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Academic Support Office, Durham University, University Office, Old Elvet, Durham DH1 3HP e-mail: e-theses.admin@dur.ac.uk Tel: +44 0191 334 6107 http://etheses.dur.ac.uk Breakup of the Gondwana supercontinent: East African perspectives from the Early Jurassic to Cretaceous



Jordan Jerad John Phethean

A thesis submitted in partial fulfilment of the requirements for the degree of Doctor of

Philosophy at Durham University

**Department of Earth Sciences** 

October, 2017

### 1 Abstract

2

3 Accurate mapping of first-order tectonic features such as oceanic fracture zones and continental 4 margins is vital for the production of reliable plate reconstructions. These reconstructions allow 5 for a better understanding of the palaeo-configuration of continental fragments within 6 Gondwana and ultimately provide insight into how and why supercontinents break apart. 7 Detection of spreading lineaments within the heavily sedimented Western Somali Basin (WSB) 8 has been achieved using a novel technique based on directional derivatives of free-air gravity. 9 This new lineament dataset allows for the construction of a high-resolution plate tectonic 10 reconstruction of the WSB, which is in good agreement with ocean magnetic data and the 11 position of the abandoned WSB spreading centre. The model also reveals a change in spreading 12 direction, from NNW-SSE to N-S, during the Late Jurassic. This controversial spreading 13 direction change places the origin of Madagascar within the Tanzania Coastal Basin (TCB), 14 inboard of the Davie Fracture Zone (DFZ), which was previously believed to be the continent-15 ocean transform margin of the WSB. This tight-fit of Gondwana fragments prior to continental 16 breakup necessitates a reassessment of both the crustal nature of the TCB, which is shown to be 17 partly oceanic in nature, and of the nature of the margins surrounding the WSB. The northern 18 margins of the WSB are likely orthogonally rifted margins. However, the western margins are 19 likely highly segmented and/or obliquely rifted margins. The model also predicts a large 20 transform offset along the Rovuma Basin.

21 Systematic gravity modelling and combined seismic investigations along the Rovuma basin 22 reveals the 'Rovuma Transform Margin', which offsets the obliquely rifted margins of northeast 23 Mozambique and Tanzania. The discovery of this transform margin confirms the initial SSE 24 plate motion predicted from gravity lineament analysis and plate reconstructions, and shows that 25 the breakup of the Gondwana supercontinent occurred not just along pre-existing lithospheric weaknesses associated with the Karoo rift system, but also along newly developed highly 26 27 oblique deformation zones as well. The final breakup of the Gondwana supercontinent, which 28 followed extensive and episodic Karoo aged rifting, was coincident with extensive magmatism 29 in Mozambique and may therefore have been triggered by the interaction of several facilitators 30 of continental breakup (i.e. oblique rifting, pre-existing weaknesses, and magmatism).

The oblique breakup of Gondwana along the TCB led to the development of a segmented midocean ridge system within this basin, offset by SSE trending fracture zones. These fracture zones were incompatible with the N-S spreading that followed the Late Jurassic change in plate motion, resulting in the abandonment of mid-ocean ridge segments and compression within this 35 basin. This compression led to the formation of the 250 km long Tanzania Coastal Basin thrust 36 belt, the largest intraplate oceanic thrust belt yet discovered. The cessation of compression 37 within the TCB followed the development of the DFZ, which propagated from south to north. This structure was subsequently dominated by transpression throughout its history, suggesting it 38 39 was not perfectly compatible with plate driving forces. Formation of the DFZ along aligned weak rifted margins and young oceanic crust may have resulted in the mismatch of plate 40 41 motions and driving forces, and also suggests a first order 'top-down' control on plate motions 42 during the breakup of Gondwana.

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## **Declaration**

I declare that this thesis, which I submit for the degree of Doctor of Philosophy at Durham University, is my own work and not substantially the same as any which has previously been submitted at this or any other university. Jordan J. J. Phethean Durham University October 2017 Signed: \_\_\_\_\_ Jordan Phethean Date: © The copyright of this thesis rests with the author. No quotation from it should be published without prior written consent and information derived from it should be

acknowledged.

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# **1. Introduction**

### 302 **1.1. Motivation and methods summary**

303 Rifted margins are of significant interest as they provide insight into the causes and 304 mechanisms of supercontinent breakup (e.g. Bercovici and Long, 2014). But besides this 305 academic interest, they have also provided a reliable worldwide source of hydrocarbon resources for over 30 years (Levell et al., 2010) and economic and political forces have 306 307 recently driven a strong shift towards the exploration of deep-water (500-2500 m below sea 308 level) regions of these margins (White et al., 2003). A wealth of research, focused on the 309 evolution of volcanic and magma-poor rifted margins (e.g. Afilhado et al., 2008; Brune et al., 2013; Geoffroy et al., 2015; Manatschal, 2004; Reston and Mcdermott, 2014), shows 310 311 that the extent of continental crust and the nature of the continent-ocean transition in these 312 deep-water settings are highly variable (e.g. Stab et al., 2016; Unternehr et al., 2010), and 313 have strong impacts on the presence and maturation of source rocks (e.g. Waples, 2002). 314 Transform margins, on the other hand, have been relatively poorly studied (Mercier de 315 Lépinay et al., 2016) and, in the past, have not presented a major target for the hydrocarbon 316 industry. The 2007 discovery of the Jubilee field along the West Africa transform margin 317 (Dailly et al., 2013), however, opened up a new play type and dramatically increased interest in transform margins. East Africa is, at a basic level, understood to represent such a 318 319 transform margin, and as such new interest in the region is rapidly growing. However, the 320 details of kinematic and dynamic margin development, as well as the extent of continental 321 crust is poorly understood along the East African margins and at transform margins in 322 general.

Reducing these uncertainties invites further interest in the exploration of the East African rifted margins for hydrocarbon resources, and is crucial for the successful discovery and development of said resources. In order for East Africa to benefit economically from its potential natural resources, and to allow associated benefits of development opportunities, stability, and prosperity to East Africa, it is therefore necessary to constrain these unknowns.

Plate motions during Gondwana disassembly, and the initial configuration of continental
fragments and deformation zones in Gondwana, are a matter of particular controversy (e.g.
Bunce and Molnar, 1977; Coffin and Rabinowitz, 1987; Davis et al., 2016; Gaina et al.,
2013; Klimke et al., 2017; Lawver and Scotese, 1987; Norton and Sclater, 1979; Pinna,
1995; Project, 2015; Reeves, 2014; Scrutton et al., 1981; Sègoufin and Patriat, 1980;

333 Shackleton, 1996; Smith and Hallam, 1970; Windley et al., 1994), and so the development 334 of accurate and reliable plate tectonic models for the region is fundamental before more 335 detailed understanding of the margin can be attempted. In order to constrain the plate 336 motions during the disruption of Gondwana we apply novel processing techniques to 337 recently available and highly accurate marine satellite gravity data. This successfully allows 338 us to develop a highly accurate plate reconstruction for the region that, due to the nature of 339 continental margins being dependant on plate motions relative to zones of lithospheric 340 deformation (e.g. Basile and Braun, 2016), indicates which styles of continental margin 341 should develop along different regions of the East Africa coast; a crucial step in the 342 development of our understanding of this enigmatic region.

The extent of continental crust along rifted margins is another key parameter in resource prospectivity, however, is not well constrained by such kinematic plate tectonic reconstructions. We therefore also perform rigorous 2D gravity modelling of the margin structure to locate the boundary between oceanic and continental crust. This confirms the proposed natures of the East African margins gained from Plate tectonic modelling.

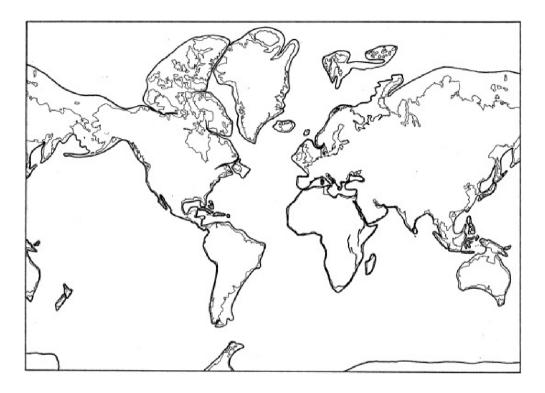
348 Finally, in order to better understand transform margin development in general, which can 349 not only reduce risk and interest in the study region but can also contribute to our 350 understanding of transform margins worldwide, we perform a detailed study of the structural 351 development of this margin through the interpretation of high quality deep imaging seismic 352 reflection data. This interpretation allows for a model of the development of the East African 353 margins to be constructed, which bears strong similarities to other transform systems 354 worldwide (e.g. Whittaker et al., 2016; Schiffer et al., 2017) and may therefore contribute to 355 a unified understanding of transform margin processes.

#### **1.2. Plate tectonics, the Wilson cycle, and what drives them**

The theory of plate tectonics was formulated in 1968 (Morgan, 1968) and is one of the fundamental principles on which the work presented herein relies. The roots of this theory, however, lie much earlier.

As early as 1596, the 'lock and key' geometry of Africa-Europe and America (Figure 1.1) was noted by the Flemish-Dutch cartographer Abraham Ortelius in 'Thesaurus Geographicus' (Ortelius, 1596). He went on to propose that the continents were "torn away...by earthquakes and floods". This remarkable insight was, however, ahead of its time and it was not until 1912 that the German meteorologist Alfred Wegener developed the hypothesis for 'continental drift'. Importantly, Wegener used observations from the geologic, climatic, fossil, and palaeo-topographic records to support the argument for an
original connection of the continents in his seminal work 'Die Entstehung der Kontinente'
('The Origin of the Continents') (Wegener, 1912).

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370

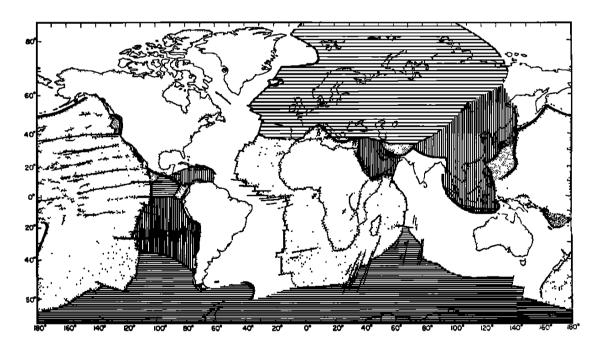
Figure 1.1. Known distribution of landmasses (thin black lines) and continental shelf edges
(thick black lines), in 1912, showing the similarity in the outlines of the east and west
Atlantic margins. Image extract from Wegener (1912).

At first, this paradigm shifting hypothesis was strongly opposed by many due to the lack of a viable mechanism by which the continents could 'plough through' solid oceanic crust (e.g. Frankel, 1990). And at the time 'Geosynclinal Theory', developed in the 1800's by James Hall and James Dwight Dana (e.g. Knopf, 1948), was able to explain the major observations from orogenic belts and sedimentary basins, as well as ocean development.

Support for continental drift, nonetheless, built throughout the 20th century, attested by the
publication of 'Our Wandering Continents: An Hypothesis of Continental Drift' by Alex du
Toit in 1937 (Du Toit, 1937). This shift was assisted by several key events that paved the
way for plate tectonics:

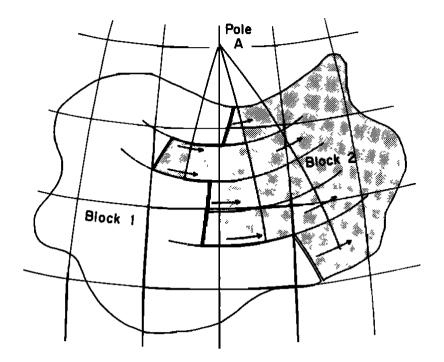
- o 1931: Arthur Holmes hypothesises the occurrence of thermal convection within the
   Earth due to the heat generated by radioactive decay, providing a driving mechanism
   for plate motions (Holmes, 1931).
- 386 o 1935: Haskell determines the fluid behaviour of the mantle from the study of glacial
  387 rebound, allowing for the drift of continents (Haskell, 1935).

- 388 o 1949: Benioff discovers Wadati-Benioff zones, suggesting a solution to space
   389 problems associated with drifting continents and shrinking oceans (Benioff, 1949).
- 390 o 1962: Harry Hess proposes 'seafloor spreading' to occur along mid-ocean rises
  391 (Hess, 1962).
- 392 o 1963: Vine and Matthews prove the occurrence of seafloor spreading using linear
   393 magnetic reversals recorded in the oceanic crust 'tape recorder' (Vine and
   394 Matthews, 1963).
- 395 o 1965: Runcorn uses polar wander to support the motion of the continents through
  396 geologic time (Runcorn, 1965), and Wilson proposes the 'transform fault', the
  397 "missing" type of plate boundary (Wilson, 1965).
- 398 o 1968: Jason Morgan coins the term 'plate tectonics', or the theory that 'the surface
  399 of the Earth is composed of several rigid plates all in relative motion about each
  400 other as a natural consequence of thermal convection in the mantle' (Figure 1.2),
  401 and describes the motion of plates on the globe about poles of rotation (Morgan,
  402 1968; Figure 1.3).
- 403



404

Figure 1.2. Configuration of the tectonic plates as initially described by Morgan (1968). The
boundaries between tectonic plates are seafloor spreading centres, ocean trenches
(subduction zones), and transform faults. Image from Morgan (1968).

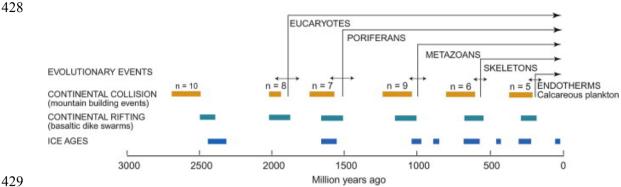


410

Figure 1.3. The motions of tectonic plates (Block 1 and Block 2) on the surface of a sphere
must follow small circles about a pole of rotation (Pole A). Image from Morgan (1968).

413 Around this time it was also discovered that not only do tectonic plates collide and rift apart 414 from each other, but that this process was cyclic. Careful analysis of the distribution of 415 faunal realms on either side of the Atlantic Ocean by Wilson (1966) showed that a 'proto-416 Atlantic' must have opened and closed prior to the opening of the present day Atlantic 417 Ocean. This cycle of rifting, continental breakup, ocean spreading, subduction initiation, 418 ocean closing, and continental collision thus became known as the Wilson Cycle.

419 An inherent periodicity to such events of mountain building, sea level change, magmatism, 420 climate, and even evolution was documented as early as the 1940's by Umbgrove in "The 421 Pulse of the Earth" (Umbgrove, 1947), however, it was not until 1984 that Worsley 422 recognised these cycles as the manifestation of supercontinent assembly and breakup 423 (Worsley et al., 1984). Regular episodes of continental collision and breakup related mafic 424 magmatism throughout Earth's history had been recognised through Rb/Sr and K/Ar dating 425 (Condie, 1976; Windley, 1977) and were shown to occur at roughly 500 m.y. intervals 426 (Figure 1.4). Subsequently, 4 or 5 supercontinent amalgamations have been proposed prior 427 to the most recent assembly of Pangea.



430 Figure 1.4. Periodic events linked to the Wilson Cycle throughout Earth's history (Worsley et al., 1984). Image from Nance and Murphy (2013). 431

432 The driving force behind the Wilson (supercontinent) Cycle is, however, still a matter of 433 debate today, and, whilst mantle convection is still generally recognised as an important 434 factor in controlling plate motions (e.g. Ziegler, 1993), other drivers have been proposed to 435 also play important roles. Common suggestions include ridge-push and slab-pull (e.g. 436 Lithgow-Bertelloni, 2014), gravitational collapse along orogenic belts (e.g. Rey et al., 2001), and mantle plumes (a special case of mantle convection; e.g Larson, 1991). The coincidence 437 438 of many continental breakups with large igneous provinces has been used to support the role 439 of hot mantle plumes in this phase of the Wilson Cycle (e.g. White and McKenzie, 1989). 440 Examples of continental breakup lacking evidence for abnormal mantle temperatures during 441 breakup, however, preclude the ubiquitous involvement of mantle plumes in this process 442 (e.g. Storey, 1995). A distinction has therefore arisen between 'active' and 'passive' rifting 443 to explain the generic observations from rifted margins: uplift, magmatism, and extension. 444 During active rifting, uplift and melting are attributed to the impingement of a mantle plume or convection cell on the base of the lithosphere resulting in extension. During passive 445 446 rifting, extension is attributed to an alternative driving force and results in the uplift and 447 melting observed during rifting (e.g. Turcotte and Emerman, 1983).

#### **1.3. Passive continental margins** 448

449 Both active and passive rifting may result in a complete thinning of the continental 450 lithosphere and the onset of seafloor spreading. This transition is termed continental 451 breakup, and the zone of recently rifted continental crust becomes a passive margin. These features form 105,000 km of the world's continental margins, almost twice the length of 452 convergent boundaries (53,000 km; Bradley, 2008), where one tectonic plate subducts 453 beneath another. The term 'passive margin' may, however, be misleading as these types of 454 margins are more and more commonly being found to undergo significant tectonic inversion, 455

456 resulting from compression, and uplift, possibly due to mantle dynamics, during their 457 'passive' stage (e.g. Japsen et al., 2012; Yamato et al., 2013). The causes of continental 458 rifting and the subsequent reactivation of passive margins are still a matter of debate and 459 form a key area of research today (e.g. Geoffroy et al., 2015; Rabineau et al., 2015)

460 Depending on the stress conditions under which they form, two endmember groups of 461 passive margins can develop. Where the trend of a deforming lithospheric zone is closely 462 orthogonal to the extension direction across it, divergent (rifted) margins are formed. When 463 the extension direction is closely parallel to the overall trend of a lithospheric deformation 464 zone, transform margins develop. Between these two endmembers, oblique margins may 465 develop and strain may be accommodated along transtensional normal faults, or may 466 become partitioned onto alternating divergent and transform segments (Basile and Braun, 467 2016).

#### 468 <u>1.3.1. Divergent Margins</u>

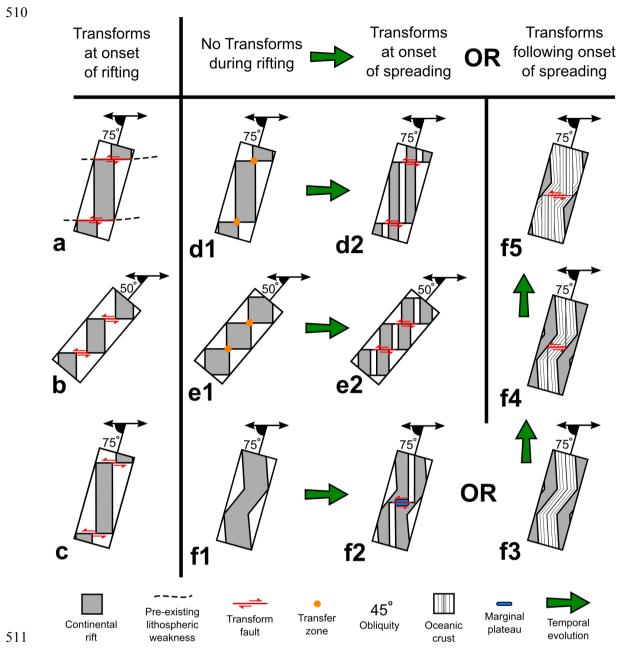
469 Divergent margins are commonly categorised by two end-member extremes. These are 470 volcanic rifted margins and magma-poor rifted margins, and are commonly distinguished by 471 the volume of syn-rift igneous rocks emplaced. Following breakup, passively upwelling 472 mantle replaces material that moves laterally away from an ocean spreading centre, and 473 consistently results in decompression melting, which produces an  $\sim$ 7 km thick crust (e.g. 474 Katz et al., 2003; White et al., 1992). Generally, an elevated mantle temperature, due to 475 continental insulation (e.g. Brandl et al., 2013) or mantle plumes (e.g. White and McKenzie, 1989), is invoked to explain the greater melt volumes observed at some volcanic rifted 476 477 margins. Several alternative hypotheses have, however, also been proposed including; rifting 478 speed (van Wijk et al., 2001), small-scale convection (SSC; Boutilier and Keen, 1999; 479 Mutter et al., 1988; Simon et al., 2009), chemical enrichment of the mantle (Lizarralde et al., 480 2007), and rifting history (Armitage et al., 2010). The development of magma-poor rifted 481 margins may therefore be explained as the result of rifting away from mantle thermal 482 anomalies (e.g. Reston, 2009), although depth-dependent stretching of the lithosphere (e.g. 483 Huismans and Beaumont, 2011) and presence of a partially depleted sub-lithospheric mantle 484 (e.g. Pérez-Gussinyé et al., 2006) have also been proposed to control their development. 485 Significant variations in magmatism and rifting style over different length scales (e.g. Behn 486 and Lin, 2000; Franke et al., 2007; Lizarralde et al., 2007) precludes one single causal 487 mechanism, and demonstrates our incomplete understanding of the development of rifted 488 margins. In reality, a spectrum between the two endmember rifted margin types likely exist 489 (Franke, 2013) as the natural result of interplay between different mechanisms.

#### 490 <u>1.3.2. Transform and oblique Margins</u>

491 In cases of continental rifting where the extension direction is not perpendicular to a 492 lithospheric deformation zone, oblique and/or transform margins develop. Following 493 breakup, mature ocean spreading centres predominantly accommodate extension in a fully 494 partitioned manner, i.e. along extension-perpendicular rifts connected by extension-parallel 495 transform faults. Exceptions are limited to slow and very slow spreading centres, e.g. the 496 Mohns (Dauteuil and Brun, 1996) and Reykjanes (Peyve, 2009) ridges, where oblique 497 oceanic accretion occurs. Strain partitioning thus occurs at the majority of plate boundaries 498 that developed in an oblique setting. The timing of this partitioning (i.e. transform fault 499 development) is, however, variable and several possible scenarios have been described by 500 Basile (2015) and are summarised below.

#### 501 **1.3.2.1. Transform development at the onset of rifting or shortly after**

502 Where extension is closely parallel to pre-existing lithospheric weaknesses, such weaknesses may be reactivated as transform faults at the onset of rifting (Figure 1.5a). This has been 503 504 proposed for several instances where oceanic transform faults are aligned with an onshore 505 tectonic structure such as in the South Atlantic (e.g. Wright, 1976) and Gulf of Aden 506 (Bellahsen et al., 2013). Alternatively, where a deformation zone is at low angles to the 507 extension direction (Figure 1.5b), or where narrow rifts predominate (Figure 1.5c), individual rift segments do not overlap and may be connected by transform faults at the 508 509 onset of rifting or shortly thereafter.

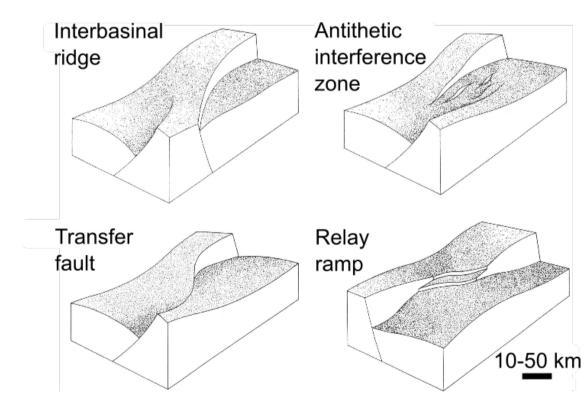


512 Figure 1.5. Development of transform and oblique margins. Left column: transform faults 513 may be present at the onset of rifting along pre-existing lithospheric weaknesses (a) or 514 between non-overlapping rift segments, which may result from a high obliquity (b) or 515 narrow rift segments (c). Other columns: transform faults may not develop during rifting if 516 rift segments overlap, which may occur in low obliquity settings (d1) or where wide rifts 517 predominate (e1), or if transtensional rifting occurs (f1). In these cases, strain partitioning 518 generally occurs at the onset of oceanic spreading (d2, e2, and f2) and new transform faults 519 cut older rift structures in the continental crust. In the case of transtensional rifting, 520 partitioning may also occur after a phase of oblique spreading (f3-5), in which case transforms may not develop in the continental crust and no transform margin will result. 521 522 Figure after Basile (2015).

#### 523 1.3.2.2. Transform development at the onset of spreading or later

524 Where a deformation zone lies at high angles to the extension direction (Figure 1.5d1), or 525 where wide rifts develop in more oblique settings (Figure 1.5e1), individual rift segments may overlap. In this setting, transfer fault zones (e.g. Basile, 2015; Milani and Davison, 526 527 1988) separate rift segments and may accommodate the offset without the need for transform 528 faults. Common structures forming interbasin transfer zones include interbasinal ridges, 529 antithetic interference zones, transfer faults, and relay ramps (e.g. Gawthorpe and Hurst, 530 1993; Figure 1.6). Transtensional rifting is also able to accommodate oblique extension 531 without the need for transform faults (Figure 1.5f1). In these cases, strain partitioning onto 532 transforms may occur at the onset of oceanic spreading (Figure 1.5d2-f2; e.g Taylor et al., 2009), in which case transform faults cut older rift structures and may result in the 533 534 development of marginal plateaus (Mercier de Lépinay et al., 2016). Alternatively, for the 535 case of transtensional rifting, partitioning may also occur after a phase of oblique spreading. 536 In this instance transform faults may not be present in the continental domain of an oblique 537 rifted margin (Figure 1.5f3-5).

538 539



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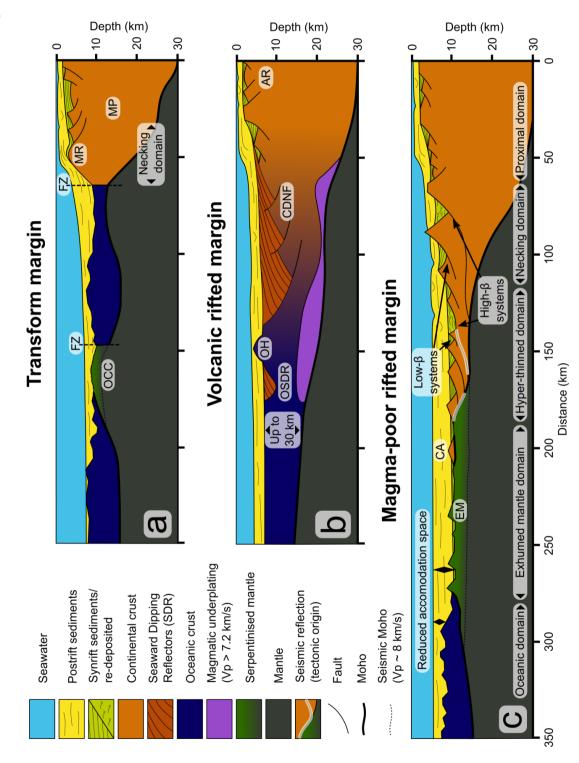
Figure 1.6. Interbasinal transfer zone geometries that may link offset, but overlapping, rift
systems without the need for transform faults. The vertical scale is greatly exaggerated.

543 Image from Gawthorpe and Hurst (1993).

#### 544 <u>1.3.3. Characterisation of passive margins</u>

Each margin type is characterised by features that develop under the different rifting regimes. Identification of these features, through the use of geophysical methods, can allow for a robust differentiation between divergent and transform type margins. The main distinguishing features of the different categories of passive margin are summarised below and are shown in Figure 1.7.

550



554 Figure 1.7. Key characteristics of each margin type as described in Section 1.2. (a) 555 Transform margins commonly exhibit: narrow necking domains of the Moho and basement, marginal plateaus, marginal ridges, adjacent thin oceanic crust and exhumed mantle at 556 oceanic core complex. (b) Volcanic rifted margins often show: continentward dipping 557 558 normal faults, adjacent abandoned rifts, SDRs, underplated high velocity lower crust, 559 initially thick ocean crust. (c) Magma-poor rifted margins generally display: up to 200 km of 560 thinned crust, both low- $\beta$  and high- $\beta$  fault systems, high amplitude seismic reflections along 561 low angle faults, exhumed and serpentinised mantle, initially thin oceanic crust. AR: 562 abandoned rift; CA: continental allochthon; CDNF: continentward dipping normal faults; 563 EM: exhumed mantle; FZ: fracture zone; MP: marginal plateau; MR: marginal ridge; OCC: ocean core complex; OH: outer high; OSDR: outer seaward dipping reflectors. 564

#### 565 **1.3.3.1. Volcanic rifted margins**

566 Where large volumes of melt are emplaced during rifting, volcanic rifted margins develop 567 (Figure 1.7b), and may indicate the breakup of the lithospheric mantle before or at the same 568 time as that of the crust (Franke, 2013). This magmatism manifests itself as thick wedges of volcanic flows, identified in seismic reflection data as seaward dipping reflectors (SDRs), 569 570 and magmatic underplating, identified in seismic refraction studies as high velocity (>7.2571 km/s) lower crustal bodies (Franke, 2013). As these margins develop and become 572 submerged, the introduction of water may results in explosive volcanism resulting in the 573 formation of an outer high, beyond which outer SDRs may develop as deep sea volcanic 574 sheet flows before the onset of normal oceanic crust production (Planke et al., 2000). The 575 enhanced melt production during breakup often also results in an initially over-thickened 576 oceanic crust up to 30 km thick (Geoffroy, 2005).

577 Crustal thinning at volcanic rifted margins often occurs adjacent to previous sedimentary 578 basins and is quite abrupt, occurring over 50-100 km (Figure 1.7b). Continentward dipping 579 normal faults may also accommodate the extension, although the large volumes of extrusive 580 volcanic rocks at these margins often masks rift structures making accurate identifications of 581 the last continental crust and the COT difficult (Stab et al., 2016). The considerable amounts 582 of intrusive magmatism and magmatic underplating also alter Moho geometries, which 583 therefore provide little insight into the structure of the continental crust as a result of rifting.

#### 584 **1.3.3.2. Magma-poor rifted margins**

585 Where exhumation of the mantle occurs and limited melt is produced during rifting, magma-586 poor passive margins develop (Figure 1.7c), indicating a complete thinning of the crust 587 before the breakup of the lithospheric mantle (Franke, 2013). These margins typically 588 display five domains. Starting from the continent and moving oceanwards, these are: 1) 589 proximal domain, where crustal thinning and accommodation space are minimal; 2) necking 590 domain, where the crust is thinned most intensively resulting in the steepest Moho slopes; 3) 591 hyper-thinned domain, where thinned continental crust (sometimes less than 10 km thick) 592 forms an extensive continental rise; 4) exhumed mantle domain, where complete thinning of 593 the crust occurs and lithospheric mantle is exhumed along detachment faults and becomes 594 partially serpentinised; and 5) oceanic domain, where following sufficient thinning of the 595 lithosphere, mantle melting occurs to produce oceanic crust (Reston, 2009; Tugend et al., 596 2015).

597 Crustal thinning at magma-poor rifted margins occurs through both low- $\beta$  and high- $\beta$ 598 extensional systems, where  $\beta$  is the crustal extension factor. Low- $\beta$  systems accommodate 599 small amounts of extension through high-angle normal faults, forming classic half-graben 600 structures with wedge-shaped sedimentary fills (Figure 1.7c). This type of extensional 601 system can occur all along the margin (Tugend et al., 2015), and most magma-poor rifted 602 margins display well defined fault blocks of this type (Reston, 2009). High- $\beta$  systems allow 603 large amounts of extension to be accommodated through long-offset normal faults. These 604 detachment faults exhume the underlying footwall over large areas forming smooth fault-605 related topography and hyperextended sag basins. These basins often contain syn-rift 606 sediments that gently onlap the low-angle basement, or that have been redeposited due to 607 significant fault block rotation (Wilson et al., 2001), losing their initial syn-rift sedimentary 608 wedge configuration. Where the low angle faults of high- $\beta$  systems are intracrustal, or form 609 the boundary between crust and mantle, they are commonly imaged as high amplitude 610 seismic reflections (e.g. Davy et al., 2016; Dean et al., 2008). High-B systems tend to develop in the necking, hyperthinned, and exhumed mantle domains alongside low- $\beta$ 611 612 systems (Tugend et al., 2015). These extensional systems lead to complete crustal thinning, 613 from a thickness of ~30 km, over a long distance of 100-200 km (Figure 1.7c), with most of 614 the crustal thinning (from ~20-10 km) occurring in the necking zone (Reston, 2009). This 615 region of sharp crustal thinning generally has average Moho slopes of between 5° - 25°, although average Moho slopes across the necking zone have reached 33 and 36 degrees 616 617 across the Porcupine Basin (O'Reilly et al., 2006) and Newfoundland margin, respectively. 618 In the latter example, however, the steepness is thought to have been influenced by lower 619 crustal gabbroic lenses localising deformation (Van Avendonk et al., 2009).

Following complete crustal thinning, partially serpentinised peridotites of the exhumed mantle domain form the OCT (Tugend et al., 2015). In these domains no clear Moho reflection is present due to the gradual downward decrease in the amount of serpentinisation (Gillard et al., 2015). The onset of magmatism and production of oceanic crust marks the beginning of the oceanic domain. Oceanic crust is often initially thin (e.g. (Davy et al., 2016)) but its presence is often marked by (sometimes weak) Moho reflections. The
boundary between the oceanic and exhumed mantle domains may correspond to a step in the
basement topography and smaller accommodation space within the oceanic domain (Figure
1.7c) due to the relative buoyancy of the oceanic crust over mantle (Tugend et al., 2015).

