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1	Gabbroic-dioritic dykes from the Sanadaj-Sirjan Zone, Zagros Orogen:
2	windows on Jurassic and Eocene geodynamic processes in western Iran
3	
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15	Abbreviated Title: Arc to Within-Plate Magmatism, NW Iran
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17	
18	Supplementary material: Field photographs, photomicrographs, additional geochemical plots,
19	descriptions of analytical methods and tables of geochemical modelling parameters are available at
20	https://doi.org/xxxx.
21	Abstract: The Sanandaj-Sirjan Zone (SaSiZ) is a magmatic-metamorphic portion of the Zagros
22	Orogen, western Iran, marking the Tethyan suture zone between the Afro-Arabian Plate and the
23	Central Iran Micro-Continent. Mafic-intermediate dyke swarms with Middle Jurassic (Group-1:
24	hornblende gabbro and diorite) and Late Eocene (Group-2: hornblende-pyroxene gabbro) ages are
25	recognized in the Malayer-Boroujerd Plutonic Complex of the northern SaSiZ. Group-1 dykes have
26	elemental and isotopic signatures consistent with melting of subduction-modified mantle which

27 relates to Neo-Tethyan subduction processes. Some Group-1 magmas evolved to intermediate 28 compositions through assimilation and fractional crystallization processes. Group-2 dykes have 29 within-plate trace element geochemical signatures, modelled as deriving from low-degree melting 30 of asthenospheric mantle without a subduction influence. Published models postulate either a 31 Cretaceous-Eocene Neo-Tethyan flat-slab scenario, or a Latest Cretaceous-Palaeogene Neo-Tethyan break-off event beneath the SaSiZ. Such models do not reconcile with the Late Eocene 32 33 presence of within-plate magmatism in westernmost Iran, very close to the Zagros Suture. We argue 34 that a period of flat-slab subduction concluded with sub-parallel subduction of ridge to the trench 35 and subsequent slab break-off in the Late Eocene, that is responsible for generation of the distinct 36 Mesopotamia and Zagros slabs in mantle tomography models. Break-off was followed by small 37 volume within-plate type magmatism before short-lived re-establishment of Tethyan subduction 38 prior to the final Arabia-Eurasia collision.

## 39 -end abstract-

40

41 Mafic rocks exposed in convergent settings yield important information about subduction zone 42 environments and magmatic evolution (e.g. Turner et al. 1997). In general, the mantle source for 43 the continental arc systems is metasomatically enriched by fluids and/or melts derived from the 44 subducting slab (Ellam & Hawkesworth 1988). Most subduction-related rocks therefore display 45 distinctive geochemical characteristics such as elevated concentrations of light or mobile 46 incompatible trace elements and relative depletions in high field strength elements (Pearce & Peate 47 1995). The lack of a subduction-related geochemical signature, in mantle-derived mafic rocks formed at or around the time of subduction, may indicate unusual chemical or geodynamic 48 49 circumstances. These situations can be melting of enriched mantle pockets (Hastie et al. 2011), the 50 presence of slab windows (Zandt & Humphreys 2008), slab tearing or break-off (Davies & von 51 Blanckenburg 1995), all of which may allow the rise of melts derived from asthenospheric sources

without the necessity for a slab-derived geochemical component. Unfortunately, in many continental arc settings, mafic rocks are at a premium, so hypotheses about igneous petrogenesis and the geodynamic conditions of past subduction events may be obscured by fractional crystallization and crustal contamination processes.

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57 This paper is an extension of previous work (Deevsalar et al. 2017) on the mafic intrusive rocks 58 from the Malayer-Boroujerd Plutonic Complex (MBPC) in the Sanandaj-Sirjan Zone (SaSiZ), W 59 Iran, which concentrated mainly on mafic-intermediate plutonic rocks. The Mesozoic-Cenozoic 60 SaSiZ (mainly plutonic) and the Cenozoic Urumieh-Dokhtar Magmatic Arc (UDMA) are two major 61 magmatic terranes of the Zagros Orogen (Fig. 1a). The orogen, part of the Cenozoic Alpine-62 Himalayan mountain belt, is a consequence of Neo-Tethyan Ocean closure and the subsequent 63 collision between the Central Iran Micro-Continent and the Afro-Arabian Plate (e.g. Stampfli & 64 Borel 2002; Agard et al. 2011; Verdel et al. 2011; Mouthereau et al. 2012; Azizi et al. 2015). The SaSiZ is a metamorphic-magmatic terrane dominated by igneous rocks that occur mostly as plutons, 65 66 stocks and irregular minor intrusions separated by metamorphic units. The subject of this present investigation, the MBPC, is in the northern SaSiZ (N-SaSiZ) and is spatially the closest magmatic 67 68 complex to the Zagros Thrust Fault (Fig. 1a) at the present level of exposure. Successive magmatic 69 events in the N-SaSiZ include: (1) Late Permian basaltic volcanic activity (Stampfli & Borel 2002), 70 (2) Middle Jurassic to Early Cretaceous calc-alkaline Andean-type granites accompanied by dioritic 71 and gabbroic intrusions (e.g. Ghalamghash 2003; Ghalamghash et al. 2009; Ahmadi-Khalaji et al. 72 2007; Sepahi 2008; Ahadnejad et al. 2010; Mahmoudi et al. 2011; Ghaffari et al. 2013; Azizi et al. 2015; Deevsalar et al. 2017); (3) Paleocene–Eocene gabbroic (Mazhari et al. 2009; Deevsalar et al. 73 2017) and granitic (Rashidnejad-Omran et al. 2002; Azizi & Moinevaziri 2009; Mahmoudi et al. 74 75 2011) intrusions. Magmatic activities in the MBPC occurred during phases (2) and (3) of magmatic 76 activity in the N-SaSiZ. Here are very rare exposures of intrusive mafic rocks which formed during the subduction of the Neo-Tethys Ocean beneath the Central Iran Micro-Continent (Deevsalar et al. 77

78 2017). Following on from a study of gabbroic plutonic rocks within the MBPC (Deevsalar *et al.*79 2017), here we focus on suites of mafic–intermediate dykes. They are of interest as they preserve
80 less-evolved compositions which may be used to model the original mantle source of magmatism
81 compared to more dominant contemporaneous felsic rocks. The dykes are also much less affected
82 by the accumulation of crystal phases compared to contemporary plutonic facies.

83

84 The geochemical study is undertaken to provide constraint on the petrogenesis of two suites of 85 mafic-intermediate dykes, one suite dating to the Middle Jurassic and one to the Late Eocene. In 86 turn, we reflect on the respective geodynamic settings of the dykes. The Late Eocene suite will be 87 emphasised, as it displays little evidence for interaction with subduction-related components and 88 can thus be used to challenge existing geodynamic models for slab geometry during the onset of the 89 Arabia-Eurasia collision (e.g. Agard et al. 2011; Verdel et al. 2011). The new geochronology, 90 petrography, and elemental and isotopic geochemistry for the mafic-intermediate dykes will thus be 91 employed to constrain the geochemical properties of the mantle source region(s), their magmatic 92 evolution and implications for geodynamic models of the region.

93

## 94 The Malayer-Boroujerd Plutonic Complex (MBPC)

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96 The MBPC comprises abundant irregularly-shaped exposures of igneous rocks that form a band striking NW-SE over an elongated area of some 1000 km<sup>2</sup> (Fig. 1b). The MBPC is surrounded by 97 98 metamorphic units like much of the N-SaSiZ. The MBPC displays a wide spectrum of magmatic 99 compositions, although these are dominantly felsic (Fig. 1b, geological map of the NW- and SE-100 MBPC entitled '8' and '9'). The MBPC felsic rocks range from syenogranite-monzogranite to 101 granodiorite-tonalite, the granitoids being the dominant lithology especially in the south-eastern 102 part of the MBPC. The MBPC granitoids intrude into the Jurassic Hamadan phyllite metamorphic 103 assemblage with well-defined, sharp contacts. Magma emplacement generated contact metamorphic

104 aureoles composed of andalusite, garnet and staurolite hornfels (Masoudi 1997). Intermediate units 105 occur as dykes and enclaves within the felsic rocks and as larger mappable 'patches' (irregularly 106 shaped intrusions) of diorites (Deevsalar et al. 2018). Mafic stocks, patches, dykes and veins form 107 exposures of various sizes and they are often found at the margins of the larger felsic bodies (Deevsalar et al. 2014) (Supplementary Item 1, Fig. S1). Zircon U-Pb dating of the felsic rocks, 108 109 which are intruded by the MBPC mafic-intermediate dykes, yielded ages of 162 to 187 Ma for the 110 NW-MBPC (Ahadnejad et al. 2010) and 169 to 172 Ma for the SE-MBPC (Ahmadi-Khalaji et al. 111 2007).

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## 113 Field and Petrographic observations

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These MBPC dykes are 0.5–5 m wide and usually exposed for between 5 and 30 m along-strike, 115 116 although lengths of up to 150 m have been observed. There are no preferred strike directions for the 117 MBPC dykes for either of the geochemically distinct groups or from dykes with different ages 118 (Supplementary Item 1, Fig. S1). They comprise samples with either a fine-grained aphyric or a 119 porphyritic texture. The dykes are dark-grey to dark-green and range from hornblende-bearing 120 pyroxene gabbros to diorites (after Streckeisen 1976). In the field, the dykes are predominantly 121 hosted in granitoids and occasionally in metamorphic country rock, but those from different groups 122 do not show any obvious difference in their locations relative to other igneous and metamorphic 123 rocks. The larger porphyritic dykes have fine-grained chilled margins and smaller dykes are mainly 124 fine-grained dolerites. The mineralogy of these two types based on outcrop size and inferred volume is similar, yet there are discernible texture and grain-size differences. 125

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127 The dykes of different geochemical groups (discussed further in following sections) cannot be 128 distinguished by difference in their constituent minerals, mineral modal proportion, alteration or 129 texture. However, in general the mafic gabbroic dykes are of two kinds: (1) a porphyritic type

130 composed of 20 - 40 vol. % medium-coarse-grained plagioclase phenocrysts (2 - 4 mm) and medium-grained clinopyroxene and amphibole (2 mm) in a 60 - 80 % groundmass of plagioclase, 131 clinopyroxene, amphibole, opaque minerals (mainly hematite), chlorite and sericite, and (2) fine-132 133 grained doleritic dykes composed of microlitic or lath-like plagioclase grains with alignment and 134 amphibole, clinopyroxene, secondary biotite and minor hematite and chlorite. In contrast, dioritic dykes have plagioclase (35 - 50 vol. %), alkali feldspar (0 - 10 vol. %), hornblende (15 - 40 vol. %)135 %), biotite (0 - 8 vol. %), pyroxene (0 - 5 vol. %), chlorite and epidote (0 - 5%) and minor phases 136 including opaque minerals and apatite are <1 vol. % (Supplementary Item 1, Fig. S2). 137

