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RESEARCH ARTICLE

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Kev Points:

- The increase of ²²⁸Ra in central Arctic surface waters points at stronger sediment water exchange due to the longer ice-free season on Arctic shelves
- The decrease in ²²⁸Th/²²⁸Ra in the Transpolar Drift may be related to the increase in the rate of the drift
- Excess ²²⁸Th in intermediate depths (100–1,500 m) may be used as tracer of export production

Supporting Information:

- Supporting Information S1
- Table S1
- Table S2
- Table S3a
- Table S3bTable S3c
- Table S4

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Radium Isotopes Across the Arctic Ocean Show Time Scales of Water Mass Ventilation and Increasing Shelf Inputs

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Abstract The first full transarctic section of ²²⁸Ra in surface waters measured during GEOTRACES cruises PS94 and HLY1502 (2015) shows a consistent distribution with maximum activities in the transpolar drift. Activities in the central Arctic have increased from 2007 through 2011 to 2015. The increased ²²⁸Ra input is attributed to stronger wave action on shelves resulting from a longer ice-free season. A concomitant decrease in the ²²⁸Th/²²⁸Ra ratio likely results from more rapid transit of surface waters depleted in ²²⁸Th by scavenging over the shelf. The ²²⁸Ra activities observed in intermediate waters (<1,500 m) in the Amundsen Basin are explained by ventilation with shelf water on a time scale of about 15–18 years, in good agreement with estimates based on SF₆ and ¹²⁹I/²³⁶U. The ²²⁸Th excess below the mixed layer up to 1,500 m depth can complement ²³⁴Th and ²¹⁰Po as tracers of export production, after correction for the inherent excess resulting from the similarity of ²²⁸Ra and ²²⁸Th decay times. We show with a Th/Ra profile model that the ²²⁸Th/²²⁸Ra ratio below 1,500 m is inappropriate for this purpose because it is a delicate balance between horizontal supply of ²²⁸Ra and vertical flux of particulate ²²⁸Th. The accumulation of ²²⁶Ra in the deep Makarov Basin is not associated with an accumulation of Ba and can therefore be attributed to supply from decay of ²³⁰Th in the bottom sediment. We estimate a ventilation time of 480 years for the deep Makarov-Canada Basin, in good agreement with previous estimates using other tracers.

1. Introduction

The Arctic Ocean is characterized by strong interactions with the surrounding continents. Over 10% of the world's river discharge finds its way into the Arctic Ocean, which has only 1% of the World's Ocean volume (McClelland et al., 2012). On the Siberian side, the shelves are especially wide and shallow. Surface waters in the central Arctic are relatively fresh, with the transpolar drift (TPD) serving as a transport route for terrestrial inputs through the central Arctic Ocean toward Fram Strait.

The Arctic Ocean is in transition. Minimum summer sea ice cover declined from 7 to 4 10⁶ km² during the past 30 years (Grosfeld et al., 2016; Serreze et al., 2007). The longer ice-free period on the shelf and increased temperatures cause permafrost thaw (Luo et al., 2016) and increased erosion (Günther et al., 2013), which will affect the transport and fate of the terrestrial inputs. While these processes may enhance the input of sediments and dissolved material, the extensive sea ice melt may cause ice rafted matter to be released and deposited at an earlier stage on its transit over the Arctic Basins. The loss of sea ice will result in light conditions that are more favorable for plankton growth (Arrigo et al., 2008; Arrigo et al., 2012), but it is questionable whether there are sufficient nutrients for an increase in export production. It is therefore important to investigate whether the transport of terrestrial material, primary production, and sedimentation are affected in response to the environmental changes taking place.

Radium isotopes are particularly suited to study the transport and fate of terrestrial inputs in the Arctic. All four naturally occurring isotopes are produced by decay of an isotope of thorium, a highly particle reactive



element. Radium is relatively soluble in seawater such that once a thorium isotope in suspended or bottom sediments decays, a fraction of the produced radium isotope may be released to solution. ²²⁴Ra (3.7-day half-life) and ²²³Ra (11.4-day half-life) are used extensively as tracers for interaction on the shelf and for submarine groundwater discharge (SGD). ²²⁸Ra is strongly enriched in shelf waters and in the TPD and its half-life (5.8 years) is well suited to study the fate of this shelf-derived water in the central Arctic (Rutgers van der Loeff et al., 1995). Finally, ²²⁶Ra (1620 y half-life) has been used to trace SGD (Moore, 1996) and inputs from deep-sea sediments (Cochran, 1980).

Primary production and the export flux of particles are very low in the ice-covered Arctic Ocean. Cai et al. (2010) reported an export production of 0.2 ± 1.0 mmol carbon m⁻² d⁻¹ in the central Arctic Ocean based on ²³⁴Th/²³⁸U disequilibria measured in August to September 2007, but observations at the seafloor in 2012 suggested a more substantial export, for example, of ice algae (Boetius et al., 2013). One reason for this discrepancy may be that the time scale of the ²³⁴Th tracer (related to the mean life of $1/\lambda = 35$ days) is too short to record export events that took place earlier in the season. The time scale of the ²¹⁰Po/²¹⁰Pb tracer system (mean life of ²¹⁰Po is 200 days) is more appropriate to cover the full productive season. Roca-Martí et al. (2016) compared export fluxes using both tracer pairs. They confirmed low export fluxes in the central Arctic (2 \pm 2 mmol C m⁻² d⁻¹ using ²³⁴Th/²³⁸U and 3 \pm 2 mmol C m⁻² d⁻¹ using ²¹⁰Po/²¹⁰Pb) and observed a depletion of ²¹⁰Po at all stations, while a significant depletion of ²³⁴Th was found in only three out of nine stations. Still, the calculated fluxes are low and have large uncertainties, and moreover, the use of ²¹⁰Po/²¹⁰Pb as tracer for export production is complicated by the particle reactive behavior of ²¹⁰Pb, which is itself scavenged. An alternative tracer pair with much promise for understanding Ocean export production is ²²⁸Th/²²⁸Ra (Luo et al., 1995; Rutgers van der Loeff et al., 2012). In the Arctic Ocean, activities of ²²⁸Ra in surface waters are very high, the time scale of ²²⁸Th can cover more than a productive season (²²⁸Th half-life 1.9 years corresponding to a mean life of 2.7 years), and even the low export fluxes create a depletion that can be measured with good precision. A complicating factor of the relatively long time scale is that the ²²⁸Th/²²⁸Ra ratio in surface waters depends not only on the export flux but also on the scavenging history of the water mass (Rutgers van der Loeff et al., 2012). Here we wish to investigate to what extent the ²²⁸Th/²²⁸Ra ratio in the subsurface water column can be used to trace

In this paper we present the distribution of ²²⁴Ra, ²²⁸Th, and ²²⁸Ra in the surface water on a section from the Barents shelf to the Bering Strait, the first full transarctic section of these parameters measured as part of the coordinated GEOTRACES Arctic study in 2015. We compare the data with earlier observations and discuss the observed changes. The observed distributions of ²²⁸Th and ²²⁸Ra in the water column are compared with a Th/Ra profile model, and we discuss what these profiles can tell us about particle fluxes. Finally, we discuss to what extent the accumulation of ²²⁶Ra that we observed in deep waters can be used as a measure of basin ventilation times.

2. Materials and Methods

In the framework of the GEOTRACES program, two expeditions were organized to the central Arctic in 2015: GEOTRACES section GN04 on RV *Polarstern* (expedition PS94, ARK XXIX/3, TRANSARC II; Schauer, 2016) and GEOTRACES section GN01 on USCGC Healy (expedition HLY1502; Kadko & Landing, 2015; Figure 1). We also report on Ra and Th analyses in surface waters collected on RV *Polarstern* expedition PS78/ARK XXVI/3 in 2011 (TRANSARC I) (Schauer, 2012). All data are presented in Tables S1–S4 in the supporting information.

During *Polarstern* expedition PS78 (5 August to 6 October 2011) large volume surface water samples were collected at 41 stations into 300-L tanks using a tap close to the ships seawater intake (first stations only) or (from North Pole Sta 218 onward) a well pump lowered over the side of the ship to a depth of 10 m. Each sample was filtered through a 0.8- μ m Supor® (polyether sulfone) filter and then passed at a flow rate of <1 L/min using a peristaltic pump through MnO₂-impregnated acrylic fiber to scavenge radium isotopes. Fibers were partially dried using compressed air, and ²²⁴Ra was measured at sea using RaDeCC delayed coincidence alpha detectors (Moore & Arnold, 1996). ²²⁴Ra supported by ²²⁸Th was derived from a second RaDeCC count after at least 20 days (5 half-lives).

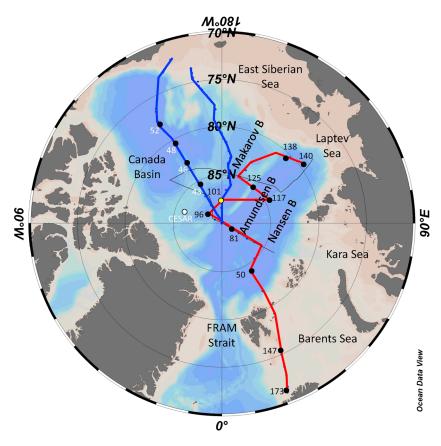


Figure 1. Cruise tracks of RV *Polarstern* expeditions PS78 (black), PS94 (GEOTRACES GN04, red, black station numbers), and USCGC Healy expedition HLY1502 (GEOTRACES GN01, blue, white station numbers) crossover station (yellow) and CESAR ice station (white dot).