#### 629 <u>1.3.4. Characterisation of Transform Margins</u>

630 When oblique extension is partitioned onto transform and divergent margins early enough, 631 transform faults cut the continental crust and lithosphere, facilitating plate separation and 632 forming transform margins (Figure 1.7a; Basile, 2015). As the transform faults of the 633 lithospheric deformation zone at transform margins are sub-vertical, extremely narrow 634 'necking zones', typically 50 km wide, develop (Mercier de Lépinay et al., 2016). This sharp 635 zone of crustal thinning is accompanied by the development of steep average Moho slopes, 636 commonly between 8° - 40°, although sub-vertical Moho slopes have been observed at the 637 Côte d'Ivoire-Ghana transform margin (Sage et al., 2000). At the oceanward edge of the 638 necking zone, a steep basement slope may also be present, the base of which often coincides 639 with the COT. The Falkland escarpment basement slope lies at 13 degrees (Lorenzo and 640 Wessel, 1997), Ghana at 12 degrees (Sage et al., 2000), the Northern Demerara and French 641 Guiana margins show slopes of 24 and 13 degrees respectively (Greenroyd et al., 2008), and 642 the Newfoundland Fracture Zone basement slope forms a slope of  $\sim 40$  degrees (Keen et al., 643 1990). Rift basins are generally not seen within this slope region, although any present may 644 not have been clearly imaged using seismic reflection techniques due to the slope steepness.

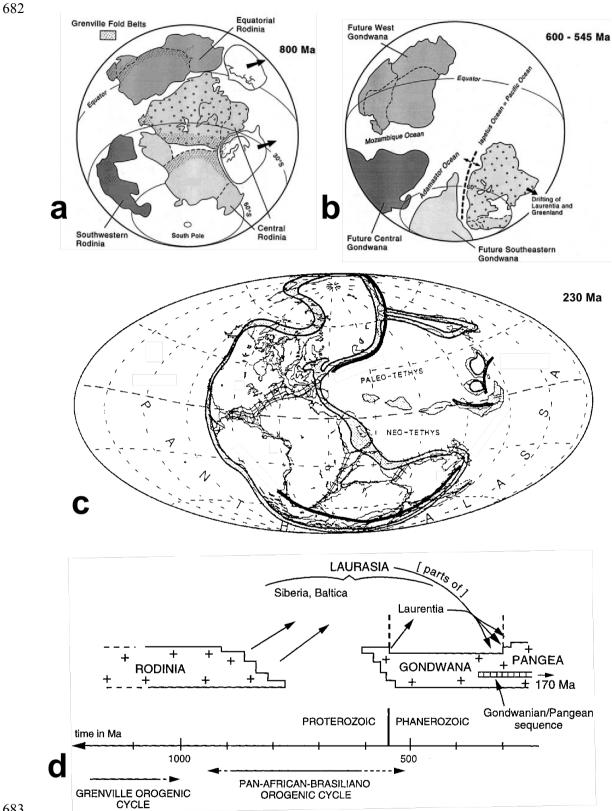
645 The presence of thick continental lithosphere adjacent to MORs at transform margins results 646 in a high thermal contrast between the continental and oceanic lithospheres. Thin oceanic crust may therefore be formed adjacent to transform margins due to heat conduction away 647 648 from the spreading centre, resulting in reduced melting (e.g., Ghana: 4 km (White et al., 649 1992); Agulas-Falkland FZ: 4 km (Becker et al., 2012)). This process is analogous to those 650 occurring at long offset transform faults in the oceanic domain. In these settings, the reduced 651 melt supply results in tectonic accommodation of extension and lower crust or mantle rocks 652 may be exhumed to the surface at oceanic core complexes up to 50 km wide (e.g. Karson, 1999). The high thermal contrast across transform margins also results in differential 653 654 subsidence, and where the oceanic and continental plates are coupled flexural downwarping 655 of the continental side may occur (Lorenzo and Wessel, 1997; Mercier de Lépinay et al., 656 2016).

Marginal plateaus, possibly inherited from continental thinning prior to the partitioning of
oblique strain onto transform faults (Basile, 2015)(Figures 1.5f1-2), are also commonly
observed adjacent to transform margins (Mercier de Lépinay et al., 2016). A marginal ridge

(e.g. Bird, 2001), where crust or pre/syn-rift sediments are elevated linearly along the margin
near the top of the continental slope, may also be present (Figure 1.7a). However, the
mechanisms for forming marginal ridges are still under debate and possibly diverse in
origin; for a full review the reader is referred to (Basile, 2015).

### 664 **1.4. Gondwana assembly and breakup**

665 The paleomagnetic, stratigraphic, and tectonic records of the continental crust span Earth's 666 almost entire history, as they are not systematically destroyed like those of the oceanic crust, 667 which is subducted comparatively soon after its formation (e.g. Roberts et al., 2015). They 668 therefore make a useful tool for reconstructing Wilson Cycles that occurred before the 669 formation of the Earth's present day oceanic crust. The formation of the supercontinent 670 Gondwana, a long lived accumulation of much of the earths continental lithosphere into one 671 landmass, represents such a cycle. Gondwana was reconstructed from the lithospheric megaplates formed during the disintegration of the previous supercontinent Rodinia between 1000 672 673 and 700 Ma (Veevers, 2004). These plates were reunited during the Pan-African orogeny 674 between 650 and 500 Ma, which brought together the present day landmasses of Africa, 675 Antarctica, Australia, India, the Middle East, North America, South America, and also 676 briefly North America (Laurentia). The return of Laurentia with Baltica and Siberia, together 677 forming Laurasia, to Gondwana at ~320 Ma resulted in the Variscan orogeny along the 678 north-eastern margin of Gondwana, and formation of the supercontinent Pangea. Gondwana 679 remained intact within Pangea and was surrounded by the near global Panthalassas Ocean, 680 which by ~300 Ma had renewed subduction beneath the southern margin of Gondwana 681 (Veevers, 2004).



684 Figure 1.8. Disassembly of Rhodinia (a) and reassembly into Gondwana (b). Later merging 685 of Gondwana with Laurasia forms Pangea, surrounded by the Panthalassas Ocean (c and d). Images a, b, and d are modified from Trompette (2000), image c is modified from Veevers 686 (2004). 687

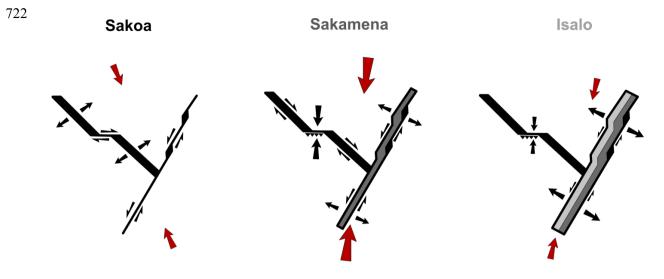
688 This subduction and accretion led to the development of foreland basins and the Cape Fold 689 Belt in South Africa between 220 and 290 Ma (Frimmel et al., 2001) with fluctuating 690 deformation intensity. This compression became more widespread, also affecting South 691 America to the west, at ~258 Ma (Veevers, 2004). The episodic and changeable 692 development of the Karoo rift system (Hankel, 1994; Schandelmeier et al., 2004) across 693 Gondwana at this time (Reeves et al., 2016) may therefore be the result of this compression 694 (Delvaux, 2001a), which varied temporally in orientation and intensity, reactivating pre-695 existing basement weaknesses along the northern parts of the Karoo rift system (Reeves, 696 2014). Studies of Karoo rifts, which predominantly follow a NW-SE trend (Delvaux, 2001b) 697 or a NE-SW trend (Schandelmeier et al., 2004) in the present day East Africa region, show 698 the general progression:

Sinistral strike-slip along NE-SW trending basins with local deposition in left lateral
step-over basins. Extension along NW-SE trending basins with strike-slip deformation
along E-W trending offsets in the rift system. This phase of rifting led to the deposition
of Sakoa aged deposits between ~265 Ma and ~300 Ma (Figure 1.9), which are capped
by the Vohitolia marine limestones (Hankel, 1994) suggesting a connection to the
NeoTethys Ocean in the north at this time.

After a brief pause in rifting, a reconfiguration occurred between ~259 and ~264 Ma
(Hankel, 1994), coincident with the broadening and intensification of compression
along southern Gondwana (Veevers, 2004). The onset of rapid subsidence in NE-SW
trending basins (Schandelmeier et al., 2004), along with strike slip deformation along
formerly extensional NW-SE trending basins and compression along their E-W trending
offsets, resulted in deposition of Sakamena aged deposits (Figure 1.9).

3) A second larger pause in rifting occurred around ~249 Ma, and was followed by a
rejuvenation of rifting along NE-SW trending rifts at ~242 Ma with little activity along
NW-SE trending rifts, which led to the development of Isalo aged deposits until ~210
Ma (Figure 1.9; Hankel, 1994).

Following the cessation of compression along the cape fold belt, a substantial period of quiescence along the Karoo rifts ensued (Geiger et al., 2004). Rejuvenation of rifting did not occur until the Early Jurassic (Toarcian), coincident with eruption of the Karoo large igneous province in Mozambique. This Early Jurassic rifting at ~182 Ma culminated in the breakup of East and West Gondwana, which had thus remained intact for over 520 Myr. The kinematics and dynamics during this drifting apart of East and West Gondwana form the main focus of this thesis.





#### 220-290 Ma Cape Fold Belt metamorphism

Figure 1.9. Schematic of Karoo rift development in East Africa, possibly in response to varying intensities and orientations of compression along the Cape Fold Belt. The distributions Sakoa aged deposits are shown in black, Sakamena deposits in dark grey, and Isalo deposits in light grey.

### 728 **1.5. Thesis outline**

Chapters 2 to 4 of this thesis are the results chapters of this thesis. They are presented as three 'journal-style' academic papers or manuscripts, and as such are designed to be standalone. Nonetheless, they contain a through-going theme and are further tied together in the conclusions chapter at the end of this thesis.

Each results chapter has either been published in, or is intended for submission to, a relevant
academic journal within the field of Earth Sciences. These papers are multi-authored and as
such my contributions to each chapter are outlined below:

Chapter 2 - 'Madagascar's escape from Africa: A high-resolution plate reconstruction for the Western Somali Basin and implications for supercontinent dispersal: this chapter has been published in *Geochemistry, Geophysics, Geosystems* in 2016. As first author of this paper my contributions included: concept design, gravity modelling, plate tectonic modelling, and manuscript preparation. Co-authors have contributed training, useful discussion, and editorial assistance.

742 Chapter 3 - 'The Rovuma Transform Margin: Pinning down the East African continent-743 ocean transform margin using seismic and gravity methods': this chapter is intended for 744 submission to *Tectonics*. As first author of this paper my contributions included: Moho and 745 bathymetry data digitisation, seismic interpretation, Matlab code development, gravity 746 modelling, manuscript preparation. Concept design was joint between myself and co-747 authors. Co-authors have contributed training, useful discussion, and editorial assistance.

748 Chapter 4 – 'Compressional consequences of complex spreading: Formation of the Tanzania 749 Coastal Basin and Davie Fracture Zone during the Mesozoic East Africa breakup': this 750 chapter is also intended for submission to Tectonics. As first author of this paper my 751 contributions included: analysis of satellite image data, seismic interpretation, regional tectonic model development, and manuscript preparation. Co-authors have contributed 752 753 training, useful discussion, and editorial assistance.

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#### Madagascar's escape from 2. 1018 high-resolution plate Africa: A 1019 reconstruction for the Western 1020 Somali Basin and implications for 1021 supercontinent dispersal 1022

## 1023 Abstract

1024 Accurate reconstructions of the dispersal of supercontinent blocks are essential for testing 1025 continental breakup models. Here, we provide a new plate tectonic reconstruction of the 1026 opening of the Western Somali Basin during the breakup of East and West Gondwana. The 1027 model is constrained by a new comprehensive set of spreading lineaments, detected in this 1028 heavily sedimented basin using a novel technique based on directional derivatives of free-air 1029 gravity anomalies. Vertical gravity gradient and free-air gravity anomaly maps also enable 1030 the detection of extinct mid-ocean ridge segments which can be directly compared to several 1031 previous ocean magnetic anomaly interpretations of the Western Somali Basin. The best-1032 matching interpretations have basin symmetry around the M0 anomaly; these are then used 1033 to temporally constrain our plate tectonic reconstruction. The reconstruction supports a tight 1034 fit for Gondwana fragments prior to breakup, and predicts that the continent-ocean transform 1035 margin lies along the Rovuma Basin, not along the Davie Fracture Zone (DFZ) as commonly 1036 thought. According to our reconstruction, the DFZ represents a major ocean-ocean fracture 1037 zone formed by the coalescence of several smaller fracture zones during evolving plate 1038 motions as Madagascar drifted southwards, and offshore Tanzania is an obliquely rifted, 1039 rather than transform, margin. New seismic reflection evidence for oceanic crust inboard of 1040 the DFZ strongly supports these conclusions. Our results provide important new constraints 1041 on the still enigmatic driving mechanism of continental rifting, the nature of the lithosphere 1042 in the Western Somali Basin, and its resource potential.

#### 1044 **2.1. Introduction**

1045 Continental breakup is a fundamental, but poorly understood, part of the plate tectonic cycle. 1046 Our understanding of the conditions needed for successful rift formation is particularly 1047 limited. Besides pre-existing weak zones (Ziegler and Cloetingh, 2004) and thermal 1048 weakening due to rifting-related magmatism (Buck, 2007), it has also recently been shown 1049 that oblique rifting is an important mechanism to facilitate the breakup (Brune et al., 2012). 1050 To further investigate these concepts, accurate reconstructions of rifting events with high 1051 spatial resolution are essential to enable detailed comparisons between models and 1052 observations (Nance and Murphy, 2013). Detailed history of rifted margin evolution is also 1053 key in hydrocarbon exploration by enabling the prediction of the petroleum potential for 1054 conjugate basins with similar tectonostratigraphic histories (Beglinger et al., 2012). Some of 1055 the most significant rifting episodes in Earth's history occurred during supercontinent 1056 breakup, e.g. rifting between East and West Gondwana, which spanned many of our present 1057 day continents.

1058 Gondwana was assembled between 600 and 500 Ma in the Pan-African orogeny (e.g. 1059 Trompette, 2000; Van Hinsbergen et al., 2011). Beginning in the Jurassic the subsequent 1060 breakup of East and West Gondwana carried Madagascar approximately southwards, as 1061 shown by ocean magnetic anomalies (Segoufin and Patriat, 1980; Rabinowitz et al., 1983; 1062 Cochran, 1988; Eagles and König, 2008; Davis et al., 2016), forming the Western Somali 1063 Basin (WSB) (Coffin and Rabinowitz 1987; Geiger et al., 2004). Knowledge of 1064 Madagascar's former position within Gondwana, and the path it followed during its 1065 southward drift, is crucial for creating accurate plate tectonic reconstructions of Gondwana's 1066 dispersal.

1067 Paleogeographic reconstructions of Madagascar's position in Africa (Figure 2.1) have a large 1068 range of fits, suggesting significantly different locations for the continent-ocean transition. 1069 This is primarily due to the lack of fracture zone expressions in bathymetry data, where the 1070 characteristic fracture zone topography is commonly obscured by over 5 km of sediment 1071 (Coffin et al., 1986). The Davie Fracture Zone (DFZ), commonly assumed to form the 1072 western transform fault (Coffin et al., 1986) or continental ocean transform margin (e.g. 1073 Gaina et al., 2013) of the WSB, is one of the few fracture zones confidently identified. 1074 However, the DFZ is overlapped by several independently generated reconstructions (e.g. 1075 Smith and Hallam, 1970; Lottes and Rowley, 1990; Reeves, 2014). This puts our 1076 understanding of this feature into doubt, and highlights the need for the comprehensive 1077 detection of fracture zones to support plate tectonic reconstructions of the basin.

1078 In this paper we present a detailed and self-consistent plate tectonic reconstruction of the 1079 WSB that can be used to further our understanding of the dynamics of continental breakup. 1080 We use a novel combination of free-air gravity, vertical gravity gradient and filtered free-air 1081 gravity directional derivatives to determine the location of the extinct mid-ocean ridge 1082 (MOR) segments and a comprehensive set of gravity lineaments related to fracture zones in 1083 the WSB. Using global gravity datasets with 1 arc-minute resolution captures the complexity 1084 of the breakup geometry and motion, beyond what can be seen in widely spaced shipboard 1085 magnetic profiles, and also constrains the history of the basin to the west of 43°E where no 1086 magnetic anomaly identifications are available (e.g. Davis et al., 2016). These spreading-1087 related features are tested against existing magnetic and seismic reflection data before being 1088 used to produce a high-resolution plate tectonic reconstruction.

The model provides new insight into Madagascar's position in Africa prior to Gondwana breakup, the nature of the DFZ, and the geometry and structure of the East African continental margin. This provides a significant advance towards a more comprehensive understanding of the nature of the margins and underlying lithosphere of the WSB, as well as a broader understanding of rifting events and continental breakup mechanisms.

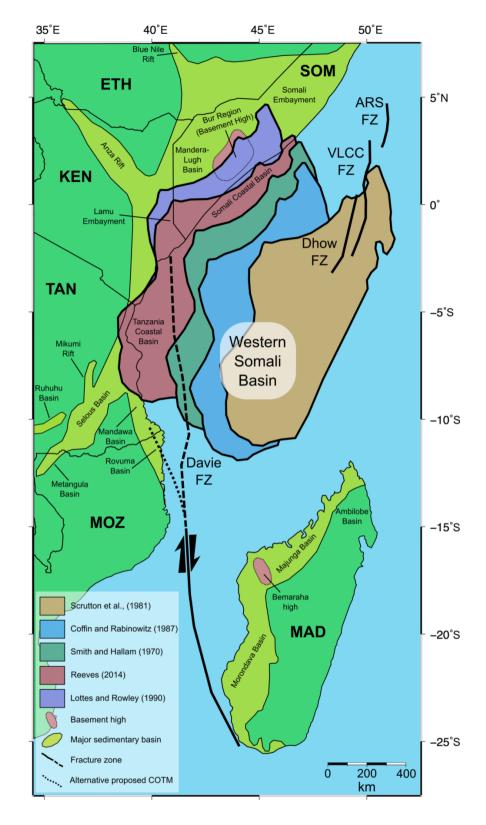




Figure 2.1. Published reconstructions for Madagascar's paleo-position in Africa are shown
alongside major continental basins, basement highs, and previously identified oceanic
fracture zones. ARS - Auxiliary Rescue and Salvage; VLCC - Very Large Crude Carrier;
ETH - Ethiopia; KEN - Kenya; MAD - Madagascar; MOZ - Mozambique; SOM - Somalia;
TAN - Tanzania; COTM – Continent-ocean transform margin.

## 1102 **2.2. Data and processing**

In this study, we used the distribution and orientation of fracture zones (FZs) and extinct
MORs to construct a plate tectonic model for the WSB. These tectonic features were
detected using a combination of gravity, magnetic and seismic data.

#### 1106 <u>2.2.1. Gravity data</u>

1107 We used version 23 of the Sandwell and Smith gravity model (Sandwell et al., 2014), which 1108 now has an accuracy of  $\sim 2$  mGal (compared to 3-5 mGal in previous versions) following the 1109 addition of retracked CryoSat-2 and Jason-1 satellite altimetry data. This improvement 1110 allows detection of many buried structures, particularly short wavelength features such as 1111 extinct MORs and FZs (e.g. Sandwell et al., 2014). In addition to the widely-used free-air 1112 gravity anomaly, gradients such as the vertical gravity gradient (VGG) amplify short 1113 wavelength gravity signals, aiding the detection of relatively small features such as 1114 seamounts and ocean spreading structures (e.g. Kim and Wessel, 2011).

1115 FZs are the inactive extensions of transform faults along the MOR, which run parallel to the 1116 spreading direction during their formation and therefore record plate motions. Oceanic plate 1117 ages, and thus seafloor depths, are offset across transform faults, causing a distinct 1118 bathymetric feature, which is permanently locked in by welding at ridge-transform 1119 intersections (Sandwell, 1984). Variations in crustal thickness due to melt supply (Blackman 1120 and Forsyth, 1991; Gente et al., 1995), transpression or transtension (Menard and Atwater, 1121 1969), and thermal contraction (Collette, 1974) can also result in bathymetric expressions 1122 along fracture zones. The resulting linear ridges and troughs can be traced in bathymetric 1123 data in order to track the spreading history, which in turn may be used in constructing a plate 1124 tectonic model.

1125 In heavily sedimented regions, however, spreading features may be completely buried. 1126 Fortunately, expressions of the spreading features are also preserved in gravity data, even 1127 when buried by sediments, due to lateral density contrasts between sediment, crust, and 1128 mantle across the structures. Gravity data therefore allows the derivation of paleo-spreading 1129 directions in heavily sedimented regions of the ocean (Sandwell et al., 2014).

#### 1130 **2.2.1.1. Bandpass filtering and gravity gradients**

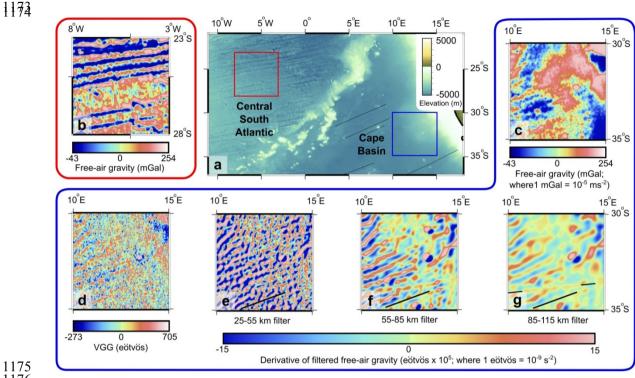
We used a combination of directional derivatives and bandpass filtering to further enhance the portion of the gravity field associated with FZs. First, a Gaussian bandpass filter was used to remove short wavelength noise and shallow features within the sediment layer, together with long wavelength signals from deep mantle heterogeneities. Mulder and Collette (1984) found that most FZs produce anomalies with intermediate wavelengths between 50 and 200 km. We further refined these bounds empirically, looking for sharp, continuous linear anomalies parallel to the overall spreading direction derived from ocean magnetic anomalies, and found that wavelengths of 55-85 km best highlighted fracture zone structure.

1140 After filtering, we exploited the linear nature of spreading features by taking directional 1141 gradients of the free-air gravity to emphasise lineations of a given orientation (e.g. Mitchell 1142 and Park, 2014). This procedure has similar advantages to illuminating a topographic map to 1143 highlight fault scarps (e.g. Arrowsmith and Zielke, 2009). In the case of the WSB, the ocean 1144 magnetic anomalies indicate an approximately N-S spreading direction. Therefore gradients 1145 taken along this strike would highlight spreading-perpendicular features, such as MOR 1146 segments, and those taken with an E-W strike would highlight spreading-parallel features, 1147 such as FZs. To account for local variability and changes in spreading direction, we took the 1148 directional gradient at 10° intervals between 30° clockwise and anti-clockwise of the chosen 1149 azimuth. This allowed near-perpendicular gradients to be sampled at all points along curved 1150 lineations, ensuring an unbiased sampling of the greatest gradient magnitudes. Examples 1151 showing the effect of changing the orientation and range of gradient sampling are provided 1152 in the supplementary material.

## 1153 **2.2.1.2. Testing the detection method**

1154 To check our methodology's usefulness in detecting spreading-related structures in buried 1155 oceanic crust, we tested it on a region of the Cape Basin, offshore South Africa (Figure 1156 2.2a). Like the WSB this basin has a thick sedimentary cover (2-5+ km), resulting in 1157 enigmatic spreading features that are difficult to interpret from free-air gravity and VGG 1158 alone. The overall spreading rate is also similar to the WSB (e.g. Eagles, 2007) likely 1159 resulting in similar FZ morphology; making this a very good natural laboratory. The free-air 1160 gravity anomaly of the test region within the Cape Basin was compared to that of an 1161 unsedimented example location from the Central Atlantic (Figures 2.2b-c), demonstrating 1162 the masking effect of sedimentation on the spreading features: clear lineations are visible in 1163 the Central Atlantic (Figure 2.2b), but lineations are only poorly distinguishable in the Cape 1164 Basin (Figure 2.2c). In this case, the VGG does little to enhance the spreading lineations 1165 (Figure 2.2d). Derivatives of the gravity anomaly, perpendicular to the spreading direction 1166 for the Cape Basin, were taken after filtering to retain different wavelengths (Figures 2.2e-g) 1167 to check which wavelengths best enhance spreading features. Shorter wavelengths (25-55 1168 km, Figure 2.2e) are noisy and reduce the continuity of spreading parallel anomalies, whilst 1169 longer wavelengths (85-115 km, Figure 2.2g) reduce the sharpness of individual lineations 1170 and can also merge anomalies into false lineations (example indicated by black arrows).

- 1171 Intermediate wavelengths of 55-85 km give the best balance between noise reduction and
- 1172 imaging of sharp FZ-related anomalies.



1175 1176

1177 Figure 2.2. (a) Test location for the processing technique within the heavily sedimented 1178 Cape Basin, and an unsedimented example location in the Central South Atlantic. Major 1179 fracture zones of the Cape basin are marked as thin black lines. (b) Free-air gravity example 1180 of non-sedimented Central Atlantic spreading features. (c) Free-air gravity from the heavily 1181 sedimented test location. (d) VGG from the test location. (e-g) Spreading parallel derivatives 1182 of gravity after filtering to retain specified wavelength bands. Black lines indicate spreading 1183 direction as constrained by major FZs from elsewhere in the Cape Basin; arrows in g 1184 indicate merged anomaly lineations.

#### 1185 2.2.2. Magnetics

1186 We used two primary magnetic data types to inform our final model: the Earth Magnetic Anomaly Grid (EMAG2 - Maus et al., 2009), and published ocean magnetic anomaly 1187 1188 interpretations from ship-track data. EMAG2 (non-directionally gridded version) shows the 1189 large scale trend of magnetic anomalies in the WSB, where many individual linear ocean 1190 magnetic anomalies are identifiable. The requirement for orthogonality between these linear 1191 ocean magnetic anomalies and identified FZs allowed us to check our interpretation.

Of the published ocean magnetic anomaly interpretations, those by Sègoufin and Patriat 1192 1193 (1980), Cochran (1988), and Davis et al. (2016) show strong similarities to each other, with the basin's centre of symmetry lying around the M0 ocean magnetic anomaly (125 Ma) and the oldest anomalies detected in the basin reaching between M21 and M24. Rabinowitz et al. (1983) and Eagles and König (2008) choose an alternative interpretation with the basin's centre of symmetry around M10, and located significantly farther south. We compare these lines of symmetry with MOR segments identified in the gravity data and used the best matching interpretations to temporally constrain our reconstruction. All interpretations suggest slow-intermediate spreading rates for the WSB

# 1201 <u>2.2.3. Seismic reflection data</u>

The East AfricaSPAN seismic reflection lines were used to identify the nature of the top basement reflector and measure the crustal thickness, thus helping determine whether the underlying basement is continental, oceanic, or transitional in nature. In addition, faults associated with the tectonic fabric at mature slow/intermediate spreading centres dip towards the MOR (e.g. Carbotte and Macdonald, 1990; Behn and Ito, 2008). Therefore fault polarity switches were used to help constrain the location of the MOR to the east of the DFZ, when spreading ceased ~50 Ma after breakup.

## 1209 **2.3. Feature identification**

Spreading features of the WSB, including the extinct MOR and FZs, must be detected in order to constrain plate tectonic reconstructions. To do this we identify the characteristics that define each group of features. We must, however, also be careful to avoid interpretation of areas modified by young rifts, such as the Quirimbas Graben, and post spreading magmatic additions, such as the Comoros Islands, Cosmoledo Group, and Wilkes Rise (e.g. Figure 2.3), which have strong expressions in the gravity data and mask the true nature of spreading features.

#### 1217 <u>2.3.1. Mid-oceanic ridge segments</u>

1218 Extinct MOR segments have orientations perpendicular to the final spreading direction, and 1219 appear as a free-air gravity low due to a persistent low density gabbroic root (Jonas et al., 1220 1991). These linear anomalies usually lie close to the basin's centre of symmetry, with the 1221 exception of basins undergoing subduction, such as the Pacific Ocean (Müller et al., 2008), 1222 or those having undergone spreading centre reorganisation and ridge jumps. To identify the 1223 extinct MOR from gravity anomalies, we therefore looked for three characteristics: 1) a 1224 linear free-air gravity low, 2) an orientation perpendicular to the approximately N-S paleo-1225 spreading direction (e.g. Sègoufin and Patriat, 1980), and 3) a location close to the axis of 1226 symmetry for the WSB. Free-air gravity and VGG maps were used for this task.

1227 Regions where only one gravity lineament with these characteristics was identified provided 1228 reliable MOR segment interpretations. These segments were therefore compared with the 1229 ocean magnetic anomaly interpretations of Sègoufin and Patriat (1980), Rabinowitz et al. 1230 (1983), Cochran (1988), Eagles and König (2008), and Davis et al. (2016) to assess 1231 confidence in these interpretations. Where multiple possible MOR anomalies existed, 1232 seismic reflection data were used to locate the MOR using basement fault polarity (e.g. Behn 1233 and Ito, 2008). If no seismic data were available, we chose the lineament that was most 1234 consistent with the ocean magnetic anomaly interpretations verified by our reliable MOR 1235 segments

#### 1236 <u>2.3.2. Fracture zones</u>

In the WSB, 2-5+ km of sediment have accumulated since the Jurassic (Coffin et al., 1986), removing most bathymetric expressions of FZs. A limited set of major lineations related to fracture zones can, however, be seen in the free-air gravity anomaly, reflecting crustal thickness variations, basement offsets, and infilling sediments along the FZs. We further highlighted FZ-related anomalies using a 55-85 km bandpass filter and an E-W directional derivative, orthogonal to the overall N-S spreading direction, as illustrated for the Cape Basin in Figure 2.2.

1244 The resulting linear anomalies relating to the FZ trends are generally of low amplitude 1245 compared with those arising from volcanic edifices or active rifts. This is primarily the result 1246 of the FZs' comparatively small scale, greater depth, and lower density contrast across the 1247 structural boundary. However, to map FZs, we were primarily interested in fairly continuous linear anomalies that form a consistent pattern, even if low in amplitude due to the thick 1248 1249 sedimentary cover. These lineations (e.g. Figure 2.2f) can be mapped along minima, 1250 maxima, or polarity changes in the gravity gradient; all have the same orientation and thus 1251 lead to the same plate tectonic model. However, for consistency, we have manually picked 1252 along the polarity change in the gravity gradient, except in cases where this is poorly defined 1253 and the maximum or minimum shows the orientation more clearly.

The Tanzania Coastal Basin, inboard of the DFZ, has previously been assumed to be underlain by continental crust. To determine whether to interpret any gravity lineations within this basin as continental shear zones or oceanic fracture zones, the East AfricaSPAN seismic reflection dataset was used to determine the nature of the crust. Oceanic crust was identified through the recognition of a rough high amplitude top reflector with a tectonic spreading fabric, lacking significant syn-rift deposits (Davies et al., 2005; Rodger et al., 2006) or by a hummocky reflector with continuous overlying sedimentary deposits (Soto et al., 2011), and, for normal oceanic crust, a two-way travel time (TWTT) between topbasement and any Moho reflections of ~2 seconds (e.g. White et al., 1992).

## 1263 **2.4. Plate tectonic reconstruction**

1264 After establishing the FZ lineations for the WSB, we used the plate tectonic reconstruction 1265 software Gplates (e.g. Williams et al., 2012), populated with the plate polygons of Seton et 1266 al. (2012), to retrace Madagascar's path back to Africa. The previously identified Dhow and 1267 VLCC fracture zones (Bunce and Molnar, 1977) were not used as input to the model 1268 because Coffin and Rabinowitz (1987) suggested they may be the result of tectonic 1269 processes other than oceanic spreading. Once the general origin for Madagascar was 1270 established, its position was refined by aligning conjugate continental shear zones and 1271 sedimentary basins, following Windley et al. (1994) and Reeves (2014). Artificial flowlines 1272 were then seeded at the deepest points of the basins between Madagascar and Africa 1273 according to the CRUST1.0 model (Laske et al., 2013), assumed to be the original centre of 1274 symmetry for spreading. In an iterative process, the motion of Madagascar away from Africa 1275 was then refined by aligning model-generated flowlines with the interpreted fracture zone 1276 trends. The plate model was temporally constrained by the ocean magnetic anomaly 1277 interpretations of Cochran (1988) and Davis et al. (2016), whose centres of symmetry 1278 around M0 most reliably matched the observed MOR segments. However, no plate velocity 1279 constraints exist between the initiation of rifting in the Toarcian (182 Ma) and the oldest 1280 ocean magnetic anomaly detected in the basin (M22, 150.5 Ma – Cochran, 1988; Gradstein, 1281 2012). Between the onset of rifting and breakup at 170 Ma in the Bajocian (Geiger et al., 1282 2004) we imposed an extensional velocity of 3.3 mm/y. This was based on well constrained 1283 present day extension rates along the East African Rift System between Malawi and Afar 1284 (Saria et al., 2014) and may therefore represent a realistic estimate for Jurassic extensional 1285 rates during the rifting phase. Following this rifting episode ~390 km remained between the 1286 spreading centre and the M22 magnetic anomaly, which was bridged with a constant 1287 velocity of 40 mm/y.

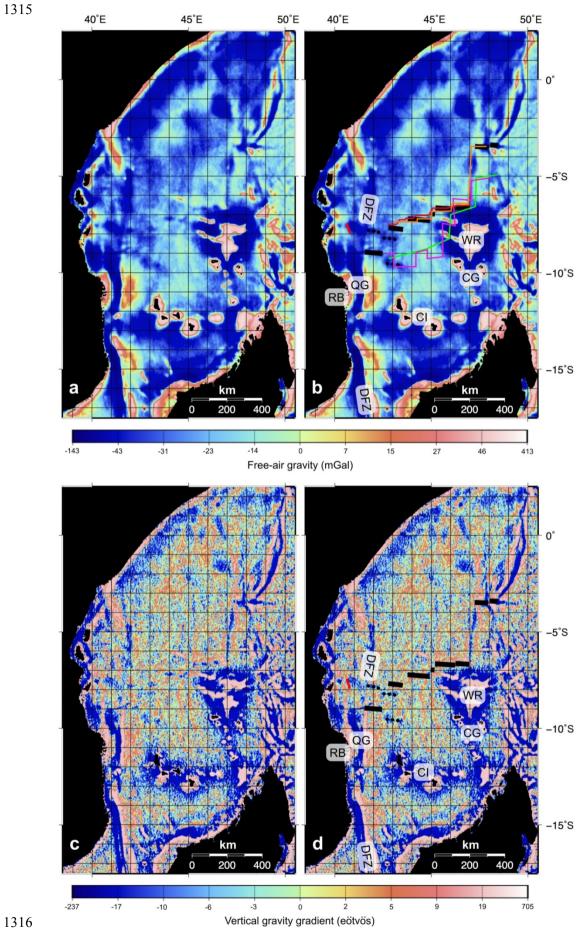
## 1288 **2.5. Results**

Using free-air gravity, VGG, and directional derivatives of filtered gravity from version 23 of the Sandwell and Smith gravity model, we have detected the extinct MOR segments and a comprehensive set of lineaments relating to spreading features in the Western Somali Basin. These features are tested against magnetic and seismic reflection data before being used to produce a high-resolution plate tectonic reconstruction of the basin.