138

- 139 Analytical results
- 140
- 141 Geochronology
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143 Using zircon U-Pb geochronology, several studies of the SaSiZ have reported Middle-Upper 144 Jurassic ages for the granitic rocks. As well as these results, two granitic intrusions from Marivan (NW-SaSiZ, area "5" in Fig. 1b) (~ 38 Ma, Sepahi et al. 2014) and Gosheh-Tavandasht (SE-145 146 MBPC, area "9" in Fig. 1b) (~ 34.9 Ma; Mahmoudi et al. 2011) demonstrate much younger magmatic activity. The presence of inherited Proterozoic-Paleozoic zircons in granitoids indicates 147 148 that this part of the SaSiZ is underlain by much older crystalline basement (Ahadnejad et al. 2010). The age of the mafic rocks from the SaSiZ have been poorly constrained thus far. Mazhari et al. 149 150 (2011) gave an age of  $96 \pm 2$  Ma for the Naqadeh gabbro-dioritic rocks (Fig. 1b, area "2" (N-K-P)) 151 and Mahmoudi et al. (2011) reported ages of 149.5 and 164 Ma for Qorveh gabbroic (Fig. 1b, area "4" (Qr)) and Alvand dioritic rocks (Fig. 1b, area "7" (Al)). The recently constrained zircon U-Pb 152 ages for the MBPC gabbroic intrusions yielded two distinct episodes of mafic magmatism in this 153 154 area including Middle Jurassic and Late Eocene ages. The result indicates the co-occurrence of m

afic, felsic and intermediate magmatism during Middle Jurassic and bimodal mafic-intermediatemagmatism during Late Eocene of the MBPC.

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158 For this study, three representative samples of gabbroic dykes (from NW-MBPC: MN<sub>2a</sub>, M<sub>29</sub> and M<sub>50</sub>) and one dioritic dyke (from SE-MBPC: BR<sub>02</sub>) from two geochemically distinct groups (see 159 160 the next section) were chosen for U-Pb zircon dating by laser ablation inductively coupled plasma 161 mass spectrometry (LA-ICP-MS). Analytical methods and procedures are fully presented in Supplementary Item 2, section A. Given that precise age measurements using <sup>207</sup>Pb/<sup>235</sup>U and 162 <sup>207</sup>Pb/<sup>206</sup>Pb ratios are feasible usually only for Precambrian zircons (cf. Ireland & Williams 2003), 163 164 the weighted mean of pooled <sup>206</sup>Pb/<sup>238</sup>U ages are taken to indicate the crystallization ages of the samples in this study. Zircon U-Pb ages for 90 analytical points on individual zircon grains are 165 166 given in Supplementary Item 3. Table S1. The separated zircons are mostly transparent, colourless 167 to slightly brown and show oscillatory zoning (Fig. 2), indicative of a magmatic origin (Hoskin & Schaltegger 2003). The results of the LA-ICP-MS analyses for twenty-four points from sample 168 169  $MN_{2a}$  yielded weighted mean age of  $169.8 \pm 2$  Ma (Supplementary Item 3, Table S1; Fig. 2a, and Fig. 3a). Forty-six grains from sample BR<sub>02</sub> yielded a mean  ${}^{206}Pb/{}^{238}U$  age of 166.93  $\pm$  0.72 Ma 170 171  $(1\sigma)$  for zircon crystallisation (Supplementary Item 3, Table S1; Fig. 2b, and Fig. 3b), after rejecting 172 three scattered or discordant ages (spots 6, 10 and 19) and one anomalous age likely due to large analytical uncertainties (spot 9). Of 9 points for sample M<sub>50</sub>, 8 points give weighted mean age of 173  $47.1 \pm 0.7$  Ma (Supplementary Item 3, Table S1; Fig. 2c, and Fig. 3c) and one point produced 174 175 anomalous age likely indicating zircon inheritance (inherited zircon age of  $1868 \pm 18$  Ma, spot 9). The 16 points analysed in sample  $M_{29}$  provide a concordant age of  $36.1 \pm 0.9$  (Supplementary Item 176 177 3, Table S1; Fig. 2d, and Fig. 3d).

178

179 The presence of inherited zircons with a Mesoproterozoic age, like those detected in MBPC 180 granitoids (e.g. Ahadnejad *et al.* 2010), indicates the presence of a crustal contaminant either as old crustal wall-rock or an old crystalline basement. The Middle Jurassic ages for gabbro-dioritic dykes are consistent with zircon U–Pb ages of subduction-related gabbroic intrusions from the MBPC (Deevsalar *et al.* 2017) and granitic rocks throughout the SaSiZ (e.g. Ahmadi-Khalaji *et al.* 2007; Ahadnejad *et al.* 2010; Shahbazi *et al.* 2010; Mahmoudi *et al.* 2011; Chiu *et al.* 2013; Sepahi 2014) and implies that syn-plutonic mafic-intermediate magmatism closely related in time and space to felsic magmatic activity. The geodynamic implications of Late Eocene gabbroic dykes (~42 Ma) and intrusive rocks (~40 Ma, Deevsalar *et al.* 2017) within the MBPC will be discussed later.

188

189 Major element geochemistry

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191 The MBPC dykes have retained much of their original igneous texture; however, they experienced 192 some sub-solidus weathering and low-grade metamorphism including growth of chlorite and epidote. Following careful sampling, the analysed MBPC dykes have chemical index of alteration 193 194 values (= molecular  $[Al_2O_3/(Al_2O_3+CaO+Na_2O+K_2O)] \times 100$ , suggested by Nesbitt & Young 1982) 195 lower than 50 (except of BR<sub>02</sub> and M<sub>23</sub>) as suggested for un-weathered igneous rocks (Ramkumar 196 2015). These dykes yielded low loss-on-ignition (LOI) values (mostly < 3 wt. %, except of M<sub>23</sub>, 197 BR<sub>02</sub>, M<sub>29</sub> and MN<sub>12</sub>) indicating a modest extent of hydrothermal alteration. Low-grade 198 metamorphism in these rocks might mobilize major and large ion lithophile elements (LILE) such 199 as Ba and Rb (e.g. Floyd & Winchester 1975; Hastie et al. 2007). However, having recognized 200 coherent primitive mantle- and chondrite-normalized patterns (Fig. 5), widespread remobilisation 201 during metamorphism is unlikely. The consistency of trends observed in Figs. 4 and 5, for unaltered 202 samples (low LOI) and those with slightly altered samples suggest that the concentrations of these 203 elements were not selectively modified. Whole-rock major and trace element data for the MBPC mafic-intermediate dykes are given in Table 1. Analytical methods and procedures are presented in 204 205 Supplementary Item 2, section B.

207 As given earlier, based upon geochemical characteristics, two set of dyes have been identified in the MBPC. The first group of MBPC dykes comprise mafic gabbros (n = 3, samples MN<sub>08</sub> and BR<sub>03b</sub> 208 209 and  $MN_{2a}$ ) and intermediate samples with dioritic compositions (n = 3, samples  $M_{23}$ ,  $BR_{07}$  and 210 BR<sub>02</sub>) (Table 1). Whole-rock SiO<sub>2</sub> in these samples varies from ~45 to 62 wt. % (Table 1). Graphs of whole-rock oxides vs. SiO<sub>2</sub> (Fig. 4a-f) show systematic variation in this group. The concentration 211 212 of CaO, TiO<sub>2</sub>, FeO<sub>T</sub>, MgO and MnO decrease with increasing SiO<sub>2</sub> whereas Al<sub>2</sub>O<sub>3</sub> and total alkalis 213  $(Na_2O + K_2O)$  increase. These dykes show a sub-alkaline character on a TAS plot (Fig. 4a). U-Pb 214 dating of two selected samples from this group defines a Middle Jurassic age. We herein refer to the 215 Middle Jurassic dykes as Group-1. In contrast to their coeval calc-alkaline gabbroic intrusions 216 (Deevsalar et al. 2017), most of the dykes fall in tholeiitic fields on FeO<sup>T</sup>/MgO vs. SiO<sub>2</sub> and AFM 217 diagrams (Supplementary Item 1, Fig. S3). The dykes have also a wider range of TiO<sub>2</sub> (0.49 to 218 3.70) than contemporaneous Middle Jurassic gabbroic intrusions (0.29 to 0.64) (Deevsalar et al. 219 2017).

220

221 The next group of MBPC dykes have gabbroic compositions (n= 16, Table 1). The SiO<sub>2</sub> concentrations of Group-2 dykes display a small range (~47 to 49 wt. %, Table 1). They have 222 relatively small range of MgO (~5 to 7 wt. %) in comparison to Group-1 dykes which range from 2 223 to 10 wt. % MgO. In Harker-style plots (Fig. 4a-f), major oxides or elements show correlation with 224 225 SiO<sub>2</sub> despite the narrow SiO<sub>2</sub> range. SiO<sub>2</sub> positively correlates with CaO and Al<sub>2</sub>O<sub>3</sub>, and negatively correlates with MnO, FeO<sup>T</sup> and TiO<sub>2</sub>. MgO either does not correlate or the correlation is slightly 226 227 positive with increasing SiO<sub>2</sub>. U-Pb dating of two samples from this group is yielded Late Eocene ages. The Late Eocene mafic dykes are herein referred to as Group-2. The Group-2 dykes have an 228 229 alkaline affinity on the TAS diagram (Fig. 4a), akin to their coeval gabbroic intrusions (Deevsalar et 230 al. 2017) and contrasting with Group-1 dykes. TiO<sub>2</sub> content in Late Eocene gabbroic dykes varies 231 within a small range (1.9 to 3.8 wt. %) compared to those coeval gabbroic intrusions (1.9 to 5.5 wt. %; Deevsalar et al. 2017). 232

## 234 Trace element geochemistry

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236 The MBPC dykes of different ages can be clearly distinguished in terms of trace element 237 characteristics. Group-1 dykes have high Sr (~300 to 600 ppm), while those from Group-2 have 238 lower Sr concentrations (~50 to 100 ppm). Group-1 dykes display systematic variations in relative 239 modal abundances of Fe-Ti oxides, including ilmenite and rutile, which reflect in their vanadium 240 concentrations. The Group-1 dykes are characterised by distinctive enrichments in LILE and light rare earth elements (LREE), and by depletion of high field strength elements (HFSE) and heavy 241 242 rare earth elements (HREE), especially in those of dioritic character (Fig. 5a). The diorites have the 243 highest LILE (La, Ba and Rb) and Th concentrations. The primitive mantle-normalized trace 244 element diagrams show negative anomalies for Nb, Ta and significant depletion in Sr (Fig. 5a), akin 245 to the Middle Jurassic non-cumulate gabbroic plutonic rocks also found in the MBPC (Deevsalar et 246 al. 2017). Aside from the distinctive negative HFSE and Sr anomalies, sample MN<sub>2a</sub> shows slightly negative Ti and Zr anomalies (Fig. 5a). The total REE concentration ( $\Sigma$ REE) in Group-1 dykes 247 248 varies from 123 to 204 ppm. These Jurassic rocks display LREE-enriched and HREE-depleted 249 chondrite-normalised (CN) patterns (Fig. 5b), including negative Eu anomalies (Eu/Eu $*_{CN}$  = 0.61 to 0.69), as well as relatively high (La/Yb)<sub>CN</sub> of 5 to 15 (Table 1). In comparison, the 250 251 contemporaneous Middle Jurassic non-cumulate gabbroic intrusions are enriched in LREE and have less variable LREE/HREE fractionation (Deevsalar et al. 2017). 252

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The Group-2 gabbroic dykes are characterised by slight and variable enrichment in the LILE relative to the LREE and HFSE (Fig. 5c). The dykes also show variable degrees of LREE/HREE fractionation with La/Yb<sub>CN</sub> ranging from 4-10 (Fig. 5d). Group-2 dykes contain slightly negative to positive Eu anomalies with Eu/Eu\*= 0.93–1.31 (Fig. 5d) and lower  $\sum REE$  (94 to 152 ppm) than the Group-1 dykes. The Group-2 dykes have primitive mantle-normalized and chondrite-normalized REE patterns consistent with their alkaline affinity and show a typical within-plate character (Winchester & Floyd 1977). There is no significant depletion in HFSE such as Nb, Ta, Zr, Hf and Ti (Fig. 5c), which is remarkably consistent with within-plate rocks from modern ocean islands.

