Fig. 6b). Based on their low-Cr character (cut-off at Cr₂O₃ <1 wt%; Schulze
511 2003), the majoritic garnets all derive from eclogitic substrates. Kiseeva et al. (2013) reported
512 that despite their overall eclogitic mineral compositions, such majoritic garnets can follow (1)

513 a substitution mechanism ($2\text{Al}^{3+}=\text{Si}^{4+}+\text{M}^{2+}$) that is characteristic for peridotitic (and
514 pyroxenitic) substrates or (2) a substitution mechanism (eclogitic trend) accommodating the
515 jadeite component of omphacitic clinopyroxene ($\text{M}^{2+}+\text{Al}^{3+}=\text{Na}^{+}+\text{Si}^{4+}$). The first substitution
516 (pyroxenitic trend) is interpreted to reflect sublithospheric diamond formation through
517 interaction of slab-derived magnesio-carbonatitic melt and adjacent pyrolytic mantle (Kiseeva
518 et al. 2016, Walter et al. 2008). At Karowe, the low-Ca majoritic garnet in diamond KW50
519 shows an excess in $\text{Mg}+\text{Ca}+\text{Fe}+\text{Mn}$ over the number of available X-sites in garnet and thus
520 falls onto the pyroxenitic trend, while the two high-Ca majoritic garnets in KW57 fall onto the
521 Na-majorite trend (Harte 2010; Kiseeva et al. 2013). The low-Ca majoritic garnet has an
522 elevated Cr_2O_3 content (0.18 wt%) and Mg# (66.2), consistent with the pyroxenitic
523 connection suggested by Kiseeva et al. (2013 and 2015), whilst the two high-Ca garnets
524 contain Cr below the limit of detection and have low Mg# (average of 42.7), indicative of a
525 typical eclogitic bulk rock composition (Harte 2010). This clear separation into two modes of
526 majorite formation is reflected in highly distinct REE_N patterns for low- and high-Ca majoritic
527 garnets (Fig.5). The low-Ca garnet exhibits a melt-like pattern characterized by strong LREE
528 enrichment and a steeply negative slope from LREE_N to HREE_N . Based on the majorite-melt
529 partition coefficients of Yurimoto and Ohtani (1992), the REE_N pattern of a calculated melt in
530 equilibrium with the low-Ca garnet would show a steep decline from 540 times chondritic La
531 to chondritic Yb abundance. The high-Ca garnet has the reverse pattern with a steep positive
532 LREE_N - HREE_N slope. Although more enriched in HREE, the high-Ca garnet REE_N pattern is
533 similar to those displayed by eclogitic majoritic garnets from Monastery (Moore et al. 1991)
534 and Jagersfontein (Tappert et al. 2005; Fig.5). The strong HREE enrichment would be
535 consistent with the high-Ca garnet crystallizing in a residuum that yielded a melt similar to
536 that forming the low-Ca majoritic garnets.

537 Karowe Mine produces large to extremely large gem-quality diamonds that are
538 nitrogen “free” Type IIa or Type IaB containing low contents of completely aggregated

539 nitrogen (D'Haenens-Johansson et al. 2017). Smith et al. (2016) studied large Type IIa
540 diamonds from several sources and found that they exclusively contained inclusions of
541 sublithospheric origin, including majoritic garnets. Inclusions of sublithospheric origin have
542 not previously been reported in diamonds from the Orapa kimberlite cluster. The two
543 majoritic garnet-bearing diamonds from Karowe reported here are not Type IIa but have very
544 low to moderate nitrogen contents of highly aggregated nitrogen (30 and 250 at.ppm with 95
545 and 86 %B respectively). Nevertheless, their discovery shows that sublithospheric diamond
546 sources were tapped by the Karowe kimberlite, an indication that the very large diamonds
547 from this mine also formed at below the lithosphere, as observed elsewhere by Smith et al.
548 (2016).