#### 1294 <u>2.5.1. MOR segment locations</u>

1295 Oceanic magnetic anomalies in the WSB show that spreading occurred in a generally N-S 1296 direction (Sègoufin and Patriat, 1980). Following this, and the NE-SW trends of the Kenya-1297 Somalia and Northern Madagascar coastlines, we expect the extinct MOR to be composed of E-W trending segments with an overall NE-SW trend following the basins centre of 1298 1299 symmetry. We identify short linear gravity lows following this pattern in both the free-air 1300 gravity anomaly (Figure 2.3a and b) and the VGG (Figure 2.3c and d). The MOR segments 1301 generally range from 30-100 km in length, with offsets between segments ranging from as 1302 little as 20 km up to 350 km between the two easternmost segments.

1303 On the eastern side of the basin, single gravity lineaments point unambiguously to the MOR. 1304 This region is therefore used as an independent check of the previous ocean magnetic 1305 anomaly interpretations. This shows that interpretations with the basin's centre of symmetry 1306 based on M0 are most reliable, and therefore the ocean magnetic anomaly interpretations of 1307 Cochran (1988) and Davis et al. (2016) are used to temporally constrain our plate tectonic 1308 model. In the western region of the basin, close to the DFZ, two segments have several 1309 possible MOR anomalies identifiable in the gravity data (Figure 2.3 – dashed lines). For the 1310 westernmost segment, seismic reflection data covers the southern gravity lineament and 1311 shows a flip in half graben polarity centred at its location. The next segment to the east is 1312 covered by magnetic data along ship tracks and the verified ocean magnetic anomaly 1313 interpretations show the northern gravity lineament to be most consistent (Figure 2.3 – solid 1314 lines).



1317 Figure 2.3. (a) Free-air gravity anomaly. (b) Figure 2.3a overlain with the picked MOR 1318 (solid black lines) and alternative segment possibilities (dashed black lines). Location of the 1319 East AfricaSpan seismic reflection line shown in this study is indicated by the thick red line 1320 in the Tanzania Coastal Basin. Previously determined basin symmetries are shown as 1321 coloured lines where constrained by ocean magnetic anomalies. The interpretations of 1322 Cochran (1988 – orange) and Davis et al. (2016 – red) are centred on M0 and lie in good 1323 agreement with the MOR defined by gravity. The interpretations of Coffin and Rabinowitz 1324 (1987 - green) and Eagles and König (2008 - pink) are centred on M10 and deviate 1325 significantly from the gravity MOR. RB - Rovuma Basin; QG - Quirimbas graben (active 1326 rift). DFZ – Davie Fracture Zone: Post spreading volcanism: CI – Comoros islands: CG – 1327 Cosmoledo Group; WR – Wilkes Rise. (c) Vertical gravity gradient. (d) Figure 2.3c overlain 1328 with MOR picks and abbreviations as for Figure 2.3b.

#### 1329 <u>2.5.2. Fracture zone trends</u>

1330 The free-air gravity anomaly shows a number of major lineaments in the Western Somali 1331 Basin, including the Davie, ARS, Dhow, and VLCC fracture zones, as well as number of 1332 more subtle lineaments with a similar trend (Figure 2.4). Fracture zones in the Indian Ocean, 1333 which has much thinner sediment cover (<1 km; Whittaker et al., 2013), are also clearly 1334 seen. These lineaments often display a significant anomaly in the gravity field, from  $\sim 20$ 1335 mGal to over 100 mGal compared to their surroundings, and can be traced for several 1336 hundreds of kilometres, including over 1000 km in the case of the DFZ. They show an arcuate spreading pattern for the Western Somali Basin. This trend can not only be seen in 1337 the north of the basin, but is also defined by a striking bend in the DFZ located at  $41^{\circ}$  E, 14 1338 <sup>°</sup> S, which appears to deflect the trend of the continent-ocean transform margin onshore 1339 1340 along the Rovuma Basin.

1341 Lineaments detected only in the filtered and directionally differentiated gravity data are 1342 generally shorter and less continuous, ranging in length from less than 100 km up to approximately 600 km (Figure 2.5). In several instances, extensions to fracture zones 1343 detected in the free-air anomaly can be made, such as at  $45^{\circ} \ge 7^{\circ}$  S, where a conjugate 1344 1345 fracture zone to one detected in the north becomes apparent in the southern half of the basin. 1346 On the whole, these lineaments align with the framework laid out by anomalies detected in the free-air gravity and provide a comprehensive record of plate spreading directions. A few 1347 1348 short lineaments, however, lie at significant angles to the general fabric. It is likely that these 1349 lineaments are the result of structures unrelated to spreading (which should produce a 1350 consistent and predictable network of FZs), such as small volcanic chains or large infilled submarine channels, producing gravity anomalies with a similar wavelength to those ofspreading features.

1353 The EMAG2 gridded magnetic dataset (Figure 2.6a) contains several linear magnetic trends 1354 within the central region of the WSB where the basement is oceanic in nature (Coffin et al., 1355 1986). Away from magmatic structures such as the Wilkes Rise and Comoros Islands these 1356 anomalies should be due to magnetization of oceanic crust during seafloor spreading, 1357 producing ocean magnetic anomalies. Their orientation appears variable and they do not 1358 seem to define a consistent spreading direction. When, however, the magnetics are overlain 1359 by the FZ trends identified in the gravity data, the linear magnetic anomalies can be seen to 1360 lie consistently perpendicular to the arcuate fracture zone lineaments (Figure 2.6b), 1361 providing independent confirmation of our proposed fracture zone structure.

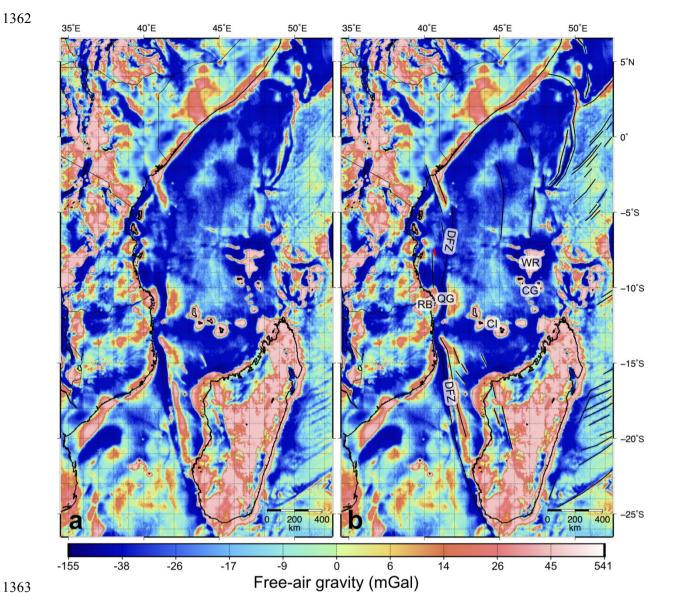


Figure 2.4. (a) Free-air gravity anomaly. (b) Figure 2.4a overlain with the linear anomalies
related to potential fracture zones. Abbreviations as for Figure 2.3. Location of the East
AfricaSPAN seismic reflection line shown in this study is indicated by the red line in the
Tanzania Coastal Basin.

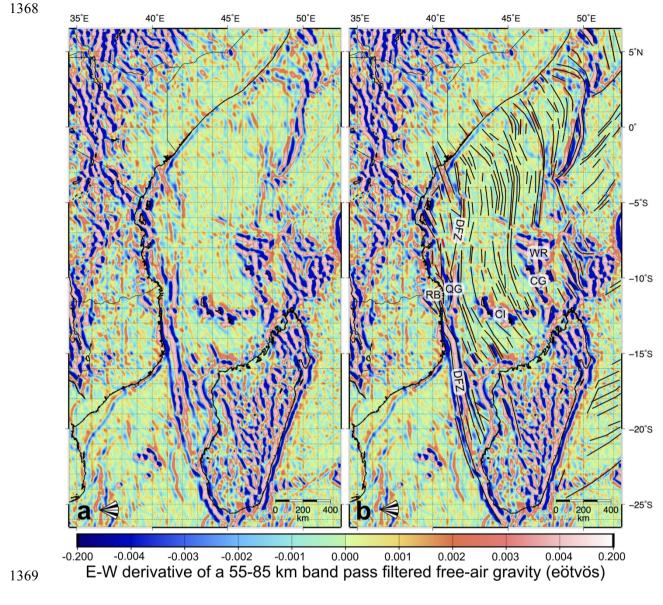


Figure 2.5. (a) E-W derivative of a Gaussian band-pass filtered free-air anomaly, 50% long and short wavelength cutoffs at 85 and 55 km, respectively, to best retain anomalies related to fracture zones. (b) Figure 2.5a overlain with the linear anomalies identified here and in Figure 2.4. Abbreviations as for Figure 2.3. Location of the East AfricaSpan seismic reflection line shown in this study is indicated by the red line in the Tanzania Coastal Basin

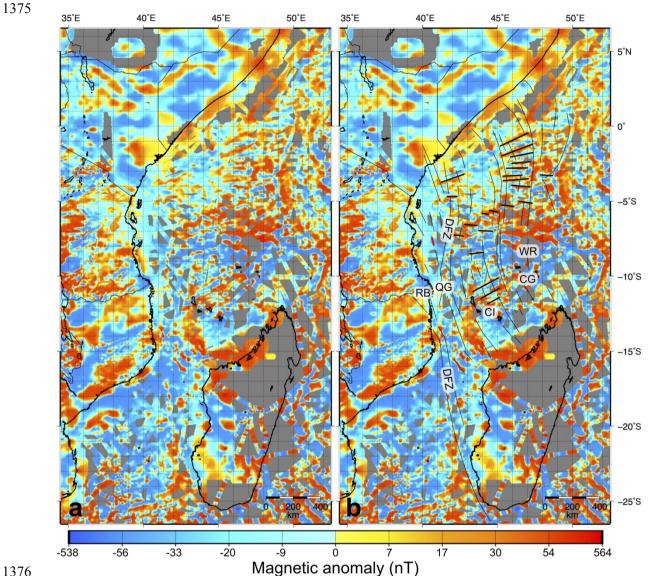




Figure 2.6. (a) The EMAG2 non-directionally gridded magnetic anomaly dataset. (b) Figure 1377 2.6a with broadly E-W linear magnetic anomalies detected from unmodified oceanic crust 1378 1379 (thick black lines) and fracture zones (thin black lines) marked, showing a consistently orthogonal relationship. Abbreviations as for Figure 2.3. Location of the East AfricaSpan 1380 seismic reflection line shown in this study is indicated by the red line in the Tanzania 1381 1382 Coastal Basin.

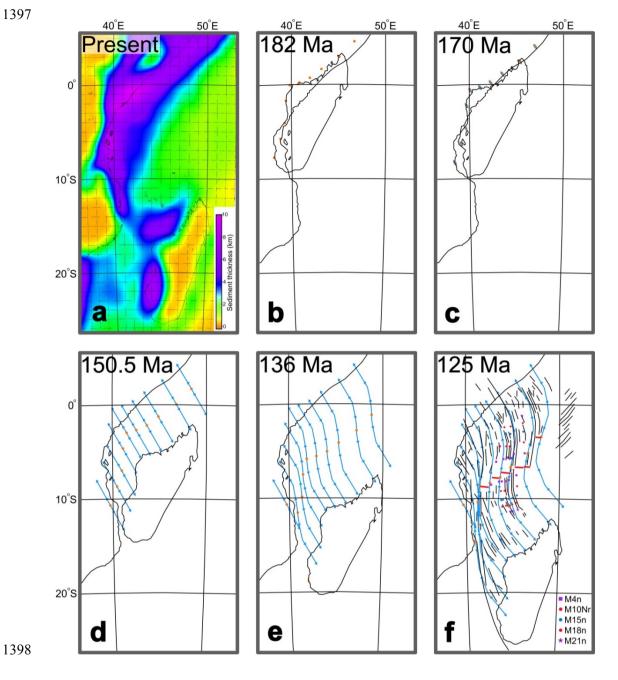
#### 1383 <u>2.5.3. Plate tectonic model</u>

1384 Using our new fracture zone lineaments, shear zone data from Reeves and De Wit (2000),

1385 and basin depth data from CRUST1.0, we developed a new plate tectonic reconstruction for

1386 Madagascar's separation from Africa (Figure 2.7).

1387 An initial phase of continental rifting from 182 Ma leads to continental break up at 1388 approximately 170 Ma (Figure 2.7b-c). Oceanic spreading commences in a NNW-SSE 1389 direction and results in strike slip tectonics between Madagascar and northern Mozambique, 1390 forming the Rovuma Basin (Figure 2.7c-d). At ~150.5 Ma the spreading direction changes to 1391 almost N-S, resulting in the near alignment of several flow lines in the west of the basin 1392 (Figure 2.7d-e). After 136 Ma the spreading direction continues to rotate causing full 1393 convergence of the flow lines in the west of the basin along the trace of the DFZ. Faster 1394 spreading in the west compared to the east also results in an anti-clockwise rotation of 1395 Madagascar to its present day position, which was reached when oceanic spreading ceased 1396 in the basin at ~125 Ma (Figure 2.7e-f).



1399 Figure 2.7. Plate tectonic reconstruction of Madagascar's escape from Africa from the Early 1400 Jurassic to the cessation of spreading in the Cretaceous. Madagascar is shown without the 1401 remainder of East Gondwana (India, Antarctica and Australia) attached. (a) Present day 1402 sediment thickness in the Western Somali Basin taken from the CRUST1.0 model. (b-e) The 1403 key stages of Madagascar's motion out of Africa. Modelled flowlines are shown as blue 1404 arrowed lines where the centre of symmetry is marked by orange circles. (f) Madagascar's 1405 present-day position, which is reached at around 125 Ma. Flowlines closely match the 1406 fracture zone pattern of the basin (additional black lines), and the basin's predicted final 1407 symmetry (orange circles) lies in good agreement with the interpreted extinct mid-ocean

ridge system (red lines). Locations of magnetic anomalies used to temporally constrain platemotions shown with symbols as interpreted by Davis et al. (2016).

## 1410 **2.6. Discussion**

#### 1411 <u>2.6.1. The nature of the WSB's margins and of gravity lineaments in the coastal</u>

1412 <u>basins</u>

1413 Modern concepts of passive margin formation define two end member types. 1) At volcanic 1414 rifted margins, crustal thinning occurs over relatively short distances (50-100 km (e.g. 1415 Franke, 2013)) and is accompanied by large volumes of magmatism. These are characterised 1416 by both thick wedges of volcanic flows that appear as seaward dipping reflectors in seismic 1417 reflection data (e.g. Planke and Eldholm, 1994; Geoffroy, 2005) and by high velocity 1418 underplating and heavily intruded crust identified in seismic refraction studies (e.g. 1419 Korenaga et al., 2000; Hirsch et al., 2009). 2) At magma-poor rifted margins, largely 1420 unthinned continental crust of the proximal domain passes into a hyperextended domain 1421 containing three sub domains (necking, hyperthinned, and exhumed mantle domains), before 1422 oceanic crust marks the onset of the oceanic domain (Tugend et al., 2015). The necking 1423 domain and hyperthinned domains accommodate most of the crustal thinning, containing 1424 continental crust of <10 km thick, and typically extend 100-200 km from the proximal 1425 domain (e.g. Reston, 2009; Sutra and Manatschal, 2012). The exhumed mantle domain 1426 forms the continent-ocean transition and is thought to consist of mantle material unroofed 1427 and serpentenised during extensional detachment faulting (e.g. Bayrakci et al., 2016; Gillard 1428 et al., 2016), which separates continental crust from oceanic crust. These margins 1429 characteristically experience limited magmatism during extension (Franke, 2013).

1430 To understand the style of margin formation in the WSB, we draw together a combination of 1431 seismic, gravity, magnetic and geological evidence. Coffin et al. (1986) confirmed the 1432 existence of oceanic crust just offshore of the Kenya-Somalia border as far north and west as 42.05° E 2.52° S, but inboard of this within the Tanzania Coastal Basin and extending 1433 onshore within the Lamu Embayment, thin crust (<13 km thick; Reeves et al., 1987) of an 1434 1435 ambiguous nature is covered by thick sediments (up to +12 km, Yuan et al., 2012 and 1436 references therein). Based on gravity and magnetic modelling, Reeves et al. (1987) proposed 1437 that this crust is oceanic in nature, consistent with observations of necking zones (as defined 1438 by Tugend et al., 2015) onshore along the western edge of the Lamu Embayment from 1439 seismic refraction data, which suggest a sharp crustal thinning from over 40 km to probably 1440 less than 15 km at this location (Prodehl et al., 1997). This implies that offshore seismic 1441 reflection data along the Tanzanian and Kenyan margins is located seaward of the necking 1442 zones. Ascertaining the margin nature is thus more difficult, since seaward dipping 1443 reflectors, which are characteristic of volcanic margins, form in vicinity to the necking zone 1444 and therefore may not be detected, whilst exhumed mantle domains, as seen at magma-poor 1445 margins, can be difficult to distinguish from oceanic crust formed at slow spreading centres 1446 when using seismic reflection data alone (e.g. Davy et al., 2016).

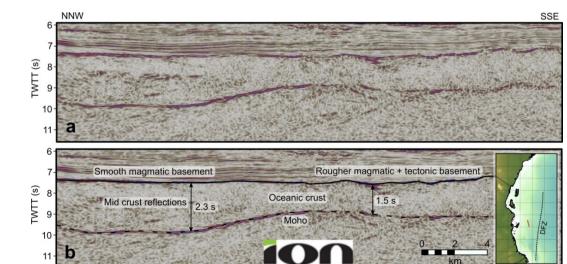
1447 Furthermore, the already thin crust onshore within the Lamu Embayment suggests that the 1448 'shelf edge' high, seen in the free-air gravity along the Somali coast east of the DFZ, is not 1449 indicative of crustal thinning, and may not coincide with the continent-ocean transition as proposed elsewhere (e.g. Bauer et al., 2000). The effects of increasing water depth and thick 1450 1451 sedimentary accumulations can also produce this pattern of gravity anomalies without an 1452 additional contribution from decreasing crustal thickness (e.g. Walcott, 1972; Watts & 1453 Stewart, 1998). Elsewhere, along the western margins of the WSB, no shelf edge gravity 1454 anomalies are present, possibly due to superimposed effects of active rifting in the area, 1455 providing little information as to the margin nature. However, as there is little evidence for 1456 Jurassic rift-related volcanic rocks exposed at the surface in Madagascar, Tanzania or Kenya 1457 (e.g. Guiraud et al., 2005), and neither have they been drilled onshore or offshore (despite 1458 the pervasive record of post-rift volcanics emplaced in the upper Cretaceous (Coffin and 1459 Rabinowitz, 1988) related to the breakup between Madagascar and India (e.g. Storey et al., 1460 1995)), significant magmatism during rifting in the WSB seems unlikely. This apparent lack 1461 of rift related volcanism, the generally thin nature of the oceanic crust interpreted elsewhere 1462 within the WSB (5.22  $\pm$  0.64 Km, Coffin et al., 1986), and the lack of any high velocity 1463 underplating interpreted around the necking zone from seismic refraction studies (Prodehl et 1464 al., 1997) make present observations from the margins of the WSB more consistent with the 1465 magma-poor endmember style of rifted margins.

1466 The ambiguous crust within the Tanzania Coastal Basins and Lamu Embayment could 1467 therefore be either oceanic, formed after breakup, or hyperextended continental crust and 1468 mantle. Whilst the present-day high heatflow along the East African margin (attested by the 1469 significant gas discoveries in the region) possibly favours the presence of radiogenic 1470 continental crust (e.g. White et al., 2003), active lithospheric thinning along the offshore 1471 branch of the East African Rift System (e.g. Delvaux and Barth, 2010; Franke et al., 2015) 1472 would also act to increase regional heatflow. In the mid-Tanzania Coastal Basin, inboard of 1473 the DFZ, the East AfricaSPAN seismic reflection lines image a strong and continuous 1474 reflector at approximately 9.5 s TWTT, which is 1.5 to 2.3 s TWTT below the top basement 1475 (Figure 2.8). This is typical of slightly thin to normal oceanic crust (White et al., 1992), and 1476 is similar to the 1.17 to 2 s TWTT derived for oceanic crust elsewhere within the WSB

1477 (Coffin et al., 1986). Several areas within the crust also have a low reflectivity character,1478 often described within oceanic crust (e.g. Bécel et al., 2015).

1479 The nature of the crust changes from NNW to SSE along the line. In the NNW, a smooth top 1480 basement reflector is imaged at 7.4 s TWTT, which is characteristic of oceanic crust formed 1481 by relatively robust magmatic accretion with little tectonic extension (e.g. Reston et al., 1482 2004). Here an additional reflector can be seen at  $\sim$ 8.3 s TWTT (0.9 seconds below top 1483 basement) which delineates an upper and lower crustal layer. Elsewhere within the WSB, the 1484 oceanic layer 2 thickness has been derived as 0.93 s TWTT (Coffin et al., 1986), and so this 1485 reflector may represent the boundary between oceanic layers 2 and 3. The smooth top 1486 basement and Moho reflectors also extend ENE along the seismic cross line, perpendicular 1487 to the spreading direction, with a consistent offset of between 1.8 and 2 s TWTT, before 1488 eventually reaching the tectonically overprinted DFZ, where their character is lost. To the 1489 SSE in Figure 2.8, following a reduction in thickness of the crust demonstrated by the 1490 shallowing of the Moho reflector, the top basement gains a weak tectonic fabric. These 1491 observations are consistent with a reduction in magma supply and resulting increase in the 1492 tectonic extensional component of oceanic spreading (e.g. Reston et al., 2004).

1493 Alternatively, the Moho reflector could represent a detachment fault formed between 1494 continental crust and mantle during hyper-extension (e.g. Tugend et al., 2015), such as the S 1495 reflector west of Galicia (Hoffmann and Reston, 1992) and H reflector in the Iberia Abyssal 1496 Plain (Dean et al., 2008). However, the smooth top basement reflector lacks the well-defined 1497 fault blocks often imaged in such hyper-extended domains (Reston, 2009). This suggests 1498 that extreme crustal extension is unlikely, especially as rift related volcanism, which could 1499 otherwise have masked fault block topography, is extremely limited during hyper-extension 1500 at magma-poor margins (Franke, 2013).



1502

Figure 2.8. (a) Seismic reflection line from the East AfricaSPAN (Ion Geophysical) inboard
of the DFZ. (b) Figure 2.8a with interpretation overlain. Inset shows location relative to the
coastline and the DFZ.

1506 All these observations thus support high levels of extension, probably including oceanic 1507 crust, in the mid-Tanzania coastal basin inboard of the DFZ. The DFZ cannot then be a 1508 simple continent-ocean transform margin. Instead, Madagascar must have originated from 1509 within the Tanzania Coastal Basins and Lamu Embayment, with the Rovuma Basin forming 1510 the continental-ocean transform margin. The onshore trend of this basin is closely aligned 1511 with the early SSE trending fracture zones detected in the gravity data (Figure 2.9b), and in 1512 fact, in our plate tectonic reconstructions, strike-slip motion of southern Madagascar along 1513 this basin is unavoidable. This is in good agreement with observations of dextral strike slip 1514 faults along the Rovuma Basin margin (Emmel et al., 2011) and an onshore sedimentary 1515 thickness of  $\sim 10$  km in the northern Rovuma Basin that rapidly thins westward to <1 km 1516 (Key et al., 2008), consistent with a continent-ocean transform margin. As noted by Reeves 1517 (2014), the passage of Madagascar along the Rovuma Basin also allows for a much tighter 1518 and consistent fit of Gondwana fragments, reducing the need for gaps and non-smooth plate 1519 motions during Gondwana's disassembly.

## 1520 <u>2.6.2. Rifting mechanisms and Gondwana breakup</u>

1521 Rifting between East and West Gondwana began in the Toarcian (Geiger et al., 2004) and 1522 was probably initiated by the eruption of the Bouvet plume, resulting in a contemporaneous 1523 volcanic passive margin in Mozambique (Klausen, 2009). Here, an 8.5 km thick suite of rift-1524 related basalts and rhyodacites defines a relatively narrow volcanic margin, where a 1525 magmatic mode of extension dominated in the lead-up to breakup (Klausen, 2009). This 1526 section of the rift system developed discordantly to the structural trend of Gondwana's 1527 sedimentary basins (Salman and Abdula, 1995), suggesting that pre-existing lithospheric 1528 structure was not a key parameter leading to breakup.

1529 In the Western Somali basin, however, there is little evidence for a magmatic breakup as 1530 discussed earlier (Section 2.6.1.). This is most likely a function of the WSB's distance from 1531 the volcanic centre in Mozambique as seen in the Gulf of Aden. Here, volcanic margins formed close to the Afar hotspot, yet farther away, east of longitude  $46^{\circ}$  E, the margins are 1532 1533 magma-poor (Leroy et al., 2012). Breakup along the Tanzanian-Kenyan and Kenyan-1534 Somalian rift sections is therefore less likely to have been influenced by magmatism and 1535 thermal weakening of the lithosphere (Buck, 2007). It is apparent from the spreading 1536 lineaments detected in the WSB that initial spreading occurred in a NNW-SSE direction, in 1537 agreement with principal extensional stresses around the Mozambique basin (Le Gall et al., 1538 2005). This is consistent with the occurrence of strike-slip tectonics along the NNW-SSE 1539 trending Rovuma Basin and oblique rifting along the N-S trending Kenya-Tanzania margin 1540 (Figure 2.9b), both of which are mechanically favourable (Emmel, 2011; Brune et al., 2012). 1541 This is similar to observations from the Gulf of California where oblique rifting assisted 1542 continental breakup through the efficient focusing of crustal thinning within pull-apart 1543 basins bounded by large offset strike-slip faults (Bennett and Oskin, 2014). If this 1544 mechanism was active during the Jurassic rifting along the Tanzania-Kenya margins, it may 1545 explain the possible margin segmentation suggested by the stepped shape of Madagascar's western coastline. Margin segmentation is common to many oblique passive margins 1546 1547 worldwide (e.g. Leroy et al., 2012; Bennett and Oskin, 2014).

1548 Conversely, the NE-SW trending Kenyan-Somalian rifted margin formed orthogonally to the 1549 breakup direction. Although the Rovuma basin shows little evidence of Karoo age rifting 1550 and sedimentation (Smelror et al., 2008), the Morondava, Majunga and Ambilobe basins of 1551 Madagascar all contain underlying Karoo sediments (e.g. Hankel, 1994). The conjugate 1552 margins on the mainland, the Tanzanian-Kenyan and Kenyan-Somalian rift systems, thus 1553 appear to have followed the pre-existing lithospheric structure of the Karoo rift system. A 1554 transition can therefore be proposed along the East African margin from dominant strike-slip 1555 tectonics and oblique rifting in the Royuma Basin, progressing northwards to oblique rifting 1556 that also follows pre-existing lithospheric structures along the Tanzanian-Kenyan section, 1557 and finally, purely orthogonal rifting along a pre-existing lithospheric structure along the 1558 Kenyan-Somalian section. This is consistent with the obliqueness of rifting (Brune et al., 1559 2012) and pre-existing lithospheric structure (e.g. Audet and Bürgmann, 2011) assisting 1560 supercontinent breakup.

1561 The rifting between East and West Gondwana therefore provide a good natural laboratory 1562 for the study of the spatially variable interplay between different rifting mechanisms during 1563 supercontinent breakup. Examples where each of the proposed facilitating mechanisms 1564 (magmatism, oblique rifting, and pre-existing structure) appears to dominate during breakup 1565 can be seen along the Gondwana rift system between Mozambique and Somalia, with 1566 predominantly magmatic breakup in the Mozambique Basin, apparent strike-slip and oblique 1567 tectonics along the Rovuma Basin, and coincident pre-existing lithospheric structure along 1568 the Kenyan-Somalian coast. Analogy can be made to the opening of the South Atlantic 1569 during breakup of the supercontinent Pangea, where evidence supports similar regional 1570 variation in breakup mechanism. Here, a south to north transition from magmatically 1571 dominated breakup in the southern South Atlantic (e.g. Gibson et al., 2006), inheritance 1572 driven rifting in the central South Atlantic (e.g. Lentini et al., 2010), and strongly oblique 1573 rifting in the Equatorial Atlantic (Heine and Brune, 2014) is seen. Together, these margins

1574 suggest that rifting during supercontinent dispersal may often be facilitated by multiple 1575 mechanisms, with regional variation along the margin due to different pre-existing 1576 geological structures and changing tectonic geometry on length scales as short as a few 1577 hundred kilometres.

#### 1578 <u>2.6.3. Plate tectonic reconstruction</u>

For the initial rifting phase we impose a plate separation rate of 3.3 mm/y, similar to that of 1579 1580 the present day East African Rift System between Malawi and Afar (Saria et al., 2014). Breakup occurred at approximately 170 Ma, as evidenced by the Early Bajocian break up 1581 1582 unconformity identified in the Morondava Basin (Geiger et al., 2004) and the overwhelming 1583 transition to marine deposits along the East Africa margins at this time (Coffin and 1584 Rabinowitz, 1992). Between breakup and the earliest magnetic anomaly constraint (M22) an 1585 average full spreading rate of 40 mm/y therefore occurred, similar to the average full 1586 spreading rate of ~49 mm/y determined by ocean magnetic anomalies for the younger 1587 oceanic crust between M22 and M0 (Cochran, 1988; Davis et al., 2016).

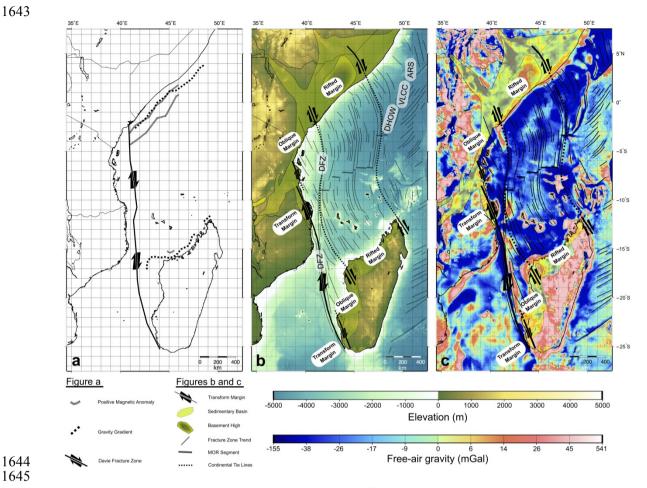
1588 Following this initial phase of spreading, which resulted in strike-slip motion between 1589 Madagascar and the Rovuma Basin, a rotation in the spreading direction occurred at  $\sim 150.5$ 1590 Ma. The oldest conjugate pair of magnetic anomalies detected, M22 (Cochran, 1988; Davis 1591 et al., 2016), constrains the age of this rotation, which is contemporaneous with Madagascars 1592 exit from the SSE trending Rovuma Basin, after which it began to follow a N-S spreading 1593 direction. This rotation began the cessation of any oceanic spreading in the Tanzania Coastal 1594 basin and offshore Morondava Basin as flow lines began to align along what was to become 1595 the DFZ (Figure 2.7d-e). This alignment suggests strike-slip tectonics began to dominate 1596 along this zone, and it is therefore possible that the DFZ formed at this point as several 1597 fracture zones coalesced into one major oceanic fracture zone with a significant accumulated 1598 offset.

Following the first rotation in plate motions at 150.5 Ma, spreading continued relatively undisturbed in the Western Somali Basin until approximately 136 Ma, when a further change in plate motion occurred contemporaneous with Madagascar's departure from the Mozambique continental transform margin. This rotation further aligned the flowlines along the DFZ as it became the dominant strike slip fault in the basin. From here until the termination of oceanic spreading at M0 (125 Ma), Madagascar underwent a gentle anticlockwise rotation to take its present day position relative to Africa.

1606 The termination points for our model flowlines lie very close to the extinct MOR identified 1607 in the gravity data. We emphasise that this striking agreement is generated only from our fracture zone trends and initial seed points for the flowlines, which were chosen independently based on the CRUST1.0 dataset, and thus provides strong confirmation of the model. No ocean magnetic anomalies have been identified to help constrain the location of the westernmost segment of the MOR. However, seismic reflection data suggests a southerly location for the MOR segment in line with the symmetry predicted from the plate tectonic reconstruction.