During *Polarstern* expedition PS94 (17 August to 15 October 2015) surface water samples (collected at a tap close to the ships seawater intake after abundant rinsing of the tubing systems) were passed over an uncoated 7.5-cm cartridge and subsequently over a series of two tubes with loose MnO_2 -coated fibers (up to Sta 96) or two MnO_2 -coated cartridges (from Sta 99 onward; Henderson et al., 2013). Flow rate was limited to 1 L/min, and volumes were measured with Kent flow meters. During 10 in situ pump (ISP) casts, MnO_2 -coated cartridges were mounted after 0.8- μ m Supor filters; average flow rates varied between 0.8 and 6.1 L/min. On 37 ISPs two cartridges were used in series, allowing the calculation of Ra absorption efficiency. There was no relationship between Ra collection efficiency and flow rate ($R^2 = 0.0013$). On other ISPs only one cartridge could be mounted, in which cases we had to use the average collection efficiency of all other cartridges (90 \pm 7%). Cartridges and fibers were partially dried using compressed air, and 224 Ra was measured at sea using RaDeCC delayed coincidence alpha detectors. Supported 224 Ra was derived from a second RaDeCC count after at least 16 days (4 half-lives).

Samples from PS78 and PS94 were analyzed for 226 Ra and 228 Ra at AWI, Bremerhaven. The fibers were leached following Elsinger et al. (1982), and the MnO₂-coated cartridges were leached by Soxhlet extraction with 6 N HCl refluxing over 10 hr. Radium in the extracts was coprecipitated with BaSO₄ (Cutter et al., 2010) and analyzed with gamma spectroscopy (Moore, 1984).

On the HLY1502 expedition, near-surface water samples were collected from a depth of 2 m using a submersible surface pump. Approximately 280 L of water was collected at each station and filtered through a MnO₂-coated fiber at a flow rate of <1 L/min. Over the shelf, the samples were first filtered through 1- μ m Hytrex cartridge to remove particles; at all other stations, the samples were not prefiltered. The fibers were rinsed with Ra-free freshwater and analyzed on RaDeCC detectors within 3 days of collection to measure 224 Ra. After 4 weeks a second RaDeCC count was performed to determine the amount of 224 Ra supported by



²²⁸Th. Fibers were then ashed in a muffle furnace at 820 °C, and the ash was packed in to polystyrene vials, sealed with epoxy (to prevent ²²²Rn loss), aged for 3 weeks, and analyzed on high purity, well-type gamma detectors in the Moore lab at the University of South Carolina. Radium-228 was measured using the lines of ²²⁸Ac (338 and 911 KeV), and ²²⁶Ra was measured using the line for ²¹⁴Pb (352 KeV). Detector efficiencies were determined using ashed fiber standards spiked with ²²⁶Ra and ²³²Th with daughters in equilibrium.

Water column samples for the HLY1502 expedition were collected using ISPs that were programmed to pump for 4 hr, typically filtering 1,200–1,600 L of seawater at an average flow rate of \sim 6 L/min. Seawater was first passed through 51 and 1- μ m filters to collect particulate isotopes, and then through MnO₂-coated cellulose cartridges (Henderson et al., 2013). After collection, cartridges were rinsed with Ra-free freshwater, partially dried with compressed air, and counted on RaDeCC detectors within 2 days of collection to measure 224 Ra. Second counts were performed after 4 weeks to determine the amount of supported 224 Ra. Cartridges were then ashed, and 228 Ra and 226 Ra were measured by gamma spectrometry in the Charette lab at the Woods Hole Oceanographic Institution using the same method as the surface samples. To determine cartridge collection efficiencies, small volume samples of 226 Ra (15–25 L) were collected using a Niskin bottle mounted either on the CTD rosette (shallow casts) or above the ISPs (deep casts) and were filtered through a MnO₂-coated fiber. The activities of Ra measured on the fibers, which quantitatively scavenge radium (Charette et al., 2012), were compared to those determined on the cartridges. For HLY1502, the collection efficiency of individual cartridges ranged from 19% to 99%, with an average of 70 \pm 19% (1 σ). Particulate 228 Th activities were measured on the 51 and 1- μ m filters using RaDeCC detectors after storing the samples for at least a month to allow 224 Ra to reach secular equilibrium with 228 Th (Black, 2018; Maiti et al., 2014).

For the intercalibration of RaDeCC counting (224 Ra), cartridges spiked with 232 Th series in equilibrium were prepared in the USC lab of Willard Moore in South Carolina and in the AWI lab in Bremerhaven. These standards were circulated among the Charette, Moore, and Rutgers van der Loeff labs and results agreed to within 9–13%, comparable to interlab differences for high 224 Ra activities in a previous intercalibration (Charette et al., 2012). For the intercalibration of gamma counting (226 Ra and 228 Ra) a fiber spiked at the WHOI lab with the WHOI 226 Ra/ 228 Ra standard was analyzed at AWI. AWI values were $4.1 \pm 1.7\%$ lower for 226 Ra and $6.4 \pm 2.2\%$ higher for 228 Ra compared to the WHOI calibration. A further check was obtained from the sampling at the crossover station: Healy HLY1502 station 30 (1 September 2015) and *Polarstern* PS94 station 101 (14 September 2015; Figure S1 in the supporting information). The offset in 226 Ra likely results from the fact that the HLY1502 analyses were performed on discrete samples, whereas the PS94 analyses depended on the estimate of the cartridge collection efficiency. In this paper we only discuss 226 Ra data of the PS94 expedition.

3. Results

3.1. Surface Distribution ²²⁸Ra

The surface distribution of ²²⁸Ra (Tables S2 and S3 and Figure 2a) shows the prominent enrichment of this isotope in the TPD. Earlier studies had shown this enrichment (Rutgers van der Loeff et al., 1995; Rutgers van der Loeff et al., 2012; Smith et al., 2003), but this is the first full transarctic section showing in a synoptic study the extension of the ²²⁸Ra plume with its limits in the Eurasian and the Canadian side of the TPD.

3.2. Surface Water ²²⁸Th-²²⁴Ra Systematics

From the two RaDeCC counts of uncoated and coated cartridges that had been collected in surface waters on the *Polarstern* expedition PS94, we calculated particulate and dissolved ²²⁴Ra and ²²⁸Th, decay/ingrowth corrected to sampling time (Table S2 and Figure 3). As expected, the particulate ²²⁴Ra activity was negligible, but particulate ²²⁸Th contributed up to 50% of total ²²⁸Th.

Excess 224 Ra (224 Ra $_{xs}$) is defined as 224 Ra not supported by 228 Th. In order to take into account the particulate 228 Th activity, we distinguish between dissolved and total 224 Ra $_{xs}$:

$$diss^{224}Ra_{xs} = {}^{224}Ra_{diss} - {}^{228}Th_{diss}$$
 (1)

$$Total^{224}Ra_{xs} = ^{224}Ra_{diss} + ^{224}Ra_{part} - ^{228}Th_{diss} - ^{228}Th_{part}$$
 (2)

Usually, ²²⁴Ra_{part} is negligible (cf. Table S2) giving

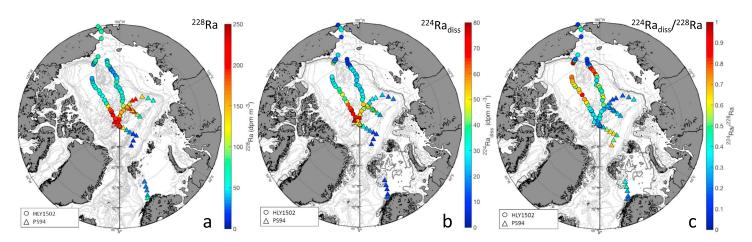


Figure 2. Distribution of (a) ²²⁸Ra, (b) ²²⁴Ra_{diss} as analogue for total ²²⁸Th, and (c) ²²⁴Ra_{diss}/²²⁸Ra as analogue for total ²²⁸Th/²²⁸Ra ratio in surface waters in 2015.

$$Total^{224}Ra_{xs} = ^{224}Ra_{diss} - ^{228}Th_{diss} - ^{228}Th_{part}$$
 (3)

Most of the dissolved ²²⁴Ra that was not supported by dissolved ²²⁸Th (i.e., dissolved ²²⁴Ra_{xs}) was actually supported by particulate ²²⁸Th (Figure 3). As expected from the shelf sources of ²²⁴Ra, total ²²⁴Ra_{xs} was a significant fraction of the total ²²⁴Ra on the Barents shelf (stations PS94–147 to 161). Although in an exceptional case Kadko and Muench (2005) once observed some excess ²²⁴Ra up to 200 km from the shelf source, we do not expect excess ²²⁴Ra in the central Arctic Ocean. Indeed, in the central Arctic, total ²²⁴Ra_{xs} is generally a small fraction of total ²²⁴Ra (Figure 3). Nevertheless, at many stations from the TPD, we find significant fractions of total ²²⁴Ra_{xs} reaching 40% of ²²⁴Ra at station PS94–96 (Figure 3). We cannot completely exclude some ²²⁴Ra release from ²²⁸Th adsorbed to the tubing of the ship, although we have tried to minimize this problem (see methods). Such a contamination would be expected to aggravate during passage in high-²²⁸Th waters. The observation that after passing the high-²²⁸Th waters of the TPD the absolute levels of total ²²⁴Ra_{xs} actually decreased (Figure 3: Stations 137, 138, and even the Barents Shelf stations 147–173) makes it unlikely that this contamination can explain the observed total ²²⁴Ra_{xs}. Such a contamination

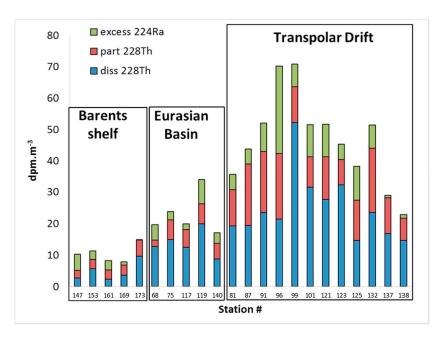


Figure 3. Dissolved (blue) and particulate (red) ²²⁸Th and total excess ²²⁴Ra activities (green) in surface waters of the PS94 expedition, grouped in the three regions Barents shelf, Eurasian Basin, and transpolar drift.