549

550

551 **Conclusions**

552

553 The major element composition of inclusions in Karowe diamonds establishes their derivation
554 from four distinct mantle lithologies present below the western edge of Zimbabwe Craton.
555 More than half of the diamond population is eclogitic (53 %), followed by peridotitic (44 %),
556 sublithospheric (2 %) and websteritic (1 %). Overall, the composition of the eclogitic and
557 peridotitic inclusions compares well with the results of previous studies conducted on other
558 localities in the Orapa kimberlite cluster (on diamonds from Orapa, Damtshaa and
559 Letlhakane). The eclogitic diamond substrates beneath Karowe are highly diverse, ranging
560 from typical basaltic to cumulate-like protolith compositions. This variety is documented in
561 the variance of calculated bulk rock Mg# (61-72) and the broad range in the jadeite
562 component (14-48 mol%) in eclogitic clinopyroxenes as well as the presence of three kyanite
563 inclusions. In combination with calculated bulk rock REE patterns, the variable inclusion

564 chemistry documents derivation of eclogitic diamonds from a range of protoliths that
565 represent both shallow and deep oceanic crust, in part modified by partial melting during
566 subduction and subsequent metasomatism. About 40 % of eclogitic clinopyroxenes are
567 augites with elevated contents of Cr, Ni and Mg at unusually low Ca contents, which may
568 represent an association transitional to pyroxenites. A single websteritic orthopyroxene
569 inclusion confirms the presence of pyroxenitic diamond substrates beneath Karowe that were
570 also recognized at Orapa.

571 The peridotitic inclusion suite documents a predominance of typical, highly depleted
572 diamond substrates of harzburgitic to dunitic paragenesis along the western margin of the
573 Zimbabwe Craton. Very high Cr₂O₃ in one of the low-Ca peridotitic garnets is indicative of a
574 depth extent of this highly depleted cratonic lithosphere to at least 220 km depth at the time of
575 diamond formation.

576 The discovery of sublithospheric (eclogitic majoritic garnet) inclusions in Karowe
577 diamonds sets this mine apart from other deposits in the Orapa kimberlite cluster and may
578 provide a key link to the regular recovery of exceptionally large diamonds at Karowe. Low-
579 Ca and high-Ca majoritic garnets follow two distinct trends of majoritic substitution
580 (pyroxenitic and eclogitic trend, respectively). These distinct origins are consistent with their
581 REE_N patterns being mirror images: (1) high LREE and low HREE for the low Ca-garnet,
582 suggested to relate to melt-aided slab-pyrolite interaction and (2) low LREE and high HREE
583 for the high-Ca garnet, which formed in substrates representing basaltic protoliths that lost a
584 partial melt similar to that reflected in the low-Ca majorites.

585

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598

599

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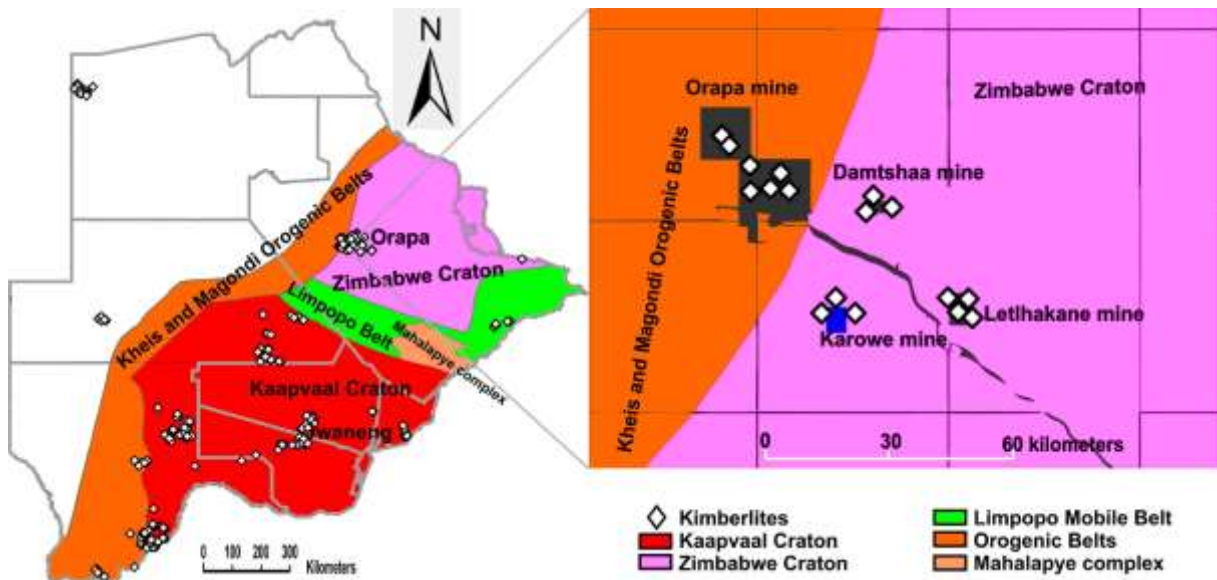
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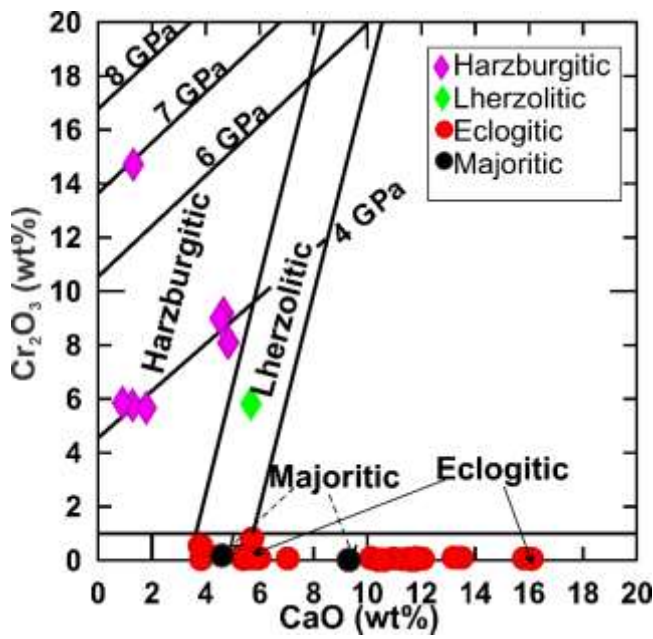
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785 **Fig. 1** Setting of the Karowe Mine in the geology of Botswana. Map modified after Key and