262

263 Radiogenic isotopes

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Whole-rock Sr, Nd and Pb isotopic compositions of representative samples of mafic–intermediate dykes from the MBPC are shown in Tables 2 and 3. Analytical methods and procedures are presented in Supplementary Item 2, section C. Initial values for the <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd ratios are calculated at 168 Ma for Group-1 and at 42 Ma for Group-2 dykes. On Figure 6, the isotopic composition of the MBPC samples are plotted against mafic rocks derived from different mantle sources beneath the modern Turkish–Iranian Plateau (Supplementary Item 3, Table S2).

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The Group-1 dykes display wide range variations of  $({}^{87}\text{Sr}/{}^{86}\text{Sr})_i$  ratios albeit at uniform  $\epsilon$ Nd<sub>i</sub> ranging 272 from 0.7077 to 0.7151 and -5.85 to -7.89, respectively. Initial <sup>87</sup>Sr/<sup>86</sup>Sr ratios in this group are 273 274 higher than those in Middle Jurassic calc-alkaline gabbroic intrusions from the MBPC, which had values ranging from 0.7035 to 0.70594 (Deevsalar et al. 2017). On Figure 6a, the Group-1 dykes 275 plot parallel with the <sup>87</sup>Sr/<sup>86</sup>Sr axis, extending from the mantle array to higher (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> ratios. All 276 the samples, except MN<sub>2a</sub>, therefore plot outside of mantle array on Figure 6a. (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> ratios of 277 the two gabbroic dykes are broadly similar to that of dioritic dykes, indicating that these rocks may 278 279 be linked through similar process on a common source. The Pb isotopic ratios are presented in Table 4 and plotted on Fig. 6b-e. The Pb isotope ratios of the gabbro-dioritic dykes from Group-1 280 281 are close to Cenozoic and Mesozoic magmatic rocks from a variety of tectonic settings across the 282 modern Turkish-Iranian Plateau. In summary, this group exhibits moderately radiogenic Sr and Pb 283 and unradiogenic Nd isotopic ratios.

284 The Late Eocene Group-2 samples have relatively lower Rb and higher Sr concentrations compared to Group-1 dykes, resulting in lower <sup>87</sup>Sr/<sup>86</sup>Sr ratios. This group has comparatively uniform 285  $(^{87}\text{Sr}/^{86}\text{Sr})_i$  values and  $\epsilon$ Nd<sub>i</sub>, 0.7049 to 0.7074 and +0.76 to +3.29 respectively, roughly overlapping 286 287 the values of Late Eocene alkaline gabbroic intrusive rocks from the MBPC (Deevsalar et al. 2017). In <sup>87</sup>Sr/<sup>86</sup>Sr vs. ɛNdi space, Group-2 dykes are less scattered that Group-1. Some samples plot 288 289 within the mantle array (Fig. 6a). In comparison with mid-ocean ridge basalts and Hawaiian alkali basalts, at comparable (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> ratios the Late Eocene dykes have lower ɛNd<sub>i</sub> values (Fig. 6a). 290 Group-2 dykes show only slightly higher (<sup>206</sup>Pb/<sup>204</sup>Pb)<sub>i</sub> and (<sup>208</sup>Pb/<sup>204</sup>Pb)<sub>i</sub> than Group-1 and similar 291 (<sup>207</sup>Pb/<sup>204</sup>Pb)<sub>i</sub> (Fig. 6b-e). 292

293

#### 294 **Discussion**

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296 Crustal contamination
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It is essential to evaluate if magmas have undergone crustal contamination before speculating on 298 their source region. Unlike the nearby felsic plutonic bodies, the MBPC gabbro-dioritic dykes do 299 300 not record macroscopic evidence for crustal assimilation such as metamorphic xenoliths or refractory metamorphic minerals. Therefore, closer inspection of elemental and isotopic data is 301 302 required. High SiO<sub>2</sub> and alkalis and low MgO, which are expected in substantially contaminated magmas (e.g., Jahn et al. 1999; Zhang et al. 2002), were only observed in Group-1 diorites. There is 303 304 minimal evidence of zircon inheritance (one inherited zircon in MN<sub>2a</sub>, Supplementary Item 3, Table S1) during Middle Jurassic dyke emplacement, although it is entirely possible that inheritance was 305 306 not picked up due to potential for biased recovery during preparation of zircon separates, or 307 dissolution of zircon within the mafic host. Sample MN<sub>2a</sub>, despite the recorded zircon inheritance, is 308 the only analysed Group-1 sample to fall within the mantle array on Figure 6a. The remaining Group-1 dykes form a parallel array towards higher <sup>87</sup>Sr/<sup>86</sup>Sr and thus upper continental crust 309

310 composition (Fig. 6a). This evidence is compounded by the overall arc-like elemental signatures of 311 the Group-1 samples, and the fact that contemporary Middle Jurassic mafic plutonic rocks almost 312 all sit within the mantle array (Deevsalar *et al.* 2017). It seems that sample  $MN_{2a}$  would be the best 313 choice for further modelling of the mantle source from the dyke suite.

314

315 In the Group-2 gabbro dykes, there is also little evidence for zircon inheritance (one inherited 316 zircon in M<sub>50</sub>, Supplementary Item 3, Table S1). There is an obvious lack of typical crustal trace 317 element signatures (i.e. there is no high LILE/HFSE signature) in most of the samples. High  $(^{87}\text{Sr}/^{86}\text{Sr})_i$  (~ 0.707) in three of the samples (M<sub>33</sub>, M<sub>38</sub> and M<sub>50</sub>; Fig. 6a) could be related to crustal 318 319 contamination, however the trend shown on Figure 6a is parallel to the y-axis and could be related 320 to later fluid-related alteration. There is no systematic LILE variation in these samples relative to 321 those within the mantle array which further suggests only limited crustal contamination. These high-(<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> samples will nevertheless be excluded from further discussion. The general lack of 322 323 crustal signatures in Group-2 remains consistent with a within-plate character, and is discussed 324 further below.

325

## 326 Fractional crystallisation (FC)

327

The generation of low-MgO, high-alumina rocks by fractional crystallization of high-MgO primary 328 mafic magma is typical of magmatic arc settings (Bartels et al. 1991; Draper & Johnston 1992). The 329 330 Group-1 dykes contain wide ranges of MgO, Cr, Ni and Co (Table 1), implying they experienced 331 varying degrees of fractional crystallisation (FC). The range in MgO values (from 10 to 2 wt. %) correlate with the wide range of Cr (83.7-904 ppm) and Ni (44.8-273.1 ppm) abundances. The 332 333 samples do not contain sufficiently high MgO to indicate the presence of near-primary compositions. In Group-1 dykes, sample MN<sub>2a</sub> with highest Mg# number (Mg<sup>#</sup> =Mg<sup>+2</sup>/(Mg<sup>+2</sup>+Fe<sup>+2</sup>): 334 0.57) and high Cr and Ni concentrations, is representative of the least differentiated magma. A 335

336 systematic change in total REE concentration ( $\Sigma REE$ ) is expected for magmas which undergo 337 fractionation of phases in which the REE are incompatible. This feature is observed for Group-1, where  $\sum$ REE positively correlate with SiO<sub>2</sub> (inset plot, Supplementary Item 1, Fig. S4a). These 338 339 dykes also have negative Eu anomalies (Fig. 5b) and low Sr concentrations (Table 1), which are 340 likely due to fractional crystallization of plagioclase or presence of plagioclase in the source rocks 341 (e.g. Hanson 1978). There are also systematic variations of SiO<sub>2</sub> and MgO with major oxides and 342 trace elements in Harker diagrams (see Fig. 4a). Increasing Al<sub>2</sub>O<sub>3</sub> contents and decreasing 343 CaO/Al<sub>2</sub>O<sub>3</sub> ratios with decreasing MgO contents (Fig. 4b) accompanied with decreasing Cr and V 344 contents with increasing Zr concentrations (Supplementary Item 1, Fig. S4a, b, c) indicate 345 fractionation of clinopyroxene. In general, the trends exhibited by the major oxides (Fig. 4a) can be 346 explained in terms of fractionation of plagioclase, clinopyroxene along with apatite and accumulation of Fe-Ti oxides. 347

348

The Group-2 gabbroic dykes exhibit only a small range of SiO<sub>2</sub> (47 to 49 wt. %) and MgO (5 to 7 wt. %), and low Ni, Co, Cr and V (Table 1), so these dykes are unlikely to represent primary mantle melts. As shown in Supplementary Item 1, Fig. S4a to S4f, general trends on Harker-style diagrams are consistent with the fractionation of olivine and clinopyroxene or chrome-spinel. The fractionation of Fe-Ti oxides is supported by a negative trend observed on the FeO<sup>T</sup> vs. SiO<sub>2</sub> graph (Fig. 4b) and the positive trend on a TiO<sub>2</sub> vs. MgO graph (Fig. 4i). The slightly positive Eu anomaly in this group rules out plagioclase fractionation.