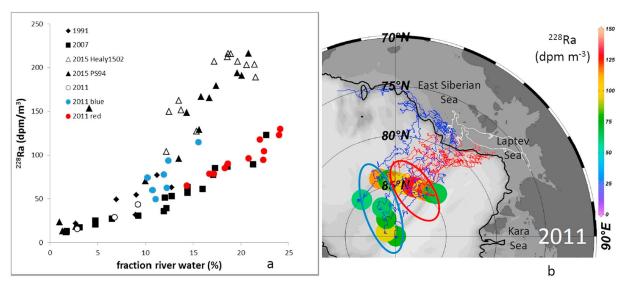


Figure 4. (a) ²²⁸Ra as function of fraction meteoric water for surface water samples collected North of 84°N in 1991, 2007, 2011, and 2015, highlighting stations in 2011 (PS78) with high ²²⁸Ra (>50 dpm m⁻³) that populate two groups with distinct slopes. Stations within the blue oval (blue dots in (a)) have higher ²²⁸Ra activities than stations within the red oval (red dots); (b) Ice back-trajectories (IBTs) of stations highlighted in panel a showing that the samples collected in 2011 with higher ²²⁸Ra activities (blue symbols in (a), blue oval and IBTs) have their origin further east than the samples with lower activities (red symbols in (a), red oval and IBTs). IBTs extending more than 2 years before sampling (Sta 224 and 227 from blue region) shown in white and 250-m depth contour highlighted in black.

cannot occur in the ISPs casts deployed on the Healy from 20 m downward. Two of the 20-m samples, station HLY1502–38 (87°N) and station HLY1502–46 (82°N), had significant total 224 Ra_{xs}. The total 224 Ra_{xs} observed in surface water at these latitudes may be due to 224 Ra released from 228 Th carried in ice-rafted particles.

3.3. Surface Distribution ²²⁸Th and ²²⁸Th/²²⁸Ra

Because our study is concentrated on the shelf signals observed in the central Arctic far offshore, we disregard total 224 Ra $_{xs}$ and estimate total 228 Th from dissolved 224 Ra (equation (3)) which is available for all cruises reported here (Tables S1–S4). For PS94 we have independent data for dissolved and particulate 228 Th, but for a consistent comparison with the other expeditions, we choose to present the distribution of dissolved 224 Ra as an analogue for total 228 Th (Figure 2b), an approach that may have caused a 17% overestimation of 228 Th (Figure S2b). The 228 Th/ 228 Ra activity ratio, which is low (order of 0.1–0.2) on the shelf as a result of scavenging (Rutgers van der Loeff et al., 2012), remains fairly low (<0.4) in the core of the TPD (Figure 2c). Only in the central Eurasian and Canada Basins are higher activity ratios reached.

3.4. Changes With Time in Surface Water 3.4.1. ²²⁸Ra

Kipp et al. (2018) observed that ²²⁸Ra in the TPD had increased from 2007 to 2015. They plotted ²²⁸Ra as a function of the fraction of meteoric water in order to show that the increase was not a result of a changed river water concentration that might have resulted from a change in freshwater accumulation in the Arctic (Rabe et al., 2014). Here we extend their plot with data from the *Polarstern* PS78 (2011) and PS94 (2015) expeditions (Figure 4a). The PS94 data from 2015 are in line with the parallel data collected on the Healy (Kipp et al., 2018), confirming the increase relative to 2007. The data from 2011 show two distinct trends: All samples collected on the detailed section over the Lomonosov Ridge closest to the Siberian shelf follow perfectly the mixing line we had observed in 2007. But the four samples with the highest ²²⁸Ra concentrations on the Canadian side of the cruise track are close to the mixing line we observed throughout in 2015, suggesting the existence of two different freshwater end-members in 2011.

3.4.2. ²²⁸Th/²²⁸Ra

In surface waters, 228 Th was more depleted with respect to its parent 228 Ra in 2015 than it was in the previous expeditions in 2007 and 2011 (Figure 5). The decrease of 228 Th/ 228 Ra is especially pronounced in the TPD: The 228 Th/ 228 Ra in the shelf-influenced waters north of 85°N with river water percentage > 13% are consistently low at about 0.3 in both the *Polarstern* PS94 and the Healy data set, compared to 0.5–0.65 in 2007.



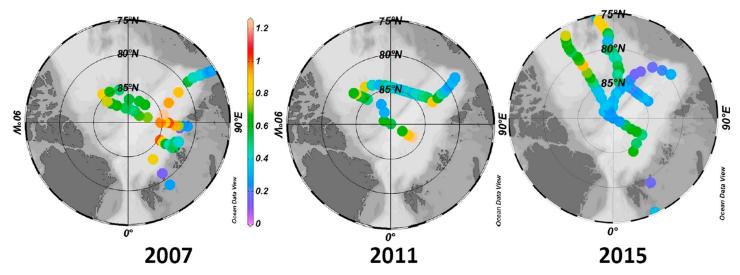


Figure 5. ²²⁴Ra/²²⁸Ra ratio as analogue for ²²⁸Th/²²⁸Ra ratio in surface waters in 2007, 2011, and 2015.

3.5. Distribution of ²²⁸Ra and ²²⁸Th in the Water Column

The distribution of ²²⁸Ra in the water column of the Eurasian and Makarov Basins (Figure 6a) is in agreement with earlier observations (Cochran et al., 1995; Rutgers van der Loeff et al., 1995; Trimble et al., 2004). The high ²²⁸Ra activities in the surface waters and especially in the TPD cause very strong gradients through the halocline. ²²⁸Ra activities in deep water (2,000–3,000 m) are low but still measurable and slightly increase toward the bottom.

We estimate the depth profiles of 228 Th from measurements of 224 Ra $_{diss}$ that we made at all stations. This implies that (1) we disregard excess 224 Ra as we did in the discussion of the surface concentrations (Figure 2b) and (2) we assume that total (dissolved + particulate) 228 Th is in secular equilibrium with dissolved 224 Ra. A comparison of dissolved 224 Ra with total (dissolved + small particulate + large particulate) 228 Th for the HLY1502 expedition confirms that 224 Ra is a reasonably good analogue for total 228 Th, although 224 Ra may underestimate total 228 Th by about 10% (Figure S2a).

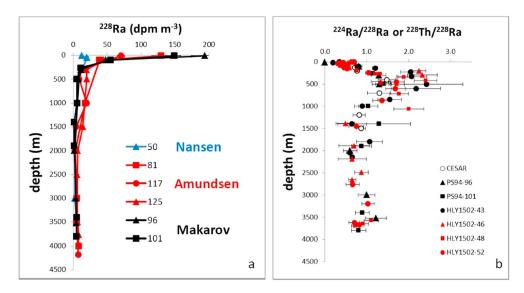


Figure 6. (a) Depth profiles of 228 Ra distinguishing PS94 stations from the Nansen (blue), Amundsen (red), and Makarov Basin (black) and (b) 224 Ra/ 228 Ra as analogue for 228 Th/ 228 Ra ratio observed in the Makarov (black) and Canada (red) basin compared with 228 Th/ 228 Ra data from ice station CESAR (Bacon et al., 1989).



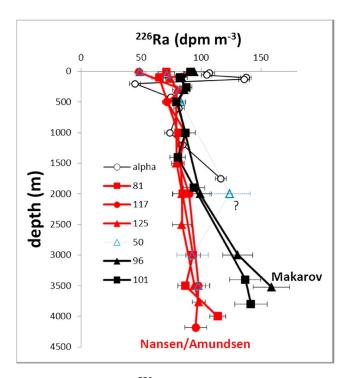


Figure 7. Depth profiles of 226 Ra distinguishing PS94 stations from the Nansen (open triangles), Amundsen (red), and Makarov Basin (black) compared to the profile measured in 1983 at CESAR ice station over the Alpha Ridge (open circles, Moore & Smith, 1986). We have no explanation for the high value at 2,000 m at Sta 50.

Two depth profiles of ²²⁸Th have been reported for the Nansen Basin by Cochran et al. (1995): one from the Barents Slope (their Sta. 287) and one from the central Nansen Basin (their Sta. 358). 228Th/228Ra ratios reported by these authors approached unity at mid depth in the central Nansen Basin but remained below 0.6 at the slope station with an average of 0.42. The only two ²²⁸Th/²²⁸Ra profiles we are aware of from the deep Canada Basin are the CESAR Ice station (Bacon et al., 1989) and station AWS2000-3 (Trimble et al., 2004). While at AWS2000-3 (75°13'N, 3,850 m) 228 Th/ 228 Ra increased gradually to a value of 1.1 \pm 0.4 at 2,000 m, the CESAR profile at the Alpha Ridge showed depletion around 0.75 in the upper 200 m, then a maximum of 1.48 \pm 0.21 at 400 m and a decrease to values below unity at 1,200 and 1,500 m. Such a subsurface maximum was also observed in the Atlantic Ocean by Li et al. (1980). Our new data (Figure 6b, ²²⁴Ra_{diss}/²²⁸Ra in Tables S1 and S3c) are in general agreement with the CESAR profile: They confirm the depleted ²²⁸Th/²²⁸Ra ratios in surface waters. Below the surface of the TPD both ²²⁸Th and ²²⁸Ra decline abruptly, but the ²²⁸Th/²²⁸Ra ratio increases to values usually in excess of unity in the 250-1,500 m depth range. In waters of 2,000 m and below, the ratio is usually below unity (0.75 \pm 0.05 standard error of mean, n=28).