786 Ayres (2000)

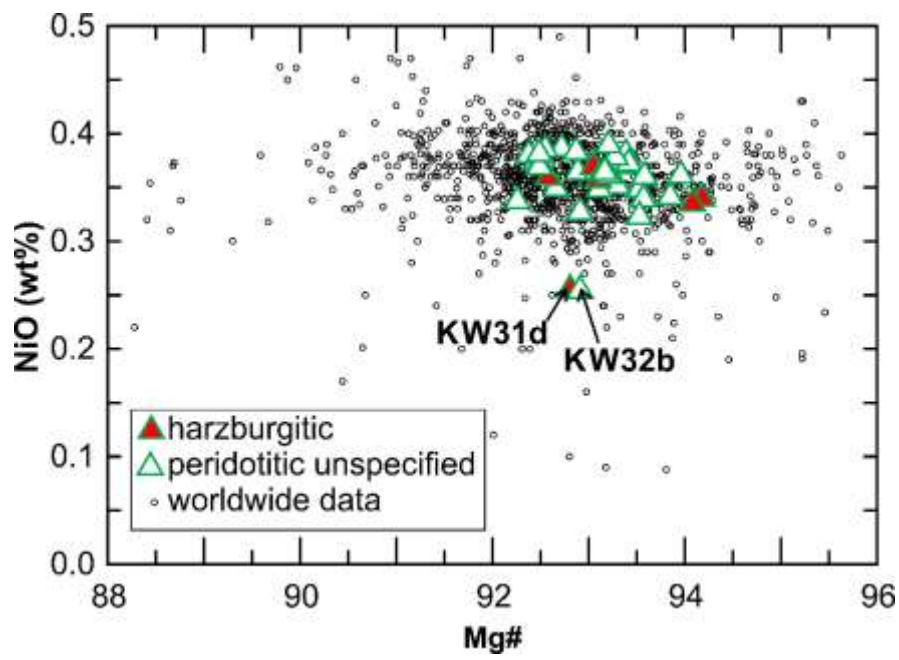


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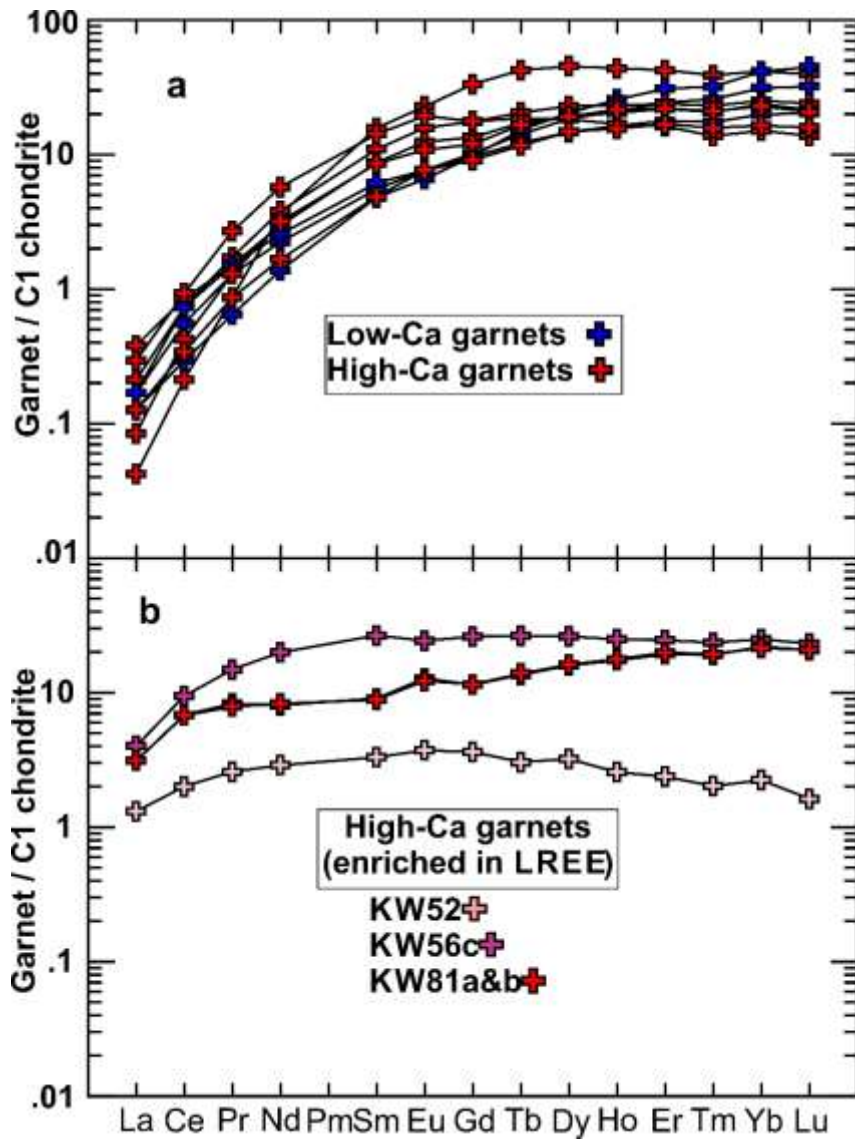
788 **Fig. 2** Cr₂O₃ vs CaO diagram for garnets inclusions in Karowe diamonds. Isobars are from

789 Grütter et al. (2006) and represent minimum pressures (as equilibrium with spinel is not

790 established)