1614 A key result of the reconstruction is that the DFZ is shown to be a major ocean-ocean FZ, 1615 where oceanic crust has formed inboard of this feature within the Tanzania Coastal Basin. 1616 This challenges many plate tectonic reconstructions which, based on the previously available 1617 literature, have defined the DFZ as the continent-ocean transform margin of the Western 1618 Somali Basin (e.g. Gaina et al., 2013), an important constraint on spreading kinematics. As 1619 the DFZ is a predominantly straight feature, treating it as the continent-ocean transition 1620 naturally results in the prediction of a less complex spreading pattern (i.e. only ~N-S without 1621 an initial NNW-SSE component) and a looser fit of Gondwana fragments due to the inability 1622 to reconstruct continents inboard of the DFZ. However, by detailed analysis of spreading 1623 lineaments on a small scale, we have been able to resolve the initial NNE-SSW spreading 1624 stage. This is in agreement with NNW-SSE principal extensional stresses during breakup 1625 around the Mozambique basin, recorded from dyke dilation in the Okavango and Limpopo 1626 dyke swarms (Le Gall et al., 2005). This spreading pattern is also strikingly similar to 1627 spreading patterns extrapolated to the WSB basin from the Mozambique basin, where they 1628 were derived from magnetic anomalies and FZs (Eagles and König, 2008). This suggests 1629 that during the earlier stages of spreading. Madagascar and Antarctica shared a similar 1630 breakup history, and moved as a cohesive unit away from Africa, as opposed to an 1631 amalgamation of continental blocks with relative motions between them. This highlights the 1632 importance of basin scale reconstructions in deciphering supercontinent dispersal 1633 mechanisms, as well as their potential for constraining the histories of neighbouring basins 1634 that lack detailed kinematic indicators and for informing larger regional reconstructions.

The Dhow and VLCC fracture zones as interpreted by Bunce and Molnar (1977) were not used as input to the plate tectonic model since they may have formed by processes other than oceanic spreading (Coffin and Rabinowitz 1987). However, their trends are independently predicted by our plate model, so they are likely to have been originally formed as the result of plate spreading after all. Reactivation of these structures may have occurred during the breakup of Madagascar and India, resulting in their more prominent expression in the gravity data compared to other fracture zones in the WSB.



1646 Figure 2.9. (a) Commonly interpreted basin configuration, where the continent-ocean 1647 transition is assumed to follow the DFZ (e.g. Bunce and Molnar, 1977; Coffin and 1648 Rabinowitz, 1987). (b) Schematic of the basin configuration suggested in this study, with 1649 strike slip tectonics dominating along the edge of the Rovuma Basin, while much of the 1650 Tanzania Coastal Basin should be considered as an obliquely rifted margin. The Davie 1651 Fracture Zone is a major ocean-ocean fracture zone, not the continent-ocean transform 1652 margin. DFZ - Davie Fracture Zone; DHOW - Dhow Fracture Zone; VLCC - Very Large Crude Carrier Fracture Zone; ARS - Auxiliary Rescue and Salvage Fracture Zone. (c) Free-1653 1654 air gravity overlain with interpretation as for Figure 2.9b.

## 1655 **2.7. Conclusions**

Using new techniques to analyse the latest Sandwell and Smith gravity datasets (V23), we have detected the location of the extinct MOR segments and, for the first time, a comprehensive set of fracture zone lineaments within the Western Somali Basin. We have used these to constrain Madagascar's position in Africa prior to breakup, validate ocean magnetic anomaly interpretations for the WSB, and construct a well constrained, high resolution plate tectonic reconstruction for the region. This plate tectonic reconstruction bears strong similarities to reconstructions from the neighbouring Mozambique Basin, and may suggest that East Gondwana broke off from West Gondwana as a cohesive unit, rather than as an amalgamation of continental blocks with relative motions between them. During this disassembly, no single parameter leads to breakup along the entire margin, with thermal weakening due to magmatism, oblique rifting, and pre-existing structure apparently dominating in turn from south to north along the Jurassic Gondwana rifts.

1668 The discovery of oceanic crust in the Tanzania Coastal Basin, fracture zone orthogonality to 1669 regional magnetic anomalies, and observations from the Rovuma Basin support this 1670 reconstruction, and show that the Davie Fracture Zone is a major ocean-ocean fracture zone, 1671 formed by the coalescence of several smaller fracture zones during changing spreading 1672 directions, and not a continent-ocean transform margin. The western edge of the basin is 1673 thus defined by a transform margin in the Rovuma Basin, whereas the Tanzanian and 1674 Kenyan margins formed in an oblique regime and are most likely segmented, magma-poor rifted margins. The change in the location and nature of the continent ocean transition has 1675 1676 important implications for the nature of the lithosphere underlying the western portion of the 1677 basin, and thus for its thermal history and resource potential.

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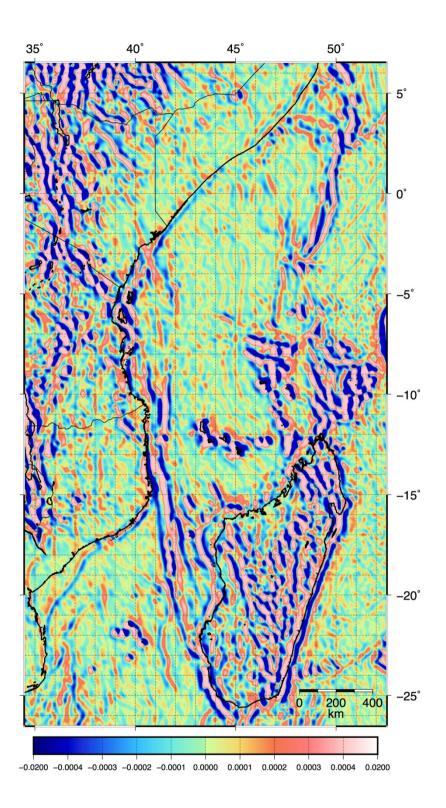
# 1875 **2.9. Supplementary material**

1876 These supplementary images show the effect of changing key parameters during the gravity

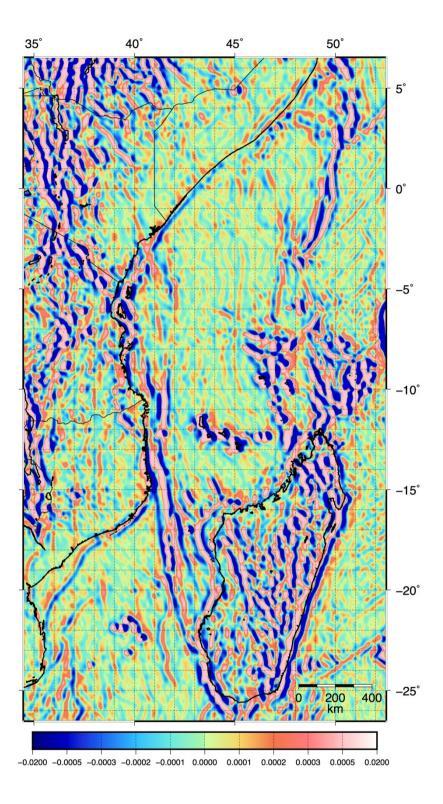
- 1877 processing described in the main text. Free-air gravity data from the Western Somali Basin
- 1878 (WSB) has been processed to enhance the expression of fracture zones. First filtering was

1879 performed using a Gaussian bandpass filter with 50% cutoffs at 55 and 85 km wavelengths. 1880 Following this, directional derivatives of the gravity were taken perpendicular to the 1881 approximate N-S spreading direction, as determined from ocean magnetic anomalies from the WSB. Derivatives were taken at 10° intervals between 30° clockwise and anticlockwise 1882 1883 of E-W. The greatest gradient magnitude recorded from this set of derivatives was stored for 1884 each grid element to form the final gradient anomaly map (Figure S2.1).

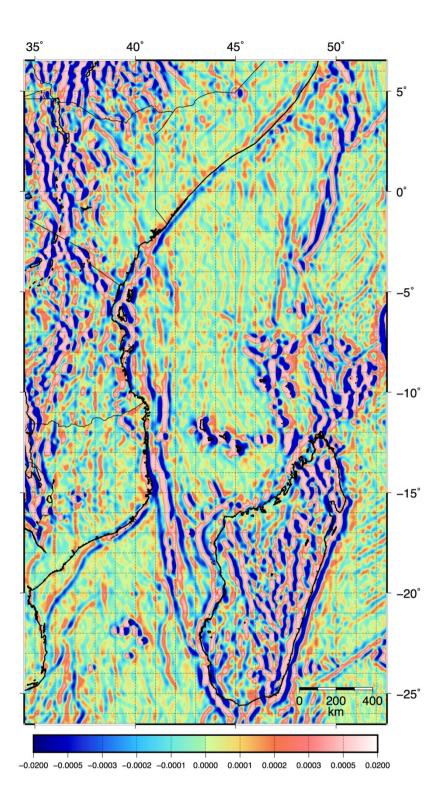
- 1885 Supplementary Figures S2.2-4 show the effects of reducing the range of azimuths over 1886 which gradients are taken; the aim being to show that the process of taking several gradients 1887 assists in highlighting the trends of curved fracture zones. In the case of the WSB, the 1888 greatest improvements can therefore be seen closer to the continents where the spreading 1889 direction early in the basin history was ~NW-SE and not N-S.
- 1890 Supplementary figures S2.5-8 show result of taking gradients over these same ranges of 1891 azimuths as in supplementary figures S2.1-4, but parallel to the spreading direction as 1892 opposed to perpendicular to it. This shows that lineations do not arise purely as a result of
- 1893 this processing without a geological driver.



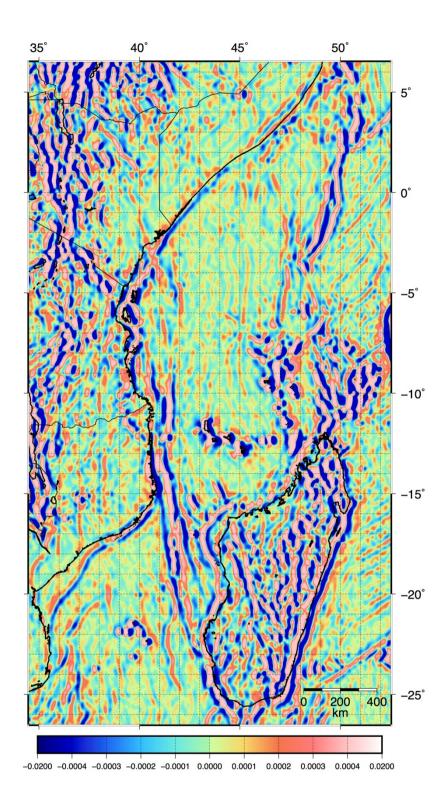
1896 Figure S2.1. Maximum free-air gravity directional derivative from azimuths of
1897 240°/250°/260°/270°/280°/290°/300°. Units of scale are eötvös.



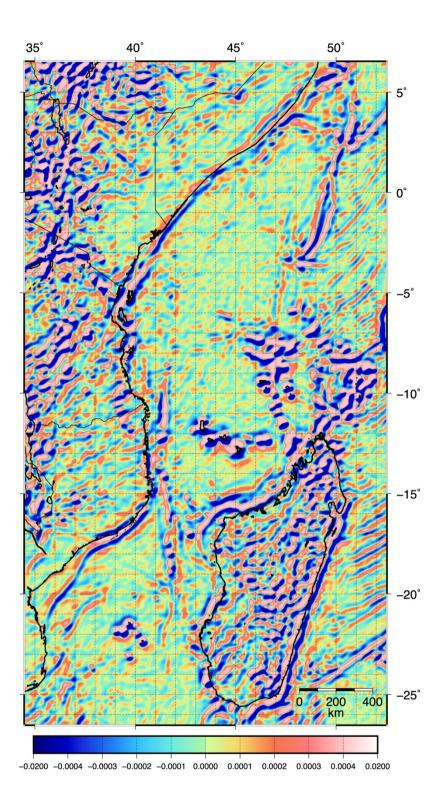
1900 Figure S2.2. Maximum free-air gravity directional derivative from azimuths of
1901 250°/260°/270°/280°/290°. Units of scale are eötvös.



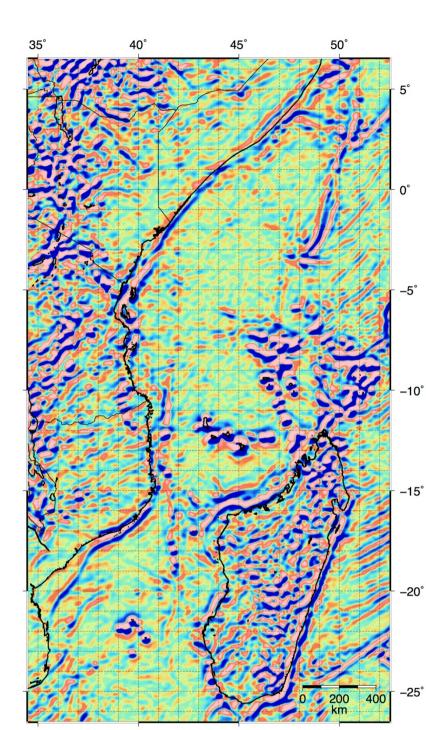
1904 Figure S2.3. Maximum free-air gravity directional derivative from azimuths of
1905 260°/270°/280°. Units of scale are eötvös.



1908 Figure S2.4. Free-air gravity directional derivative at an azimuth of 270°. Units of scale are
1909 eötvös.

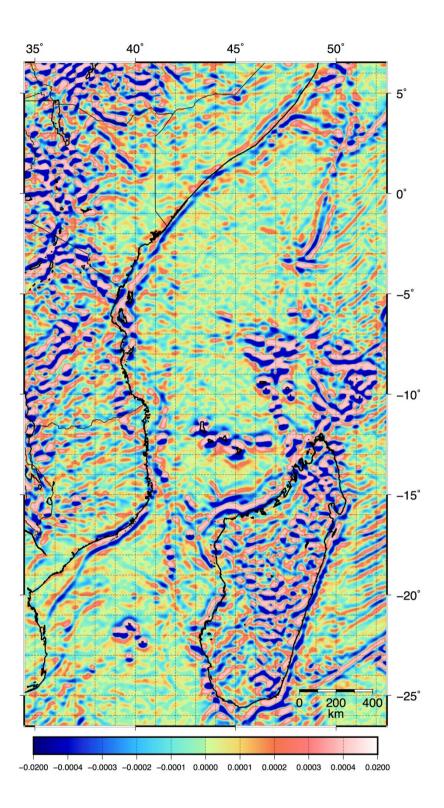


1912 Figure S2.5. Maximum free-air gravity directional derivative from azimuths of
1913 330°/340°/350°/0°/10°/20°/30°. Units of scale are eötvös.



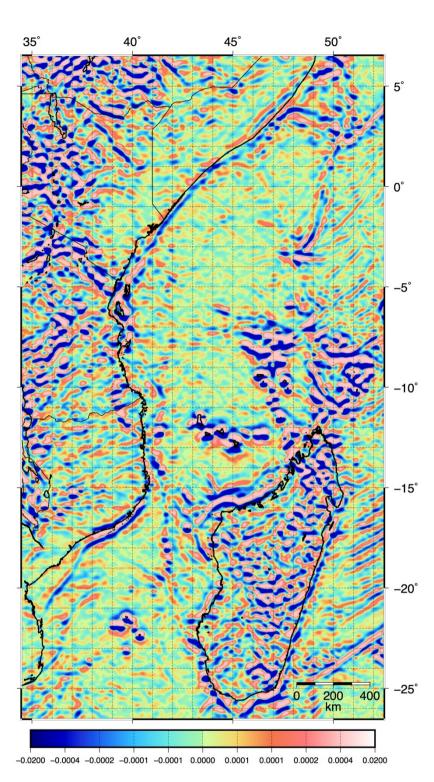
-0.0200 -0.0004 -0.0003 -0.0002 -0.0001 0.0000 0.0001 0.0002 0.0003 0.0004 0.0200

1916 Figure S2.6. Maximum free-air gravity directional derivative from azimuths of
1917 340°/350°/0°/10°/20°. Units of scale are eötvös.



**Figure S2.7.** Maximum free-air gravity directional derivative from azimuths of 350°/0°/10°.

1921 Units of scale are eötvös.



1924 Figure S2.8. Free-air gravity directional derivative at an azimuth of 0°. Units of scale are
1925 eötvös.

# 3. The Rovuma Transform Margin: Pinning down the East African continent-ocean transform margin using seismic and gravity methods

# 1930 Abstract

1931 A firm knowledge of the nature and location of the continental margins of East Africa is 1932 crucial for producing accurate plate tectonic reconstructions of Gondwana, hydrocarbon 1933 resource exploration, and for developing our understanding of how supercontinents break 1934 apart. Here, we present evidence for a newly identified continent-ocean transform margin 1935 along the Rovuma Basin, identified using seismic and gravity data. The 'Rovuma Transform 1936 Margin' lies landward of the Davie Fracture Zone, which has previously been interpreted as 1937 the continent-ocean transform margin of the Western Somali basin, and trends in a NNW 1938 direction, following the onshore basement outcrop of the Rovuma Basin. The >400 km long 1939 transform margin connects the obliquely rifted Tanzania Coastal Basin in the north, with the 1940 Nacala and Mozambique basins in the south. This configuration supports an origin for 1941 Madagascar within the Tanzania Coastal Basin, and tight-fit reconstructions of Gondwana 1942 fragments prior to the Jurassic.

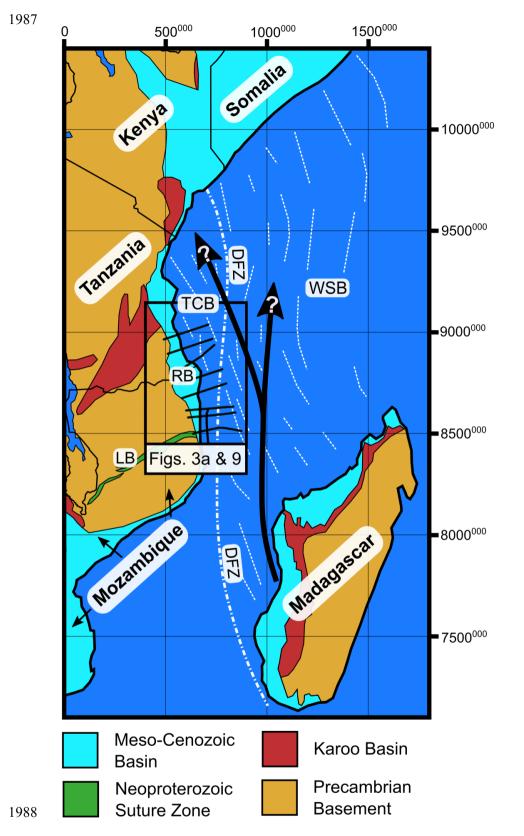
# 1943 **3.1. Introduction**

1944 Continental breakup occurs when a zone of continental lithosphere is thinned to the point of 1945 rupture due to the separation of tectonic plates (Ebinger, 2005). This results in partial 1946 melting of the asthenospheric mantle, leading to oceanic crust production and the 1947 localisation of strain at the mid-ocean ridge (MOR). At this time, the region where 1948 continental lithosphere is thinned (from ~30 to 0 km) may be referred to as a passive 1949 continental margin. Due to the relation between extension and accommodation space 1950 (Tugend et al., 2015), continental margins may be readily identified between the continents 1951 and ocean by using seismic reflection data to identify the great increase in sedimentary 1952 accommodation space that occurs across them. This increase is primarily due to the subsidence that occurs across passive margins as the result of isostatic compensationfollowing thinning of the relatively buoyant crust.

1955 The East Africa margins developed during the Early Jurassic, following the cessation of 1956 older Karoo rifting episodes (Geiger et al., 2004), by the breakup of East and West 1957 Gondwana (e.g. Reeves et al., 2016). The subsequent drift of East Gondwana (Madagascar, 1958 India, Antarctica, and Australia) away from West Gondwana (Africa, South America) 1959 generated the Western Somali Basin (WSB; e.g. Rabinowitz et al., 1983). Figure 3.1 1960 summarises the two main groups of plate tectonic models have been proposed to describe 1961 plate motions during the formation of the WSB. 1) Direct N-S motion of East Gondwana 1962 away from West Gondwana, often assumed to follow the Davie Fracture Zone (DFZ) and 1963 requiring an initial 'loose' fit of Gondwana fragments (e.g. Coffin and Rabinowitz, 1987). 2) 1964 A similar pattern, but where an initial NW-SE (Klimke et al., 2017; Reeves et al., 2016) or 1965 NNW-SSE (Phethean et al., 2016) motion precedes N-S spreading allowing for a 'tight' 1966 initial configuration of Gondwana fragments.

Depending on the orientation of a margin relative to the extensional direction under which it forms, margins develop under different modes (i.e. divergent, transform, or oblique). The resulting geometries and structures are characteristic of each margin type, allowing us to differentiate between them. Due to the different initial extension directions used by the two groups of plate models, each group predicts that the margins of the curved WSB will develop differently. For example, models using N-S spreading predict transform margins with a N-S trend, whereas those using NNE-SSW spreading predict a NNE-SSW trend.

1974 Understanding the mode of rifting and margin structures of the WSB is not only crucial for 1975 our understanding of how supercontinent dispersal occurs, but also has significant 1976 implications for the nature of the basement rocks surrounding the margins of East Africa and 1977 the location of the continent-ocean transition (COT), both important factors for hydrocarbon 1978 exploration. In this study we use seismic reflection data and gravity modelling to investigate 1979 the geometry, structure, and trend of the Rovuma continental margin (Northern 1980 Mozambique) in order to decipher: 1) the margin type and its mode of formation (divergent 1981 vs strike-slip), and 2) which group of plate tectonic models are best able to predict the 1982 observed structures and margin trend. For the first time, we unambiguously identify a 1983 transform continental margin in seismic reflection data from the southern Rovuma Basin and 1984 support our observations with rigorous gravity modelling. The identified margin is not 1985 coincident with the DFZ, which is often assumed to form the continent-ocean transform 1986 margin (COTM) of the WSB.



**Figure 3.1.** Map of the Western Somali Basin (WSB) showing the location of gravity profiles (black lines) studied in the Rovuma Basin (RB). Karoo and Jurassic aged sedimentary basins used to reconstruct the conjugate margins are shown. White dashed lines depict gravity lineaments related to the ocean spreading fabric of the basin (Phethean et al.,

2016). Large black arrows indicate the two main groups of reconstruction paths proposed for
Madagascar. DFZ, Davie Fracture Zone; LB, Lurio Belt; TCB, Tanzania Coastal Basin.

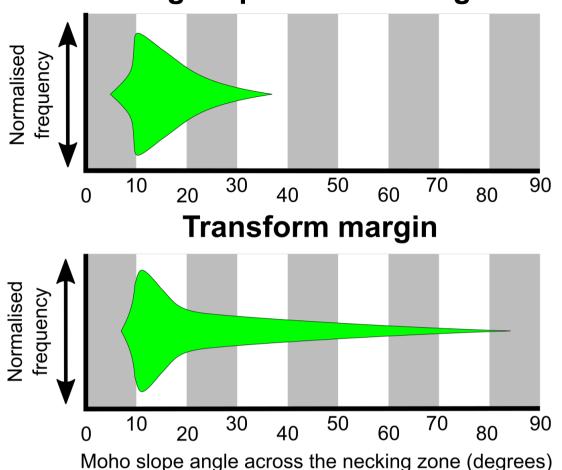
# 1995 **3.2. Characterisation of passive continental margins**

1996 The angle between the lithospheric deformation zone and the regional extension direction 1997 (previously discussed in Chapter 1.2) controls which of the two endmember types of passive 1998 margins, divergent and transform margins, develops. Where the deformation zone lies almost perpendicular (~ >  $80^{\circ}$ ) to the extension direction, divergent margins develop. 1999 2000 Decreasing the obliquity between the extension direction and the deformation zone to 2001 between  $\sim 60^{\circ}$  and  $\sim 80^{\circ}$  results in the development of divergent margin segments linked by 2002 transfer zones. A further reduction to between  $\sim 20^{\circ}$  and  $\sim 50^{\circ}$  leads to the development of 2003 alternating divergent and transform margin segments, possibly connected through horsetail 2004 splay faults. Below  $\sim 20^{\circ}$ , transform margins develop (Basile and Braun, 2016).

Harry et al., (2003) devised a method of differentiating between margin endmembers by comparison of crustal thickness variations across a margin to a compilation of well-studied examples. This compilation has since been updated by Mercier de Lépinay et al., (2016) and allows for a rapid identification of margin type and, therefore, mode of development.

2009 This method of margin interpretation can also be supplemented through the identification of 2010 features characteristic of each margin type to provide a more robust differentiation between 2011 divergent and transform type margins. We use the main characteristics of passive margin 2012 endmembers summarised in Chapter 1.2 for this purpose. Furthermore, average Moho slopes 2013 across the necking zones of magma-poor divergent margins and transform margins show 2014 differing distributions (Figure 3.2) and may therefore be used to differentiate between the 2015 two margin types. Whilst the bulk of Moho slopes for both transform and magma-poor 2016 divergent margins overlap, transform margins with slopes less than 7° are yet to be 2017 identified, as are magma-poor divergent margins with slopes greater than 36°. Average 2018 Moho slopes across necking zones that are lower or higher than these values therefore imply 2019 a divergent or transform margin, respectively. As the crustal necking zones of volcanic 2020 divergent margins are modified by magmatic additions, the slope of the Moho across them 2021 may not be used to identify the margin as divergent or transform in nature. These margins are, however, readily identifiable by other characteristic features in seismic reflection data 2022 2023 (e.g. Chapter 1.2) and do not require such Moho slope analysis for identification.

Magma-poor rifted margin



2025

Figure 3.2. Normalised frequency distribution of average Moho slope angles across the necking zones of magma-poor rifted margins and transform margins respectively. Data is derived from: Reston (2009) and Van Avendonk (2009) for magma-poor rifted margins; and, Todd et al., (1988), Todd and Reid (1989), Barton et al., (1990), Keen et al., (1990), Edwards et al., (1997), Sage et al., (2000), Greenroyd et al., (2008), and Parsiegla et al., (2009) for transform margins.

# 2032 3.3. Data and Methods

Seismic reflection data and 2D gravity modelling are used to investigate the Rovuma Basin margin geometry, structure, and trend in order to discern the margin type and therefore strain regime under which it formed. This allows for the comparison of predictions from different plate tectonic models with observations in order to determine which best accounts for the evolution of the WSB.

#### 2038 <u>3.3.1. Seismic reflection data and interpretation</u>

2039 Multichannel seismic reflection data from the ION East AfricaSPAN and VEMA (leg 3618) 2040 cruises are used for the geological interpretation of basement structures around the 2041 continental margin of the Rovuma Basin (Figure 3.1), and to constrain 2D gravity models.

Three lines from the ION East AfricaSPAN PSDM (pre-stack depth migrated) dataset cross the continental margin of the Rovuma Basin at near perpendicular angles (Figure 3.3a). These lines are used to identify features of the Rovuma margin that are characteristic of the different endmember margin types (as described in Chapter 1.2). This assists in the interpretation of the margin type, helps to delineate the COT, and identifies the lateral position of the bottom of the basement slope, which is compared with those predicted by gravity modelling.

A fourth line from the ION East AfricaSPAN PSDM dataset crosses the continental margin at a highly oblique angle and captures continental Moho reflections. Moho slope angles measured across the necking zone on this line are therefore corrected to their true dip. For this purpose, the strike of the margin is assumed to follow the onshore basement outcrop of the Rovuma Basin.

# 2054 <u>3.3.2. Gravity data</u>

2055 Three different gravity datasets are used to constrain the gravity anomaly for gravity 2056 modelling. These include the Trident Bouguer onshore, free-air offshore anomaly dataset 2057 (Trident BAFA) provided by Getech (Fairhead et al., 2009), the Sandwell and Smith V24 2058 free-air anomaly (SS24; Sandwell et al., 2014; Figure 3.3a), and the DTU15 free-air 2059 anomaly (Stenseng et al., 2015). Each dataset has its own strengths. For instance, the Trident 2060 BAFA dataset contains additional proprietary data onshore; SS24 has a 'rich' frequency 2061 content due to its use of residual slopes of sea surface height during gravity estimation 2062 (Pavlis et al., 2012); and DTU15 has been argued to have an increased accuracy in coastal 2063 regions due to its use of simply the residual sea surface heights in its estimation of gravity 2064 anomalies (Pavlis et al., 2012). Using the datasets together also allows a qualitative measure 2065 of uncertainty by comparing their variability and reduces the error produced by local 2066 deviations of any one dataset during modelling.

To reduce the errors associated with modelling a 3D interface in 2D, and to make all three gravity datasets BAFA and thus comparable, we perform Bouguer and terrain corrections on the SS24 and DTU15 datasets before 2D modelling. Those corrections were performed in GMT using the Okabe method, which is able to handle arbitrarily complex geometries. A 2071 resolution of 1 km was used for SRTM elevation data, and we assume a density for the 2072 upper crust of 2.67 g cm<sup>-3</sup> to be consistent with the Trident BAFA dataset.

2073 Following BAFA conversion, the three datasets can be directly compared. Onshore, both the 2074 DTU15 and SS24 datasets use the Earth Gravitational Model (EGM2008; Pavlis et al., 2012) 2075 solution, but where the additional proprietary data included in Trident result in significant 2076 local differences between this dataset and the others, we consider Trident more reliable. 2077 Offshore, the Trident dataset is produced primarily as a stacked solution of older versions of 2078 the SS and DTU anomalies, and so this dataset generally follows the trend of the DTU15 and 2079 SS24 anomalies. Where it does differ, it is deemed less reliable as it was produced before the 2080 addition of new CryoSat-2 and Jason-1 altimetry data. As, however, the DTU15 and SS24 2081 datasets are produced using different methodologies, differences between these datasets 2082 represent real uncertainty in the gravity anomaly. To mitigate this problem, gravity profiles 2083 for 2D modelling are located away from any significant disagreements so that differences 2084 are generally < 10 mGal. Where onshore sections of gravity profiles must cross regions of 2085 significant disagreement between the Trident and SS24/DTU15 datasets, we replace the SS 2086 and DTU onshore data with that of the more reliable Trident dataset. A spline covering 10 2087 km either side of the coastline is used to produce a smooth join between the datasets. A map 2088 showing the variations between datasets is provided in the supplementary material (Figure 2089 S3.1).

#### 2090 <u>3.3.3. Gravity modelling</u>

Three types of gravity modelling have been performed during this study in order to determine the margin type, the margin's trend, and the detailed crustal structure of the margin, respectively.

2094 Firstly, a rigorous parameter space search investigating the preferred slope angle and 2095 location of a simple ramp-style Moho geometry is performed for the three profiles where 2096 coincident seismic data crossing the margin is available to constraint the seabed and 2097 basement interfaces (Figure 3.3b). This provides an understanding of the Moho geometry 2098 unbiased by human interpretation and is used to generate profiles of crustal thickness 2099 variation across the margin. These profiles are then compared to the global margin crustal 2100 thickness variation compilations of Harry et al., (2003), and Mercier de Lépinay et al., 2101 (2016), providing an independent interpretation of the Rovuma Basin margin type.

2102 Secondly, the preferred Moho slope angles and locations derived for profiles 2 and 3 are 2103 used to produce a simplified average geometry of the margin. We then use this simplified 2104 geometry to find the best fitting margin location along gravity profiles 4-8, which have no seismic constraints. Profile 1, south of the Lurio belt, is used as an independent test profile for this method. These best fit locations constrain the margin trend and, in combination with the interpretation of the margin type, allow us to determine which group of plate tectonic models best predicts observations from the Rovuma Margin.

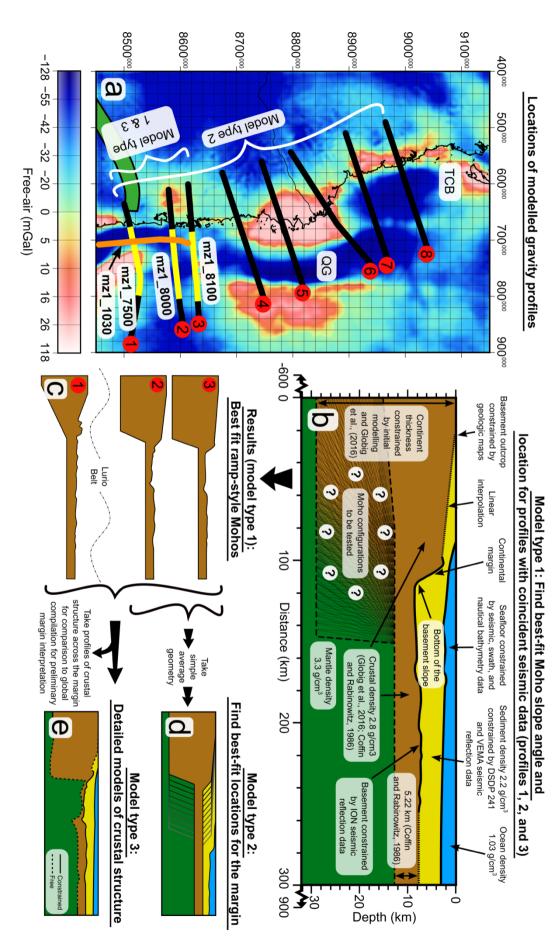
Following the first-order determination of the margin type, a final procedure involves detailed 2D forward gravity modelling of the three gravity profiles with coincident seismic data, using the commercial Interpex IX2D-GM software. This modelling allows for the addition of geologically plausible refinements to reduce misfit between the modelled gravity profile and data, where required. The resulting detailed crustal structure of the margin is then used to identify features characteristic of the different margin types, providing additional constraints for the interpretation of the margin.

# 2116 **3.3.3.1. Modelling type 1: Systematic investigation of Moho geometry**

2117 Gravity profiles with coincident seismic data crossing the margin's basement slope (Profiles 2118 1,2, and 3; Figure 3.3a) were used to investigate the underlying Moho geometry using the 2119 accurate seismic constraints on the remaining density interfaces. Simple ramp-style Moho 2120 geometries were tested for slopes between 5 and 85 degrees over a 65 km distance around 2121 the bottom of the basement slope as identified from seismic data. An initial search at 5 2122 degree and 5 km intervals covers the plausible parameter space with over 250 Moho 2123 geometry combinations (Figure 3.3b). This is then followed by a refined search around the 2124 initial parameter minimum at 1 degree and 1 km intervals. The root-mean-square fit between 2125 the calculated gravity anomalies and the three gravity datasets used in this study is then 2126 determined.