356

357 Mantle source characteristics

358

The presence of Middle Jurassic gabbroic plutonic rocks in the MBPC, with only slightly less isotopically enriched signatures to  $MN_{2a}$ , also lends credence to using  $MN_{2a}$  to at least qualitatively discuss the Middle Jurassic mantle source. These rocks are characterized by enrichment in LILE

362 and depletion of HFSE with all the distinctive features of continental subduction zone setting. The 363 significant negative anomaly for Nb and Ta in all the Middle Jurassic dykes (Fig. 5a) indicates the influence of a relatively Nb-Ta depleted aqueous fluid derived from oceanic sediments as opposed 364 365 to the influence of a silicate melt in which Nb and Ta have high solubility (Ionov & Hofmann 1995). Furthermore, the position of Group-1 dykes in Th/La vs. (Ce/Ce\*)<sub>Nd</sub> plot (Hastie et al. 2013; 366 367 Supplementary Item 1, Fig. S5), including that of MN<sub>2a</sub>, suggest that the source region was 368 contaminated with sedimentary components equivalent to continental detritus/GLOSS II. This is 369 consistent with geological evidence for Tethyan subduction on the Eurasian margin throughout the 370 Mesozoic and Paleogene (Dilek & Sandvol 2009; Agard et al. 2011; Verdel et al. 2011; Chiu et al. 371 2013). Furthermore, the isotopically enriched Sr-Nd-Pb signatures of even the least contaminated 372 samples, suggest partial melting of metasomatised mantle was probably a major contributor to their petrogenesis, like the coeval gabbroic intrusions. However, the slightly lower ENdi values in Middle 373 374 Jurassic gabbro-dioritic dykes compared to gabbroic plutonic rocks seem to indicate a higher 375 volume of subducted materials (or greater extents of contamination) were involved in the genesis of 376 the former.

377

378 For the Group-2 Late Eocene gabbro dykes, the narrow range of  $SiO_2$ , MgO > 4 wt. %, high  $TiO_2$ 379 and low La/Nb ratios (OIB-like signatures), as well as high concentrations of typical basaltic 380 components (i.e. Al<sub>2</sub>O<sub>3</sub>  $\approx$  14 to 18 wt. %, FeO<sup>T</sup>  $\approx$  10 to 12 wt. %, and CaO  $\approx$  9 to 11 wt. %) reflect a fertile mantle source. As the ratios of non-mobile elements, including Nb, Ta, Zr and Hf are not 381 382 usually affected by aqueous fluids and remain constant during magmatic processes, they can be used to investigate the original composition of the mantle source (Pearce & Peate 1995). As a result 383 of their variable Hf/Yb and relatively constant Nb/Zr ratios, the Group-2 dykes plot within the 384 385 enriched mantle field (EM-type) in Fig. 7a (John et al. 2004). In Th/Yb vs. Ta/Yb space they plot 386 within the mantle array and show within-plate affinity (Fig. 7b). The enriched mantle source for

387 Group-2 is further highlighted on Figure 7c, where samples plot close to the OIB compositional388 field and

above Western Anatolian Mantle compositions (WAM; Aldanmaz et al. 2000), E- and N-MORB 389 390 (Enriched and Normal Mid Oceanic Ridge Basalt). However, on Supplementary Item 1, Fig. S5, the 391 Group-2 dykes sit above the N-MORB field. An OIB-like signature is also supported by Zr/Ta vs. 392 Nb/Hf variations (Fig. 7d) and the primitive mantle normalized multi-element diagram (Fig. 7e), in 393 which Group-2 dykes show OIB-type trajectories. A lack of HFSE depletion, minor enrichment of 394 LILE and LREE and no significant Eu anomalies (Fig. 5d), indicate these are within-plate rocks 395 originating either from the asthenosphere (Sun & McDonough 1989; Wittke & Mack 1993), or an 396 enriched upper mantle source (e.g. Hastie et al. 2011) without contemporary subduction influence. 397 However, the trace element composition may necessitate incorporation of enriched mantle in the 398 genesis of group-2 dykes (e.g., Lassiter et al. 2000; Lundstrom et al. 2003), but given the depleted 399 isotopic signatures, these are most likely small-batch melts formed by low degree melting of depleted upper mantle asthenosphere (e.g., Ramos & Key, 1992; Gorring et al. 1997). This outcome 400 401 is more consistent with the rare and small outcrops of Late Eocene exposed mafic rocks (in both 402 forms of dykes and intrusions) within the MBPC.

403

### 404 Mantle melting process

405

The occurrence of less fractionated samples with near primary melt composition among the samples from Group-1 and Group-2 indicates melting of mantle lithologies. It may still be possible to model the extent of partial melting responsible for the generation of the least-evolved samples. As postulated by Jaques & Green (1980), the composition of primary mafic magma produced by melting of given source region under certain P-T conditions, is in part controlled not just by pressure and temperature, but by specific modal mineralogy. For the MBPC, the conditions of partial melting almost certainly will have changed between the subduction-related magmatism of the Middle Jurassic to the within-plate setting of the Late Eocene, concurrent with the models forcontemporary gabbroic intrusions (Deevsalar *et al.* 2017).

415

416 Because of different phase/melt partition coefficients of Yb and Gd for garnet and spinel (McKenzie and O'Nions, 1991), a mantle source with residual spinel will produce (Gd/Yb)<sub>CN</sub> ratios 417  $\sim$  1.2, whereas those melts originated from garnet-bearing source are expected to have higher values 418 (Allen et al. 2013; Kelemen et al. 2003). The, (Gd/Yb)<sub>CN</sub> ratio (MN<sub>2a</sub>), least fractionated sample 419 420 from Group-1, is 1.49., However,  $(Dy/Yb)_{CN}$  ratio of this sample (= 1.9) is in the range of values 421 suggested for garnet-spinel (~ 1.5) and garnet-bearing mantle sources (> 2.5) (e.g. Chang et al. 2009; Jiang et al. 2010). This observation may indicate a changing depth of melting within the 422 garnet-spinel transition zone, which is consistent with the result for contemporary gabbroic 423 intrusions. Given the high Ba/Rb ratio in  $MN_{2a}$  (= 3.91), amphibole is also likely to have been 424 425 present as a residual phase of melting mantle (Allen et al. 2013). Other possible processes include 426 mixing between melts originated from garnet-bearing and spinel-bearing peridotites, which will be 427 further examined using trace element models.

428

## 429 Melting, Mixing, FC and AFC models

430

To constrain the formation conditions of the Group-1 dykes, we modelled non-modal batch melting (Shaw 1970) of an assumed source, i.e. metasomatically enriched mantle lherzolite from the garnetspinel transition zone, including maxima of 2 % amphibole, 7 % garnet and 3 % spinel (the parameters used in models and the results are given in Supplementary Item 3, Table S3, S4, Mode-I). The result for Mode-I indicates that 15 % partial melting of such a composition gives a best fit trajectory on primitive mantle-normalized trace element patterns of Group-1 dykes (Fig.8b; Table 437 4), consistent with the tholeiitic affinity of the Group-1 dykes (Jaques & Green 1980).

As already discussed, the closed-system fractional crystallization cannot be the only process 439 440 involved in the magmatic evolution of Group-1 dykes, owing to the isotopic evidence 441 for crustal contamination. The possible mechanism is open-system fractional crystallization plus 442 assimilation of crustal materials either through AFC processes in shallow crustal magma chambers 443 or mixing with crustal-derived melt. However, lower compatible element concentrations (such as Cr, Ni, Co) and higher (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> ratios in dioritic dykes than those in the most primitive gabbroic 444 445 sample (MN<sub>2a</sub>) indicate a possible role for AFC-process in Group-1 dykes, but this is not supported 446 by the results of trace element modelling of AFC (Figs. 8a, c). Alternatively, simple binary mixing model and multi-element modelling favours a two-component mixing process between mafic 447 magma and felsic crustal melts. The results are shown as trajectories on (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> vs. 1/Sr (Fig. 448 8b) and PM-normalized multi-element plot (Fig. 8d) (the parameters used in models and the results 449 450 are given in Table 4 and Supplementary Item 3, Tables S2 to S4). Subsequent ascent will involve 451 further cooling and fractional crystallization of these crustally-contaminated mafic magmas to produce the final Group-1 chemical composition (Fig. 8d). 452

453

For Group-2 dykes, the geochemical data attest of an origin by extraction of small melt batches 454 455 from the asthenosphere, again within the garnet-spinel transition zone. In Fig. 9a, the melting model of an enriched mantle source — similar to Western Anatolian Mantle (WAM) — with garnet-spinel 456 457 mineralogy (point M) for the MBPC Late Eocene Gabbros (Deevsalar et al. 2017) has been compared to that of depleted mantle with garnet-lherzolite and spinel-lherzolite compositions (point 458 459 M'). However, melting trajectories of garnet-lherzolite and spinel-lherzolite sources on a La/Sm vs. La plot (Fig. 9a) are almost the same, but suggest they can be considered as the potential sources of 460 magmatism. The La/Yb vs. Zr/Nb plot (Fig. 9b) indicate that the Group-2 dykes can be modelled by 461 462 mixing of magmas generated by low degree melting of depleted mantle with garnet- and spinel-463 lherzolite mineralogy rather than enriched mantle. The geochemical models also suggest a more depleted composition of the asthenospheric source (M' in Fig. 9a, b) than PM, but somewhat more 464

465 enriched than N-MORB.Trace element modelling of non-modal batch melting (Table 5) 466 demonstrates that small batch melts of garnet lherzolite source ( $F_{melting} = 0.05$ , Mode-III; Supplementary Item 3, Table S3, S4), probably mixed with melts originated from spinel lherzolite 467 468 in a shallower depth of mantle ( $F_{melting} = 0.15$ , Mode-IV; Supplementary Item 3, Table S3, S4; 469  $F_{\text{mixing}} = 0.6$ ). This model could explain the trends for the Group-2 dykes after the effects of FC are 470 included (F<sub>c</sub>: 0.45). Similar scenarios are frequently documented in the literature for the mafic 471 alkaline rocks (e.g., Jaques & Green 1980; DePaolo & Daley 2000; Macpherson et al. 2006). The 472 FC model conducted for the magmas indicates  $\sim$ 45 % fractional crystallisation of 0.5 Ol + 0.35 Cpx  $+ 0.1 \text{ Am} \pm 0.05 \text{ Apt}$  could explain the geochemical variations observed in the Group-2 dykes (Fig. 473 474 9c).

475

## 476 Implications for regional geodynamic models

477

478 As discussed above, the geochemical and isotopic characteristics of Middle Jurassic and Late 479 Eocene gabbro-dioritic dykes from the MBPC imply changing magma sources. Such changes are likely to be associated with changing subduction dynamics, related to Tethyan ocean closure and 480 481 the eventual onset of the wider Arabia-Eurasia collision. Such dynamics might include the rate, angle, depth and continuity of the subduction, all of which will control the location and chemistry 482 of magmatic activity. We have little comment to pass on the Middle Jurassic Group-1 dykes: they 483 484 formed at a time when there is unequivocal evidence for subduction of Neo-Tethyan oceanic crust 485 beneath the Central Iran Micro-Continent.