792 **Fig. 3** NiO (wt%) versus molar Mg# ($100 \times \text{Mg} / [\text{Mg} + \text{Fe}]$) in olivine inclusions from this study
 793 and localities worldwide (n=1,306; Stachel and Harris 2008)

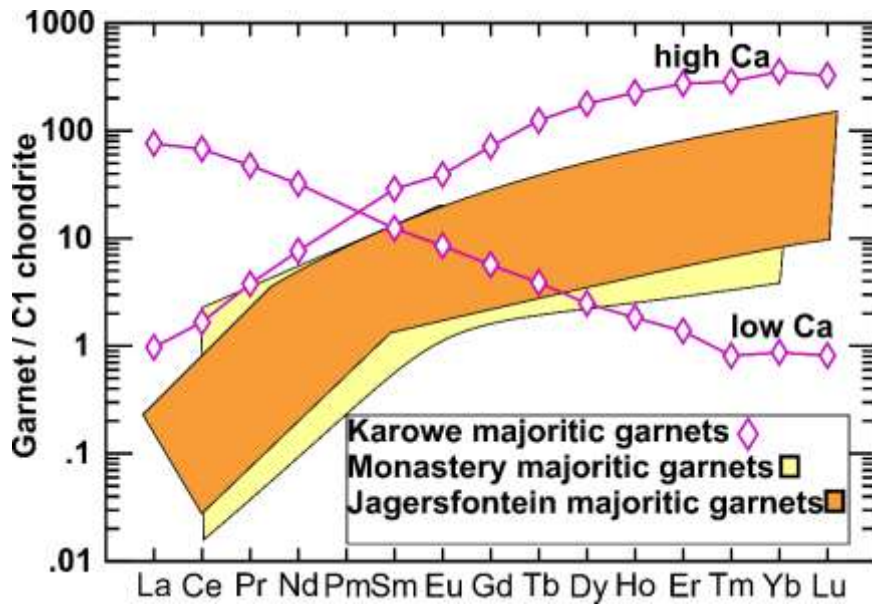


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795 **Fig. 4** Chondrite normalized (McDonough and Sun 1995) REE patterns for Karowe eclogitic