The 2D forward modelling was performed in Matlab using the analytical expression for the gravity effect of line elements with horizontal and vertical dimensions of 100 m and 20 m respectively. The model domain was extended 600 km beyond the extent of the gravity profile to mitigate boundary effects, whilst linear smoothing of density contrasts was performed laterally for 500 m on either side of a density interface to reduce artefacts resulting from gravity calculations of the rectangular grid elements across extremely shallow interfaces, such as outcropping basement.

The densities and geometries of the sediment and crustal layers used for all types of gravity modelling are constrained using a variety of Industry and academic data, as well as information from the literature. These are described below.



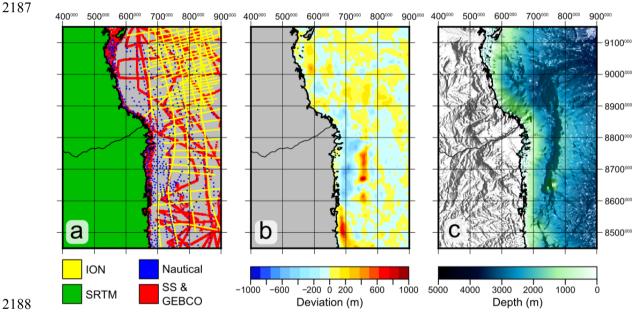
2139 Figure 3.3. Locations of seismic and gravity profiles, and summary of gravity modelling 2140 methods (a) Sandwell and Smith free-air gravity anomaly map showing modelled gravity 2141 profiles, which are indicated by black lines with corresponding identification number in red 2142 circles. Coincident seismic profiles that cross the continental margin are marked in yellow. 2143 Seismic profile mz1 1030 (orange) runs highly obliquely to the margin trend. The Lurio 2144 Belt is marked by the green band. QG, Quirimbas Graben; TCB, Tanzania Coastal Basin. (b) 2145 Summary of type 1 modelling – Moho geometry investigations along seismically 2146 constrained profiles. Constrained density interfaces are shown as solid lines, assumed 2147 interfaces are shown as short dashed lines, and the region of Moho configurations to be 2148 tested around the margin is marked as long dashed lines. (c) The results from type 1 2149 modelling on profiles 2 and 3 (situated to the north of the Lurio Belt) are used to generate an 2150 average margin geometry for use in type 2 models, with profile 1 used to test the method. 2151 Results from profiles 1, 2, and 3 are used as starting models for type 3 modelling. (d) 2152 Summary of type 2 modelling. The average margin geometry is modelled across the margin 2153 for all profiles (1-8) to find the best fit locations and determine the margin trend. (e) 2154 Summary of type 3 modelling. Seismically constrained lines are modelled in detail using 2155 Interpex IX2D-GM software. The Moho and unconstrained regions of the basement are 2156 allowed to move freely. Crustal thicknesses of 29 km and 5.22 km are used as continental 2157 and ocean boundary conditions, respectively.

# 2158 **3.3.3.1.1.** Bathymetry surface

2159 The seabed is a shallow interface with a high density contrast across it. Because of this, it is 2160 one of the major contributors to the gravity anomaly, and it must be well constrained to 2161 perform accurate gravity modelling. The ION seismic reflection lines provide an extensive 2162 and high quality seabed constraint for the region. However, as they sometimes do not cross 2163 the shelf edge, they are unable to fully constrain the geometry of this major shallow 2164 structure. We therefore supplement the ION seabed constraints with available ship-based 2165 singlebeam and multibeam bathymetry data extracted from the GEBCO and Sandwell and 2166 Smith global elevation datasets, as well as additional depth soundings taken from nautical 2167 charts for the region, which significantly increase the seabed constraints around the shelf 2168 break (Figure 3.4a). In places these additional constraints deviate from global elevation 2169 datasets by more than 500 m and represent a significantly more accurate constraint on the 2170 seabed for gravity modelling (Figure 3.4b).

As ship-based measurements are generally quite widely spaced, they can provide good constraints on the long wavelength information of the seabed, but provide little information on short wavelengths between data points. Following surfacing of the combined dataset (performed using Generic Mapping Tools (GMT; Wessel et al., 2013) with a tension factor 2175 of 0.25), we therefore apply a 15 km low-pass Gaussian filter and replace the short 2176 wavelengths with those of the Sandwell and Smith gravity-derived seabed dataset, which has 2177 full spatial coverage.

2178 The seabed interface will thus inherently account for the vast majority of the very short 2179 wavelength (< 15 km) content of the gravity anomaly, allowing us to produce realistic 2180 models with a full range of wavelengths without affecting the modelling of the basement and 2181 Moho structures of interest, which inherently have longer wavelength gravity signatures due 2182 to upward continuation. A 15-km filter is used as it provides the best trade-off between 2183 increased continuity of structural seabed trends derived from the SS data and maximised 2184 accuracy in regions of data constraint. The resulting seabed surface is shown in Figure 3.4c. 2185 Examples using alternative filter cut-off wavelengths are presented in the supplementary 2186 material (Figures S3.2-5).



2188

2189 Figure 3.4. (a) Input data for the new bathymetry surface. (b) Difference between the new 2190 bathymetry surface and global elevation models. Offsets of almost 1 km can be seen and 2191 represent a significant improvement in the predicting capabilities of the dataset. (c) The new 2192 bathymetry surface, illuminated from the east.

#### 2193 3.3.3.1.2. Sediment density

2194 We constrain the density of the sediment layer using seismic velocities derived during post 2195 stack migration processing of the multichannel seismic reflection line 84v, shot during the 2196 R/V Vema leg 3618. This line is situated in the north of the WSB, but is chosen due to the 2197 availability of coincident density measurements from the DSDP 241 borehole. The Nafe-2198 Drake relation provides a good fit of densities derived from velocity conversion to densities 2199 measured at DSDP 241, where converted densities range from 1.62 to 2.38 g cm<sup>-3</sup>, with a 2200 mean value of  $\sim$ 2.2 g cm<sup>-3</sup>. Further information and figures on the derivation of the sediment 2201 density are provided in the supplementary material (Figures S3.6-8).

#### 2202 3.3.3.1.3. Sediment thickness

2203 Offshore, where ION PSDM seismic reflection lines are coincident with gravity profiles, the 2204 top basement interface constrains the sediment thickness. Where possible, crosslines are 2205 used to check the interpretation of the top basement reflector. Where young volcanic edifices 2206 within the sediment layer obscure the basement interface, we linearly interpolate between 2207 basement picks.

2208 Onshore, the basement outcrop is constrained using regional geological maps. Between the 2209 onshore basement outcrop and the seismically constrained basement offshore, we linearly 2210 interpolate the basement horizon. This is likely a gross simplification of the true geometry of 2211 the basement interface; however, it avoids the introduction of unjustified complexities in the 2212 gravity modelling. Sensitivity to this simple linear geometry was also tested by modelling 2213 alternative imposed geometries, such as simple basins or rises, within the unconstrained 2214 region.

### 2215 **3.3.3.1.4.** Crustal density and thickness in the ocean and continent

In order to investigate the nature of the transition from continent to ocean, densities and thicknesses of the continental and oceanic crust on either side of the transition must first be constrained.

2219 Offshore, (Coffin et al., 1986) used sonobuoy experiments to determine the thickness and 2220 velocity structure of oceanic crust in the WSB. They showed that the crust has an average 2221 thickness of 5.22 km, which we use to constrain the depth of the Moho below top basement 2222 at the seaward edge of the coincident seismic data. The crust has an average layer 2 2223 thickness of 2.73 km with a velocity of 5.83 + -0.27 km/s, and a thin layer 3 averaging just 2224 2.62 km thick with a velocity of  $7.03 \pm 0.25$  km/s. Using the Nafe-Drake relationship, we 2225 convert these seismic velocities to densities for the upper and lower bounds of the velocity 2226 estimate. We then average the densities of upper and lower crust weighted by the thickness 2227 of each layer to calculate an average crustal density for the oceanic crust. Due to the dense 2228 layer 3 being thin compared to most estimates from normal oceanic crust, the average 2229 oceanic crustal density is also low at between 2.76 and 2.89 g/cm3.

Onshore, a recent study by Globig et al. (2016) incorporating a comprehensive dataset of
 Moho depths showed that the average crustal density of the African continent is 2.79 g/cm<sup>3</sup>.
 The continent-scale study also predicted crustal thicknesses along the Rovuma basin as low

as 30 km. To determine the final crustal thickness to be used, we initially perform modelling phase 1 using crustal thicknesses between 20 and 35 km, with testing at 1 km intervals for all profiles between 27 and 31 km, and selected the thickness consistently resulting in the lowest RMSD for use in the final models. The full results from this testing are presented in the supplementary material (Figure S3.9).

2238 As the average density of the continental crust falls within the range for the oceanic crust, 2239 we model both using a single density of  $2.8 \text{ g/cm}^3$ . This provides several advantages: Firstly, 2240 there is no imposed lateral density contrast between the ocean and continent. This allows us 2241 to avoid any circularity in determining the location of the COB from crustal thickness, which 2242 would be influenced by the chosen location of a lateral density contrast. Secondly, by not 2243 differentiating between upper and lower crusts, we avoid the introduction of an additional 2244 unconstrained interface into the model. A disadvantage of this method is that the basement-2245 sediment interface may be modelled using a slightly higher density contrast than might be 2246 present at these depths.

## 2247 **3.3.3.2. Model type 2: Determining the margin trend**

To determine the along-strike trend of the margin, we generate a simplified average margin geometry and model its best fit location along gravity profiles that cross the continental margin to the north of the Lurio Belt (Figure 3.3d). We use the Matlab code previously described in section 3.3.3.1, and the best fit location is again determined by calculating the RMSD between the modelled gravity anomaly and the three gravity datasets.

2253 The simplified average margin basement geometry is generated from seismic and geological 2254 map constraints of the basement interface along profiles 2 and 3, and is defined by 3 2255 inflection points along the basement interface (Figure 3.3c-d). From onshore seaward the 2256 inflection points are: 1) the onshore basement surface outcrop, where the basement interface 2257 begins to dip gently seaward; 2) the top of the basement slope, at a depth of 2.8 km, after 2258 which the basement interface dips steeply at 21°, until; 3) the bottom of the basement slope, 2259 at a depth of 7.8 km, after which the basement interface returns to horizontal. A ramp-style 2260 Moho geometry, determined from the type 1 modelling, defines the base of the crust and has 2261 a slope of 67°, the top of which is positioned 8 km landward and 5.22 km below the bottom 2262 of the basement slope.

2263 Only profiles 2 and 3 are used for the generation of the average margin geometry for two 2264 main reasons. Firstly, it allows us to use profile 1 to test the search algorithm's ability to 2265 correctly locate the base of the basement slope on a profile with coincident seismic data. 2266 Secondly, profile 1 is located to the south of the Lurio Belt, an ENE-WSW trending 2267 Neoproterozoic suture zone separating the northern and southern basement complexes of 2268 northern Mozambique (Emmel et al., 2011), across which basement structure differs (Franke 2269 et al., 2015). The geometry of profile 1 is therefore less likely to be representative of the 2270 margin to the north, making it a poor contribution to the average geometry of the margin. It 2271 does, however, therefore serve as a demanding test for the search algorithm allowing an 2272 assessment of the techniques robustness. We quantify the accuracy of type 2 modelling by 2273 measuring the offset between its predicted margin location and the Moho and basement 2274 slopes determined from seismic data and type 1 modelling. We also perform further testing 2275 of the search algorithm's sensitivity to the inputted average margin geometry, along with 2276 checks on the geological plausibility of required adjustments to type 2 model output in order 2277 to satisfy gravity data.

# 2278 3.3.3.3. Model type 3: Detailed 2D gravity models

Following systematic investigation of preferred Moho geometries using a simple ramp-style Moho slope, more complex 2D forward modelling is performed with full freedom of the Moho interface and, within its unconstrained region, top basement interface. This modelling is performed using Interpex's IX2d-GM software.

2283 This investigates the detailed crustal structure along the profiles, allowing for the 2284 identification of features characteristic of different margin types. Key features that could be 2285 identified by this modelling include: a) marginal ridges or complex rift structures within the 2286 region of unconstrained basement, characteristic of some transform margins and rifted 2287 margins respectively; b) complex Moho geometries, some of which may influence Moho 2288 slope angles determined across the necking zone from type 1 modelling; c) margin flexure, 2289 found at transform margins with mechanically coupled ocean and continental domains, and 2290 d; hyper-thinned continental crust and exhumed mantle found at magma-poor rifted margins, 2291 or exhumed mantle and thin oceanic crust found at transform margins and ocean fracture 2292 zones.

The same layer densities as for type 1 and 2 modelling are used, although we also test layered sediment densities and two layer crustal densities. The same crustal thicknesses as for type 1 and 2 modelling are also imposed at the ends of the gravity profiles as boundary conditions.

# 2297 **3.4. Results**

# 2298 <u>3.4.1. Seismic reflection observations</u>

The location of the four seismic reflection lines from the ION East AfticaSPAN dataset are shown in Figure 3.3. Three of the lines cross the continental margin perpendicularly. Results for these lines (from north to south: mz1\_8100, mz1\_8000, and mz1\_7500) are shown in Figure 3.5. A fourth line, mz1\_1030, runs perpendicular to the three previous lines and also crosses the continental margin at a highly oblique angle, and results for this line are shown in Figure 3.6.

# 2305 3.4.1.1. Mz1\_8100

2306 At the western edge of line mz1 8100, a massive 6 km increase in accommodation space 2307 occurs over an extremely short distance of  $\sim 17$  km (Figure 3.5a and d). The large increase in 2308 accommodation space indicates that the continental crust has been thinned and that this 2309 structure represents the continental margin. No faults are imaged within the steep 24° 2310 basement slope, which is instead onlapped by steeply dipping parallel sequences of post rift 2311 sediments. Immediately at the base of the margin slope a crossline confirms the top 2312 basement pick, which becomes obscured slightly eastwards due to the St Lazare volcanic 2313 edifice. East of St Lazare and of the DFZ, which is obscured by the above volcanics, the top 2314 basement is again imaged and has a hummocky character commonly associated with oceanic 2315 crust. No continuous reflections define the Moho surface for any region of this line and no 2316 evidence for syn-rift volcanics is present, despite extensive post breakup volcanism.

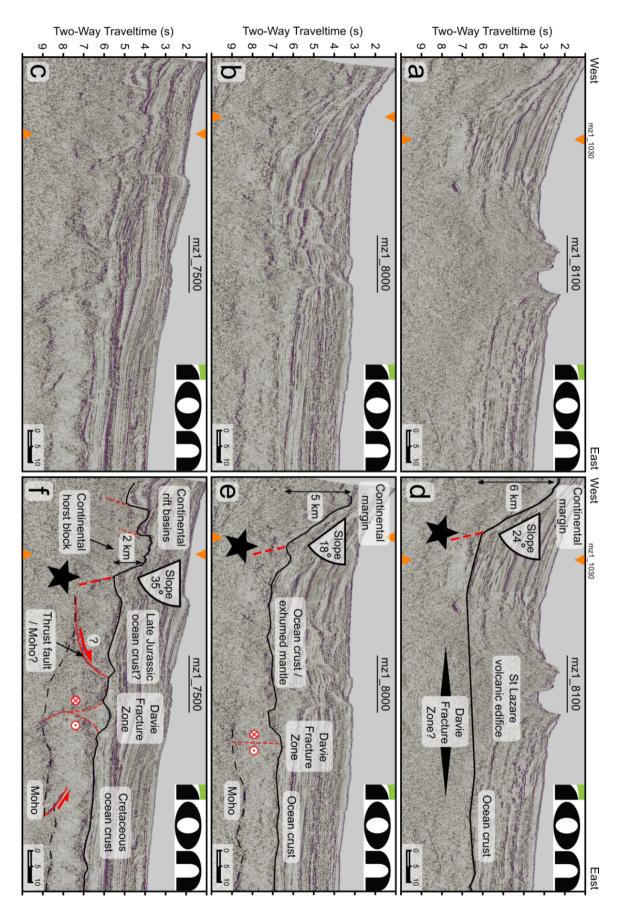
# 2317 3.4.1.2. Mz1\_8000

2318 Approximately 35 km to the south a similar setting is observed (Figure 3.5b and e). Again a 2319 steep basement slope, here 18°, defines a large 5 km increase in accommodation space and is 2320 void of any faults. The presence of limited post rift volcanics allows the top basement to be 2321 traced across the entirety of this line. To the west of the DFZ this surface lies at  $\sim$ 7.3 s 2322 TWTT, and shows a rough and hummocky character. No Moho reflections are visible 2323 anywhere to the west of the DFZ, but recent faulting of the overlying sediments may 2324 partially obscure the deep structure of some of this region due to the disruption of seismic 2325 energy by such faults before the basement is imaged and by tectonic overprinting of the original basement structure. The DFZ itself is defined by a rise in the top basement, and 2326 2327 possible weak Moho reflections define a crustal thickness of ~2.4 s TWTT. To the east of 2328 the DFZ, the top basement is well defined and lies at ~6.8 s TWTT, 0.5 s above that to the west, and is defined by a hummocky character. In the vicinity of the DFZ, Moho reflections 2329 2330 have a sudden onset and clearly define an average crustal thickness of  $\sim 1.8$  s TWTT until the 2331 end of the profile.

Aligning the base of the basement slopes (black stars on Figure 3.5) of these two similar margin profiles (mz1\_8100 and mz1\_8000) defines a slightly NNW-SSE trend of the margin, at 172°.

#### 2335 **3.4.1.3. Mz1\_7500**

2336 Moving 70 km farther southwards, and crossing the offshore extension of the Lurio Belt, a 2337 change in the margin architecture occurs. On the west, rift grabens filled with divergent 2338 wedges of synrift sediment are bounded by continentward-dipping normal faults, possibly 2339 forming the northern continuation of the Angoche Basin (e.g. Mahanjane, 2014; Figure 3.5c 2340 and f). These half grabens sit at a depth of approximately 4 s TWTT, substantially above the 2341 basement to the east, which lies at between 5 and 7 s TWTT. This is consistent with a 2342 thicker crust beneath this region providing isostatic support, and with a continental origin of 2343 this basement. This faulted crust is bounded seaward by a large horst block, sitting up to 1.5 2344 km above the half grabens. A large (2 km), and extremely steep, basement slope then dips 2345 35° towards the east. This structure lies along the continuation of the 172° trend of the 2346 margin seen on the northern lines, and is steeply onlapped by sediments lacking a divergent 2347 nature. The similar steep slope, sedimentation history, and the along-strike alignment of this 2348 structure with the basement slope seen farther north suggests a genetic relationship between 2349 them. Seaward of this slope, the top basement can be seen to gain a smooth character, 2350 possibly indicating the presence of oceanic crust. This smooth character continues eastwards 2351 of the DFZ. The DFZ at this location is expressed by a positive flower structure, indicating a 2352 compressional component to the strike-slip tectonics along it. Moho reflections beneath the 2353 DFZ, which are stronger on the eastern flank, define a crustal thickness of over 3 s TWTT, 2354 attesting to the transpressional nature of the DFZ in this location. On the western flank of the DFZ, westward-dipping low angled reflectors can be seen to cut across the crust and are 2355 2356 associated with a slight shallowing of the basement and counter-clockwise rotation of the 2357 upper crust. These structures are associated with crustal thickening around the DFZ and may 2358 represent thrust faults that developed in response to the compression. A similar structure, not 2359 cutting the entire crustal section, also offsets the Moho to the east of the DFZ and is again 2360 associated with a thickening of the crust. In this region, strong Moho reflections define a crustal thickness of generally between 2.2 and 1.7 s TWTT, and a more seismically 2361 2362 transparent nature of the crust is seen to the east of the DFZ, indicative of oceanic crust.



2365 Figure 3.5. ION seismic reflection profiles crossing the continental margin are shown 2366 unmarked (a-c) and interpreted (d-f). Line intersections with the perpendicular mz1-1030 2367 profile are marked by orange triangles. For the location of the seismic lines, see Figure 3.3a. 2368 (d-f) The basement interfaces are marked as solid black lines and highlight the steep 2369 continental margin. Basement slopes across the margin are labelled, as is the depth-2370 converted vertical offset from top to bottom. Possible Moho reflections are indicated by 2371 dashed black lines. Positions of the base of the margin slope as interpreted from seismic are 2372 shown as black stars.

# 2373 3.4.1.4. Mz1\_1030

Running perpendicular to the above seismic lines, and at a highly oblique angle to the trend of the identified basement slope, line mz1\_1030 also crosses the continental margin (Figure 3.6.). A deep set of strong reflections, with a southerly apparently dip, lie well below the top basement and are interpreted as the Moho. At the southern edge of the line these reflections lie at a depth of ~30 km. Moving northwards, the reflections rise consistently shallower to where they outcrop at the basement surface just north of line mz1\_8000. To the north of this outcrop, exhumed mantle may therefore form the basement surface.

2381 The rise of the Moho reflections northward indicates that the transition from continent to 2382 ocean occurs in this direction. In map view, the Moho reflections are oriented at 176 (i.e. 2383 almost due N-S, see Figure 3.3a). To transition from continent to ocean in a northerly 2384 direction, the margin must trend more NW-SE than the seismic line (i.e. <176°). Onshore, the basement outcrop along the southern Rovuma Basin has a trend of 172°, the same as the 2385 2386 basement trend of the slope identified from margin-perpendicular seismic lines. Therefore, 2387 assuming a margin trend of 172°, we can correct the apparent dip of the Moho slope from 2388 line mz1 1030 to its true dip. This shows that it dips extremely steeply at  $\sim 74^{\circ}$  towards the 2389 continent.

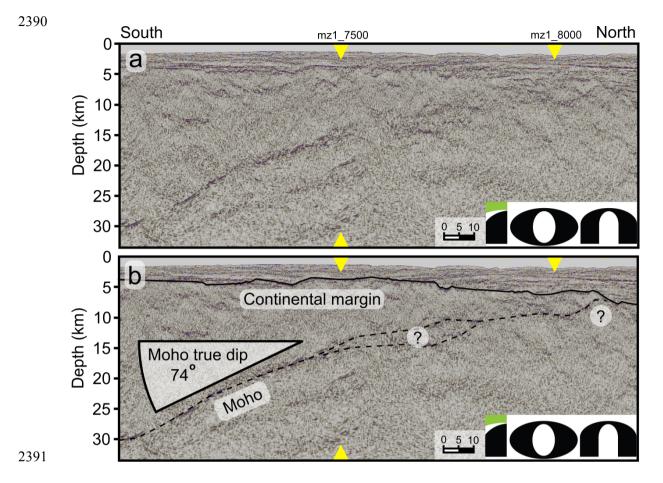


Figure 3.6. Depth-converted ION seismic reflection profile mz1\_1030, crossing the margin at a highly oblique angle. Basement and Moho interfaces are marked as for Figure 3.5 and intersections with perpendicular seismic lines are marked by yellow triangles. For location, see Figure 3.3a. The true dip of the Moho, after correction for the obliquity of the profile to the margin trend, is 74°. The downward dip of the Moho reflections to the south indicates that the margin must trend in a more NW-SE direction than the seismic profile.

# 2398 <u>3.4.2. Gravity</u>

# 2399 3.4.2.1. Model type 1: Best fit ramp-style Moho geometries across the margin

2400 Using constraints on the remaining density interfaces, gravity profiles with coincident 2401 seismic data crossing the margin are used to systematically determine best fit Moho 2402 geometries across the margin. An initial set of models is used to test a range of continental 2403 crustal thicknesses and consistently shows minimum RMSD fits to gravity data for 29 km 2404 thick continental crust. Visual best fits of the landward trend of the calculated gravity 2405 anomaly to data also consistently use a crustal thickness of 29 km. This value is in close 2406 agreement with Globig et al., (2016) and all final models therefore use a 29 km thick 2407 continental crust.

Model sensitivity to the geometry of the basement interface within the unconstrained region, between the basement outcrop and the onset of seismic constraint (e.g. Figure 3.3b), is also tested by the introduction of basins and ridges. Whilst the inclusion of basins sometimes results in lower RMSDs, it has little impact on the best fit Moho geometry. The introduction of complexities in this region is therefore unnecessary to determine best fits for the Moho, and is thus avoided.

All three profiles (mz1\_8100, mz1\_8000, and mz1\_7500) show a similar pattern of results with a strong preference for steep Moho slope angles (15-85°) and a relatively narrow band (~15-30 km) of acceptable slope locations (Figure 3.7). The arcuate pattern of low RMSD fits reflects the natural trade-off between these two parameters in achieving a good model fit to the gravity data.

The northernmost profile with coincident seismic data, profile 3 (mz1\_8100), is well constrained by type 1 modelling. Only Moho slopes greater than 25° result in low RMSD fits and the top of the Moho slope location is also constrained to within a 15 km band, centred on 107.5 km from the profile's western edge. Minimum RMSD values are achieved using an extremely steep Moho slope of 69° and a top of Moho slope location 105 km from the start of the profile. The top of the best-fit Moho slope is located 9.5 km to the west of the bottom of the basement slope identified in seismic data.

Profile 2 (mz1\_8000) shows similarly well constrained Moho interface. Again, only Moho slopes greater than 25° result in low RMSD fits and the top of the Moho slope location is also constrained to within 15 km, centred on 102.5 km from the profile's western edge. Minimum RMSD values are again achieved using an extremely steep Moho slope of 65°, with a top of Moho slope location 99 km from the start of the profile. This places the top of the Moho slope 6.5 km to the west of the bottom of the basement slope identified in seismic data.

Profile 1 (mz1\_7500) is the least well constrained of the three profiles by type 1 modelling and yet low RMSD fits are still only achieved using Moho slopes of 15° or greater. A 30 km band of low RMSD fits is achieved using Moho slope locations centres around 95 km from the profile's western edge. Minimum RMSD values are achieved for this profile using steep Moho slopes of 28°, with the top of the slope positioned at 95 km from the start of the profile. This places the top of the Moho slope 6 km to the east of the bottom of the basement slope identified in seismic data.

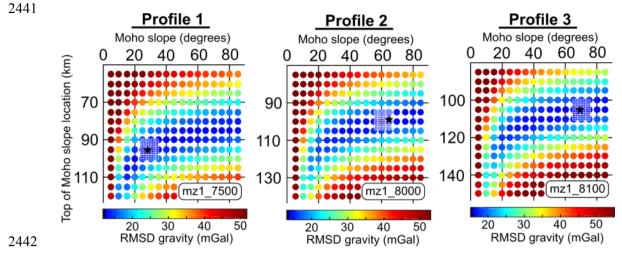


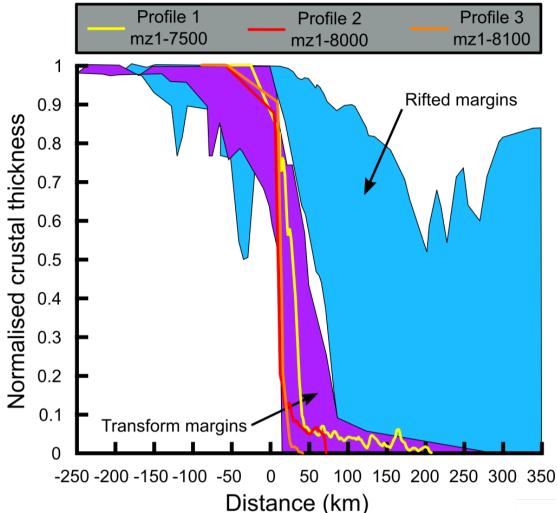


Figure 3.7. RMSD fits of Moho location and angle configurations for gravity profiles 1-3 as determined from type 1 modelling. All profiles show a similar trend, with the majority of the preferred parameter space covering steeper Moho slopes. Initial models at 5° and 5 km intervals are followed by models at 1° and 1 km intervals around the best-fitting parameter space determined during 5° and 5 km modelling. The minimum RMSD combination for each profile is marked by a black star.

# 24503.4.2.1.1. Comparison of crustal thickness profiles to a global compilation of2451margins

Using the best fit Moho parameters for the three lines we generate crustal thickness profiles across the margin for comparison to global compilations of margin profiles by Harry et al., (2003) and Mercier de Lépinay et al., (2016) (Figure 3.8). All profiles fall firmly within the transform margin domain, with profiles 2 and 3 lying along its most extreme edge. These profiles are incompatible with observed profiles from rifted margins, and thus support an initial interpretation of the Rovuma Margin as a transform margin.





2459

Figure 3.8. Normalised crustal thickness profiles across transform (purple) and rifted (blue)
margins after Harry et al., (2003) and Mercier de Lépinay et al., (2016). Normalised crustal
thickness profiles derived from type 1 modelling of gravity profiles 1, 2, and 3 are overlain.
All profiles across the Rovuma margin lie within the transform margin domain.

# 3.4.2.2. Model type 2: Best-fit locations for average margin geometries and margin trend

2466 We first test the search algorithm on profile 1, where it locates the bottom of the basement 2467 slope just 3.5 km seaward of its location in seismic data, despite the very different 2468 geometries of profile 1 and the average of profiles 2 and 3. The midpoint of the Moho slope 2469 is also in good agreement with that derived from type 1 modelling (Figure 3.9), validating 2470 the search algorithm's effectiveness. The resulting best fit gravity profile, however, differs 2471 substantially from the gravity data. This is due to complexities in the margin's real geometry 2472 not present in the tested average geometry. For the effectiveness of the model it is only 2473 required that the RMSDs of all tested locations are similarly affected by such complexities. 2474 Due to the simplicity of the input geometry this is to be expected and the only way to further

remove the misfit between modelled and observed gravity profiles is to change thegeometries in the gravity model (which will be the focus of modelling type 3).



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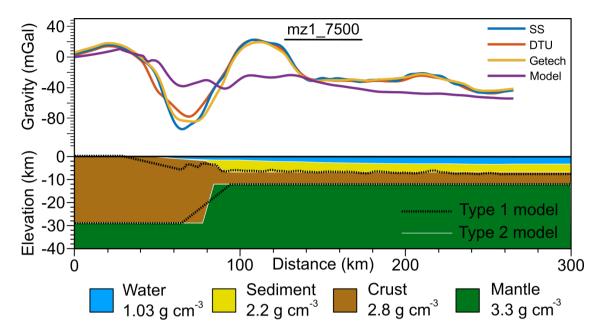


Figure 3.9. Calculated gravity profile (purple line) for the best fit location of the average margin geometry along profile 1 as determined from type 2 modelling. Despite the significant differences between the modelled average geometry and that determined from seismic data and type 1 modelling, there is a good agreement of the resulting margin locations. Misfits between the modelled gravity anomaly and data that result from simplifications in the modelled average geometry derived from the type-1 modelling, therefore, have not affected the model's ability to accurately locate the margin.

2486 Application of the methodology on all profiles consistently results in a margin location with 2487 a sharp, single RMSD minimum (Figure 3.10), and defines a NNW-SSE trend of the margin. 2488 There is no indication that the continental margin follows the DFZ, and north of profile 3 the 2489 margin trends slightly more NW (at  $\sim 160^{\circ}$ ) than farther south (172°). The margin generally 2490 follows the trend of the basement outcrop on the western edge of the Rovuma Basin along 2491 its length. Small deviations in best fit locations of the margin from the average trend, of up 2492 to 25 km, may reflect complications in the margin structure and resulting shifts in the 2493 location predicted by the model. They may also be due to real variations in the margin's 2494 trend, and we note that the most significant deviation, across the border of Mozambique and 2495 Tanzania, is in alignment with the major Sea Gap Fault, which runs ~NNE-SSW through the 2496 TCB and may be related. Nonetheless, reasonable RMSD fits of the margin (RMSD <= 2497 120% minimum RMSD) can be linked with a smooth curve through the margin, representing 2498 the margin's general trend.

Additional testing of the search algorithm's sensitivity to the input geometry is performed on all eight profiles. Steeper and shallower Moho slopes, of 87° and 47°, shift the predicted margin's location on average by 7.5 km landward and 8.5 km seaward, respectively. Seaward and landward shifts of the Moho slope location by 8 km relative to the basement slope shift the predicted margin's location by 16 km and 19 km, respectively. Whilst these variations are not insignificant, they are small compared to the length scale of the margin and do not affect the general interpretation of the margin's trend.

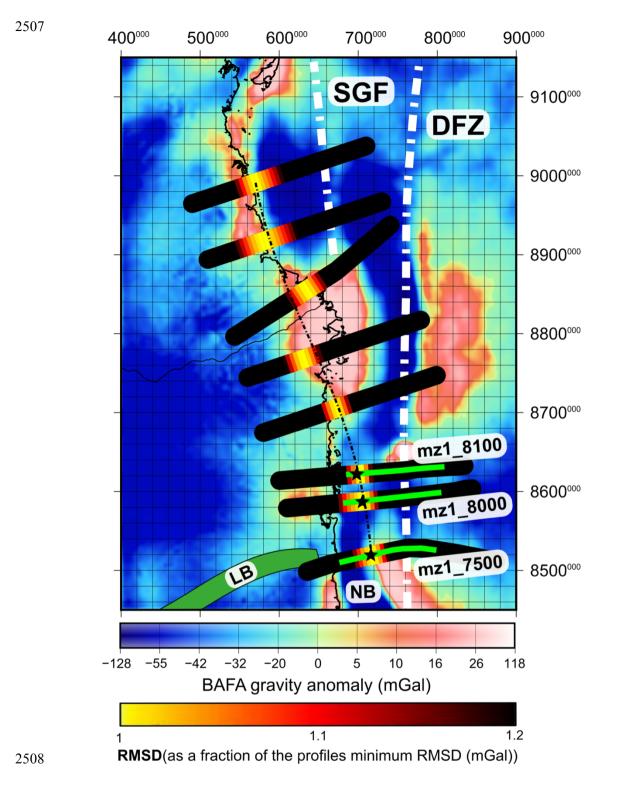


Figure 3.10. Sandwell and Smith free-air gravity anomaly map overlain with RMSD fits of the average margin geometry across the margin for profiles 1-8, as determined from type 2 modelling. RMSDs show a sharp convergence along each profile and define a NNW trend of the margin, marked by the dashed black line. The best fit locations are in close agreement with those determined from seismic data (black stars). The margin is significantly offset from the DFZ, which runs N-S farther to the east. Coincident seismic lines are shown in green. SGF; Sea Gap Fault.