3.6. ²²⁶Ra

Previous studies of Ra in the deep Arctic Ocean have mostly reported $^{228}\mbox{Ra}/^{226}\mbox{Ra}$ ratios where $^{226}\mbox{Ra}$ was estimated from relationships with Si (Cochran et al., 1995; Rutgers van der Loeff et al., 1995). Moore and Smith (1986) reported a profile at station CESAR over the Alpha Ridge and observed that the deepest sample (1,750 m) was enriched in $^{226}\mbox{Ra}$ by 40 dpm m $^{-3}$ relative to the samples at 500–1,000 m, much more than

the about 3 dpm m⁻³ expected from the increase in silicate of only 3 μ mol/kg. The depth distribution of ²²⁶Ra on cruise PS94 confirms an enrichment of ²²⁶Ra in deep waters (Figure 7). In the Nansen and Amundsen Basins the deep waters below approx. 2,000 m are enriched by about 10–15 dpm m⁻³, whereas the deep waters in the Makarov Basin are enriched by up to 80 dpm m⁻³ relative to the average concentrations in intermediate waters of 500–1,500 m (Figure 7 and Table S1). Integrating these excess activities with depth we find an excess inventory of 20,000 dpm m⁻² in the Nansen and Amundsen Basin and 70,000 dpm m⁻² in the Makarov Basin.

4. Discussion

4.1. Reasons for Change in ²²⁸Ra

Some ²²⁸Ra is supplied by rivers, but the major source of ²²⁸Ra in the Arctic is released from the seafloor where it is produced by decay of ²³²Th. The accumulation of ²²⁸Ra in shelf waters depends on the depth of the shelf sea and on the time the water resides over the shelf. The high ²²⁸Ra activities in surface waters of the TPD have traditionally been explained by the wide extent of Siberian shelf seas with depths decreasing eastward from the Barents and Kara Sea toward the very shallow East Siberian Sea. The correlation of ²²⁸Ra in offshore surface waters with salinity and with the fraction of meteoric water is due to the common nearshore source of ²²⁸Ra and meteoric water. This results in a virtual freshwater end-member that is not actually observed in river waters (Rutgers van der Loeff et al., 2003).

Our results from *Polarstern* 2015 (PS94) confirm the conclusion of Kipp et al. (2018) that the ²²⁸Ra concentration in the virtual freshwater end-member of the ²²⁸Ra on the Siberian shelf has roughly doubled since 2007 (Figure 4a). If this change had occurred gradually, we would have expected the results for 2011 to be intermediate between 2007 and 2015. It is very surprising that they are not, but indeed follow the mixing lines of 2007 and 2015 in two discrete geographical regions represented in Figure 4b by red and blue ovals, respectively. Ice back-trajectories (IBTs; Krumpen, 2017) show that the ice found in these regions was formed in distinct shelf sources (Figure 4b). To the extent that ice follows surface water flow, this implies that the



surface water at the stations with high meteoric water content and a lower virtual freshwater end-member ²²⁸Ra activity (indicated by red symbols in Figure 4a and by red ovals and red IBTs in Figure 4b) has a more westerly origin in the Laptev Sea region, whereas the high-²²⁸Ra water at the stations in the blue oval comes from a source further east, influenced by the East Siberian Shelf. This demonstrates that we cannot assume that all ²²⁸Ra values observed offshore in the Arctic Ocean have a similar virtual freshwater end-member. The ice at two stations in the blue oval (PS78–224 and 227) was more than 2 years old. The ice was formed in the Laptev Sea (white lines in Figure 4b) but then spent 15 months on the shelf north of the North Siberian Islands where the water can have accumulated additional ²²⁸Ra before being carried northward. The existence in 2011 of two source regions with distinct virtual ²²⁸Ra end-member activities raises the question whether the apparent differences in ²²⁸Ra activities between 2007 and 2015 (Figure 4a) can also be due to sources with distinct end-member concentrations resulting from changes in flow paths and/or shelf residence times.

Some insight into the possible role of circulation changes in the interannual differences between 228Ra at the surface can be gained from model experiments. We have used the coupled sea ice ocean model NAOSIM to disperse a conservative tracer released in Bering Strait (e.g., Aksenov et al., 2016) and the anthropogenic radionuclide ¹²⁹I, released from the European reprocessing facilities in La Hague and Sellafield (Karcher et al., 2012). While the former serves as a tracer for Pacific Water on its way across the Chukchi and East Siberian Seas, the latter, due to the dominant advection of the ¹²⁹I along the Siberian shelves from the Barents Sea via the Kara Sea into the Laptev Sea, can be used as an indicator for the pathway of Kara and Laptev Sea water into the central Arctic. Typically, the bulk of the waters in the Amerasian Basin stem from Pacific derived waters. The borderline between waters dominantly derived from the Pacific (via the East Siberian Sea) and the Atlantic (via the Kara and Laptev Sea), however, varies, due to changes of the surface circulation, for example, as a consequence of changing atmospheric conditions. Figure 8 shows the distribution of ¹²⁹I for September 2007, 2011, and 2015 in comparison with the distribution of ²²⁸Ra in these years. The Pacific water tracer (not shown) gives a largely complementary picture, as expected. For 2007 the position of stations with elevated ²²⁸Ra was over the Alpha Ridge and in the Makarov Basin close to the North Pole. The model simulation suggests that this is an area where the front between the water stemming from the East Siberian shelves and the Laptev/Kara Sea shelves was located in this period. Water from both source regions would be similarly likely to be found here. In 2011 the situation was different. For this year the model suggests, in agreement with the IBT analysis, that the stations closer to the New Siberian Islands (red oval in Figure 4b) are mostly derived from water stemming from the Laptev/Kara Sea. Only the stations closer to the North Pole (blue oval in Figure 4b) are likely to incorporate water stemming from the East Siberian Sea. For the year 2015 finally, the model simulation indicates a source region to be entirely from the Laptev/Kara Sea region for those stations with elevated ²²⁸Ra found in the Makarov Basin.

Essential for the accumulation of ²²⁸Ra is how long the water passed over shallow shelf areas. That is why we need actual flow paths that can be based on modeling (Figure 8) and on IBTs. Kipp et al. (2018) concluded from IBT analyses that the shelf origin of the high-²²⁸Ra waters found in the central Arctic was not different for the stations visited in 2015 by USCGC Healy or those from the 2007 study. This is confirmed by an IBT analysis of the Polarstern 2015 (PS94) study (Figure S3a). The history of the water masses before the time span of the IBTs was nevertheless quite different, as the NAOSIM modeling (Figure 8) suggested that the 2007 samples were from both sides of the Atlantic-Pacific divide, whereas the 2015 samples had a large Atlantic component. We have seen in the 2011 data that different histories can cause different ²²⁸Ra concentrations and a temporal increase in the residence time over a shelf could produce a water mass with high ²²⁸Ra concentrations. It would be of great value to have data on the variability of residence times of water over the Siberian shelves. But the increase in ²²⁸Ra concentrations in 2015 compared to earlier years is so widespread (Figure 8) that it is unlikely that it is caused by such a stagnation event and we conclude that indeed the input of ²²⁸Ra has increased. There are several possible explanations for this increase (Kipp et al., 2018). SGD is in many ocean areas a major pathway for inputs of ²²⁸Ra to the ocean (Moore et al., 2008). Due to the wide extent of permafrost it has been doubted whether SGD plays a role in the Arctic Ocean, but ²²⁸Ra inputs from this source have been identified by a recent study in the Laptev Sea (Charkin et al., 2017). However, because of the ubiquitous presence of ²³⁸U and ²³²Th in soils, SGD supplies ²²⁶Ra along with ²²⁸Ra. If an increase in SGD had resulted from an increase in the flow of old groundwater or from the exposure of recently thawed subsea permafrost to seawater circulation, we would expect their 228 Ra/ 226 Ra ratios to be closer to the crustal ratio of \sim 1.



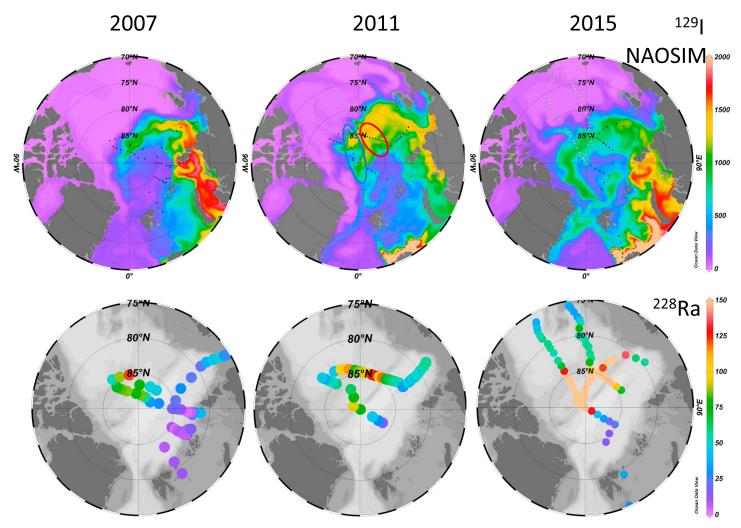


Figure 8. NAOSIM model simulation of the distribution of ¹²⁹I (10⁷ at/L) for the Septembers in 2007, 2011, and 2015 in comparison with the distribution of ²²⁸Ra (dpm m⁻³) in these years. The blue and red ovals in the middle upper panel represent regions where ²²⁸Ra and IBTs suggest that the surface water stems from the East Siberian and Kara/Laptev shelves, respectively (Figures 4a and 4b).

The relatively small enrichment of 226 Ra in the TPD (Kipp et al., 2018) is an indication that SGD is not a major contributor to the increase in 228 Ra.