796 garnet inclusions. **a:** Low-Ca garnets (n=3) and high-Ca garnets (n=6). **b:** Second group of

797 high-Ca garnets with nearly flat REE patterns



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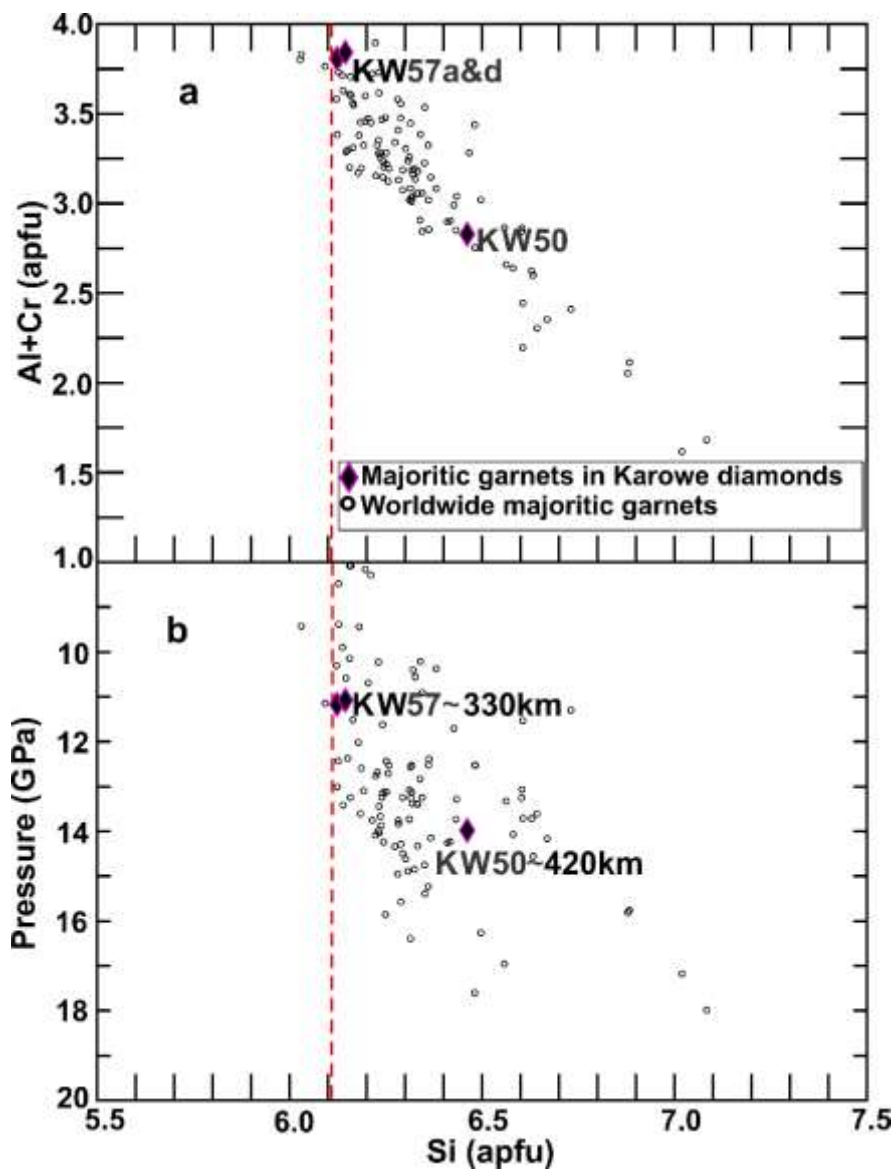
Fig. 5 Chondrite normalized (McDonough and Sun 1995) REE patterns of low- and high-Ca

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majoritic garnets from Karowe compared to majoritic garnets from Monastery (Moore et al.

801

1991) and Jagersfontein (Tappert et al. 2005).



802

803 **Fig. 6 a:** Al+Cr versus Si as atoms per formula unit (apfu, with [O]=24) in majoritic garnets
 804 inclusions from Karowe and worldwide sources (n=69; database of Stachel and Harris 2008).
 805 Red dashed line indicates threshold value of > 6.12 apfu Si used to define majoritic garnets. **b:**
 806 Si apfu plotted against pressure (GPa), calculated using the barometer of Beyer and Frost
 807 (2017)