# 2516 **3.4.2.3. Model type 3: Detailed 2D gravity modelling**

We perform detailed 2D gravity modelling of profiles 1, 2, and 3, which have coincident seismic data to assist in constraining the seabed and top basement interfaces, using the software IX2D-GM (Figure 3.11). Using a layered sedimentary package and/or a density contrast between the upper and lower crusts has little impact on the final result, so we avoid unnecessary complication and use constant densities for the sediments and crust in our final models.

- 2523 The resulting density models from all three profiles are in good agreement with the 2524 preliminary results of type 1 modelling and, in addition, reveal many more aspects of the 2525 margin structure that help to determine the margin type. Moho slope angles and locations 2526 follow the same pattern as seen from type 1 modelling, with extremely steep slopes of up to 82° and 83° for profiles 2 and 3 (mz1 8000 and mz1 8100), respectively, and a shallower 2527 Moho slope of up to 45° for the southernmost profile 1 (mz1 7500). The top of Moho slope 2528 2529 locations can also be seen to lie slightly to the west of the bottom of the basement slope for 2530 profiles 2 and 3, and to the east for profile 1, similar to the type 1 models.
- 2531 Profiles 2 and 3 again show a great similarity in their detailed structure. Both profiles require 2532 a deepening of the basement interface within the unconstrained region to the west of the 2533 basement slope, defining a marginal ridge structure. These profiles also show a downward 2534 flexure of the margin which increases seawards. This flexure is not only defined by the 2535 Moho architecture, but also by the down-flexed top basement interface. This basement is 2536 covered by an extensive (up to 50 km wide) thin sedimentary layer directly adjacent to the 2537 basement outcrop, which has been deposited within accommodation space generated by the 2538 flexure. On both of these profiles the basement slope is immediately bounded seaward by 2539 extremely thin crust, just 0-3 km thick, or exhumed mantle. This thin crustal layer extends 2540 seawards until the DFZ, although on profile 3 a slight thickening of the crustal layer to a 2541 consistent 4 km also occurs before the DFZ. This interpreted crustal thickening to 4 km is, 2542 however, beneath the St Lazare volcanic edifice, and may therefore be the influenced by 2543 intruded dense volcanic rocks above, which would require a low density crustal root to 2544 compensate them. Seaward of the DFZ, these two profiles show fairly consistent crustal 2545 thicknesses of ~6 km, likely representing oceanic crust of the WSB.

Profile 1 does not exhibit the same margin flexure as seen to the north. However, the faultbounded blocks of continental crust landward of the basement slope do define a basement high that could be related to the marginal ridge seen on profiles 1 and 2. Seaward of the basement slope, the crust is ~4 km thick, thicker than the 0-3 km of profiles 2 and 3, but still slightly thinner than the 5.2 km typical of the WSB (Coffin et al., 1986). On this

- southernmost profile, significant crustal thickening to ~11 km, with a width of ~ 20 km,
- 2552 defines the DFZ. A slightly thicker than normal oceanic crust also extends ~ 50 km seawards
- 2553 of the DFZ on this profile, before normal crustal thickness values for the WSB return.

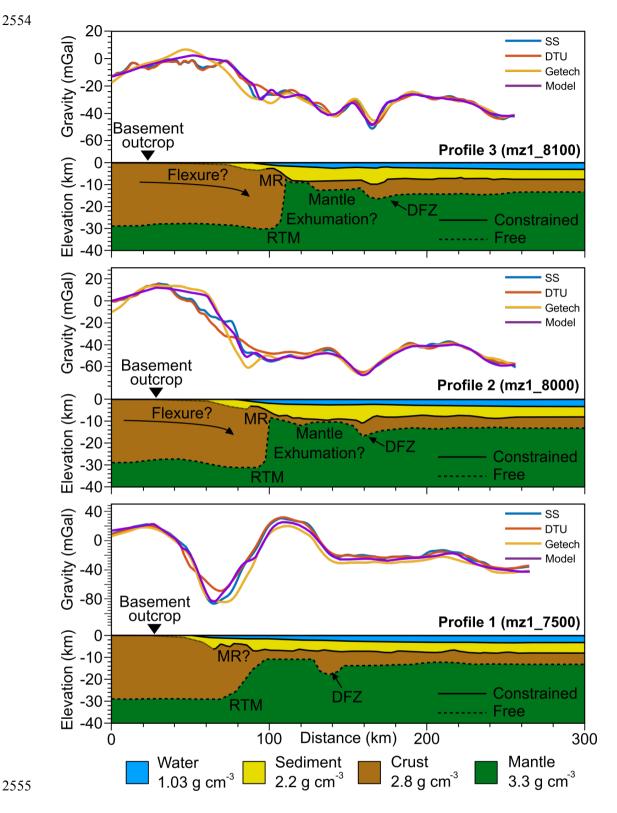


Figure 3.11. Detailed crustal structure as determined from type 3 modelling along profiles 1-3 with coincident seismic data. All profiles show possible marginal ridges adjacent to the margin. Profile 1 displays a large thickening of the crust around the DFZ. Profiles 2 and 3 indicate possible exhumation of the mantle immediately adjacent to the steep basement slope of the continental margin, which may have undergone downward flexure.

# **3.5. Discussion**

#### 2562 3.5.1 The continental margin of northern Mozambique and southern Tanzania

Following type 1 gravity modelling, we generated crustal thickness profiles across the continental margin for comparison with the global compilation of Mercier de Lépinay et al., (2016; Harry et al., 2003). All three profiles fall well within the transform regime, and provide a robust and unbiased preliminary interpretation of the margin as a transform margin. Following this preliminary interpretation, we use additional observations from seismic reflection data and gravity modelling as an independent check.

2569 Moho slopes across the necking zone identified in profiles 2 and 3 lie between 65-82°, with 2570 the more detailed type 3 modelling preferring the upper end of this estimate. These slope 2571 angles are steeper than any recorded from rifted margins across the globe, and comparably 2572 steep Moho slopes have, to date, only been inferred across the Côte d'Ivoire-Ghana 2573 transform margin, where they are thought to be sub-vertical (Sage et al., 2000). Along 2574 profile 1, type 1 modelling predicts Moho slope angles of an ambiguous 28°, which may be 2575 found across necking zones at both magma-poor divergent margins and transform margins. This simple ramp-style Moho geometry is, however, a simplification of true necking zone 2576 2577 geometries. More detailed type 3 modelling of this profile reveals Moho slope angles possibly reaching up to  $43^{\circ}$ , with an average Moho slope of  $32^{\circ}$  across the necking zone. 2578 2579 Such Moho slope angles are near to the upper limit of those found at divergent margins 2580 (Figure 3.2) yet lie well within those observed at transform margins worldwide. This supports the interpretation of this section of the margin as having strong transform affinities. 2581 2582 However, the large reduction in Moho slope angles along profile 1, compared to profiles 2 2583 and 3, may indicate a transition from transform to rifted margin at this location.

Steep basement slopes accommodating a large step in basement depth, without evidence for normal faulting, are common features of transform margins (e.g. Lorenzo and Wessel, 1997; Sage et al., 2000). These are identified in all three margin-perpendicular seismic lines and dip seaward at between 18° and 35°, similar to the equivalent slopes of the Northern and Newfoundland transform margins, respectively (Greenroyd et al., 2008; Keen et al., 1990).

Landward of the basement slopes identified in seismic data, marginal ridges, common features of transform margins (e.g. Bird, 2001), have also been identified. Along profile 1, this feature is represented by a continental horst block, as identified in seismic data (Figure 3.5f), rising ~1.5 km above continentward-dipping half grabens, which bound it to the west. Along profiles 2 and 3, the tectonic structure of the marginal ridge is unknown as identification is only possible through gravity modelling. It can, however, be seen to rise a similar 1-2 km above the basement to the west of the ridge. Seaward flexure of the margin,
seen along profiles 2 and 3, is also indicative of transform margins. It probably developed as
the result of mechanical coupling across the transform margin and thermal subsidence of the
oceanic domain following the cessation of transform tectonics (Mercier de Lépinay et al.,
2016).

2600 Seaward of the marginal ridge and basement slope, thin crust and exhumed mantle is 2601 predicted by detailed 2D gravity modelling (type 3 models) and is supported by observations 2602 from seismic data. Line mz1 1030 shows the Moho outcropping at the surface in the vicinity 2603 of the base of the basement slope between profiles 2 and 3, and exhumed mantle is also 2604 consistent with the lack of Moho reflections adjacent to the basement slope along both of 2605 these profiles. Whilst exhumed mantle may be present at both magma-poor rifted margins 2606 and oceanic fracture zones (e.g. Doré and Lundin, 2015; Tucholke et al., 1998), the 2607 extremely narrow margin width of <50 km, predominantly lacking in rift structures, is 2608 incompatible with other observations of magma-poor rifted margins worldwide, supporting 2609 continent-ocean transform tectonics.

These observations overwhelmingly support the interpretation of the continental margin of Northern Mozambique and Southern Tanzania as a transform margin. Type 2 modelling and observations from seismic data constrain the trend of this margin as  $\sim 172^{\circ}$  to the south of profile 3 and  $\sim 160^{\circ}$  to the north. The margin follows the onshore trend of the Rovuma Basin's basement outcrop, and the northern section of the transform margin runs onshore, where it is enveloped beneath the Rovuma Delta which has prograded into the oceanic domain, similarly to the Niger Delta (e.g. Dickson et al., 2016).

In light of the margin's trend, and its newly recognised status, we term this margin the Rovuma Transform Margin (RTM) so that it may be distinguished from the DFZ. The DFZ is an ocean-ocean fracture zone that lies to the east of the RTM and is often coincident with the Davie Ridge. In the past, the DFZ and Davie Ridge have together been inferred to form the continental margin of the WSB (e.g. Coffin and Rabinowitz, 1987; Gaina et al., 2013).

### 2622 <u>3.5.2 Changes in margin style across the Lurio Belt</u>

The RTM can be split into northern and southern sections, roughly bisected by the Lurio Belt, across which a change in margin geometry, trend, and adjacent ocean domain may occur. In the southern section, the Moho slope angle is much shallower, dipping at  $\sim 30^{\circ}$ compared to  $\sim 70^{\circ}$  farther north. The vertical offset across the margin's basement slope is also reduced, here only  $\sim 2$  km compared with  $\sim 5$  km farther north. Furthermore, continentward-dipping normal faults, forming syn-rift half grabens, also appear to the south of the Lurio Belt, and may form part of the Angoche Basin. The northward extent of thisbasin, however, is not delineated by seismic reflection data.

These observations suggest an increased extensional component may have been present during continental breakup south of the Lurio Belt, faulting and thinning this lithosphere prior to plate separation. This is consistent with the change in margin trend that occurs just to the north of the Lurio belt, in the vicinity of profiles 2 and 3. In this northern section of the margin, which trends at ~160°, the occurrence of dextral strike-slip tectonics would lead to highly oblique plate separation along the margin to the south of the Lurio Belt, which trends at 172°.

2638 Therefore, during plate separation, strike-slip dominated transtensional rifting may have 2639 occurred along the Angoche Basin to the south of the Lurio Belt. This highly oblique 2640 southern section of the margin was likely linked to the RTM through horsetail splay faults, 2641 requiring normal faults along the edge of West Gondwana to dip towards East Gondwana. 2642 Subsequent isolation of these faults from West Gondwana during the onset of seafloor 2643 spreading and mid-ocean ridge propagation (e.g. Basile, 2015) in the Mozambique Basin 2644 may provide an explanation for the presence of continentward-dipping normal faults inboard 2645 of the margin slope in seismic line mz1 7500. In this case, this region of rifted continent 2646 between the transform margin and unrifted continental crust may represent a marginal 2647 plateau (e.g. Mercier de Lépinay et al., 2016). It should not be ruled out, however, that the 2648 continentward-dipping normal faults may have developed in response to volcanic rifted 2649 margin formation (e.g. Geoffroy, 2005) to the south within the Mozambique Basin (e.g. 2650 Mueller and Jokat, 2017).

2651 Differences between the northern RTM (profiles 2 and 3) and the Angoche Basin (southern 2652 RTM; profile 1) also occur within the oceanic domain. The northern section of the RTM is 2653 immediately abutted by exhumed mantle or extremely thin crust, whereas to the south 2654 adjacent to the Angoche Basin, gravity modelling predicts crust ~4 km thick. The northern 2655 region also shows less crustal thickening in the vicinity of the DFZ, around  $\sim 0.5$  s TWTT 2656 thickening, compared to seaward of the Angoche Basin where a larger crustal thickening of 2657  $\sim 1$  s TWTT around the DFZ is accompanied by compressional thrusting of the crust on 2658 either side. It should be noted, however, that observations of compressional tectonic 2659 structures are less likely in the north due to the obscuring of basement structures by post 2660 breakup volcanics and recent faulting.

Three scenarios may have led to exhumed mantle along the northern margin, yet 4 km thick crust adjacent to the continental margin along line 7500. Firstly, mantle exhumation may be local to profiles 2 and 3. This may be the case if oceanic core complexes have formed along the transform margin, as is common along ocean-ocean fracture zones. The scale of the
exhumed mantle zone, ~50 km wide, is similar to core complexes found along fracture zones
today.

Secondly, mantle exhumation may be widespread along the northern portions of the transform margin. In this case, the onset of oceanic crustal production along the COTM may be directly linked with the change in margin trend around the Lurio Belt and corresponding thinning of the continental lithosphere which occurred during rifting to the south of this feature. In this case, the lithospheric thinning may have increased heatflow from the mantle along the Angoche Basin, reducing the amount of heat lost from the MOR as it passed along the COTM, thereby allowing greater mantle melting and production of a thin oceanic crust.

2674 Finally, it should be considered that the more intense manifestation of compressional 2675 tectonics to the south of the Lurio Belt may have acted to thicken any crust present. It may 2676 also have allowed the introduction of water into the mantle along thrust faults, resulting in 2677 its serpentinisation. The resulting reduction in mantle density at this location would thus 2678 result in the interpretation of a thicker crustal layer when using gravity methods. The top 2679 basement is, however, quite smooth and is not typical of exhumed mantle. Furthermore, the 2680 presence of a possible Moho reflection, although potentially also the manifestation of thrust 2681 faulting, also makes this final scenario is less likely.

### 2682 <u>3.5.3 Possible plate tectonic models</u>

2683 North of the Lurio Belt, the RTM appears to be void of extensional normal faults and forms 2684 a > 400 km transform margin separating the highly oblique Angoche Basin from the TCB. 2685 This large offset between the TCB and Angoche Basin suggests that these rift segments did 2686 not overlap at their ends, and instead a transform fault would have been necessary to link the 2687 two from the onset of rifting (e.g. Basile, 2015). We therefore assume that the strike of this 2688 transform fault formed parallel to the initial extension direction between East and West 2689 Gondwana, supporting the initial SSE spreading direction proposed in Chapter 2, and, to a 2690 lesser extent, the NW-SE spreading directions of Klimke et al., (2017) and Reeves et al., 2691 (2016). This continental transform fault would go on to form the Rovuma Transform Margin 2692 following the passing of the southern Morondava Basin (Madagascar) along it to the SSE.

This initial phase of SSE drift, followed by a switch to a more southerly drift of East Gondwana, is also reflected in the outcrop pattern of the volcanic margin of Southern Mozambique. Here, the Lebombo Monocline shows a similar change in trend, which may in part have developed in response to the change in spreading direction at ~150 Ma. Furthermore, offshore of the Mozambique volcanic margin, a possible small microcontinent, 2698 the Biera High, has been identified (e.g. Mueller and Jokat, 2017). Several instances of 2699 microcontinent release have been documented during changes in plate motion, resulting 2700 from the build-up of transpressional stress along long-offset fracture zones (e.g. Schiffer et 2701 al., In Press; Whittaker et al., 2016). It is therefore possible that the cleaving of the Biera 2702 High microcontinent resulted from compressional stress build-up across left-stepping 2703 fracture zones during this clockwise plate rotation near the end of the Jurassic. Recent 2704 evidence for continental crust beneath the Comoros islands, through identification of Pan 2705 African age zircons (533 Ma) within xenoliths of the Grande Comore (e.g. Roach et al., 2706 2017), points to the occurrence of a similar microcontient-cleaving event in the WSB, 2707 possibly driven by the same plate rotation.

The DFZ, which sits to the east of the Rovuma Transform Margin, also developed as the result of this plate rotation (e.g. Reeves et al., 2016; Chapter 2), and forms a major largeoffset ocean-ocean fracture zone. The eastward offset of this structure from the RTM, and the presence of oceanic crust landward of this fracture zone within the TCB (e.g. Chapter 2; Sauter et al., 2016), preclude it from forming the COTM of the WSB. Loose-fit plate tectonic reconstructions, utilising the DFZ without an initial phase of SSE spreading, are therefore unable to predict the observations of this study.

### 2715 <u>3.5.4 Transform margin development and impacts</u>

2716 The RTM also crosscuts the pre-existing tectonic fabric of Gondwana (e.g. Reeves and de 2717 Wit, 2000; Windley et al., 1994) and whilst some Karoo-aged sedimentary deposits have 2718 been postulated to be present below parts of the Rovuma Basin (e.g. Salman and Abdula, 2719 1995), these are not exposed at the surface as they are along the Kenya, Tanzania, and 2720 Madagascar margins. Late Triassic to Early Jurassic sediments, drilled within the Mandawa 2721 Basin (northernmost Rovuma Basin; Hudson and Nicholas, 2014), are thought to have been 2722 deposited in response to the Jurassic breakup of East and West Gondwana, and the oldest 2723 succession within isolated early rift grabens along the Rovuma Basin were found to be 2724 Jurassic in age (Smelror et al., 2008). This suggests that, whilst isolated Karoo basins may 2725 have been present beforehand, the development of the RTM was largely not controlled by 2726 pre-existing structures and instead formed as a new tectonic feature in direct response to the 2727 Jurassic breakup of Gondwana.

This new transform fault offset the volcanic Mozambique Basin (e.g. Leinweber et al., 2013) from the magma-poor Tanzania Coastal Basin (e.g. Chapter 2). Similar variations in magmatism between rift segments have been observed at other locations including: the Gulf of California (e.g. Lizarralde et al., 2007), the South Atlantic (e.g. Franke et al., 2007), and Central Afar (e.g. Stab et al., 2016). It is possible that a causal link between magmatism and 2733 margin segmentation exists. In the case of the Jurassic East Africa breakup, it may be 2734 postulated that the lack of lithospheric thinning along the RTM may have acted as a barrier 2735 to the lateral flow of anomalously hot mantle (postulated beneath the Mozambique basin due 2736 to the upwelling of the Bouvet mantle plume; e.g. Reeves, 2014) into the TCB, resulting in a 2737 lack of magmatism during the breakup of this basin.

2738 Following breakup and an initial phase of SSE spreading, the change in plate motion near 2739 the end of the Jurassic would have resulted in transpression along the RTM. It is possible 2740 that these transpressional tectonics resulted in the development of the marginal ridge along 2741 the RTM inferred from gravity modelling. Such transpressional uplift has been found to 2742 control the development of marginal ridges elsewhere, such as at the Ivory Coast-Ghana 2743 transform margin (Huguen et al., 2001). Recent studies have also shown that a previously 2744 unexpected amount of along-strike variation in transform margin geometry may be common 2745 (Mercier de Lépinay et al., 2016). The basement slope angles and vertical offset across the 2746 margin seen on profiles 2 and 3, change from 18° and 5 km to 24° and 6 km, respectively, 2747 over a distance of ~50 km. Misfits between modelled gravity profiles and gravity data during 2748 type 2 modelling, which uses an average margin geometry, also suggest that variations in the 2749 margin geometry continue along strike, supporting the observations of Mercier de Lépinay et 2750 al., (2016). Such variations likely arise from small deviations in the margin trend along 2751 strike, leading to increased transpression and transtension, or from inherited features from 2752 the continental crust (e.g. Mercier de Lépinay et al., 2016). The seaward flexure of the 2753 margin supports suggestions that mechanical coupling between the continental and oceanic 2754 domains occurred post-transform motions. At this point, thermal subsidence of the oceanic 2755 domain would have induced the observed downward flexure of the continent adjacent to the 2756 margin, as seen along other COTMs (e.g. Lorenzo and Wessel, 1997).

### 2757 **3.6. Conclusions**

2758 Seismic reflection data reveal a newly identified COTM in the southern Rovuma Basin. The 2759 margin lies landward of the DFZ, which has previously been interpreted as the COTM of the 2760 WSB. We term this newly identified COTM the 'Rovuma Transform Margin' (RTM), to 2761 distinguish it from the DFZ.

The presence of a marginal ridge: steep (18° to 35°) basement slopes with large vertical offsets of up to 6 km, lacking evidence for internal rift structures; rapid crustal thinning from ~28 km to 2 km over a distances of less than 20 km; seaward flexure of the margin; extremely steep Moho slopes across the necking zone of up to 83°; and presence of exhumed mantle and thin oceanic crust adjacent to the margin support its interpretation as a transform margin. Furthermore, comparison of crustal thickness profiles across the margin to the
global margin compilations of Mercier de Lépinay et al., (2016) confirms the transform
nature of this continental margin.

Inverse gravity modelling shows that the RTM runs NNW-SSE along the Rovuma Basin,
supporting an origin for Madagascar within the Tanzania Coastal Basin and tight-fit
reconstructions of Gondwana fragments.

A slight change in the trend of the margin from 160° in the north to 172° in the south occurs approximately at the Lurio Belt. The slight bend in the margin controlled a change from pure transform tectonics to the north of the Lurio belt, to the highly oblique opening of the Angoche Basin to the south. The oblique nature of the margin is reflected by shallower Moho slopes across the necking zone and the appearance of half grabens in continental crust.

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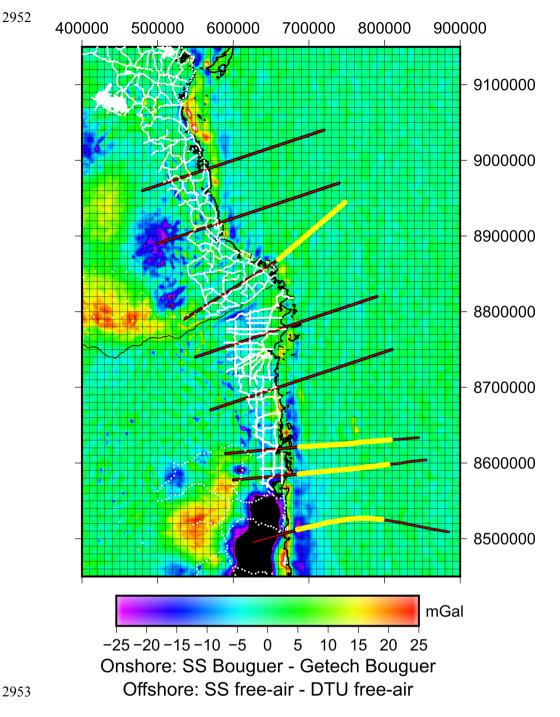
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### 2933 **3.8. Supplementary material**

### 2934 <u>3.8.1. Misfit between gravity datasets</u>

2935 Onshore, the difference between the bouguer anomaly generated from the World Gravity 2936 Map (WGM; used for the onshore SS and DTU anomalies) and Getech's Trident bouguer 2937 anomaly allows for an assessment of the data uncertainty (Figure S3.1). Gravity 2938 measurements included within the Getech Trident dataset are shown as white dots. Gravity 2939 profiles for 2D modelling are located to avoid areas of significant misfit between the 2940 datasets where possible. The largest disparity between the anomalies occurs onshore of the 2941 southern Rovuma Basin (500000-660000, 8450000-8620000). Here, the misfit exceeds 25 2942 mGal over a distance of +100 km. The extent of this misfit coincides with a proprietary 2943 gravity survey included only within the Trident dataset and is therefore likely caused by this 2944 additional data. Therefore, where profile 1 cannot avoid the region of large disagreement 2945 between datasets, Getech data is used to replace data from the SS and DTU bouguer 2946 anomalies

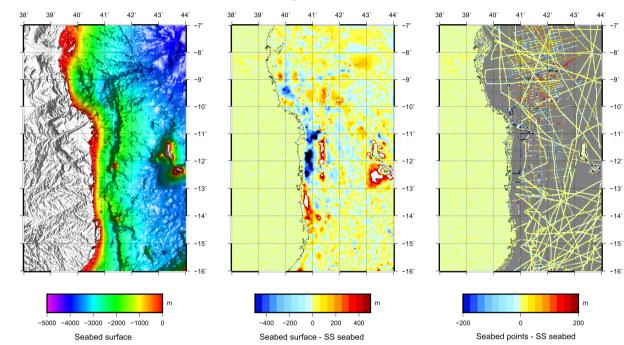
Offshore, comparison of the SS and DTU free-air anomalies is shown. Generally, misfits are concentrated in regions of shallow water along the coastline and are smaller than in the onshore regions. Nonetheless, where possible, profiles are located to avoid regions of larger disagreement.



**Figure S3.1.** Comparison of different gravity datasets

2955 <u>3.8.2. Examples of seabed surfaces generated using alternative filter cut-off</u>

### 2956 <u>wavelengths</u> 2957



# No wavelength replacement

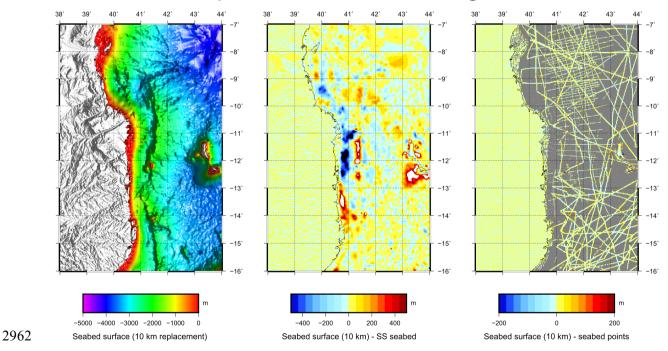
### 2959 Figure S3.2. Seabed surface without short wavelength replacement. N.B. The panel on the

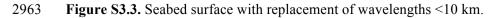
right of this figure is not comparable with those in Figures S3.3-5.

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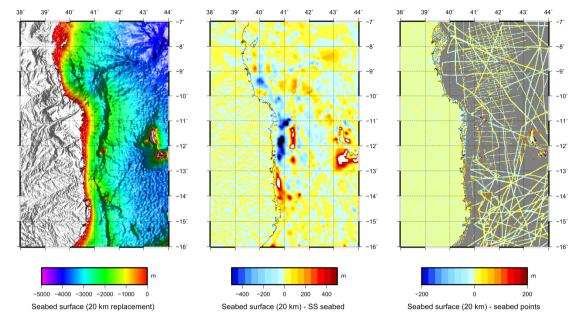
# Replace < 10 km wavelengths

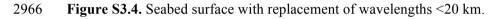


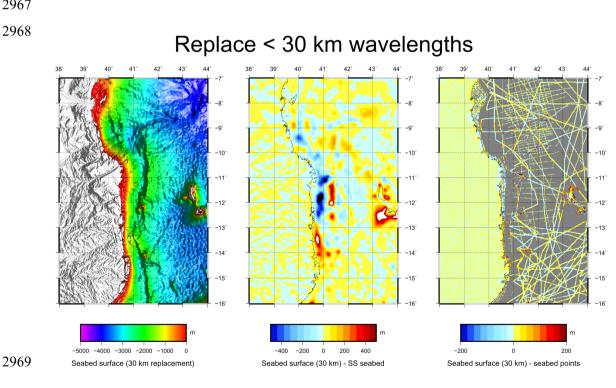


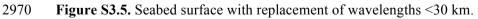


# Replace < 20 km wavelengths



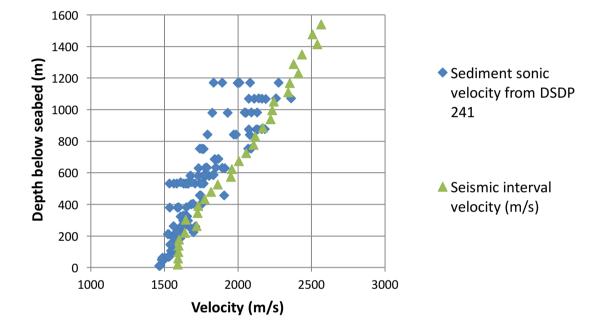






# 3.8.3. Deriving an average sediment density for the WSB for use in 2D gravity models To constrain the density of the sediment layer we use coincident data from the DSDP 241 borehole and the 84v multichannel seismic reflection line shot during the R/V Vema leg 3618, providing a chance to quality-check derived seismic velocities and densities. Seismic velocities derived during post stack migration processing of line 84v envelope the velocities measured at DSDP 241 (Figure S3.6).

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# VEMA 3618 and DSDP 241 velocities

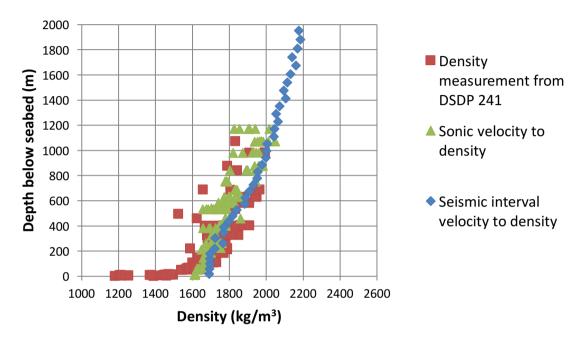
2980

2981 Figure S3.6. Measured and seismic-derived sediment velocities from DSDP 241.

We use the Nafe-Drake relationship between velocity and density to convert the seismic velocities (Figure S3.7) as it provides a superior fit to density data derived at DSDP 241 than Gardner's relationship.

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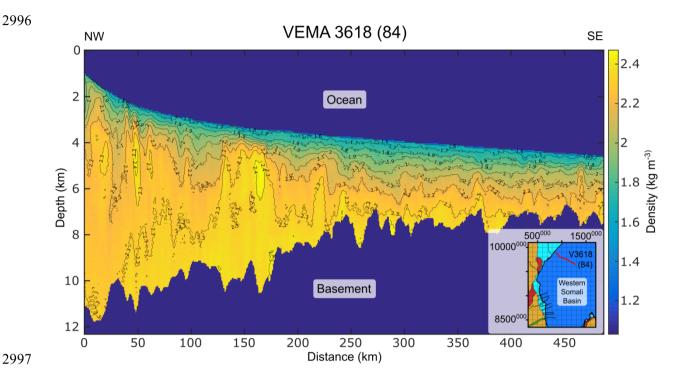
# VEMA 3618 and DSDP 241 densities

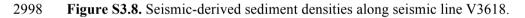
2988

2989 Figure S3.7. Measured and seismic-derived sediment densities at DSDP 241.

As the VEMA 3618 (84) line has similar sediment thicknesses to the ION PSDM lines used for gravity modelling, average sediment densities are calculated between the seabed and top basement using the entirety of the line (Figure S3.8). Densities range between 1.62 and 2.38 g cm<sup>-3</sup>, with a mean value of  $\sim$ 2.2 g cm<sup>-3</sup>.

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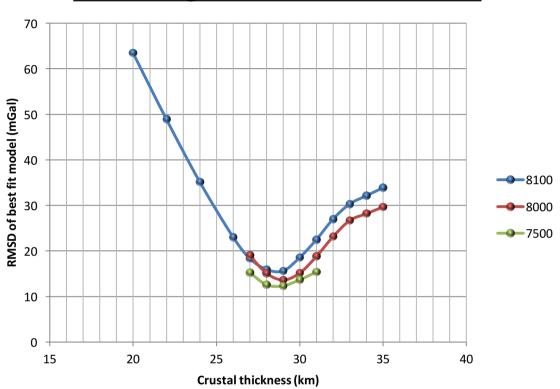




### 3000 <u>3.8.4. Determination of the thickness of continental crust</u>

To determine the crustal thickness for use in all final models, phase 1 modelling was initially performed using crustal thicknesses between 20 and 35 km. A thickness of 29 km consistently resulted in the lowest RMSD was used in all final models (Figure S3.9).

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### Determining the best fit crustal thickness

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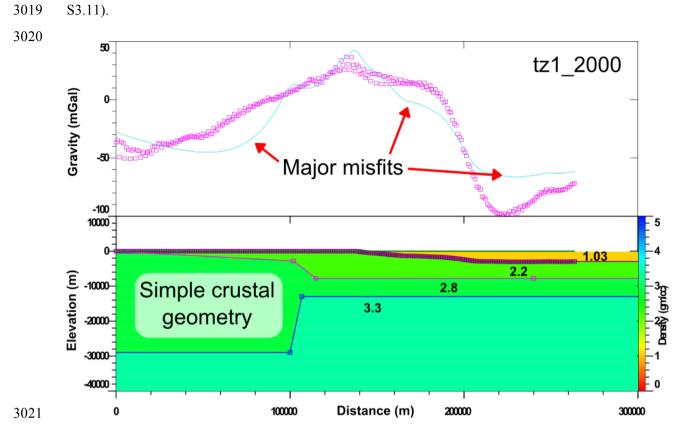
3006 Figure S3.9. Type 1 modelling results using different thicknesses for the continental crust.

### 3007 <u>3.8.5. Resolving gravity misfits of type 2 models</u>

3008 IX2D-GM software was used to model misfits along profile 6 (in addition to profiles 1, 2, 3009 and 3) as it also had coincident seismic constraint on the top basement interface along a 3010 portion of its length (tz1\_2000). Unfortunately, as the seismic line does not cross the 3011 continental margin, it was not suitable for type 1 or type 3 modelling.