Coastal erosion is increasing in the Arctic, enhanced by temperature rise and permafrost thaw. This will certainly contribute 228 Ra to coastal waters, but this source is insufficient to explain the increase of 228 Ra in the TPD (Kipp et al., 2018). Moreover, just like SGD, coastal erosion is a source of 226 Ra along with 228 Ra, and a parallel increase in 226 Ra is not observed.

The most likely explanation for the increase in ²²⁸Ra is therefore increased resuspension and pore water exchange due to the enhanced exposure of shelf sediments to wave action. As a result of sea ice retreat, larger areas of the continental shelf are ice-free for a longer period of the year. And such a sediment source would supply relatively higher amounts of ²²⁸Ra relative to ²²⁶Ra due to the faster regeneration time of the former.

4.2. Changes in Other Terrigenous Inputs in the TPD

In the TPD, waters with a high meteoric contribution flow from the Siberian shelves toward Fram Strait. The extent of this ~50–100 m thick shelf-influenced layer is clearly seen in the surface distribution of ²²⁸Ra (Figure 2a) but also is clearly visible in the distribution of various terrigenous components like Fe, Mn, and



DOM (Klunder et al., 2012; Middag et al., 2011; Slagter et al., 2017). Middag et al. (2011) and Klunder et al. (2012) argue that rivers are the main source of Mn and Fe to the TPD. Along with these trace metals, the TPD also transports ligands and other DOM.

Just like ²²⁸Ra (Figure 4a), other terrigenous components like dissolved Fe (Klunder et al., 2012; Slagter et al., 2017) and DOM (Stedmon et al., 2011) are correlated with the fraction of meteoric water, while Fe and iron-binding organic ligands are correlated with CDOM and humic substances (Slagter et al., 2017). This raises the question of whether the increase in ²²⁸Ra is accompanied by a similar increase in these other terrigenous constituents. In contrast to ²²⁸Ra, where shelf inputs are predominant and we defined a virtual freshwater end-member, DOM and other terrigenous components like REE (Laukert et al., 2017) have primarily freshwater sources. The correlation between ²²⁸Ra and such river borne terrigenous components (not shown) is therefore in part fortuitous because the sources of these terrigenous inputs do not coincide. Inputs of Fe, Fe ligands, Mn, and CDOM to the TPD will only increase along with ²²⁸Ra to the extent that their sources are on shelf sediments, not in the rivers.

4.3. Reasons for Change in ²²⁸Th/²²⁸Ra

The development of 228 Th/ 228 Ra in the TPD with time t since the water mass lost contact with the shelf follows (Rutgers van der Loeff et al., 2012, note typo corrected here)

$$A_T = A_{T^0} e^{-(\lambda_T + \lambda_s)t} + \frac{\lambda_T}{\lambda_T + \lambda_s - \lambda_R} A_{R^0} \left(e^{-\lambda_R t} - e^{-(\lambda_T + \lambda_s)t} \right)$$

$$\tag{4}$$

where A_T and A_R are the activities of ²²⁸Th and ²²⁸Ra, λ_T and λ_R the decay constants of ²²⁸Th and ²²⁸Ra, respectively, λ_s is the Th scavenging rate, and superscript ⁰ denotes the situation at t=0 when the water leaves the shelf. In equation (4) we assume that lateral mixing is unimportant relative to advection in the TPD. In our 2007 study we found that ²²⁸Th/²²⁸Ra was about 0.15 on the shelf and increased to 0.4–0.6 in the TPD in the central Arctic. In 2011, ²²⁸Th/²²⁸Ra in samples in the TPD ($f_r > 15\%$) ranged from 0.31 to 0.49, while in 2015, the ²²⁸Th/²²⁸Ra values were significantly lower (0.2–0.35) than in 2007. In all these surface water data the use of dissolved ²²⁴Ra as analogue for total ²²⁸Th may have caused an overestimation of ²²⁸Th (Figure S2b), but the procedure was the same in all three expeditions and cannot have caused the observed trend. The decrease could be due to one or a combination of several reasons:

Recent increase in ²²⁸Ra: equation (4) assumes a constant source concentration of ²²⁸Ra while we have found an increase over the past 8 years. The ingrowth of ²²⁸Th will lag behind. However, because the time change (8 years) is slow relative to the half-life of ²²⁸Th, we disregard this lag effect.

Lower depletion on the shelf: In 2015 we have few data of 228 Th_{part + diss}/ 228 Ra from the eastern Arctic shelves. On the Barents shelf we found 0.18–0.19, and the lowest values on the approach of the Laptev shelf were 0.14–0.16. We assume that 228 Th/ 228 Ra has remained depleted to about 0.15 before leaving the shelf.

A higher scavenging rate (lower λ_s) would counteract the ingrowth and explain lower ²²⁸Th/²²⁸Ra values. If sufficient nutrients are available, the reduction in ice cover could cause increased productivity (Arrigo et al., 2008) and consequently increased export flux and scavenging. Net community production and export in the central Arctic Ocean have been estimated using several approaches. We have already mentioned that export production measured with ²³⁴Th/²³⁸U and with ²¹⁰Po/²¹⁰Pb was very low in 2007 and 2011. New ²³⁴Th data from 2015 show essentially no increase in export productivity (Black, 2018). In 2011, Ulfsbo et al. (2014) found enhanced net community production values on the shelves and ice margin, but in the central Arctic their O₂/Ar, pCO₂, and nutrient drawdown estimates are very low and do not appear to have increased in 2015 (Ulfsbo, pers. comm.).

Alternatively, the ice melt could have caused an increase of scavenging rate through the release of ice rafted particles that previously used to be carried on to Fram Strait.

Finally, the lower ²²⁸Th/²²⁸Ra ratios could result from a more rapid transport of the TPD. The IBTs of the Healy stations crossed the shelf break (200-m isobaths) 8–18 months before sampling (Kipp et al., 2018). For the *Polarstern* study this was 5–11 months, clearly shorter than the corresponding time for the IBTs of the 2007 study (7–27 months, Figure S3b). The observed trend (1992–2009) toward increasing ice drift speed has



been related to the thinning of the ice cover and to changed wind forcing (Spreen et al., 2011). The roughly 9 months shorter transport time in 2015 compared to 2007 would explain much less ingrowth of ²²⁸Th into parent ²²⁸Ra during transit (equation (4)) and 9% higher ²²⁸Ra values as a result of less decay.

4.4. Depth Distribution of ²²⁸Ra

The major source of 228 Ra in the central Arctic Ocean is the supply of shelf-derived water to the low-salinity surface layer. As a first approximation of the depth distribution we therefore use a one-dimensional model with a source of high 228 Ra concentrations in the surface mixed layer (depth MLD). Thereafter, the depth distribution is controlled by vertical diffusion (eddy diffusion coefficient k_z) and radioactive decay and follows an exponential decay.

$$A_R = A_{R^0} e^{-\sqrt{\frac{2}{R_2}}(z-MLD)}$$
 (5)

With this equation and using k_z of order 10^{-4} m² s⁻¹ (Munk, 1966), ²²⁸Ra falls below 1 dpm m⁻³ at 1,000 m and would be below our detection limit in deeper layers. ²²⁸Ra is also released from the deep-sea floor. Close to the seafloor we observe activities up to 8.2 dpm m⁻³ (Sta. PS94–81). Even with a 10 times higher vertical diffusivity in the benthic boundary layer, this input would fall below 1 dpm m⁻³ at 1,100 m above the seafloor. Nevertheless, all intermediate water values are above our detection limit. The lowest activities observed on the PS94 expedition are down to 1.14 ± 0.31 dpm m⁻³ in the Makarov Basin (Sta. PS94–101 and Sta. PS94–96), but in the Eurasian Basin, all values are above 1.9 ± 0.4 dpm m⁻³. The higher activities must be supplied by horizontal diffusion and advection from the slopes, similar to the situation observed during the GEOSECS program in the NW Atlantic (Sarmiento et al., 1982). In the Arctic Ocean, this process is enhanced by slope convection resulting from brine rejection during sea ice formation (Rudels et al., 2000). Analysis of the distribution of Al, Si (Middag et al., 2009), and Ba (Roeske et al., 2012) showed that shelf sources were a major contribution to the enrichment of these elements in the deep basins. We expect that slope convection contributes to Ra enrichment at all depths in a similar way.

Intermediate depth waters—particularly high subsurface ²²⁸Ra activities—are observed in the Amundsen Basin down to approximately 1,500 m. The circulation of intermediate depth waters in the Arctic Ocean has been reviewed by Rudels (2015). Based on hydrographic (T, S, and Si) and CFC data, Rudels et al. (1994) showed how the Atlantic input splits into a Fram Strait Branch and a Barents Sea Branch. These branches meet north of the Kara Sea, and the resulting intermediate depth water mass is strongly affected by inputs from the shelves, via inputs to the Barents Sea Branch itself and through further sinking shelf plumes. These authors found a return flow of Atlantic water at 200–1,700 m in the Amundsen Basin with a residence time on the order of a decade ("the deeper layers being the oldest"). ²²⁸Ra data collected in 1987 (Rutgers van der Loeff et al., 1995) gave the first evidence that this return flow had enhanced ²²⁸Ra concentrations through its contact with the shelves. Samples taken from a submarine cruising at 132 m (Kadko & Aagaard, 2009) showed a maximum in the Amundsen Basin (their Sta. 6), thought to be derived from the Laptev or Kara Sea after a transit of at least 2 years based on estimates of Woodgate et al. (2001) that the water speed along the Eurasian side of the Lomonosov ridge is 2 cm/s.