486

Three key observations are relevant in discussion of the geodynamic situation in Iran during the Late Cretaceous to Eocene. 1) Termination of typical subduction-related magmatism prior to the end of Eocene is consistent with the within-plate character of the MBPC gabbroic intrusions (~ 40 Ma, Deevsalar *et al.* 2017) and confirms other recent studies of Eocene rocks in the UDMA (e.g.

Ballato et al. 2011; Verdel et al. 2011). Recently, Zhang et al. (in press) also described ~40 Myr old 491 492 within-plate type rocks around Piranshahr around 500 km NW of Boroujerd (area "3" in Fig. 1b), 493 the only other recorded Eocene magmatic rocks within the SaSiZ. 2) Separately, Agard et al. (2011) 494 and Angiboust *et al.* (2016) tied exhumation of high pressure rocks in the Zagros Mountains at  $\sim 65$ 495 Ma to a possible break-off event of the Neo-Tethyan slab, some 25 Myr prior to any such within-496 plate type magmatism. 3) This break-off model was adopted by van der Meer et al. (2018) in 497 ascribing an age of  $65 \pm 5$  Ma to the 'top' of their Mesopotamia slab which can be seen in global 498 mantle tomography model, with the 'top' age reflecting the time of detachment of the slab from the 499 overlying lithosphere (Fig. 10). van der Meer et al. (2018) also designated this as the 'base' age of 500 the Zagros slab seismic anomaly, where the Zagros slab was ascribed to renewed Tethyan 501 subduction shortly after break-off of the Mesopotamia slab. We can now discuss a challenge to the 502 timing or nature of these geodynamic scenarios, based on the scattered occurrence of Eocene 503 within-plate type magmatism in the SaSiZ and wider geodynamic considerations. Firstly, the onset of voluminous magmatic activity in the UDMA and further east during the Eocene has already been 504 505 ascribed to rollback-induced slab steepening (e.g. Verdel et al. 2007; 2011; Mouthereau et al. 506 2012). Verdel et al. (2011) argued that slab-rollback may have been preceded by a period of flat-507 slab subduction lasting from the mid-Cretaceous, as evidenced by a regional unconformity 508 separating Cretaceous sediments from overlying Paleogene volcanic rocks (Stöcklin 1968). In this 509 model, there need not have been a slab break-off event around the Cretaceous-Palaeogene boundary. Late Cretaceous deformation in the SaSiZ (Tillman et al. 1981) and a north-eastward 510 511 shift in the locus of subduction-related magmatism from the SaSiZ to northern Iran and the UDMA 512 also support the conclusion that a Tethyan slab was continuously attached to the surface during the mid-Cretaceous to Eocene (e.g., Guest et al. 2006; Omrani et al. 2008), albeit perhaps becoming 513 514 shallower in angle at this time. Rapid exhumation in the Zagros mountains at ~65 Ma (Agard et al. 515 2011) could have several potential causes, including changes in collision rate or in the thickness and/or density of the subducted crust (Iannace et al. 2007; Brun & Faccenna 2008; Rosenbaum & 516

517 Mo 2011). Furthermore, an upsurge in mantle-derived magmatism is commonly ascribed in the 518 literature to slab break-off. However, this upsurge is not recorded until 10's of Ma after the 519 proposed break-off event of Agard *et al.* (2011), and as a corollary, nor did any extensive 520 magmatism occur close to the Zagros Suture around ~65 Ma ago (Verdel *et al.* 2011).

521

522 Our understanding of the Eocene mafic rocks in the MBPC and the Piranshahr area provides a 523 potential constraint on the issue described. The implication of the model of Verdel et al. (2011) is 524 that during the Eocene, the Neo-Tethyan flat slab was rolling back towards the SW but did not roll 525 back so far as to allow upwelling asthenosphere to flow beneath the SaSiZ. It seems difficult in this 526 scenario for asthenosphere-derived magmas to be emplaced during the Late Eocene in the SaSiZ, 527 given the proposed lack of a mantle wedge so close to the suture. Therefore, we feel that the most probable mechanism for igniting Late Eocene mafic magmatism in SaSiZ is a slab-tearing or 528 529 detachment event concurrent with roll-back, perhaps related to subduction of a ridge or transform (Fig. 10). In such a scenario, a slab window would be generated whereby upwelling asthenosphere 530 531 could rise past the slab, largely free of subduction-related fluids and/or melts, and thus partially 532 melt at the base of the SaSiZ lithosphere. This slab-tearing, we propose, would lead to a Late 533 Eocene break-off event rather than one at ~65 Ma. Widely-discussed models for slab window 534 formation such as subduction of a spreading ridge orthogonal to the trench and/or lateral slab-535 tearing (Zandt & Humphreys 2008), seem unlikely processes. The first reason is that trench-536 orthogonal ridge subduction would lead to a narrow slab window which cannot explain the spread 537 of SaSiZ magmatism, at sites some 500 km apart at similar times. Secondly, lateral slab-tearing is 538 modelled to progress at hinge migration rates of ~25-155 km/Myr (Hale et al. 2010); implying there should be both a gap of ~3-20 Myr between magmatism in the MBPC and Pi, which does not exist 539 540 as far as can be judged from the available geochronology. Trench-parallel subduction of a spreading 541 ridge (see Zhang et al., in press) seems a more likely model. A brief lapse in 'normal' subduction 542 processes during descent of the proposed spreading ridge might explain the limited occurrence of within-plate type rocks in the SaSiZ. Immediately after this short period of SaSiZ magmatism, subduction of the Zagros slab (after van der Meer *et al.* 2018) became established, and the locus of arc magmatism returned to the UDMA from the Latest Eocene to Oligocene (Verdel *et al.* 2011). This resumption of magmatism was short-lived, however, as it was followed by the Oligocene onset of continental collision between Arabia and the Central Iran Micro-Continent (e.g., McQuarrie & van Hinsbergen 2013) which resulted in a new magmatic shut-down.

549

550 In summary, we conclude that, whilst we support the flat-slab model of Verdel et al. (2011), we 551 caution that the slab may not have been continuously subducting, but instead segmented and 552 experienced break-off to form the Mesopotamia slab of van der Meer et al. (2018) during the Late 553 Eocene. Subduction of Neo-Tethys then briefly resumed during the mid-Cenozoic and thus formed both the mid-Cenozoic UDMA and the young, shallow Zagros slab. Our new model requires 554 555 revision of the 'top' age of the Mesopotamia slab and the 'base' age of the Zagros slab to ~45-40 Ma, and for the trigger of rapid exhumation of high pressure rocks in the Zagros around ~65 Myr 556 557 ago (Agard et al. 2011) to be re-considered, perhaps as a consequence of a change in the nature of 558 subducted crust at this time (e.g., Spikings & Simpson 2014).

559

#### 560 **Conclusions**

561

562 Mafic magmatism in the N-SaSiZ includes suites of gabbroic dykes and plutons. In this study, we 563 identified and studied two distinct generations of mafic-intermediate dykes in the MBPC. The first 564 occurred in the Middle Jurassic and the second in the Late Eocene. These dykes and plutons 565 preserved critical records of mantle source compositions and magma evolution from the MBPC, 566 records which are not accessible from coeval felsic rocks.

567

568 Based upon geochemical and isotopic data, Middle Jurassic Group-1 dykes originated from 569 metasomatized mantle peridotite containing garnet and spinel during active subduction of Neo-Tethys and evolved through fractional crystallization of clinopyroxene, plagioclase and apatite. 570 571 They experienced degrees of mixing with crust-derived melt. The Late Eocene Group-2 dykes are 572 products of low degree melting of asthenospheric mantle in the garnet-spinel transition zone. They 573 appear to have undergone fractional crystallisation involving olivine, clinopyroxene, amphibole and 574 apatite, but evidence for crustal assimilation is limited. Based upon U-Pb zircon ages, geochemical 575 and isotopic data, it is concluded that during the Cretaceous and Paleocene, a period of flat-slab 576 subduction took place, and the mantle wedge beneath the SaSiZ was removed from overriding 577 plate. We concluded that the sub-parallel subduction of Neo-Tethys spreading ridge to the trench 578 during Late Eocene, roll-back induced slab-tearing along the ridge axis, and subsequent break-off to generate the Mesopotamia slab, identified in the lower mantle by tomography studies. Break-off 579 580 was followed by localised asthenospheric upwelling, which in turn fed within-plate magmatism in 581 the N-SaSiZ, before short-lived re-establishment of Tethyan subduction and the resumption of arc 582 magmatism in the UDMA, prior to the final Arabia-Eurasia collision. This final episode of subduction generated the Zagros slab which can be identified today in the mantle transition zone. 583 584 Such a model is at odds with the continuous present of a subduction zone during the Cretaceous to 585 Late Cenozoic or with a slab break-off event at ~65 Ma, but fits the magmatic record and could be adopted in current mantle tomography models (van der Meer et al. 2018). 586

587

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#### Table 1

Whole rock chemical composition of the MBPC dykes (major oxides in wt. %, trace elements in ppm).