In order to resolve the misfit inherent to models using the simple margin geometry, profiles 1, 2, and 3 commonly required modifications of the flat basement surface, introduction of a region of thin crust adjacent to the margin, and/or small adjustments to the slope angle of the Moho. No adjustment to the margin location as determined from type 2 modelling was necessary for any profile. Profile 6 (Figure S3.10) was no exception and only required simple adjustments of the basement interface to resolve the majority of misfit and

3018 modifications follow the basement interface where constrained by seismic data (Figure



3022 Figure S3.10. Best-fit result for profile 6 following Type 2 modelling.

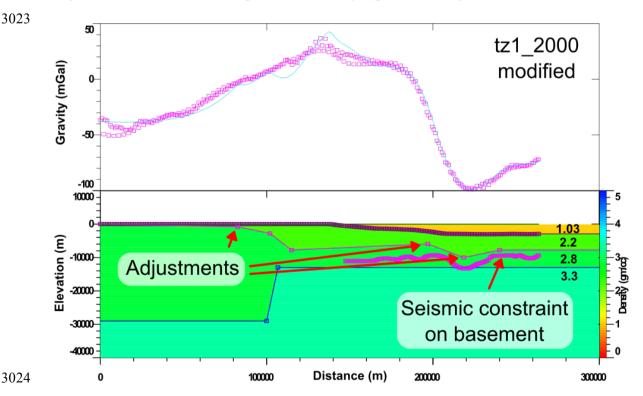


Figure S3.11. Small geologically plausible alterations are able to remove the majority ofmisfits and follow the trend of seismic constraints.

# 4. Compressional consequences of complex spreading: Formation of the Tanzania Coastal Basin and Davie Fracture Zone during the Mesozoic East Africa breakup

### 3032 Abstract

Changes in plate motion often lead to transpressional deformation along oceanic fracture zones, due to their incompatibility with the new spreading direction. Here, we present evidence for compression, which localised along the trend of fracture zones, within the Tanzania Coastal Basin (TCB) that was coincident with a change in spreading direction near the end of the Jurassic. This compression of the TCB probably led to the extinction of midocean ridge segments, development of the 250 km long Tanzania Coastal Basin thrust belt (TCBtb), and short wavelength (50 km) buckle folding within young (<3 Ma) oceanic crust.

3040 We argue that this change in plate motion was triggered by a coincident alignment of weak 3041 lithosphere, in the form of rifted margins and mid-ocean ridge segments, along the future 3042 trend of the Davie Fracture Zone (DFZ), and was thus not driven by changes in plate driving 3043 forces, but by a reduction in resisting forces along the strong Rovuma Transform Margin. 3044 The cessation of compression within the TCB, possibly related to the development of the 3045 DFZ, migrated from south to north, suggesting a similar northward propagation of the 2000 3046 km DFZ to join the spreading ridges of the Mozambique and Western Somali Basins. Since 3047 its formation, the DFZ was dominated by transpressional deformation, suggesting that the 3048 new plate motions were not fully aligned with plate driving forces. This supports the top-3049 down concept that the DFZ formed where and when plate configurations 'allowed' it, as 3050 opposed to where and when plate driving forces 'preferred' it, indicating a first order control 3051 of lithospheric strength on plate motions.

The subsequent drift of East Gondwana along this 2000 km fracture zone led to the collision of southern Madagascar with the oblique rifted margin of northeast Mozambique, forming the 'Davie Compression' and the 'Davie ridge'. The Davie Ridge has a similar geometry tothe TCBtb and may also represent a rotated oceanic thrust.

### 3056 4.1. Introduction

Changes in spreading direction have been attributed to both top-down (plate-derived; e.g. Anderson, 2001) and bottom-up (mantle flow-derived; e.g. King et al., 2002) driving mechanisms, and can result in transpression or transtension along existing fracture zones due to their incompatibility with the new spreading direction (Whittaker et al., 2016). Where strong, long offset transform boundaries exist between plates, plate motion may be restricted orthogonally to the transform due to the large force required to deform this lithosphere, and can lead to anisotropic coupling of the plates (Silver et al., 1998).

3064 Where transpressional forces along fracture zones exceed the lithospheric strength, 3065 compressional forces may fold (e.g. McAdoo and Sandwell, 1985; Müller and Smith, 1993), 3066 and thrust (e.g Briggs et al., 2009a; Jiménez-Munt et al., 2010; Massell et al., 2000), oceanic 3067 crust, generally along pre-existing fracture zones (e.g. Briggs et al., 2009b). The strength of 3068 oceanic lithosphere is strongly age dependent, increasing approximately linearly with age 3069 (Mammerickx and Sandwell, 1986), due to cooling and accretion of thicker mantle 3070 lithosphere with time. After 20 Myr, it may be equal in strength to normal continental 3071 lithosphere; however, before this time it may also be preferentially susceptible to 3072 deformation (Vauchez et al., 1998). McAdoo and Sandwell (1985) demonstrated a strong 3073 rheological control on folding of the oceanic lithosphere, with the shortest wavelength folds 3074 occurring in the youngest oceanic lithosphere. Where thrusting occurs, it may be thin or 3075 thick-skinned, with detachment occurring at the Moho or deep in the mantle, respectively. 3076 Generally, the direction of overthrusting has been observed to be age-dependent, where 3077 thrusting exploits pre-existing fracture zones with an age offset across them, with the 3078 younger plate overriding the older (e.g. Gorringe Ridge; Owen Ridge).

3079 If the new plate motion cannot be accommodated gradually by ocean spreading and the 3080 formation of progressively curved fracture zones, plate shearing may be favourable, and can 3081 lead to the calving of microplates and the formation of new spreading axes (e.g. Nunns, 3082 1983; Schiffer et al., In Press; Whittaker et al., 2016). Due to thinning of the crust, 3083 associated development of structural weaknesses and fluid permeation, mechanical and 3084 thermal thinning of the lithospheric mantle, increases in the geotherm, and thermal 3085 blanketing by post-rift sedimentary sequences (Cloetingh et al., 2008), young (<25 Ma) 3086 rifted margins are also inherently weak, and may preferentially succumb to deformation.

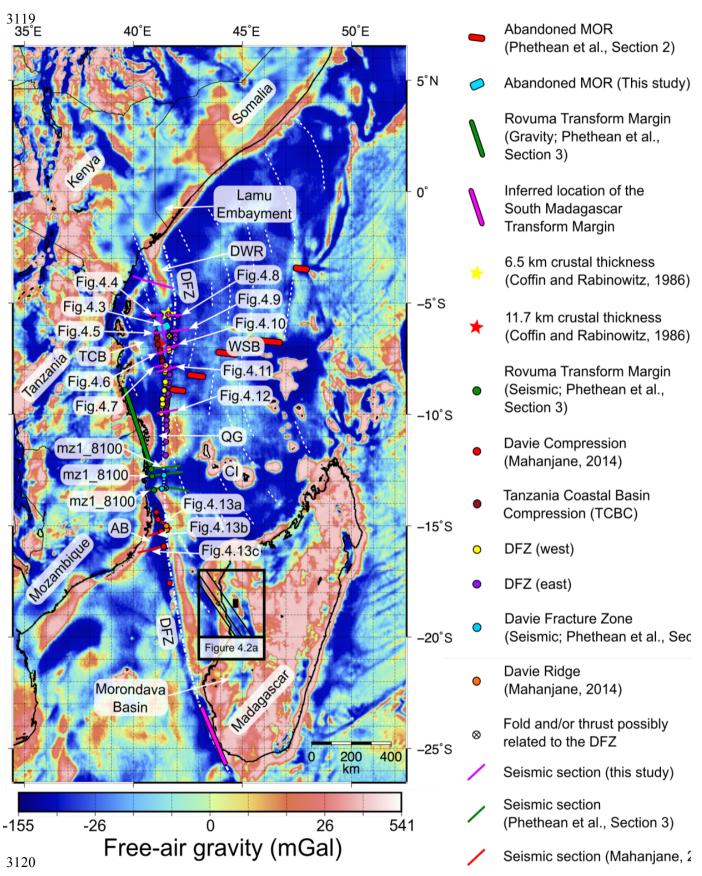
The Early to Middle-Jurassic breakup of East and West Gondwana was coincident with the impact of the Bouvet plume and eruption of the Karoo Large Igneous Provence in SE Africa. Following breakup, oceanic spreading resulted in the drift of Madagascar to the SSE and the formation of the TCB (Tanzania Coastal Basin; Figure 4.1) (e.g. Phethean et al., Section 2), before a switch to N-S spreading around the end of the Jurassic resulted in the southward drift of East Gondwana, carrying Madagascar to its present day position (Reeves et al., 2016).

Here, we present evidence for an episode of NE-SW compressional deformation within the Tanzania Coastal Basin (TCB), which we interpret as the result of plate motion changes near the end of the Jurassic. Our findings present a rare glimpse into thrust tectonics within oceanic crust and provide an opportunity to further study controls on plate motion changes. They also have large consequences for our understanding of the tectono-thermal history of the TCB and adjacent basins, with widespread implications for paleo-heatflow and trap development within the associated petroleum province.

### 3101 **4.2. Database**

3102 The principal data used in this study consist of a subset of the ION East AfricaSPAN deep 3103 imaging seismic reflection dataset, including data from the ke1, tz1, tz3, tz4, and mz1 surveys. The data cover an area of ~450,000 km<sup>2</sup> from offshore Kenya to offshore north-east 3104 Mozambique, including much of the TCB and DFZ. In light of our findings, we also 3105 3106 reinterpret seismic reflection data along the DFZ, immediately to the south of the East 3107 AfricaSPAN, which were presented in Mahanjane (2014). We also supplement our dataset 3108 with reprocessed seismic reflection data from the VEMA 3618 cruise, and two crustal 3109 thickness measurements from the TCB derived from sounobuoy data collected during this 3110 cruise. The locations of seismic sections and crustal thickness measurements, together with 3111 structural picks made from these data, are shown in Figure 4.1. Free-air gravity data, also 3112 shown in Figure 4.1 (SS24; Sandwell et al., 2014), are used for the identification of 3113 structural elements within the basin, and to assist in the determination of their strike and 3114 continuity where seismic data alone are insufficient.

3115 Satellite imagery from the northern Morondava Basin, Madagascar, was used for the 3116 mapping, and determination of the origin of, structural lineaments within this basin. The 3117 geological maps of Besairie (1964) are used to determine the ages of these structures.



3121 Figure 4.1. Free-air gravity anomaly map of East Africa highlighting locations of figures 3122 and seismic lines used in this study. AB, Angoche Basin; DFZ, Davie Fracture Zone ; 3123 DWR, Davie-Walu Ridge; OG, Ouirimbas Graben; CI, Comoros Islands; TCB, Tanzania 3124 Coastal Basin; WSB, Western Somali Basin. Inset shows the location of Figure 4.2a, and 3125 within this, the internal solid black box shows the location of Figures 4.2b and c. Figure 3126 numbers to seismic references: 4.3, tz3 1300; 4.4, tz1 4000 (west); 4.5, tz3 3600 (west); 3127 4.6, tz4 3300; 4.7, tz4 2950; 4.8, tz1 4000 (east); 4.9, tz3 3600 (east); 4.10, tz4 3350; 4.11, tz4 2850; 4.12, tz3 2101; 4.13a, Mahanjane 3a; 4.13b, Mahanjane 2c; 4.13c, 3128 3129 Mahanjane 2b.

### 3130 **4.3. Pre-breakup position of Madagascar**

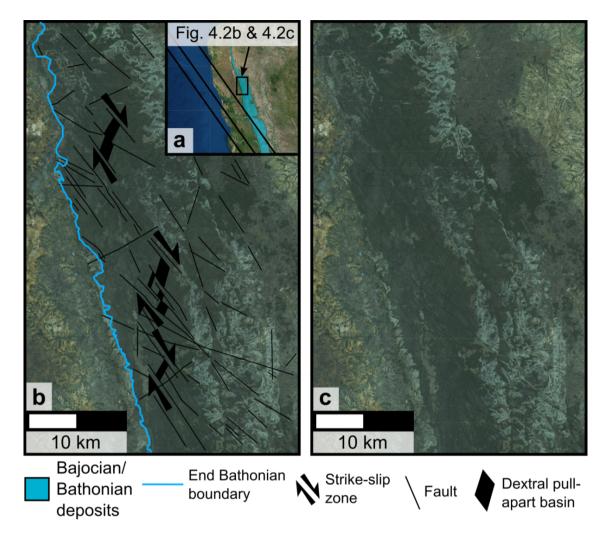
3131 The initial configuration of Gondwana fragments prior to supercontinent disassembly has 3132 been the source of much debate for over 30 years, and Madagascar forms a key piece in this 3133 puzzle. An accurate determination of the origin of this continental fragment has large 3134 repercussions for the fit of the surrounding continents due to its central position and the 3135 presence of several major intercontinental shear zones commonly used to align conjugate 3136 margins. Suggestions for the origin of Madagascar generally fall into one of two groups, and 3137 may be described as 'tight-fit' or 'loose-fit' reconstructions. Loose fit reconstructions 3138 assume that the N-S trending DFZ represents the continent-ocean transform margin, where 3139 to the west of this feature no oceanic spreading has occurred. These models use the DFZ to 3140 guide Madagascar back to a northern position seaward of the TCB. Tight-fit reconstructions, 3141 on the other hand, allow for an initial NNW-SSE phase of spreading before the formation of 3142 the DFZ and subsequent N-S drift. This results in an initial position of Madagascar within 3143 the TCB (Reeves et al., 2004). This latter model has recently been supported by the 3144 discovery of a transform margin along the Rovuma Basin and oceanic crust within the TCB, 3145 both inboard of the DFZ (Sauter et al., 2016; Chapter 2), which necessitates an initial phase 3146 of oceanic spreading to the west of the DFZ, and thus an origin for Madagascar within the 3147 TCB.

This initial phase of oceanic spreading within the TCB sent Madagascar in a SSE direction, constrained by the analysis of gravity lineaments related to spreading features (Phethean et al., 2016), and resulted in strike-slip tectonics between Southern Madagascar and Southern Tanzania/Northern Mozambique. This strike-slip motion led to the development of the Rovuma Transform Margin, the location of which has recently been accurately constrained through combined seismic reflection and gravity modelling studies (Chapter 3). This transform margin generally follows the trend of the onshore basement outcrop along the coastlines of the Southern TCB and the Rovuma Basin, and is located ~80 km seaward of
the basement outcrop along its length.

The location and trend of this transform margin further constrains the initial motion of 3157 3158 Madagascar relative to Africa, as the conjugate transform margin in southern Madagascar 3159 must have been in continuous contact along it during the active transform phase. 3160 Ascertaining the location and trend of the transform margin along southern Madagascar is, 3161 however, difficult due to tectonic overprinting and disruption of this margin during the 3162 formation of the DFZ at  $\sim$ 150 Ma (Reeves et al., 2016; Chapter 2). The present day location 3163 of the basement outcrop along southern Madagascar, however, is unlikely to have been 3164 affected by this tectonic overprinting. Therefore, we assume the transform margin of 3165 southern Madagascar also runs parallel to, and ~80 km seaward of, the basement outcrop as 3166 for the conjugate transform margin. This allows for an accurate reconstruction of 3167 Madagascar and Africa by aligning the two transform margins, which not only constrains the 3168 orientation of Madagascar relative to Africa, but also its south-westerly position. The north-3169 westerly position may then be derived by aligning the eastern edges of the once-adjoining 3170 Selous and Morondava basins to give the absolute location of Madagascar relative to Africa 3171 prior to breakup. Imposing these simple constraints on the position and orientation of 3172 southern Madagascar results in a good alignment of structural features in northern 3173 Madagascar with those of the conjugate East African margin.

Newly recognised strike-slip faults are preserved in the syn-breakup Bajocian-Bathonian 3174 3175 limestones of the northern Morondava Basin (the Karoo-recent aged basin spanning the 3176 entire west coast of Madagascar) and are highly oblique to the trend of this basin (Figure 3177 4.2). The strike-slip nature of these faults is indicated by their highly linear nature and 3178 interlinking pull apart basins, which are similar to others found within Karoo aged deposits 3179 of the Morondava Basin (Schandelmeier, 2004) and indicate a dextral sense of motion across 3180 the faults. These faults closely follow the trend of NNW-SSE coastline segments (Figure 3181 4.2a), and terminate at the End Bathonian boundary (blue line; Figure 4.2b), possibly 3182 indicating the localisation of strike-slip deformation onto oceanic transform faults at this 3183 time. Following the reconstruction of the southern Madagascar and Rovuma Transform 3184 Margins, these strike-slip faults are closely aligned with the SSE trend of the Davie-Walu 3185 ridge (DWR), a prominent gravity high just offshore the conjugate Lamu Embayment 3186 (Figure 4.1). They define the same SSE initial plate motion constrained by the Rovuma 3187 Transform Margin, and together, these features provide an additional robust constraint on 3188 the Mid- to Late-Jurassic relative plate motions of East and West Gondwana.

3189



3190

Figure 4.2. (a) Location of Figures 4.2b and c within the Morondava Basin. Black lines show the extrapolated trend of structures identified in Figures 4.2b and c. The extent of Bajocian/Bathonian deposits is taken from the geological maps of Besairie (1970). The location of Figure 4.2a is shown in Figure 4.1. (b) Landsat Satellite imagery of Bajocian-Bathonian syn-breakup deposits from the northern Morondava Basin, with fault lineament interpretation as black lines. (c) Figure 4.2b without fault interpretation. The location of the Morondava Basin is shown in Figures 4.1 and 2.1.

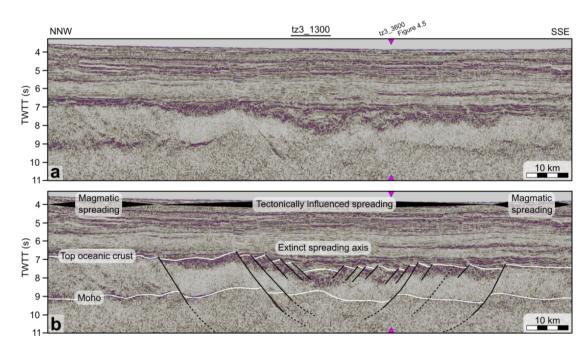
# 3198 4.4. Major tectonic signatures of the TCB, DFZ, and Western

### 3199 Somali Basin

The interpretation of seismic reflection data from the Western Somali Basin (WSB) and surrounding coastal basins of East Africa reveals evidence for several major tectonic events. Here, we assess the nature of these events and constrain them in relation to the plate tectonic framework of (Chapter 2).

### 3204 <u>4.4.1. Cessation of spreading in the TCB</u>

3205 Within the TCB, at 6.2°S 41.4°E, the SSE trending tz3 1300 line displays a change in 3206 character of the oceanic crust (Figure 4.3; location shown in Figure 4.1). The northernmost 3207 section of the line displays a smooth and continuous top basement reflector, which changes southwards to a rough and internally reflective layer. Coincident with this change, the mid-3208 3209 lower crustal layers become more seismically transparent, and the onset of extensional 3210 faulting of the basement, which generates tilted half grabens that dip symmetrically about a 3211 central graben, also occurs. Some of these normal faults offset the Moho by up to 0.5 s 3212 TWTT. In the northern and southern regions of Figure 4.3, Moho reflections define a crustal 3213 thickness of  $\sim 2.1$  s TWTT, and in places bound the base of the seismically transparent crust. 3214 Where Moho reflections are faint or not present, the base of the seismically transparent body 3215 is, therefore, taken as an indication of the Moho level. This and the high amplitude 3216 reflections of the top basement allow the definition of a roughly symmetrical thinning of the crust from  $\sim 2.1$  s at the edges of the central graben to  $\sim 1$  s TWTT at the graben's centre 3217 3218 over a distance of  $\sim 40$  km.



### 3220

3219

**Figure 4.3.** Extinct ocean spreading segment identified within the TCB. (a) Uninterpreted section. (b) Interpreted section. Extensional faults (black lines) sometimes offset the Moho and become increasingly dominant approaching the spreading centre. The crustal thickness decreases from  $\sim 2.1$  s TWTT outside of the axial rift, to  $\sim 1$  s TWTT at the rift centre. An associated change in the character of the top basement (top white line) occurs alongside crustal thinning. Moho reflections and the basal termination of the seismically transparent crust define the base of the oceanic crust (bottom white line). Location shown in Figure 4.1. 3228 Smooth oceanic crust is generated by magmatically dominated spreading, whereas rough 3229 oceanic crust results from an insufficient supply of melt to the spreading axis and the onset 3230 of tectonically accommodated plate separation (Louden et al., 1996; Smith, 2013). As 3231 melting at mid-ocean ridge (MOR) systems is primarily controlled by spreading rate, where 3232 slower spreading results in greater conductive cooling of the upwelling mantle and therefore 3233 less melt production, lower melt supplies generally reflect lower spreading rates (White et 3234 al., 1992). Mid to lower oceanic crust may display a transparent character in seismic 3235 reflection data (e.g. Bécel et al., 2015; Morris et al., 1993), and half grabens within oceanic 3236 crust have been shown to generally dip towards the spreading centre about which they form 3237 (Behn and Ito, 2008). The axial valleys at such spreading centres generally range in width from 16-62 km, averaging at 35 km (Malinverno, 1990), as measured along the Mid-Atlantic 3238 3239 Ridge.

3240 On line tz3 1300, the reduction in crustal thickness from  $\sim 2.1$  s to  $\sim 1$  s TWTT over similar 3241 distances to the width of axial spreading valleys and the symmetrical dip of half grabens 3242 about this thinning suggest that this structure represents an abandoned spreading centre 3243 within the TCB. Similar observations of crustal thinning and an associated broad zone of 3244 rotated fault blocks have been observed at the extinct Labrador spreading centre (Louden et 3245 al., 1996). Therefore, a reduction in spreading rate probably occurred in the lead up to the 3246 extinction of the spreading centre, resulting in greater cooling of the upwelling 3247 asthenosphere and less melt production (e.g. Niu and Hékinian, 1997). The resulting 3248 deepening of the brittle-ductile transition may also have allowed normal faults to extend into 3249 the upper mantle (e.g. Harper, 1985), offsetting the Moho.

3250 The abandoned spreading centre is situated  $\sim 210$  km away from the coastline in the paleo 3251 spreading direction; according to the plate tectonic model of (Chapter 2), this places the age 3252 of abandonment at  $\sim 150$  Ma, close to the end of the Jurassic. At roughly the same time, a 3253 major clockwise rotation of East Gondwana relative to West Gondwana, and the subsequent 3254 formation of the DFZ, occurred (Reeves et al., 2016). Subsequent oceanic spreading 3255 between 150 and 125 Ma, forming the Western Somali Basin, transposed the remaining 3256 MOR segments farther south offsetting the abandoned ridges of the Tanzania Coastal Basin 3257 and younger Western Somali Basin.

### 3258 <u>4.4.2. Compression post SSE spreading</u>

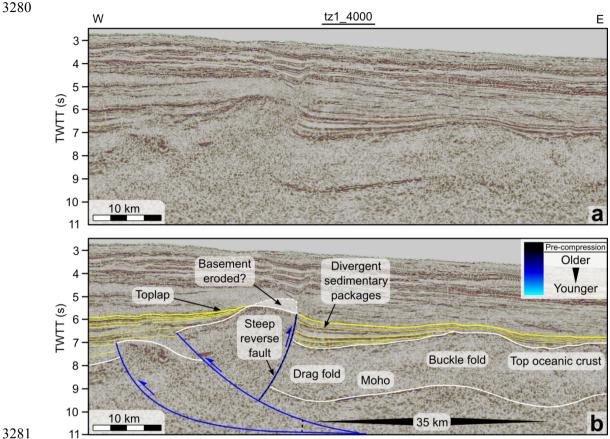
3259 Several E-W trending seismic lines within the TCB show evidence for crustal thickening and

3260 compression of the basement, and are described from north to south below.

### 3261 4.4.2.1. Line tz1\_4000 (west)

The central and eastern portions of Figure 4.4 display a smooth top basement and clear Moho reflection, defining oceanic crust with a uniform thickness of  $\sim 2.2$  s TWTT. This crust forms an open buckle fold, which has a trough to trough wavelength of  $\sim$ 35 km. The crest of the buckle fold sits at a depth of 6.6 s TWTT, and the troughs at 7.3 s TWTT. Overlying the buckle fold, sedimentary packages onlap, and then overlap, the folded crust, and diverge towards the troughs of the fold.

3268 At the western edge of the buckle fold, reflections from the sediments, top basement, and 3269 Moho, are bent sharply upwards before terminating abruptly in a steep west dipping line, 3270 perpendicular to the folded crust. To the west, an adjacent basement block with sporadic 3271 internal westward dipping reflections sits at a depth of 5.4 s TWTT. This block has been 3272 uplifted by at least 1.2 s TWTT relative to the surrounding crust, and the rotated internal 3273 reflections are truncated by an erosion surface along the top of the eastern flank that 3274 continues to form top of the divergent sedimentary fill to the east. Above the western side of 3275 this block, toplap of sedimentary packages against the same erosion surface also occurs, but 3276 diminishes westward (Figure 4.4). Between the west-dipping uplifted block, and west-3277 dipping upper crustal reflectivity at the far west of the line, an  $\sim 10$  km wide block of east-3278 dipping upper crustal reflectivity is present, and is delineated by offsets in the higher 3279 amplitude top basement reflections.



3282 Figure 4.4. Thrusting and folding of the basement along seismic line tz1 4000 (west) in 3283 response to compression. (a) Uninterpreted section. (b) Interpreted section. Yellow lines 3284 show the top (thick lines) and internal structure (thin lines) of syn-compressional sediments. 3285 Upper white lines indicate the top basement, where dashes indicate eroded sections, and the 3286 lower white line indicates the Moho. The interpreted relative ages of faults in this section are 3287 shown in the key. Location shown in Figure 4.1.

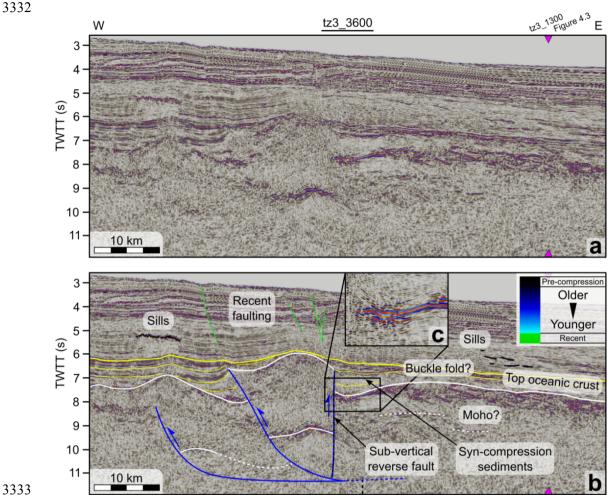
3288 The buckle fold in the eastern half of Figure 4.4b likely developed in response to 3289 compression, and the divergence of overlying sedimentary packages towards the troughs of 3290 the buckle suggests that they were deposited syn-compression. The onlap of sediments 3291 against the crest of the buckle suggests either 1) the presence of an original basement high, 3292 or 2) the onset of deposition after the onset of compression (e.g. Burbank and Vergés, 1994). 3293 Assuming a predominantly terrestrial sedimentary source, this onlap and the subsequent 3294 sedimentary overlap of the fold suggest that uplift rates of the fold crest were less than the 3295 rate of sedimentary deposition.

3296 To the west of the buckle fold, the uplift and rotation of the basement block has resulted in 3297 crustal thickening from 2.2 s up to 3.5 s TWTT where it overlies the edge of the buckle fold. 3298 This crustal thickening also strongly supports compressional tectonics. The abrupt 3299 termination of sedimentary, top basement, and Moho reflections against the western end of 3300 the buckle fold, and the upward bending of reflections immediately before their termination, 3301 suggest reverse faulting in this location (Figure 4.4b). The orthogonality of the reverse fault 3302 to the restored basement section indicates the reactivation of a sub-vertical structure during 3303 compression, and the folding associated with this faulting is therefore interpreted as drag 3304 folding as opposed to fault propagation folding associated with thrust fault initiation. 3305 Erosion of the tip of the uplifted block and toplap of sedimentary packages against this 3306 erosion surface (which may represent a depositional surface farther to the west) suggest that 3307 uplift of this block was faster than the rate of sedimentary deposition (Burbank and Vergés, 3308 1994; Hardy and Ford, 1997). Near the western edge of Figure 4.4, the rotation and offset of 3309 basement reflections may be the result of low angle thrust faulting during the compression, 3310 which allows for the uplift of the basement block along the steep reverse fault.

### 3311 4.4.2.2. Line tz3\_3600 (west)

3312 Similar crustal geometries to those along tz1-4000 (west) are seen 90 km to the SSE in Line 3313 tz3 3600 (west), and are shown in Figure 4.5. The eastern half of this figure crosses the 3314 heavily faulted extinct spreading centre of the TCB (i.e. Figure 4.3) at a highly oblique angle 3315 (see Figure 1 for line locations), possibly resulting in the sporadic reflective character of the 3316 top basement and lack of clear Moho reflections in this region. Identification of the 3317 geometry of the top basement, by enveloping top of basement reflections, reveals a possible 3318 folding of the crust, the top of which sits at a depth of 7.2 s TWTT. Above the western flank 3319 of the fold, the sedimentary sequence can be seen to thicken into the trough, although the 3320 internal structure of these sediments is poorly defined and does not allow for an analysis of 3321 their possibly divergent nature.

3322 Before the top basement reflections terminate abruptly to the west, they bend back upwards. 3323 Immediately west of this, a rotated and uplifted basement block sits at a depth of 6 s TWTT, 3324 and west-dipping top basement and upper crustal reflections define its western limb. Strong 3325 Moho reflections, which have been similarly tilted, define the base of the crust, and bend 3326 sharply downwards before their abrupt termination at the eastern edge of the block. Farther 3327 west, east-dipping top basement and upper crustal reflections define a second crustal block, 3328 along the eastern flank of which reflections bend back upwards, as can be seen particularly 3329 clearly in the overlying sedimentary packages (Figure 4.5). West of this block, close to the 3330 western edge of the figure, reflections from the upper crust and top basement are sub-3331 horizontal.



3333

3334 Figure 4.5. Thrusting and folding of the basement along seismic line tz3 3600 (west) in 3335 response to compression. (a) Uninterpreted section. (b) Interpreted section. (c) Close up of 3336 the inset region without interpretation showing the curved top basement reflections. Line 3337 descriptions are as for Figure 4.4. The interpreted relative ages of faults in this section are 3338 shown in the key and continuation of the structures at depth is speculative. Location shown 3339 in Figure 4.1.

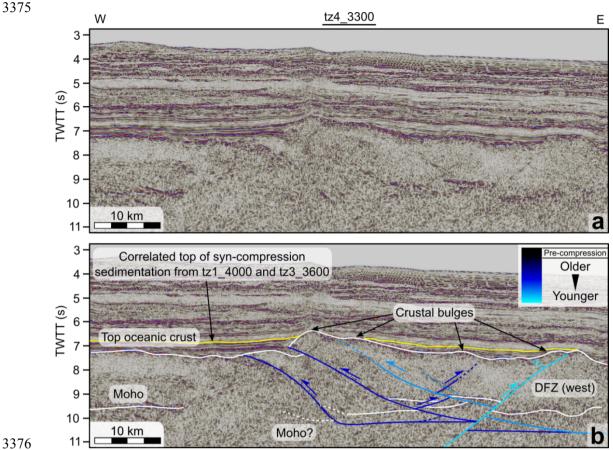
3340 The overall geometry of line tz3\_3600 (west) is similar to that seen along line tz1\_4000 3341 (west), despite the 90 km offset between the two lines, and may be interpreted in a similar 3342 way. Again, buckling and thrusting of the crust have occurred in response to compression, and in places the resulting crustal thickness may reach 4 s TWTT. In tz3\_3600, the sudden 3343 3344 termination of reflections against the eastern edge of the uplifted basement block defines 3345 another sub-vertical reverse fault, similar to the one seen in tz1 4000, again possibly 3346 reactivating a pre-existing structure. Drag folds have developed along this reverse fault and 3347 are defined by the top basement reflections to the east, and Moho reflections to the west. 3348 Similarly to line tz1 4000 (west), uplift along this reverse fault was probably facilitated by

the development of thrusts to the west, which have also rotated and uplifted the crust in thislocation.

### 3351 **4.4.2.3. Line tz4\_3300**

3352 Another basement high is imaged within the TCB another 100 km farther to the SSE. Here, 3353 smooth top basement reflections, underlain by a more transparent mid-crust and strong 3354 Moho reflections, are prevalent over the west and east portions of the section (Figure 4.6). 3355 They define a crustal thickness of between 1.7 s and 2.3 s TWTT within undeformed 3356 regions, generally sitting at a depth of around 7.3 s TWTT. The smooth top basement of the 3357 eastern portion of Figure 4.6 shows slight bulges on an approximately 10 km length scale, 3358 and the tips of bulges are commonly aligned with dipping reflections within the crust, and 3359 sometimes also offsets in the Moho. In this eastern portion of the line, a pair of reflective 3360 bands is sometimes present instead of a single band of Moho reflections. Both bands are 3361 offset in several locations, show a similar seismic waveform and amplitude, and do not 3362 always have a consistent vertical separation from one another.

3363 In the central region of the line, the top of the basement has been uplifted to a depth of  $\sim 6.4$ 3364 s TWTT, and in places the crust possibly thickens to 3.7 s TWTT. Below and to the east of 3365 the crest of this basement uplift, reflections within the crust dip gently eastwards, whereas below and to the west of the basement's crest, an  $\sim 5$  km wide block of seismically 3366 3367 transparent crust sits directly above a band of east-dipping high amplitude reflections. Strong Moho reflections are sparse in the central region, but a short high amplitude reflector dips 3368 3369 eastward  $\sim 3.5$  s below the crest of the thickened crust. Above the basement, the deep 3370 sedimentary section, correlating to the folded and divergent packages of the syn-3371 compressional sediments along lines tz1\_4000 (west) and tz3\_3600 (west), shows little sign 3372 of deformation, and follows the regional sedimentary tilt of the basin. On this line, unlike 3373 along lines tz1 4000 (west) and tz3 3600 (west), no buckle folding has occurred adjacent to 3374 the crustal thickening.