Our new data collected in 2015 show clearly enhanced 228 Ra activities in intermediate waters up to about 1,500 m in the Amundsen Basin compared to the Nansen and Makarov Basin (Figure 6). At 1,000 m, we find 18 and 20 dpm m $^{-3}$ at Stas PS94 $^{-81}$ and PS94 $^{-117}$, respectively, compared to only 5 dpm m $^{-3}$ at Sta. PS94 $^{-101}$ in the Makarov Basin. SF $_6$ data on the PS94 transect (Figure 9, full section in Figure S4, data in Smethie, 2017; Smethie & Swift, 2018) across the Amundsen Basin show indeed that at the depth horizon of 1,000 m this basin is locally enriched at Sta. PS94 $^{-81}$ to $^{-81}$ fmol/kg, corresponding to a ventilation age of 15 $^{-81}$ 8 years, if before ventilation the surface water had reached equilibrium with the atmosphere. We assume that the Atlantic water mass carrying 20 $^{-30}$ dpm m $^{-3}$ when entering the Arctic through the Barents Sea Opening and Fram Strait (Key et al., 1992) is enriched up to concentrations of order 150 dpm m $^{-3}$ in the Kara and Laptev Sea (Rutgers van der Loeff et al., 2003), while the SF $_6$ content equilibrates with the atmosphere. In a purely advective transport of this water mass as intermediate water in the Amundsen Basin, the SF $_6$ content would remain unchanged, while 228 Ra would decay to the observed 18 dpm m $^{-3}$ in 18 years, which means that the time scale of the two tracers matches. A further independent support for this time scale

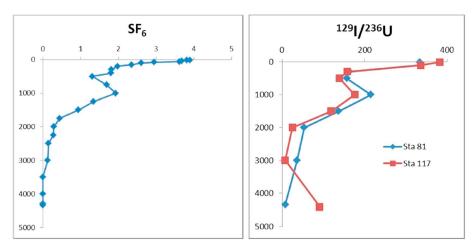


Figure 9. Depth profile of SF₆ (left, fmol/kg) at station PS94–81 and of 129 l/ 236 U (right, at/at) at station PS94–81 (blue) and PS94–117 (red) in the Amundsen Basin showing ventilation of the intermediate water up to about 1,500 m.

comes from the $^{129}I/^{236}U$ signal (Figure 9). The relatively high $^{129}I/^{236}U$ ratios of 215 and 176 for Stas PS94–81 and 117 observed at 1,000 m, respectively, could only be explained by waters that were released after 1998 by the European Reprocessing Plants of Sellafield and La Hague (Casacuberta et al., 2016; Christl et al., 2015). Although transit times using the $^{129}I/^{236}U$ atom ratio cannot yet be calculated as precisely as for SF₆, they would still generally agree with the ones estimated based on SF₆ and ^{228}Ra .

4.5. A Model of the Distribution of ²²⁸Th, ²²⁸Ra, and the ²²⁸Th/²²⁸Ra Ratio

A soluble parent with a particle reactive daughter with half-life of 1.9 years, the ²²⁸Th/²²⁸Ra tracer pair appears ideally suited as a tracer for export production in the Arctic where ²²⁸Ra is abundant. Compared to the ²³⁴Th/²³⁸U tracer pair, the longer half-life has the advantage of averaging out short-term variations within the productive season, and it creates a much larger depletion relative to the soluble parent that can be determined with better precision. Here we model the vertical distribution of ²²⁸Ra and ²²⁸Th in order to

TPD shelf

TR

Intermediate water

Sc

4000 F

B

TPD

Shelf

Figure 10. Profile model of the distribution of 228 Ra and 228 Th with rapid mixing down to the mixed layer depth (MLD), lateral input at the surface in the TPD, ventilation of the upper 1,500 m with shelf water from the Kara Sea (time scale $\tau_{\rm K}$), exchange at all depths with 228 Ra-enriched bottom waters at the slope (time scale $\tau_{\rm M}$) with 228 Ra-enrichment maintained by slope convection (sc), and an input F_B from the seafloor. The vertical particle flux (black arrow) carrying particulate 228 Th is modeled using data of 234 Th and 230 Th.

investigate to what extent these tracers can be used to determine particle flux and export production.

We use a profile model (Figure 10, equations provided in the supporting information) where ²²⁸Ra and ²²⁸Th are introduced in the surface layer. We represent the supply in the TPD by setting the activities of ²²⁸Ra and ²²⁸Th in the surface mixed layer (depth MLD) at fixed values ²²⁸Ra° and ²²⁸Th°. Below the surface mixed layer, vertical eddy diffusion is set at 1.3 10^{-4} m² s⁻¹ (Munk, 1966). ²²⁸Ra release from the seafloor sediment causes an input F_B of 250 dpm m⁻² yr⁻¹ corresponding in steady state to a deepwater inventory of 2,000 dpm m⁻². Moreover, the supply of ²²⁸Ra from slope sediments is represented as an exchange at all depths with a water mass at the margin with ²²⁸Ra_M = 25 dpm m⁻³ and ²²⁸Th_M = 10 dpm m⁻³ with time constant τ_M . To represent the special case of the Amundsen Basin, the water mass from MLD to 1,500 m is renewed with Kara Sea shelf water with composition ²²⁸Ra_K and ²²⁸Th_K with time constant τ_M . This input is balanced by outflow through FRAM Strait.

In the water column below depth MLD, 228 Ra decays with a half-life of 5.75 years to 228 Th, which decays with a half-life of 1.9 years. For 228 Th we use the reversible scavenging model of Nozaki et al. (1981) and Bacon and Anderson (1982) in which the particulate and dissolved forms are in continuous exchange with adsorption rate k_1 and desorption rate k_{-1} . We calculate these rate constants from observations of the



Table 1Parameters of the Profile Model Representing ²²⁸Ra and ²²⁸Th

Parameter	Symbol	Value		Unit	Sensitivity analysis	
		Nansen/Amundsen	Makarov	Canada		
Mixed layer (ML) depth	MLD		100		m	
Eddy diff coeff	k_z		4,100		$m^2 yr^{-1}$	410-4,100
Benthic ²²⁸ Ra flux	Fo		250		$dpm m^{-2} yr^{-1}$	
Ra in ML	²²⁸ Ra°	157		55	$dpm m^{-3}$	
²²⁸ Th in ML	²²⁸ ть°	47		25	dpm m ⁻³	
²²⁸ Ra at margin	228 _{D2}		25		dpm m ⁻³	
²²⁸ Th at margin	²²⁸ Th _M		10		dpm m ⁻³	0-25
Exchange time with margin	τ		50		yr	1-inf.
²²⁸ Ra Kara Sea	228 _{Pau}		150		$dpm m^{-3}$	
²²⁸ Th Kara Sea	²²⁸ Th _K		0		$^{\cdot}$ dpm m $^{-3}$	
$C_{\rm p}/C_{\rm d}$ ²³⁰ Th ^b	K ₂₃₀	0.5	0.152	0.152 ^a		
C _p /C _d ²³⁴ Th ^b	K ₂₃₄	0.12	0.026	0.026 ^a		
Adsorption rate constant	k ₁	1.59	0.324	0.324 ^a	yr ⁻¹	
Desorption rate constant	k_1	3.18	2.132	2.132 ^a	yr^{-1}	
C _d ²³⁰ Th at 4,000 m	²³⁰ Th _{d,4000}	0.35	1.55	1.55 ^a	$dpm m^{-3}$	
Particle settling rate	S S	582	434	434 ^a	m yr ⁻¹	
Exchange time with Kara Sea	τ_{K}	50	inf	inf	yr	1-inf.

^aAssumed identical to Makarov. ^bValk, pers. comm.

distribution of the Th isotopes 230 Th and 234 Th (Valk, pers. comm.). For every isotope, the ratio of particulate to dissolved Th is given by $C_p/C_d = k_1/(\lambda + k_{-1})$; equation (13) in Bacon & Anderson, 1982). From the ratio of particulate to dissolved activities of 230 Th and 234 Th in the deep Nansen Basin of 0.5 and 0.12, respectively, we find values for k_1 and k_{-1} of 1.6 and 3.2 yr $^{-1}$, respectively. Applying the reversible exchange model to the distribution of the long-lived isotope 230 Th, C_p increases linearly with depth z as $C_p = Pd/S^*z$, where Pd is the 230 Th production rate and S is the particle sinking rate. From the distribution of 230 Th, with dissolved 230 Th increasing to 7.5 fg/kg (0.35 dpm m $^{-3}$) at 4,000 m in the Nansen/Amundsen Basin, we find a value for the average particle sinking rate S of 582 m yr $^{-1}$. For the Makarov Basin, where dissolved 230 Th increases to 33 fg/kg (1.55 dpm m $^{-3}$) at 4,000 m, we find rate constants k_1 and k_{-1} of 0.3 and 2.1 yr $^{-1}$, respectively, and a particle settling rate S of 434 m yr $^{-1}$.

The differential equations in the model (supporting information) were solved for ²²⁸Ra and ²²⁸Th in the water column below the surface mixed layer using programming language R based on scripts presented by Soetaert and Herman (2008). All parameters are listed in Table 1.

4.5.1. ²²⁸Ra Model

The time scales of the lateral exchange processes with the slope (τ_M) and by ventilation with the Kara Sea (τ_K) are modified (Table 1) in order to obtain the best representation of the observed 228 Ra profiles. Without input from the slope, 228 Ra becomes practically zero from 2,000 to 4,000 m depth. With a margin exchange time of 50 years, 228 Ra in this depth range becomes 3.5 dpm m $^{-3}$ (Figure 11a), which fits with observations. The best fit for 228 Ra in the intermediate water in the Amundsen Basin is found for a renewal time of intermediate water with Kara Sea surface water of 50 years. This yields a 228 Ra activity of 22 dpm m $^{-3}$ at 1,000 m (Figure S5). This renewal time is longer than the pure decay time of 18 years found above because here we do not use a complete replacement of the water in a purely advective model but rather a continuous flushing of a well-mixed water mass.