	Group-1: Middle Jurassic dykes							Group-2: Late Eocene dykes					
		Gabbro			Diorite				Gal	obro			
Location	MN2a	MN8	BR03b*1	BR02*1	BR07	M23*1	M33*1	MN04	M29*1	MN12	MN13	MN07	
Х	48.76006	48.80782	48.81201	48.80916	48.93265	48.85036	48.81733	48.80776	48.75991	48.76612	48.77025	48.80722	
Y	34.14333	34.10255	33.97396	33.97642	33.74269	34.06678	34.10433	34.12035	34.14344	34.08282	34.08297	34.11624	
						wt	.%						
SiO <sub>2</sub>	51.84	45.45	46.93	59.56	52.25	59.39	49.34	48.40	47.04	46.91	47.53	47.33	
TiO <sub>2</sub>	0.49	3.70	1.90	0.97	1.62	0.96	1.91	2.21	2.44	2.27	2.58	2.58	
Al <sub>2</sub> O <sub>3</sub>	12.61	15.16	16.39	19.6	17.19	19.78	17.36	17.67	14.88	16.01	16.45	16.16	
FeOT	7.25	12.26	11.15	8.07	9.07	7.41	9.98	10.55	12.05	10.76	11.56	12.08	
MnO	0.18	0.21	0.18	0.09	0.15	0.09	0.17	0.21	0.20	0.20	0.19	0.20	
MgO	9.75	5.01	6.89	2.31	4.00	2.36	5.08	5.15	4.90	5.10	5.02	5.62	
CaO	11.51	8.93	9.79	0.35	7.21	0.34	9.70	9.56	9.55	9.25	10.27	9.39	
Na <sub>2</sub> O	1.51	3.22	2.98	0.69	2.29	0.68	3.52	3.22	3.02	2.87	3.34	3.14	
K <sub>2</sub> O	1.15	1.05	0.92	4.65	1.90	4.74	1.38	1.33	0.80	1.35	1.05	1.38	
$P_2O_5$	0.04	0.69	0.31	0.14	0.34	0.14	0.46	0.52	0.55	0.60	0.71	0.77	
	2.86	3.02	2.13	3.62	2.60	4.03	0.68	0.85	3.52	3.67	1.10	1.26	
I otal	99.19	98.7	99.57	100.05	98.62	99.92	99.58	99.67	98.95	98.99	99.8	99.91	
Mg"	0.57	0.29	0.38	0.22	0.29	0.24	0.34	0.33	0.29	0.32	0.30	0.32	
C	2.62	12.04	8.12	14.08	20	14 30	2.80	1 77	2 20	1.50	1.28	1.8/	
Rh	34 70	147	151.40	216.1	141.60	205.7	2.00 40 1	44 04	17.64	50.26	31.27	29.79	
Ro	136	219.40	427.60	478	341	533.0	332	273.00	142	265.40	187.50	79.70	
Th	675	14 07	16 55	16.67	15.68	17 10	4 50	1 42	1.67	2 12	1 41	1 43	
U	1 41	1 93	1 60	1 76	2.02	2.70	0.90	0.30	0.36	0.48	0.34	0.31	
Nb	7.07	14.36	17.47	17.91	15.56	16.10	29.30	16.55	19.64	16.99	19.62	17.17	
Ph	14.36	22.42	28.35	8.87	22.04	2.95	1.87	7.75	10.83	3.76	3.63	5.74	
Sr	82.7	83.52	102.6	93.31	89.22	48.10	625.00	415.1	482.5	438.3	453.1	445.9	
Hf	3.78	5.10	5.03	4.97	4.90	4.90	1.00	1.11	1.05	1.36	1.16	0.77	
Zr	84	263.2	149.50	200.5	161.8	196.9	189.50	144.8	151.6	160	147	152	
Та	0.60	1.29	1.41	1.49	1.51	1.20	1.70	1.34	1.45	1.20	1.48	1.37	
Y	31.66	11.16	9.81	9.59	20.46	31.10	24.70	26.07	28.89	29.60	29.03	26.69	
Cr	904.00	131.10	140.10	153.40	119.40	83.70	85.61	79.07	85.61	101.50	54.96	79.62	
Ni	273.10	44.81	48.21	59.45	47.62	57.10	32.91	32.09	32.91	39.25	26.53	29.83	
Со	37.06	17.11	17.42	19.81	18.95	17.60	38.61	35.59	38.61	37.69	32.49	34.42	
V	198.80	142.20	149.80	179.20	143.90	115	241.50	227.20	241.50	217.70	180.90	230.50	
La	19.79	38.84	45.12	46.3	41.82	46.3	33.90	18.39	20.04	20.67	25.04	15.71	
Ce	44.39	81.92	93.03	93.2	87.2	91	61.70	39.19	42.45	44.13	51.13	34.07	
Pr	6.058	9.03	10.20	10.12	9.769	10.11	6.77	5.032	5.441	5.702	6.334	4.373	
Nd	25.38	23.4	37.43	19.9	19.1	18.6	25.20	21.79	23.87	25.09	27.31	19.06	
Sm	5.977	5.19	6.69	4./8	4.76	4.35	5.24	5.188	5.634	5.938	6.225	4./9/	
Eu	1.274	2.21	1.37	1.05	1.0/	1.55	1.81	1./49	2.104	2.343	2.208	1.552	
Gu	3.328	3.48	3.09	5.55	3.5	4.70	3./1	3.187	0.0026	3.808	0.0822	4.919	
10 Dv	5.624	0.8	0.02	0.79	0.8	5.06	0.85	5 097	5 2 2 2	5 572	5 627	5.054	
Бу	1 1 2 3	0.85	2.08	0.88	4.09	0.80	4.57	0.0676	1.045	1.056	1.087	0.0727	
Fr	3 207	2 48	0.40	2 79	2 63	24	2 41	2 532	2 808	2 765	2 811	2 612	
Tm	0.4683	0.33	0.10	0.34	0.36	0.34	0.38	0.3406	0.375	0.3679	0.3694	0.3537	
Yh	2,957	2.01	0.53	2.22	2.09	2.21	2.37	1 971	2.228	2.157	2.12	2.067	
Lu	0.4164	0.29	0.07	0.31	0.34	0.29	0.34	0.2563	0.3032	0.2844	0.2693	0.2803	
ΣREE	122.90	177.27	204.17	193.06	181.53	188.56	152.21	108.52	117.96	122.89	137.53	96.64	
Nb/Y	0.22	1.29	1.78	1.87	0.76	0.52	1.19	0.63	0.68	0.57	0.68	0.64	
Zr/Nb	11.88	18.33	8.56	11.19	10.40	12.23	6.47	8.75	7.72	9.42	7.49	8.85	
K/Nb	1350.7	607.04	437.19	2235.2	1013.7	2548.5	389.9	667.16	338.16	659.66	444.29	667.25	
Ba/Nb	19.24	15.28	24.48	26.69	21.92	33.11	11.33	16.50	7.23	15.62	9.56	4.64	
La/Nb	2.80	2.70	2.58	2.59	2.69	2.88	1.16	1.11	1.02	1.22	1.28	0.91	
Ce/Ce*	1.00	1.07	0.85	1.03	1.05	1.01	0.98	0.98	0.97	0.96	0.97	0.99	

 $\frac{\text{Ce/Ce}^{*} \quad 1.00 \quad 1.07 \quad 0.85 \quad 1.03 \quad 1.05 \quad 1.01 \quad 0.98 \quad 0.98 \quad 0.97 \quad 0.96 \quad 0.97 \quad 0.99}{\text{LOI: loss on ignition}}$ EO?: Total Fe is represented as FeO. Ce/Ce\* is calculated by Ce<sub>CN</sub>/La<sub>CN</sub><sup>2/3</sup>× Nd<sub>CN</sub><sup>1/3</sup>. \*<sup>1</sup> Major elements determined by XRF in Naruto University, Japan; and trace elements by ICP-MS in ACME laboratory, Canada. Other data were produced in University of the Ryukyus.

	Group-2: Late Eocene dykes											
					Gat	obro						
Location	M34* <sup>2</sup>	M37* <sup>2</sup>	M50* <sup>2</sup>	M25b*2	M54* <sup>2</sup>	M38* <sup>2</sup>	M58* <sup>2</sup>	M22* <sup>2</sup>	BR04*2	BR03d*2		
Х	48.81394	48.81497	48.67419	48.86703	48.57739	48.82069	48.69875	48.87164	48.24472	48.81139		
Y	34.10503	34.10214	34.22044	34.08203	34.28564	34.09997	34.23411	34.07728	33.56667	33.97397		
					wt	.%						
SiO <sub>2</sub>	47.99	48.67	46.97	48.03	46.68	47.97	48.77	46.98	48.06	48.11		
TiO <sub>2</sub>	2.10	1.89	2.63	2.03	3.80	1.96	2.08	3.20	1.87	1.88		
$Al_2O_3$	16.78	16.97	15.61	16.75	14.33	16.31	17.12	15.46	17.02	16.88		
FeO	10.67	10.76	11.65	10.36	13.90	10.96	9.80	12.17	10.19	10.20		
MnO	0.17	0.15	0.21	0.16	0.22	0.17	0.15	0.19	0.17	0.18		
MgO	5.96	6.65	5.21	6.13	4.06	/.1/	5.08	6.30	7.09	/.16		
	10.70	9.50	11.18	10.78	9.48	10.21	10.41	9.93	10.04	10.07		
	3.16	2.81	1.6/	3.25	3.98	3.00	3.82	2.96	3.20	3.13		
	0.91	0.34	5.00	0.98	0.80	0.09	0.52	0.96	0.94	0.94		
P <sub>2</sub> O <sub>5</sub>	0.38	0.54	0.50	0.30	1.10	0.33	0.32	1.05	0.27	0.31		
Total	0.91	99.5	99.35	0.87	99.57	90.78	0.95	99.68	99.81	0.87		
Mo#	0.36	0.38	0.31	0.37	0.23	0.40	0.34	0.34	0.41	0.41		
1115	0.50	0.50	0.51	0.57	0.25 nr	0.40	0.54	0.54	0.41	0.41		
Cs	1 90	5.00	1 70	1 90	1 30	1 30	2.20	3.00	1.60	1.80		
Rb	10.90	30.00	118.00	17.90	30.10	13.70	20.30	22.50	34.10	34.30		
Ba	152	183	312	175	170	107	298	228	107	116		
Th	1.70	2.60	1.90	1.60	2.50	1.30	1.50	2.50	1.30	1.60		
U	0.50	0.70	0.60	0.40	1.10	0.30	0.50	0.70	0.30	0.40		
Nb	18.10	16.60	20.50	15.20	27.30	14.20	15.20	21.70	11.50	11.90		
Pb	2.92	3.30	11.29	2.43	5.30	2.76	2.81	1.38	1.49	1.88		
Sr	467.70	474.40	427.20	524.80	368.50	464.00	619.60	450.40	424.10	444.00		
Hf	1.49	1.60	1.51	1.72	1.23	1.42	0.71	1.26	0.81	1.31		
Zr	141.60	152.40	144.60	139.70	197.60	130.50	124.60	201.60	156.70	159.10		
Та	1.30	1.20	1.30	1.00	2.00	0.90	1.00	1.30	0.90	0.80		
Y	24.00	26.30	24.80	27.10	37.20	22.00	23.50	34.80	27.00	27.40		
Cr	9.30	38.40	43.10	13.50	1.70	3.50	10.20	9.70	23.70	111.10		
Ni	17.80	38.60	25.90	19.60	5.40	26.40	19.40	20.60	27.90	18.20		
Co	17.90	27.30	24.80	18.10	21.80	17.90	16.00	10.80	17.20	13.90		
V	3 10 1	115	131	17.7	25.2	54 16.6	42	26.4	54 14 5	44		
La	36.7	30.0	23.4 14.7	36	72.2	31.7	43.6	20.4 53.1	32.6	32.5		
Pr	4 75	4 81	5 48	4 55	8.81	4 12	5 33	6 3 9	4 17	43		
Nd	19.9	19.1	23.4	18.6	36.6	16.3	20.3	26.1	18.7	19.3		
Sm	4.78	4.76	5.19	4.35	8.26	4.24	5.01	6.09	4.52	4.43		
Eu	1.65	1.67	2.21	1.53	3.19	1.63	2.33	2	1.49	1.63		
Gd	5.35	5.3	5.48	4.76	9.07	4.66	5.88	6.77	5.31	5.15		
Tb	0.79	0.8	0.8	0.72	1.32	0.72	0.81	1.04	0.81	0.8		
Dy	4.43	4.69	4.44	5.06	8.05	4.32	4.61	6.19	4.86	5.16		
Ho	0.88	1	0.85	0.89	1.49	0.84	0.87	1.29	0.91	1.07		
Er	2.79	2.63	2.48	2.4	3.96	2.44	2.39	3.85	2.87	2.98		
Tm	0.34	0.36	0.33	0.34	0.49	0.33	0.32	0.51	0.4	0.41		
Yb	2.22	2.09	2.01	2.21	3.17	1.98	1.84	3.44	2.76	2.72		
Lu	0.31	0.34	0.29	0.29	0.43	0.32	0.28	0.51	0.39	0.41		
∑REE NL (M	103.99	106.45	121.06	99.40	192.24	90.20	115.07	145.68	94.29	96.46		
Nb/Y 7/Nb-	0.75	0.63	0.83	0.56	0.73	0.65	0.65	0.62	0.43	0.43		
LT/IND K/NIL	/.82	9.18 520 7	1216.2	9.19 536 0	1.24	9.19 402 2	0.20 636 7	9.29 360 0	13.03	13.37		
N/IND Ro/Nh	913.9 8.40	11 02	1210.5	11 51	6 22	403.3	10.61	10 51	075.0	030.5		
La/Nh	1.06	1 1 4	1 1 1 4	1 16	1 29	1 17	1 41	1 22	1.30	1 31		
Ce/Ce*	2.38	1.02	0.93	0.97	0.98	0.93	1.00	0.98	1.00	0.94		

 $Ce/Ce^*$ 2.381.020.930.970.980.931.000.981.00\*<sup>2</sup>Major elements determined by XRF in Naruto University, Japan; and trace elements by ICP-MS in ACME laboratory, Canada<br/>reported in Deevsalar *et al.* (2014).Ce/Ce\* is calculated by  $Ce_{CN}/La_{CN}^{23} \times Nd_{CN}^{1/3}$ .

 Table 2

 Sr-Nd isotopic composition of the mafic-intermediate dykes from MBPC.

	<sup>87</sup> Sr/ <sup>86</sup> Sr	Rb	Sr	Error	<sup>87</sup> Sr/ <sup>86</sup> Sr	143Nd/144Nd	Sm	Nd	143Nd/144Nd	a (T)	Error
	(measured)	(ppm)	(ppm)	×10-6	(initial)	(measured)	(ppm)	(ppm)	(initial)	$\epsilon_{\rm Nd}(1)$	×10 <sup>-6</sup>
	Group-1 ~ 168 Ma										
BR02	0.72948	216.1	93.1	6	0.71362	0.51216	7.03	38.28	0.51204	-7.54	4
BR07	0.72464	141.6	89.22	6	0.71329	0.51218	7.13	37.15	0.51206	-7.18	6
MN2a	0.70831	34.7	82.7	7	0.70770	0.51228	5.98	25.38	0.51212	-5.86	15
MN8	0.72679	147	83.52	6	0.71427	0.51223	6.25	33.60	0.51211	-6.10	5
BR03b	0.72560	151	98.52	6	0.71510	0.51214	6.691	37.43	0.51202	-7.92	3
					Group-	2 ~ 42 Ma					
M38	0.70712	13.7	464.0	7	0.70707	0.51281	27.71	34.90	0.51268	1.8	6
M22	0.70546	22.5	450.4	7	0.70537	0.51273	39.80	55.89	0.51261	0.6	5
BR04	0.70551	34.1	424.1	6	0.70538	0.51279	29.54	40.04	0.51267	1.6	7
M50	0.70780	118	427.2	6	0.70732	0.51278	33.92	50.11	0.51267	1.7	5
M25b	0.70503	17.9	524.8	8	0.70498	0.51279	28.43	39.83	0.51267	1.7	6
M33	0.70754	40.1	625.0	7	0.70743	0.51273	34.25	53.96	0.51262	0.8	4
MN7	0.70602	29.79	89.22	6	0.70591	0.51280	4.80	19.06	0.51276	3.5	5
M29	0.70500	17.64	482.5	6	0.70494	0.51279	5.63	23.87	0.51275	3.3	10

#### Table 3

Lead isotope composition of the mafic-intermediate dykes from MBPC.

	<sup>206</sup> Pb/ <sup>204</sup> Pb (measured)	<sup>207</sup> Pb/ <sup>204</sup> Pb (measured)	<sup>208</sup> Pb/ <sup>204</sup> Pb (measured)	U	Th	Pb	<sup>206</sup> Pb/ <sup>204</sup> Pb(i)	<sup>207</sup> Pb/ <sup>204</sup> Pb(i)	<sup>208</sup> Pb/ <sup>204</sup> Pb(i)
	( )	(	G	roup-1	~ 168 N	Ia			
BR02	18.76	15.69	39.56	1.41	6.75	14.36	18.44	15.68	38.57
BR07	18.58	15.65	38.96	2.01	15.68	22.04	18.43	15.64	38.58
MN2a	18.79	15.70	39.03	1.41	6.75	14.36	18.63	15.70	38.78
MN8	18.57	15.64	38.84	1.93	14.07	22.42	18.43	15.63	38.51
BR03b	18.55	15.68	38.97	1.6	16.55	28.35	18.46	15.68	38.66
			0	Group-2	~ 42 M	а			
M38	18.46	15.65	38.69	0.3	1.3	2.76	18.417	15.645	38.629
M22	18.46	15.59	38.61	0.7	2.5	1.38	18.251	15.583	38.373
BR04	18.53	15.66	38.75	0.3	1.3	1.49	18.447	15.654	38.636
M50	18.60	15.65	38.89	0.6	1.9	11.29	18.578	15.650	38.869
M33	18.71	15.67	38.98	0.9	4.5	1.87	18.519	15.662	38.657
M29a	18.44	15.63	38.65	0.36	1.66	10.83	18.425	15.634	38.633
MN13	18.73	15.67	38.93	0.34	1.41	3.63	18.693	15.664	38.875
MN12	18.77	15.67	38.97	0.48	2.12	3.76	18.713	15.664	38.899
MN4	18.48	15.66	38.74	0.30	1.41	7.75	18.468	15.658	38.719
MN7	18.68	15.67	38.84	0.31	1.43	5.74	18.654	15.668	38.809
-	Standard	l Sample: Mean	(n=20)						
	16.9437	15.5010	36.7236	-					
		2SD		-					
	0.0023	0.0028	0.0053						

		Mantle	e wedge		High crustal level							
	Source	Mix	ing	NMBM*4	Modal Frac	tional M	lelting	AI	FC	Mixing	and FC	
	PM*1	A*2	EM* <sup>3</sup>	f <sub>m</sub> =	Source	M par	Melting parameter		R*9	Mixing*10	FC*11	
				0.15*5	Greywacke*6	Kd* <sup>7</sup>	$f_m = 0.25$	0.4	0.3	$f_{mix} = 0.4$	$f_c = 0.6$	
Rb	0.63	71.04	4.16	27.28	72	1.43	54.87	110	0.83	37.90	62.20	
Ba	6.99	1019.20	57.60	379.26	426	2.34	214.48	1082	2.61	310.23	490.64	
Th	0.09	5.10	0.34	2.22	9	0.10	7.17	11.	.07	6.35	10.46	
U	0.02	1.11	0.08	0.50	2	0.05	0.14	2.	51	0.50	0.84	
Nb	0.71	4.64	0.91	6.01	8.4	1.42	6.44	20.	.31	6.13	10.20	
Κ	249.98	21038.38	1289.40	8164.83	16602.8	0.87	18301.9	29504.28		12291.53	20001.93	
La	0.69	16.72	1.49	9.44	34	0.16	47.62	13.18		32.19	27.13	
Ce	1.77	34.55	3.41	20.76	58	0.16	78.69	22.	.60	57.01	45.69	
Pr	0.28	1.64	0.34	2.01	6.1	0.12	6.25	1.	63	5.20	3.29	
Sr	21.10	473.80	43.73	264.40	201	0.96	206.88	384	1.42	241.61	295.96	
Р	94.98	486.66	114.56	654.94	567.3		-	197	8.63	-		
Nd	1.35	15.51	2.06	11.41	25	0.14	29.83	5.	30	24.76	12.36	
Zr	11.20	254.15	23.35	129.59	302	0.06	68.68	465	5.25	148.81	235.87	
Sm	0.44	2.79	0.56	2.77	4.6	0.14	5.75	0.	76	5.05	1.94	
Eu	0.17	0.61	0.19	0.90	1.2	0.60	1.65	0.	19	1.23	0.43	
Ti	1300.24	1922.51	1331.36	5682.87	4315.17	0.02	0.27	1538	32.34	3411.99	5455.17	
Dy	0.74	1.60	0.78	3.11	3.4	0.17	4.88	0.	73	4.55	1.96	
Y	4.55	8.81	4.77	15.10	26	0.77	31.05	38.	.16	21.73	30.79	
Yb	0.49	0.72	0.50	1.34	2.1	0.16	2.93	1.	10	2.44	2.02	
Lu	0.07	0.10	0.08	0.18	0.37	0.16	0.51	0.	21	0.40	0.36	

# Table 4 The results of melting and mixing model calculations for Middle Jurassic gabbro-dioritic dykes.

\*1 PM: primitive mantle composition (Sun & McDonough 1989)

\*<sup>2</sup> Subduction sediment-derived fluids/melts: 0.55SF/0.45SM. The parameters used in modeling are: normalizing values from Sun & McDonough (1989), Bulk subducted sediment composition (BOS) are from Plank & Langmuir (1998), sediment fluid partition coefficients (DSF) from Johnson & Plank (1999). Sediment melt (SM) calculated by 5% modal fractional melting of bulk subducted sediment, sediment fluids (SF) composition calculated using BOS and DSF(C<sub>BOS/DSF</sub>).

\*<sup>3</sup> Middle Jurassic Enriched Mantle (80% PM + 20% A)

\*4 NMBM: Non-Modal Batch Melting (Mode-I, Supplementary item 3, Table S3)

 $*^{5}$  fm: degree of melting = 15%; partial melting of metasomatised garnet-spinel Lherzolite (Supplementary Item 3, Table S3)

\*<sup>6</sup>Upper crust-greywacke composition is from Wedepohl (1995)

\*7 Bulk partition coefficient calculated for greywacke (35%Qtz + 30%Bio + 15% Plg + 5% AF+ 15% Mus)

\*<sup>8</sup> F<sub>c</sub>: degree of fractional crystallization; Bulk partition coefficient calculated for 70.66% Cpx + 26.77% Pl + 0.44% Ap \*<sup>9</sup> R-value: EM/UCC-greywacke

\*10 Mixing of EM and Greywacke-melt. fmix: degree of mixing

\*<sup>11</sup> Fractional crystallization of 70.66% Cpx + 26.77% Pl + 0.44% Ap from crustally-contaminated EM

#### Table 5

The results of melting, mixing and fractional crystallization model calculations for Late Eocene gabbroic dykes.