4.5.2. ²²⁸Th Model

If 228 Ra followed equation (5), that is, in the absence of lateral inputs, and if thorium were unreactive, that is, in the absence of scavenging in the water column ($\lambda_s = 0$) the activity of 228 Th (A_7) would be described by equation (6) in Rutgers van der Loeff et al. (2012)

$$A_{T} = \frac{\lambda_{T} A_{R^{0}}}{\lambda_{T} - \lambda_{R}} e^{-\sqrt{\frac{\lambda_{R}}{k_{z}}}(z - MLD)} + \left(A_{T^{0}} - \frac{\lambda_{T} A_{R^{0}}}{\lambda_{T} - \lambda_{R}}\right) e^{-\sqrt{\frac{\lambda_{T}}{k_{z}}}(z - MLD)}$$

$$\tag{6}$$

and at great depth the activity ratio would approach



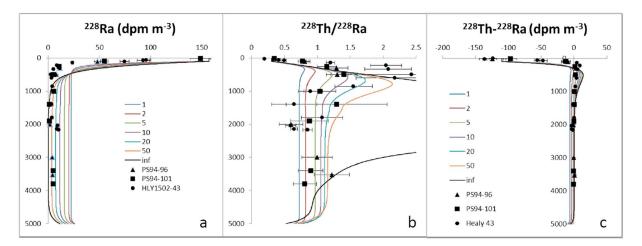


Figure 11. Model run for 228 Ra, 228 Th/ 228 Ra, and 228 Th- 228 Ra in the Makarov Basin varying the time scale $\tau_{\rm M}$ (years) for exchange with the margin. High 228 Ra values near 2,100 m at station HLY1502–43 are due to proximity to seafloor.

$$\frac{A_T}{A_R} \rightarrow \frac{\lambda_T}{\lambda_T - \lambda_R} = 1.49 \tag{7}$$

(cf. Rutgers van der Loeff et al., 2012, their Figure 11). Using our model for ²²⁸Ra and the reversible exchange model for Th with rate constants determined for the Makarov Basin (Table 1), we find that the release of ²²⁸Th from particles settling out of the mixed layer causes an enrichment of ²²⁸Th over ²²⁸Ra with a maximum AR of 2.14 at a depth of 780 m (Figures 11b and 12). At greater depth, the AR stabilizes at a value of 1.11. This result is in reasonable agreement with observations in the Makarov Basin (Stas PS94–96 and 101, HLY1502–43, Figures 11b and 12a).

We find that the distribution of 228 Th/ 228 Ra and of the excess 228 Th (228 Th $_{xs} = ^{228}$ Th – 228 Ra) are very sensitive to the assumed parameters for the exchange with the margin (Figures 11b and 11c and 13). If no exchange is taken into account (exchange time scale infinity), ²²⁸Ra will tend to zero at middepth and the ²²⁸Th/²²⁸Ra ratio increases to very high values. ²²⁸Th_{xs} reaches the highest values at 510 m depth and is zero in the deep water. A rapid exchange with the margin (time scale 1 year) causes higher ²²⁸Ra concentrations and lower 228 Th/ 228 Ra ratios. If we select the exchange rate of 50 years that best fits the observed 228 Ra profile and vary the degree of 228 Th/ 228 Ra disequilibrium (activity ratio AR_M) in the water mass at the margin that is the source for the lateral supply of ²²⁸Ra, we find that especially the deepwater ²²⁸Th/²²⁸Ra ratio is strongly dependent on the ²²⁸Th in the source water (Figure 12a). If the water at the margin is depleted in ²²⁸Th by scavenging at the slope (AR_M = 0), the modeled 228 Th_{xs} at depth becomes negative (228 Th/ 228 Ra = 0.95). If however 228 Th is in equilibrium with 228 Ra in the margin water (we use AR_M = 1 and not 1.49 because of the proximity of 228 Ra to its 232 Th parent), the 228 Th_{xs} remains positive (228 Th/ 228 Ra = 1.31). The result for AR = 0.4, the value found at the Barents Slope by Cochran et al. (1995), is intermediate (228Th/228Ra = 1.15). This sensitivity analysis shows that the ²²⁸Th/²²⁸Ra ratio in the deepwater column (especially >1,500 m) is a delicate equilibrium between the lateral supply of ²²⁸Ra and ²²⁸Th and the vertical rain of particulate ²²⁸Th. In the upper 1,000 m, the modeled ²²⁸Th/²²⁸Ra ratio is strongly dependent on the proper choice of the Ra inputs (Figure 11), but if the ²²⁸Ra profile is properly resolved with good measurements of ²²⁸Ra, the measured 228 Th_{xs} is only slightly dependent on the assumptions of the composition of the margin water (Figure 12b). This implies that the cumulative 228 Th_{xs} in the upper 1,500 m does not depend on lateral supply. The excess does depend on the inherent excess resulting from a parent/daughter pair where the decay of the parent is significant (equations (6) and (7)). This contribution can be estimated by setting the settling rate S at 0 (dashed line in Figures 12b and 12c).

In our model, the 228 Th_{xs} integrated from the depth where it becomes positive down to 1,500 m amounts to 7,700 dpm m $^{-2}$ (varying between 6,900 and 8,400 dpm m $^{-2}$ if AR_M is varied between 0 and 1), while in the absence of particle flux, the excess is only 1,500 dpm m $^{-2}$. At Stas PS94–96, PS94–101, and HLY1502–43 we

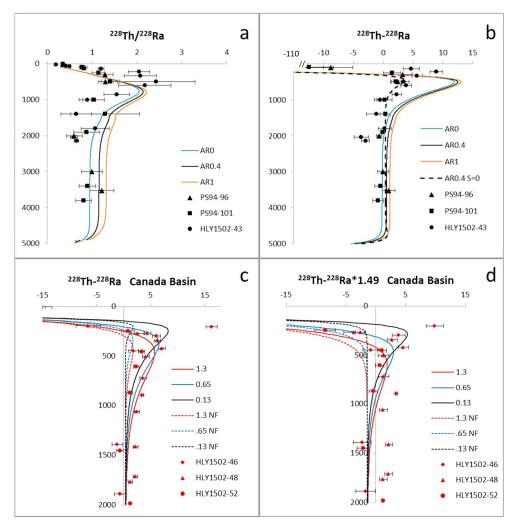


Figure 12. Depth distribution of (a) 228 Th/ 228 Ra ratio and (b) 228 Th/ 228 Ra activity ratio at the margin (AR_M) varying between 0 and 1. Dashed line in (b): modeled 228 Th/ 88 Ra (AR_M = 0.4) without particle flux. (c) 228 Th- 228 Ra (based on 224 Ra- 228 Ra) and (d) 228 Th-1.49* 228 Ra (dpm m $^{-3}$) in the Canada Basin compared with model results with (full lines) and without (NF, dashed lines) particle flux, for the standard (1.3 10 m 2 s $^{-1}$) and two reduced values (0.65 10 and 0.13 10 m 20 s $^{-1}$) of the vertical diffusion coefficient k_z , using surface activities of 228 Ra and 228 Th of 55 and 25 dpm m $^{-3}$, respectively, and for further parameters the same values as in (a) and (b).

found an integrated $^{228}\text{Th}_{xs}$ of 2,040 \pm 900, 1,160 \pm 550, and 3,400 \pm 440 dpm m $^{-2}$ (Figure 12b), not or not much exceeding the zero-particle-flux estimate.

In the Canada Basin, south of the core of the TPD, 228 Ra in surface water is much lower, which reduces the theoretical diffusive (i.e., zero-particle-flux) contribution to the 228 Th excess. At GN01 Stas HLY1502–46, 48, and 52, 228 Ra and 228 Th are on average 55 and 25 dpm m $^{-3}$, respectively. Using for the other parameters the same choice as in the Makarov Basin, we expect a 228 Th excess of 3,650 dpm m $^{-2}$ compared to a zero-flux excess of 910 dpm m $^{-2}$ (Figure 12c). The 228 Th excess actually observed agrees at Stas HLY1502–46 and 48 (3,710 \pm 850 and 4,260 \pm 210 dpm m $^{-2}$) with this estimate, while the excess observed at Sta HLY1502–52 (1,510 \pm 330 dpm m $^{-2}$) only slightly exceeds the zero-flux estimate.

The maximum excess is observed at shallower depth than our model predicts. We may have overestimated the vertical diffusion coefficient, which is strongly reduced in the highly stratified surface layers of the Arctic Ocean. Figure 12c shows the sensitivity of the model results for the vertical diffusion coefficient k_z . A reduction of k_z by a factor of 2 or 10 does not affect the total predicted ²²⁸Th excess, but it shifts the maximum

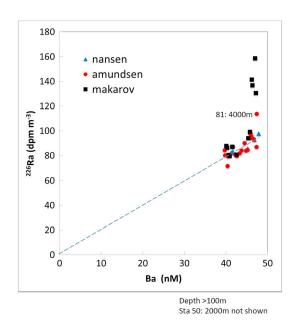


Figure 13. 226 Ra as function of Ba for all PS94 data in excess of 100 m depth showing that the accumulation of 226 Ra in deep water of the Makarov Basin is not associated with an accumulation of Ba. One unlikely data point of station PS94–50 (2,000 m) not shown.

excess to shallower depth. At the same time, it reduces the zero-flux-estimate of the cumulative excess from 910 to 780 or 604 dpm $\,\mathrm{m}^{-2}$, respectively (Figure 12c).

An alternative way to take the inherent ²²⁸Th excess in the unsupported ²²⁸Ra decay chain into account is by plotting ²²⁸Th-1.49*²²⁸Ra (Figure 12d). The zero flux model results no longer show a subsurface maximum. The asymptotic value of (228Th-1.49*228Ra) approached at depth in the model results of Figure 12d is negative as a result of the lateral supply included in the model. When we integrate (228Th-1.49*228Ra) from the depth where it becomes positive down to 1,500 m, we find 2,048 \pm 614, 2,094 \pm 296, and $192 \pm 170 \text{ dpm m}^{-2}$ for Stas HLY1502–46, 48, and 52, respectively. In the Makarov Basin we find 1,476 \pm 306 dpm m⁻² at Sta HLY1502–43, while ²²⁸Th never exceeds 1.49*Ra at stations PS94–96 and 101 (Table 2). These excess inventories at depth can be compared with the depletion of ²²⁸Th in the surface water, roughly (228Ra-228Th)*MLD, which amounts to 11,000 and 3,000 dpm m⁻² in the Makarov and Canada Basin, respectively (Table 1). We conclude that the ²²⁸Th depletion in the Makarov Basin is largely advected in the TPD and not related to local particle export, whereas at Stas HLY1502-46, and 48 in the Canada Basin local export accounts for a larger part of the depletion and advection is of less importance.

Assuming steady state, the excess inventories can be converted to the vertical 228 Th flux (multiplying by λ_T) and then to the scavenging rate in the

mixed layer (dividing by the ²²⁸Th inventory in the MLD, Table 2). These results will be used in a subsequent paper dealing with carbon export fluxes.

4.6. ²²⁶Ra and Deep Water Ventilation

It is tempting to use the accumulation of 226 Ra as an independent approach to estimate residence times of deep waters in the Arctic Basins. Broecker and Peng (1982) discussed the accumulation of 226 Ra in the deep waters of the World Ocean, based on data collected in the GEOSECS program. They considered that the accumulation was due to remineralization of 226 Ra from settling particles and to inputs from decay of 230 Th in deep-sea sediments. In principle, the two sources could be separated using Ba as analogue for a stable isotope of Ra, but they showed that in a two-layer ocean model the expected accumulation of the 226 Ra/Ba ratio in the deep sea due to the bottom source of 226 Ra was very small. The observed changes with very high accumulations in the deep NW Pacific were interpreted as result of inhomogeneities in the bottom source and in regeneration, making its use as deepwater ventilation tracer problematic. These problems are less restrictive in the Arctic Ocean with its relatively small and enclosed deep basins. Limited data on 230 Th_{xs} activities in surface sediments (Hoffmann et al., 2013; Moran et al., 2005) show a general pattern of higher activities in the low productivity, permanently ice-covered interior Arctic, with average 230 Th_{xs} in the upper 10 cm of the sediment of 9.9 dpm g $^{-1}$ for the Nansen Basin (AOS94 core BC 32, 3,471 m) and 13.2 dpm g $^{-1}$ for the Makarov

Table 2Integrated Values of $\binom{228}{1}$ Th- $\binom{228}{1}$ Ra*1.49) and Corresponding Vertical $\binom{228}{1}$ Th Flux and Scavenging Rate λ_s in the Mixed Layer

3	•		, ,	3 3		•		
Station	MLD	²²⁸ Ra _{MLD}	²²⁸ Th _{MLD}	$\Sigma (^{228}\text{Th-}^{228}\text{Ra*}1.49)^{a}$ 100–1,500 m	±	Flux	τ_{S}	λ_{S}
	m	${\rm dpm}~{\rm m}^{-3}$	${\rm dpm}~{\rm m}^{-3}$	${ m dpm\ m}^{-2}$		$\mathrm{dpm}\;\mathrm{m}^{-2}\mathrm{yr}^{-1}$	yr	yr ⁻¹
PS94-96	100	157	47	- 		-,-		
PS94-101	100	157	47			-,-		
1502-43	100	157	47	1,476	306	538	8.7	0.11
1502-46	100	55	25	2,048	614	747	3.3	0.30
1502-48	100	55	25	2,094	296	764	3.3	0.31
1502–52	100	55	25	192	170	70	35.7	0.03

Note. -.-: no positive values.

aPositive values only.



Basin (BC 20, 3145 m; Hoffmann et al., 2013). The influence of remineralization on the accumulation of 226 Ra in the water column can be judged from a plot of 226 Ra versus Ba (Figure 12, Rember, 2017). Apart from a single data point at 4,000 m in the Amundsen Basin at Sta PS94–81, data from the deep Eurasian Basin do not stand apart from the general 226 Ra/Ba correlation and we conclude that in the Eurasian Basins the accumulation of 226 Ra cannot be unequivocally attributed to sediment sources. In the Makarov Basin, however, the accumulation of 226 Ra is not associated with a corresponding accumulation of Ba. In this case, as in the case of the 226 Ra accumulation observed at the Alpha Ridge by Moore and Smith (1986), we conclude that 226 Ra accumulation is due to release from bottom sediments. Following Cochran (1980) we estimate the 226 Ra release rate $F_{B,226}$ from the 230 Th inventory in the upper layer of the sediment. Using a 230 Th concentration of 14.8 dpm g $^{-1}$, a mixed layer of 8 cm as used by Cochran (1980), and a dry bulk density of 0.5 g dry material per cm 3 of sediment, we find a 226 Ra flux of 176 dpm m $^{-2}$ yr $^{-1}$. This flux estimate varies between 156 and 180 dpm m $^{-2}$ yr $^{-1}$ if we allow for the range of mixed layer depths reported for AOS94 cores (2–10 cm, Smith et al., 2003). This flux must be balanced by decay and ventilation:

$$F_{B,226} = (\lambda_{226} + 1/\tau_{MD}) I_{226}$$
 (8)

where λ_{226} is the decay constant of ²²⁶Ra, τ_{MD} is the residence time of water in the deep Makarov Basin, and I_{226} is the inventory of excess ²²⁶Ra in the deep water. With the inventory derived above (70,000 dpm m⁻², Figure 7) we find a deepwater residence time of ~480 years, which is consistent with the estimated isolation age for the deep Canada and Makarov Basin of 450 years based on ¹⁴C and ³⁹Ar (Schlosser et al., 1994; Schlosser et al., 1997).

5. Conclusions

When plotted as function of fraction meteoric water, ²²⁸Ra has increased in the TPD from 2007 through 2011 to 2015. This change is most likely due to increased wave action on the shelves as a result of the longer ice free season. Whether the increase in ²²⁸Ra flux is associated with increased fluxes of other terrigenous components requires continued studies of the ratio of these components to ²²⁸Ra. The concomitant decrease of the ²²⁸Th/²²⁸Ra ratio likely results from a more rapid transit of surface waters depleted on the shelf and consequently less time for ingrowth of ²²⁸Th.

In the TPD, particulate 228 Th accounts for most of the dissolved 224 Ra $_{xs}$. This implies a concept of excess 224 Ra that is very different from common use, that is, in most studies of coastal mixing 224 Ra $_{xs}$ is assumed to be derived only from the shelf such that (particulate) 228 Th is rarely considered of importance.

From the 228 Ra activity in the intermediate water in the Amundsen Basin (200–1,500 m) we derive a time scale for ventilation with shelf water derived from the Barents Sea Branch Water of about 18 years, in good agreement with estimates based on transient tracers SF₆ and 129 I/ 236 U.

The ratio ²²⁸Th/²²⁸Ra can be used in addition to ²³⁴Th/²³⁸U and ²¹⁰Po/²¹⁰Pb as a measure of export production in the Arctic Ocean. In the surface ocean, the depletion of ²²⁸Th relative to its parent ²²⁸Ra is widespread and it can be measured with higher precision than disequilibria of ²³⁴Th/²³⁸U and ²¹⁰Po/²¹⁰Pb, which are often very small in Arctic surface waters. The long time period has the advantage that removal over more than a full season is integrated, but it has the disadvantage that a large part of the disequilibrium is advected. The interpretation thus requires consideration of both advection and export (Rutgers van der Loeff et al., 2012).

In deep waters (>1,500 m) the 228 Th/ 228 Ra ratio is a delicate balance between the horizontal supply of 228 Ra and 228 Th by mixing with the slope and the vertical supply of 228 Th on sinking particles. In this depth range the ratio is an inappropriate tool to determine particle fluxes.

In subsurface waters, down to approx. 1,500 m, 228 Th accumulates up to a 228 Th/ 228 Ra ratio of about 2. This 228 Th_{xs} is in part due to the small difference between the half-lives of 228 Ra and 228 Th (equations (5) and (6)), leaving barely any significant export flux of 228 Th in some stations (PS94–96, PS94–101, and HLY1502–52), but in others (HLY1502–43, 46, and 48), we find a significant export flux. Comparison with the depletion in the surface water shows that in the Makarov Basin this depletion is largely advected and must represent



export on the shelf where the water originated. At two stations in the Canada Basin such advection is less important and the export flux explains the larger part of the surface water depletion.

However, when this intermediate water mass is influenced by ventilation with the shelf (as is the case in the Amundsen Basin), the 228 Th/ 228 Ra signal in this depth range (100–1,500 m) depends critically on the 228 Th/ 228 Ra ratio in the ventilating shelf water mass, reducing its suitability as a tracer for export production.

The lack of coherence between ²²⁶Ra and Ba in the deep Makarov Basin shows that the accumulation of ²²⁶Ra here is not due to vertical particle flux but can be associated with release from bottom sediments. From a ²²⁶Ra balance we estimate a deepwater residence time of about 480 years, which is consistent with the published estimates based on other tracers.

Data

The data used in this paper are provided in the supporting information and, where indicated, in the PANGAEA database (www.pangaea.de). ²²⁸Ra and ²²⁶Ra data of the HLY1502 expedition can be found at https://www.bco-dmo.org/dataset/718440.

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