		Crustal level								
	Source	Source NMBM*2 Mixing*6								
	Source	Mode-II	Mode-III	Mode-IV	witxing	crystallization				
	PM*1	$F_{melting} =$	$F_{melting} =$	$F_{melting} =$	Grt Lhz Melt <sub>(5%)</sub> + Spl	Residual melt* <sup>7</sup>				
		0.03***	0.03**	0.15**	Lhz Melt(15%)	F= 0.45				
Rb	0.63	12.39	12.40	4.20	7.48	16.59				
Ba	6.99	136.61	136.74	46.29	82.47	182.71				
Th	0.09	1.67	1.66	0.56	1.00	2.17				
U	0.02	0.41	0.41	0.14	0.25	0.55				
Nb	0.71	14.05	14.06	4.74	8.47	18.79				
Κ	249.98	4908.51	4902.47	1659.08	2956.44	6517.45				
La	0.69	12.50	12.41	4.48	7.65	14.99				
Ce	1.77	30.00	29.68	11.34	18.68	35.84				
Pr	0.28	5.27	5.23	1.82	3.19	5.80				
Sr	21.10	347.09	342.83	133.91	217.48	468.20				
Р	94.98	1268.97	1255.22	549.71	831.91	1842.80				
Nd	1.35	18.64	18.18	8.17	12.17	20.96				
Zr	11.20	161.90	157.07	69.41	104.48	222.93				
Sm	0.44	4.68	4.43	2.49	3.27	5.22				
Eu	0.17	2.62	2.40	1.10	1.62	2.57				
Ti	1300.24	11573.47	11427.96	6672.09	8574.44	18366.38				
Dy	0.74	3.19	2.50	3.79	3.27	5.22				
Y	4.55	18.35	14.32	22.89	19.46	37.54				
Yb	0.49	1.16	0.84	2.47	1.82	3.29				
Lu	0.07	0.14	0.10	0.34	0.24	0.46				

\*1 PM: primitive mantle composition (Sun & McDonough 1989)

<sup>\*2</sup> NMBM: Non-Modal Batch Melting
<sup>\*3</sup> 5% partial melting of Garnet-Spinel Lherzolite (Mode-II, Supplementary Item 3, Table S2, S3)

<sup>x4</sup> 5% partial melting of Garnet Lherzolite (Mode-III, Supplementary Item 3, Table S2, S3)
 <sup>x5</sup> 15% partial melting of Spinel Lherzolite (Mode-IV, Supplementary Item 3, Table S2, S3)
 <sup>x6</sup> EM: mixing between melts originated from garnet lherzolite and spinel lherzolite source

 $*^{7}$  Residual melt of fractional crystallization of 0.5 Ol + 0.35 Cpx + 0.1 Am ± 0.05 Apt from enriched mantle magma

Fig. 1



## **Fig. 2**







Fig. 4



Fig. 5









**Fig.** 7













#### **FIGURE CAPTIONS:**

**Fig. 1** (a) Simplified distribution of the main tectonic subdivisions of Iran. (b) Distribution of magmatic complexes in the N-SaSZ and the position of the MBPC. The MBPC is marked by two rectangles (8 and 9). 1- Oshnavieh (Os); Urumieh; 2- Naghade-Khalfe-Pasveh (N-K-P); 3- Piranshahr (Pi), Saqez, Takab (Ta); Miandoab, Mahabad (Mh); Sanandaj (Sn) and Baneh (Ba); 4- Qorveh (Qr); 5- Kamyaran (Ka); 6- Almogholagh (Al); 7- Alvand (Al), 8- Malayer (Ma); 9- Boroujerd (Br). (c) The areas marked by (8) and (9) respectively contain geological maps of the MBPC indicating the distribution of mafic–intermediate dykes. The locations of dyke samples with corresponding geochemical data are labelled.

Fig. 2 Representative CL images of zircon grains.

**Fig. 3** Conventional U–Pb concordia plots (Tera and Wasserburg, 1972) and weighted mean zircon ages for MBPC magmatic dykes. (a) Middle Jurassic diorite. (b) Middle Jurassic gabbro. (c, d) Late Eocene gabbro.

**Fig. 4** (a-f) Harker diagrams of major oxides (wt. %) vs.  $SiO_2$  (wt. %) for the gabbro–dioritic dykes from the MBPC. Arrows indicate changes in composition relative to magma evolution. Thin dashed and solid lines are respectively mineral fractionation and accumulation vectors from an assumed parental magma (MN<sub>2a</sub>). (g-i) TiO<sub>2</sub> (wt. %) vs. Mg<sup>#</sup>, CaO/Al<sub>2</sub>O<sub>3</sub> and Al<sub>2</sub>O<sub>3</sub> vs. MgO (wt. %).

**Fig. 5** (a, b) Primitive Mantle (PM)-normalised trace element diagrams and Chondritenormalised REE patterns for for the Middle Jurassic dykes, (c, d) PM-normalised trace element diagrams and Chondrite-normalized REE patterns for Late Eocene gabbroic dykes. PM compositions are from Taylor and McLennan (1985). Chondrite compositions are from Nakamura (1974).

**Fig. 6** Isotope ratio diagrams for the MBPC dykes. (a) (<sup>87</sup>Sr/<sup>86</sup>Sr)<sub>i</sub> vs. εNd. 1- Hawaii (PREMA): (Hofmann 2005); 2- Upper Jurassic – Lower Cretaceous Kapan arc (Mederer *et al.* 2013); 3- Late Cenozoic Georgian and NW Armenian mafic alkaline rocks (Neill *et al.* 2013, 2015); 4- MBPC

Late Eocene gabbroic intrusions (Deevsalar *et al.* 2017); 5- MBPC Middle Jurassic gabbroic intrusions (Deevsalar *et al.* 2017); 6- OIB (Zhang *et al.* 2002); 7- MBPC granitoids (Ahmadi-Khalaji *et al.* 2006; Ahadnejad *et al.* 2010; 2011); 8- Alvand granite (Shahbazi *et al.* 2010; Sepahi *et al.* 2017). (b, c) Initial Sr and Nd isotopic values plotted against <sup>206</sup>Pb/<sup>204</sup>Pb. (d, e) <sup>206</sup>Pb/<sup>204</sup>Pb vs. <sup>207</sup>Pb/<sup>204</sup>Pb and <sup>208</sup>Pb/<sup>204</sup>Pb for the MBPC dykes. Labelled rectangular fields in each plot show mantle reservoirs (EMI, EMII, HIMU, and DMM) taken from Shimuda (2009). Also shown are Late Cenozoic lavas from Karacadag, S Turkey (Ekici *et al.* 2014) and E Iran (Kheirkhah *et al.* 2015); Upper Jurassic-Lower Cretacous mafic rocks from the Kapan Arc (Mederer *et al.* 2013); and Late Cenozoic mafic lavas of NE Armenia (Neill *et al.* 2015). The gray triangles are used for screening the proportion of each of these mantle reservoirs in the source region.

**Fig. 7** (a) Hf/Yb vs. Nb/Zr plot (John *et al.* 2004). PM: primitive mantle, EM: enriched mantle, DM: depleted mantle. (b) Th/Yb vs. Ta/Yb plot (Pearce 1983). Enclosed area in this figure: 1-Upper Jurassic- Late Cretaceous rocks from the Kapan arc (Mederer *et al.* 2013), 2- alkaline rocks from East Iran (Kheirkhah *et al.* 2015). VAB: volcanic arc basalts, Th; tholeiitic, Alk: alkaline, Tr: transitional, CA; calc-alkaline. (c) La/Yb vs. Zr/Nb (Aldanmaz *et al.* 2006). WAM (Western Anatolian Mantle); (d) Ta/Zr vs. Nb/Hf. Green filed shows the MBPC Late Eocene Gabbros (Deevsalar *et al.* 2017). (e) Comparison of the concentration of selected trace elements in the Late Eocene gabbroic dykes with E-MORB, N-MORB and OIB (values from Sun & McDonough 1989).

**Fig. 8** Trace element modelling of magmatic processes for Group-1 dykes. (a) AFC trajectory on a ( ${}^{87}$ Sr/ ${}^{86}$ Sr)<sub>i</sub> vs. Sr (ppm) plot. Parameters are R: 0.3, D<sub>Sr</sub>: 0.6, and contaminant is upper crustagreywacke (taken from Wedepohl 1995; 35% Qtz + 30% Bio + 15% Plg + 5% Alk + 15% Mus). AFC calculated using the equations of DePaolo (1981). (b) Mixing model trajectory plotted on ( ${}^{87}$ Sr/ ${}^{86}$ Sr)<sub>i</sub> vs. 1/Sr (ppm). End-members are metasomatised mantle-derived magma (Supplementary Item 1, Table S4) and greywacke-derived melt (modal fractional melting of upper crustal-greywacke, Table 5; F<sub>melting</sub>: 0.2). The degree of mixing (F<sub>mixing</sub>) is 0.4. Each point on the curves represents 10% assimilation. Mixing equation is from Langmuir *et al.* 1978. PM- normalized trace element pattern of: (c) metasomatically-enriched mantle magma ( $F_{melting}$ : 15%) and AFC-trajectory, and (d) mixing and FC model trajectories modelled on a multi-element plot of Group-1 dykes. Parameters for non-modal batch melting and Rayleigh fractional crystallization (FC) are given in Supplementary Item 3, Tables S2 and S3.

**Fig. 9** Non-modal batch melting vectors are plotted on (a) La/Sm vs. La (ppm), and (b) Zr/Nb vs. La/Yb logarithmic plots for Group-2 dykes. M: enriched mantle source was suggested for MBPC gabbros (Deevsalar et al. 2017), M': depleted asthenospheric source. (c) Trace element modelling of melting, mixing and fractional crystallization. The parameters used in this model are in Supplementary Item 3, Table S2 and S3. The results of melting (calculated in three modes) and mixing event in the source region, and fractional crystallization model in crustal level are given in Table 5. Normalising values are from Sun & McDonough (1989). The mixing equation is from Langmuir *et al.* (1978).

**Fig. 10** (a) Schematic representation of the proposed geodynamic scenario for emplacement of within-plate type igneous rocks in the SaSiZ and the UDMA during the Late Eocene. The model involves roll-back of the Mesopotamia slab, combined with detachment beneath the SaSiZ. Next, subduction of the Zagros slab removes the last remaining Neo-Tethys lithosphere between Iran and Arabia and shifts the locus of renewed arc magmatism east to the UDMA. Modified after Zhang *et al.* (in press). The location of the MBPC and Piranshahr (Pi) are shown in Fig. 1b. (b) Summary of the present-day location of the two slab remnants from Neo-Tethyan subduction beneath Iran. Based on seismic tomography results published in van der Meer *et al.* (2018).