3377 Figure 4.6. Thrusting of the basement along seismic line tz4 3300 in response to 3378 compression. (a) Uninterpreted section. (b) Interpreted section. Dual reflector bands around 3379 the depth of the Moho result from Moho stacking in response to thrusting and backthrusting. 3380 Line descriptions are as for Figure 4.4, where the top of the syn-compressional sediments is 3381 now interpreted to correlate with the top of deformed sediments on lines tz1 4000 (west) 3382 and tz3 3600 (west). The interpreted relative ages of faults in this section are shown in the 3383 key and continuation of the structures at depth is speculative. Location shown in Figure 4.1.

3384 Along line tz4 3300, the clear Moho reflections on either side of the crustal thickening, the 3385 smooth character of the top basement, the undeformed crustal thickness of  $\sim 2$  s TWTT, and 3386 the seismically more transparent mid-crust indicate oceanic crust. As this oceanic crust 3387 surrounds, and therefore likely also comprises, the region of uplifted basement, tectonic 3388 thickening of the crust is likely the cause of the basement uplift. Beneath this, the eastward-3389 dipping reflections within the crust and the juxtaposition of crustal blocks, which sometimes 3390 have seismically transparent internal characteristics, suggest repeated stacking of the crust 3391 along eastward-dipping thrust faults (Figure 4.6b). The short, high amplitude reflector, seen below the crest of the crustal thickening, may therefore form a faulted contact between crust 3392 3393 and mantle rocks, where shearing of this boundary may have increased its reflectivity.

3394 To the east, bulges in the top basement, as well as east- and west-dipping reflections within 3395 the crust and offsets in the Moho (where offsets occur with the top to both the east and 3396 west), suggest further thrust and back-thrust development in this region. The complex 3397 geometry of the pair of Moho reflection bands in the east may therefore be the result of 3398 Moho stacking along such thrusts (Figure 4.6b). It should be considered, however, that the 3399 lower band of reflections may also have developed in response to early thrusting at this 3400 level, which was subsequently offset by later thrusts. This interpretation is presented in the 3401 supplementary material, but, without sufficient justification for the abandonment of an early 3402 detachment fault in favour of developing a new deeper one, this scenario is less attractive.

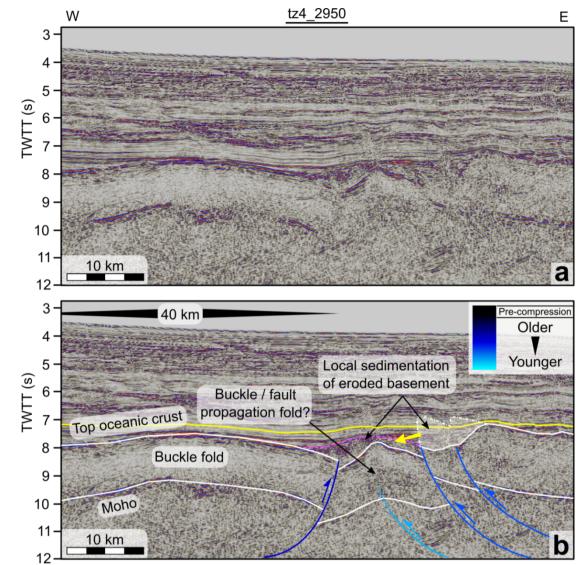
3403 Above the thrust stack, sediments correlating to the syn-compressional sedimentary 3404 packages seen farther north along lines tz1 4000 (west; Figure 4.4) and tz3 3600 (west; 3405 Figure 4.5) have been identified (yellow line, Figure 4.6b). These sediments onlap both the 3406 basement thrust stack and the DFZ (west), and show little sign of deformation beyond the 3407 regional background, following the same tilt as the overlying strata. This suggests that 3408 either, 1) the stratigraphic correlation of the sediments does not correspond to a temporal 3409 correlation, or 2) the sediments were deposited following thrusting and development of the 3410 DFZ, in which case the cessation of thrusting and formation of the DFZ must have occurred 3411 earlier farther south.

At the eastern edge of the line, a west-dipping thrust (cyan colour on Figure 4.6b) offsets the Moho by +0.5 s TWTT, and the lack of counter-clockwise rotation of Moho reflections within the hanging wall suggests that the thrust does not curve and sole out in the immediate vicinity. This fault may therefore cut the older west verging thrusts, in which case it would be younger. This fault tips out in the DFZ (west), where it contributes to crustal thickening.

### 3417 **4.4.2.4. Line tz4\_2950**

3418 Approximately 65 km farther to the SSE, just before intersecting the main trace of the DFZ, 3419 basement uplift and folding of the crust has again occurred, and is shown in Figure 4.7. 3420 Here, as to the north, smooth and high amplitude reflections define the top and bottom of 3421 oceanic crust, which also commonly displays a seismically transparent mid-crust. In the 3422 western part of the section, the crust forms an open buckle fold, with a trough to trough 3423 wavelength of ~40 km. The crest of the buckle fold sits at a depth of 7.6 s TWTT, and the 3424 troughs at  $\sim 8$  s TWTT. At the eastern edge of the buckle fold, the east-dipping smooth top 3425 basement reflections, low reflectivity mid-crust, and Moho reflections step down, and a 3426 sharp change to a westward dip of these reflections occurs. It should be noted, however, that 3427 below 10 s TWTT, patches of west-dipping reflections occur sporadically in several places 3428 on this line, and may interfere with the Moho interpretation. Less than 10 km to the east, the

- west-dipping top basement reflections are tightly folded to dip back eastwards, before
  terminating abruptly ~6 km on. The basement topography immediately above the tight fold
  has been partially filled by a layer of discontinuous, moderate to high amplitude reflections,
  which roughly follow the dip of the top basement.
- Along the eastern edge of the line, a block of uplifted basement sits at 7.2 s TWTT, and clear top basement and Moho reflections dip eastwards. Between this uplifted block, and the tightly folded crust to the west, additional east-dipping reflections are slightly offset below the trend of Moho reflections beneath the uplifted block. These reflections are bounded above by a region of reduced reflectivity, and below by chaotic reflections.
- 3438 Similarly to line tz4\_3300, sediments that correlate with deformed or divergent sedimentary 3439 packages farther north (along lines tz1\_4000 (west) and tz3\_3600 (west)), onlap the 3440 basement and do not form large divergent growth wedges as seen farther north. They have
- 3441 also not been subject to deformation beyond the regional background, and follow the same
- trend as the overlying strata.



### 3444

**Figure 4.7.** Folding and thrusting of the basement along seismic line tz4\_2950 in response to compression. (a) Uninterpreted section. (b) Interpreted section. Line descriptions are as for Figure 4.4, where the top of the syn-compressional sediments is now interpreted to correlate with the top of deformed sediments on lines tz1\_4000 (west) and tz3\_3600 (west), and the dashed pink line shows the top of locally sourced sediments deposited during compression. The interpreted relative ages of faults in this section are shown in the key. Location shown in Figure 4.1.

The termination and step down of top basement and Moho reflections at the eastern end of the buckle fold imply a faulted contact between the buckle fold and the tight fold to the east. The eastward step from the Moho terminations to the top basement terminations suggests this fault is west-dipping and, therefore, a reverse fault, consistent with other observations of basement compression. The sharp folding of oceanic crust immediately to the west is uncharacteristic of the buckle folding elsewhere, and may have required the influence of apropagating fault to localise the deformation.

3459 Immediately east of the fault-propagation fold, the presence of a lower reflectivity area 3460 above the east-dipping reflections is compatible with the interpretation of these reflections as 3461 the Moho, with overlying seismically transparent oceanic crust. The relative uplift of this 3462 basement during compression, with top to the east offset, implies the presence of east-3463 dipping thrusts, as interpreted elsewhere along the compressional structures of the TCB. To 3464 the east, Moho reflections of the uplifted block are offset slightly upwards. A second thrust 3465 fault, offsetting the Moho and further uplifting the eastern block, is therefore likely. Between 3466 the uplifted block and the fault-propagation fold, the shallow level of the Moho and 3467 relatively low top basement, therefore, imply a large amount of erosion of this crust during, 3468 and following, its uplift. It is possible that local deposition of the derived sediment occurred 3469 above the developing limbs of the fault-propagation fold, where continuous steepening of 3470 the fold limbs resulted in deformation and re-deposition of these sediments, resulting in their 3471 chaotic nature. The later deposition of onlapping and undeformed sediments, which correlate 3472 to syn-compressional sediments in the north of the basin, may again support the earlier 3473 cessation of compression in this more southern location, as compared to the north of the 3474 TCB.

#### 3475 **4.4.2.5. Summary of compressional deformation within the TCB**

Regions of highly tilted and uplifted basement associated with 1) coincident offsets in the Moho, 2) thickening of oceanic crust, and 3) in some cases, large adjacent drag folds provide strong evidence of thrust tectonics within the TCB. Furthermore, the development of buckle folds is compatible with these observations of compressional tectonics.

Observations of crustal thickening from Figures 4.4 - 4.7 suggest that the crust has been locally thickened from approximately 2 s TWTT to, from north to south, 3.5 s, 4.0 s, 3.7 s, and, before significant erosion, 3.3 s TWTT. This thickening is very similar to a crustal thickness measurement of 3.6 s TWTT (converted here to TWTT assuming a crustal velocity of 6.5 km/s) performed by Coffin et al., (1986), which sits between the observed thrusts on seismic lines tz3\_3600 (west) and tz4\_3300, and independently confirms the observations from seismic reflection data.

This similar level of crustal thickening across each of the thrust structures, with an apparent continuity between seismic observations; consistent west vergence of thrust structures; and alignment of all thrust structures along a SSE trending gravity anomaly that also follows other gravity lineaments within the TCB (Figure 4.1) suggest that the observed thrusting 3491 within the TCB may all form part of a single structure. This structure, which we term the Tanzania Coastal Basin thrust belt (TCBtb), runs from at least 5.5°S to 7.8°S along a SSE 3492 3493 trend, spanning over 250 km. This makes the TCBtb, to the best of our knowledge, the 3494 longest intraplate oceanic thrust complex known on the globe. Its SSE trend, similar to the 3495 Rovuma Transform Margin, spreading lineaments of the TCB, and strike-slip faults related 3496 to initial plate separation in Madagascar, in conjunction with evidence for the reactivation of 3497 sub-vertical faults during thrusting and termination of an extinct MOR segment against this 3498 structure suggests that the TCBtb developed along a pre-existing fracture zone in the oceanic 3499 lithosphere of the TCB.

3500 Buckle folds that have developed alongside the TCBtb have short wavelengths, between 35 3501 and 40 km. As the wavelength of buckle folds in oceanic lithosphere is strongly dependent 3502 on the age of the lithosphere, where younger lithosphere results in shorter fold wavelengths, 3503 the compression of the TCBtb must have occurred < 3 Myr after of the formation of the 3504 buckled crust (McAdoo and Sandwell, 1985). This may account for the cessation of buckle 3505 folding on the east of the TCBtb with increasing lithospheric age to the south of the extinct 3506 MOR. The appearance of buckle folds to the west of the TCBtb yet farther south may 3507 therefore indicate an approach towards an offset MOR segment, which lay in the vicinity of 3508 the present day DFZ.

3509 The distance between the abandoned MOR of the northern TCB (Figure 4.3), and the 3510 northernmost observed buckle folding along line tz1 4000 (west; Figure 4.4), is ~80 km. 3511 Assuming a palaeospreading rate of  $\sim 20$  mm/y (half rate; Phethean et al., Section 2), a 3512 rough relative age of this crust may be calculated, which is  $\sim 4$  Myr older than the extinct 3513 MOR. This suggests that the compression of this basin predated and overlapped the 3514 abandonment of spreading. Due to several assumptions made during this calculation (i.e. 3515 wavelength of folding not affected by decoupling along thrusts, constant spreading rate 3516 before extinction, etc.), however, we prefer to simplify this finding to say that the 3517 compressional event and extinction of the MOR occurred at roughly the same time.

As changes in plate motion can lead to the build-up of transpressional stress along fracture zones (Section 4.1), which is seen in the TCB, we propose that the plate motion change near the end of the Jurassic (e.g. Chapter 2; Reeves et al., 2016) may have been the ultimate cause of compression along the TCBC. This change may have also resulted in the extinction of MOR segments incompatible with the new spreading direction, accounting for the contemporaneity of compression and MOR extinction. The absence of a sedimentary record of the compressional event along the southern TCBtb, despite the presence of sediments equivalent to those that record such a compression farther north along this structure,suggests diachronous cessation of compression from south to north.

## 3527 <u>4.4.3. The DFZ</u>

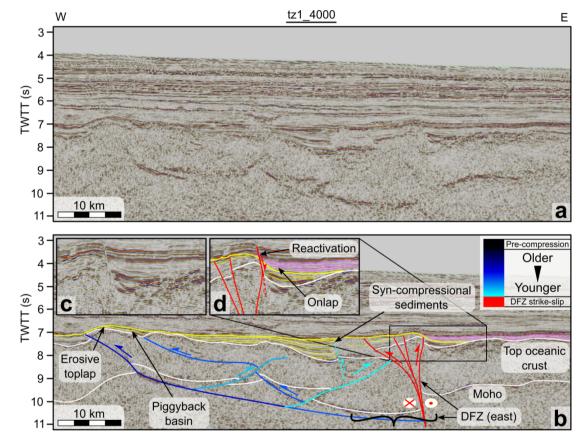
3528 The DFZ forms the eastern boundary of the TCB, and is readily recognisable in free-air 3529 gravity data as an arcuate low gravity anomaly, which spans from just offshore Kenya at 3530 ~5°S, to the southwestern tip of Madagascar at -25°S (Figure 4.1). This large-offset fracture 3531 zone separates the Jurassic oceanic crust of the TCB from the younger crust of the central 3532 WSB. The DFZ developed in response to a change in plate motion from NNW-SSE 3533 spreading to ~N-S spreading near the end of the Jurassic, after which East Gondwana was 3534 translated southwards along this major transform fault (e.g. Chapter 2; Reeves et al., 2016). 3535 To the north of the WSB's spreading centre, which was abandoned when Madagascar 3536 reached its present day latitude at ~125 Ma, the DFZ experienced a variable cumulative 3537 offset along its length. This is due to the cessation of transform motions following the 3538 southward passage of the MOR, which means that the cumulative dextral offset along the 3539 DFZ to the north of the extinct spreading centre increases from north to south.

### 3540 **4.4.3.1. Line tz1\_4000 (east)**

3541 Near the northern limit of the TCB, just to the south of the Davie-Walu Ridge, line tz1 4000 3542 (east) crosses the DFZ (Figure 4.8). Smooth or hummocky top basement reflections and high 3543 amplitude Moho reflections define crustal thicknesses of between 1.5 s and 2.4 s TWTT in 3544 regions of undeformed crust. To the west of the DFZ, within the TCB, the crust has been 3545 folded and buckled, sometimes in association with offsets in the top of the basement and 3546 Moho. High amplitude convex downward reflections, which are also present beneath the 3547 level of Moho reflections, are aligned with basement offsets and the steepening of crustal 3548 folds. Above the basement, sedimentary deposits that correlate with syn-deformational 3549 sediments elsewhere along line tz1 4000 (west; i.e. Figure 4.4), show both onlap and toplap 3550 relationships, and diverge into the trough of the central fold. In the centre of this trough, a 3551 small step up in the basement occurs moving east and above this the layered reflections of 3552 the sedimentary package has been disrupted. Both of the edges of the disrupted region dip to the east, but the eastern edge propagates farther through the stratigraphy than the western, 3553 3554 which is onlapped by later deposits. East of this, a block of uplifted crust in the shape of an 3555 inverted triangle sits adjacent to a small basin on the east, and is shown in detail in Figures 3556 4.8c and d. The western edge of this crustal block is defined by reflections in the mid crust 3557 which become more steeply dipping downwards, and may align with the offset in the top of 3558 the basement along the eastern edge of the disrupted sediments. The eastern edge of the 3559 block is defined by the termination of top basement and mid-crustal reflections from the

3560 east, and may also steepen downwards to align with an offset of Moho reflections. An 3561 erosion surface, truncating the sedimentary packages that overlie the uplifted block, bounds 3562 local deposits to the east of the block, which have a high amplitude reflectivity (Figure 4.8c 3563 and d). Above this erosion surface, sedimentary deposits onlap the palaeobathymetric high, 3564 and a small normal fault has developed along the edge of the uplifted block. East of the 3565 uplifted block, the Moho locally deepens before flattening to the east, and the top basement 3566 surface does the same. This crust, to the east of the DFZ, shows a slightly more seismically 3567 transparent nature than crust to the west, and, away from the DFZ, remains at a relatively 3568 consistent depth.

3569





**Figure 4.8.** Folding, thrusting, and strike slip deformation along seismic line tz1\_4000 (east) in response to compression and development of the DFZ. (a) Uninterpreted section. (b) Interpreted section. (c) Close up of the inset region without interpretation. (d) Close up of the inset region with interpretation. Line descriptions are as for Figure 4.4, and pink lines depict the structure of the post-deformation sequences. The relative ages of the faulting are shown in the key and continuation of the structures at depth is speculative. Location shown in Figure 4.1.

The thickness of the crust, and smooth high amplitude top basement reflections, indicate an oceanic nature of the basement in this region. To the west of the DFZ, within the TCB, the 3580 folding of crust and offsets in the top basement and Moho reflections are similar to 3581 observations along the TCBtb. In Figure 4.8, as along the TCBtb, buckle folding and 3582 thrusting have accommodated compressional deformation, resulting in some crustal 3583 thickening, particularly around the DFZ. The tightening of the folding to the west of the 3584 DFZ, resulting in the deepening of the central syn-deformation basin, is coincident with a 3585 convex downward reflector in the mid crust, and is likely influenced by a propagating thrust. 3586 The western edge of the band of disrupted stratigraphy in the middle of this basin is 3587 coincident with an offset in the top of the basement, and likely results from back thrusting 3588 along a large Moho-offsetting thrust. The matching tops of divergent sedimentary packages 3589 here, and syn-compressional stratigraphies elsewhere in the north of the TCB, suggests that 3590 compression occurred at the same time here as farther west around the northern TCBtb.

3591 The western fault boundary of the uplifted triangular block crosscuts the entire overlying 3592 syn-compressional sedimentary sequence, and is therefore one of the latest structures to form 3593 in this region. The downward steepening of these reverse faults is characteristic of positive 3594 flower structure geometries that develop during transpression, suggesting strike-slip 3595 deformation across this structure. These observations are consistent with satellite gravity 3596 data that suggest that the DFZ passes through this region, and the change in the character of 3597 the oceanic crust across the strike-slip zone further supports the presence of a major crustal 3598 discontinuity. Crust to the east of the DFZ has, therefore, only been accreted adjacent to the DFZ following the passing of the MOR and cessation of strike-slip deformation. The very 3599 3600 limited and localised deposition of sediments below the erosional surface bounding the syn-3601 deformation sedimentary packages on the east of the DFZ (Figure 4.8c and d) suggests that 3602 this crust formed after, or at a similar time to, the cessation of compressional deformation. It 3603 is likely that these localised deposits were sourced from the eroding DFZ flower structure, 3604 and rapidly deposited in new accommodation space being generated at the passing spreading 3605 centre. In this case, the intercalation of these sediments with volcanic deposits at the MOR 3606 may have resulted in their high amplitude nature, and the loading of this zero-age crust may 3607 have contributed to the local subsidence and depression of the Moho. Following this, further 3608 loading of this crust by post-compressional sediments, which onlap the unconformity 3609 bounding the syn-deformation sediments, and thermal subsidence may have resulted in the 3610 partial reactivation of the DFZ as a normal fault (Figure 4.8d).

# 3611 **4.4.3.2. Line tz3\_3600 (east)**

3612 75 km to the south, line tz3\_3600 (east) crosses the DFZ and two other deviations of the top 3613 basement interface (Figure 4.9). At the western end of the section, the top basement, upper 3614 crustal reflections, and a band of less reflective mid-lower crust dip to the west, defining a 3615 small (~15 km) fold, which uplifts the top basement at its crest by 0.2 s TWTT. The top of 3616 this fold is onlapped by the syn-compressional sediments identified to the west (i.e. Figure 3617 4.5). The available data do not constrain the continuation of this package to the east; 3618 however, there is little evidence that it continues to the east of the DFZ, with the possible 3619 exception of local deposits as seen in Figure 4.8. Just east of the fold, the crust, which has a 3620 thickness of approximately 2.3 s TWTT, has flat top basement and Moho reflections until 3621 near the DFZ. At this point, a thickening of the crust to 3.2 s TWTT occurs as the top 3622 basement becomes tilted to the west and the Moho dips to the east, and the reflective nature 3623 of the upper crust becomes subdued. On the other side of the DFZ, the opposite occurs, 3624 returning the crust to a normal thickness.

Beneath, and to the east, of the DFZ, straight low-angle reflections of low-mid amplitude are common in the crust and mantle. Within the crustal thickening of the DFZ, however, a high amplitude convex downward reflector dips to the west, and may coincide with an offset in the Moho. The upper termination of this reflector coincides with the downward continuation of an east dipping reflector that steepens downwards.

3630 Crust to the east of the DFZ shows a generally consistent thickness, with flat Moho and top 3631 basement interfaces spanning  $\sim$ 50 km. The eastern limit of this flat crust is defined by an 3632 offset in the top of the basement down to the east and an  $\sim$ 3 km wide inverted triangular 3633 trough, containing flat-lying reflective material, which separates it from more flat-lying 3634 oceanic crust to the east.

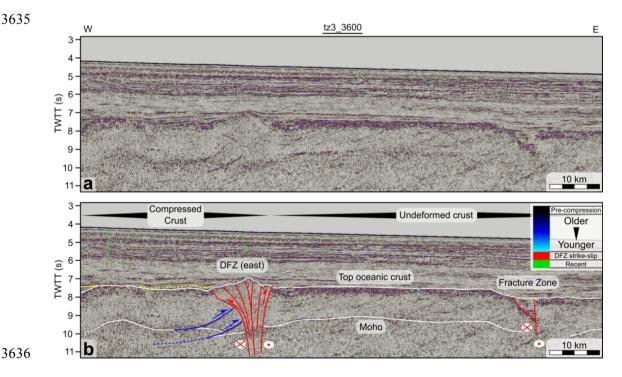


Figure 4.9. Folding, thrusting, and strike slip deformation along seismic line tz3\_3600 in response to compression and development of the DFZ. (a) Uninterpreted section. (b)

Interpreted section. Line descriptions are as for Figure 4.4, where the dashed yellow line is
the inferred top of syn-compressional sediments. The relative ages of the faulting are shown
in the key. Location shown in Figure 4.1.

3642 The small fold at the western end of the profile, with a slightly uplifted crest, is to the west 3643 of the DFZ, and therefore within the TCB. The same compressional event that resulted in the 3644 thrusting and folding of the basement elsewhere within the TCB may, therefore, have also 3645 generated this fold structure. The wavelength of the fold is shorter than other buckle folds 3646 seen within the TCB, and may have been influenced by a propagating thrust fault, as thought 3647 to result in tighter folds elsewhere within the TCB. The lack of clear Moho reflections, 3648 however, precludes the determination of any Moho offsets that would confirm this 3649 interpretation. The oceanic crust adjacent to this fold has a thickness typical of that elsewhere in the TCB, at ~2.3 s TWTT, and does not appear to have been thickened. 3650 3651 Approaching the DFZ, however, the tilting of the crust and the presence of high amplitude 3652 convex downward reflections, which may correspond to sheared crust and the sheared crust-3653 mantle boundary, suggest that east verging thrusts have developed in this location and 3654 contribute to the crustal thickening of the DFZ. The truncation of these thrusts by later faults 3655 that steepen downwards and have contributed to the uplift of the top basement interface 3656 suggests that a positive flower structure developed at a later stage in response to an increasing strike-slip component of deformation here. These steeply dipping strike-slip faults 3657 3658 may have destroyed the initially reflective structure of the upper oceanic crust within the 3659 DFZ.

To the east of the DFZ, the flat oceanic crust has a thickness of  $\sim 1.8$  s TWTT, typical for the oceanic crust of the WSB (Coffin et al., 1986). This crust shows no sign of compressional deformation, despite the transpression experienced locally along the DFZ here, and farther north along line tz1\_4000 (east). This suggests that this crust formed later than both the compressional episode that affected the TCB and the transpression along the DFZ, and is consistent with its accretion on the north of the MOR in the WSB.

As line tz3\_3600 (east) is oriented perpendicular to the palaeo-spreading direction of the WSB, and would therefore be expected to crosscut fracture zones related to oceanic spreading, the step down in the top of the oceanic crust farther east and accompanying basin likely represent a locally transtensional fracture zone that developed between spreading segments of the WSB.

#### 3671 **4.4.3.3. Line tz4\_3350**

3672 Approximately 80 km farther south, seismic line tz4 3350 images both the DFZ and the SSE 3673 trending TCBtb as they come into proximity (Figure 4.10). At the western edge of the 3674 section, the basement has been uplifted to ~6.5 s TWTT, above the adjacent crust at 7.7 s TWTT, and forms part of the TCBtb. On the eastern flank of the TCBtb, an ~0.2 s TWTT 3675 3676 offset in the top basement uplifts crust to its east. This offset may align with east dipping 3677 reflections in the mantle via a disrupted zone of Moho reflectivity, across which the Moho is 3678 possibly offset, and weak east dipping reflections within the crust. In the same location, mid 3679 and low amplitude reflections in the crust and mantle, respectively, may be aligned and dip 3680 to the west, but are not associated with any resolvable offset. In this western region of Figure 3681 4.10, clear top basement and Moho reflections and a seismically transparent mid crust define 3682 an oceanic crustal thickness of  $\sim 1.7$  s TWTT. Moving eastwards, the crust thickens to 2.7 s 3683 TWTT and mid-crustal reflectivity increases. The top basement and Moho reflections form a 3684 bulbous shape, similar to that which forms the DFZ (east) farther north. This thickened crust 3685 contains several prominent crosscutting reflections, dipping to both the east and west, which 3686 may align with hummocks in the top of the basement. This crustal thickening is offset to the 3687 west of the trend of the DFZ (east) by  $\sim 20$  km.

3688 East of the thickened crust, a sharp onset of chaotic and high amplitude reflections in the top 3689 of the crust and seismic transparency in the mid-lower crust occurs across a sub-vertical 3690 boundary, which possibly also connects offsets in the top basement and Moho. East of this 3691 boundary, a clear Moho reflection defines the base of flat lying oceanic crust. The top of this 3692 crust lies at a depth of 7.7 s TWTT and has a thickness of 1.8 s TWTT. This crustal strip, 3693 with a width of approximately 10 km, loses its character to the east beneath an intruded sill, 3694 which disrupts imaging of the basement structures. A rise in the top basement to 7.5 s 3695 TWTT, however, is imaged just to the east of the sill. This basement uplift has a width of 3696 between 10 and 15 km, similar to the bulbous crustal thickening to the west and the DFZ 3697 (east) to the north (i.e. Figure 4.9). It is also aligned with the trend of the DFZ (east) as 3698 derived from lines tz1 4000 and tz3 3600. Beneath, and on either side of, the uplifted top 3699 basement interface, mid-crustal reflections bend upwards before terminating against 3700 boundaries that possibly steepen downwards. To the east of this uplifted basement, an 3701 intruded sill complex obscures detailed investigation of the basement structures.

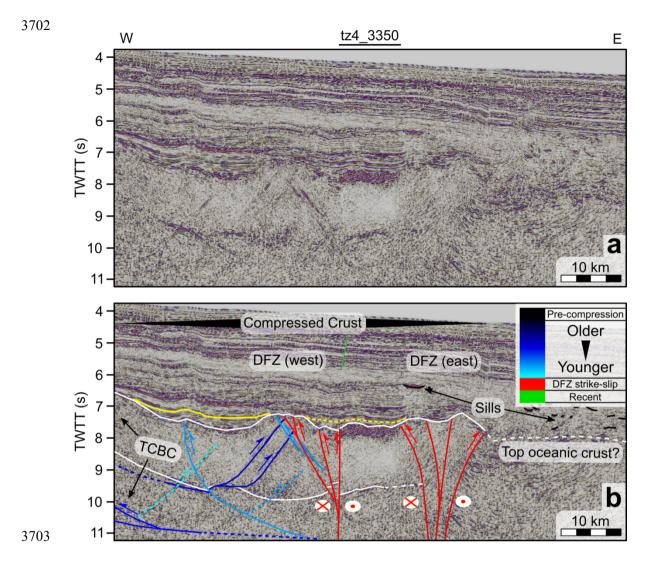


Figure 4.10. Folding, thrusting, and strike slip deformation along seismic line tz4\_3350 in response to compression and development of the DFZ. (a) Uninterpreted section. (b) Interpreted section. Line descriptions are as for Figure 4.4, where the top of the syncompressional sediments is now interpreted to correlate with the top of deformed sediments on lines tz1\_4000 (west) and tz3\_3600 (west) and the dashed yellow line is inferred. The relative ages of the faulting are shown in the key. Location shown in Figure 4.1.

3710 The TCBtb, at the western edge of Figure 4.10, has a similar geometry to the TCBtb in 3711 Figure 4.6, located 10 km to the south, and likely formed in a very similar way. The top 3712 basement offset in its flank, coincident with dipping reflections in the crust and mantle and a 3713 possible offset in the Moho, likely developed in response to a west verging thrust. As this 3714 structure does not offset crosscutting, east verging reflections, it likely developed at an 3715 earlier stage. The bulbous crustal thickening in the western half of Figure 4.10 contains 3716 dipping reflections that may be coincident with hummocks in the top of the basement, and 3717 likely results from thrusting of the crust, similar to that seen along line tz3 3600 (east). Possible offsets of the dipping reflections may indicate a time progression of thrusting. Sub-3718

3719 vertical faults, offsetting the basement and bounding areas of different crustal characteristics, 3720 likely separate the crustal thickening from undeformed crust to the east and offset earlier 3721 thrusts. The steepness of these faults and inconsistent offsets of Moho and top basement 3722 interfaces along them indicate a strike-slip nature. These faults may therefore form part of a 3723 transform system related to, but offset from, the DFZ (east), which we therefore term the 3724 DFZ (west). 20 km to the east, the second crustal thickening, with reflections possibly 3725 dragged upwards into it along downward steepening faults, likely forms part of the DFZ 3726 (east). This is supported by its similar positive flower geometry and alignment with the trend 3727 of this structure. The two branches of the DFZ are separated by a thin 10 km strip of 3728 undeformed oceanic crust.

# 3729 **4.4.3.4. Line tz4\_2850**

3730 Another 110 km to the south, the DFZ nearly intersects the TCBtb (Figure 4.11). In the west, 3731 folded top basement and Moho interfaces are separated by a seismically transparent mid 3732 crust. This oceanic crust has a thickness of 1.8 s TWTT. The eastern termination of the fold 3733 is marked by the termination of upper crustal reflectivity and mid crustal seismic 3734 transparency along a sub-vertical boundary. A sudden flip in the dip angle of the top 3735 basement, which is also possibly upthrown on the east, occurs just above this and lies along 3736 the trend of the TCBtb. In this location, beneath the west dipping top basement, short, 3737 segmented mid-crustal reflections also dip to the west.

3738 To the east by  $\sim 10$  km, the top basement interface and internal reflections of the crust are 3739 segmented and folded on a short ( $\sim$ 1-3 km) length scale. Beneath this and to the east, 3740 segmented and offset Moho reflections, which dip to the east before reversing to dip to the 3741 west, define the base of a bulbous or lens-shaped thickening of the crust. Several offsets and 3742 hummocks in the top of the basement also define this crustal thickening, which in places 3743 reaches a thickness of  $\sim 2.4$  s TWTT. This structure lies  $\sim 20$  km to the west of the trend of 3744 the DFZ (east), and, as seen farther north, therefore likely represents the DFZ (west). Along 3745 the trend of the DFZ (east), a second lens-shaped segment of thickened crust is present, 3746 likely representing this fracture zone. Along the top basement interface of the DFZ (east), 3747 variations in reflector amplitude, dip, and continuity, occur. Below this, reflections which 3748 dip into the central region of the DFZ (east) may align with some of these variations in the 3749 top of the basement. Between the two regions of thickened crust, the structure of the 3750 basement is obscured by sills intruded into the overlying sediments, which precludes a 3751 determination of the crustal thickness in this region. East of the DFZ (east), within the WSB, 3752 strong top basement reflections are underlain by a reflective crustal layer, below which a 3753 high amplitude reflector, possibly the Moho, defines a thin (~1.1 s TWTT) oceanic crust.

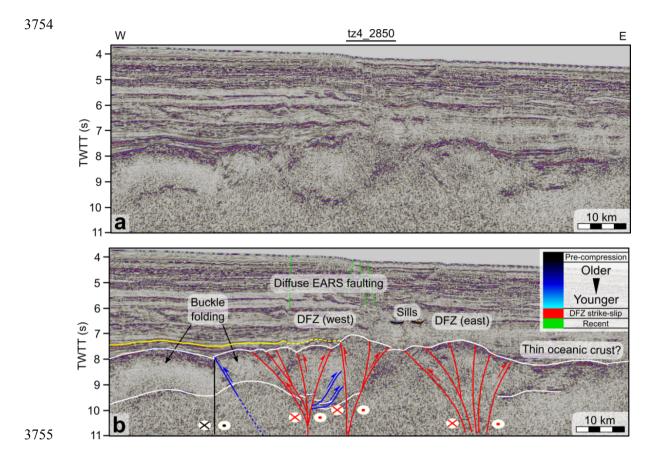


Figure 4.11. Folding, thrusting, and strike slip deformation along seismic line tz4\_2850 in
response to compression and development of the DFZ. (a) Uninterpreted section. (b)
Interpreted section. Line descriptions are as for Figure 4.4. The relative ages of the faulting
are shown in the key. Location shown in Figure 4.1.

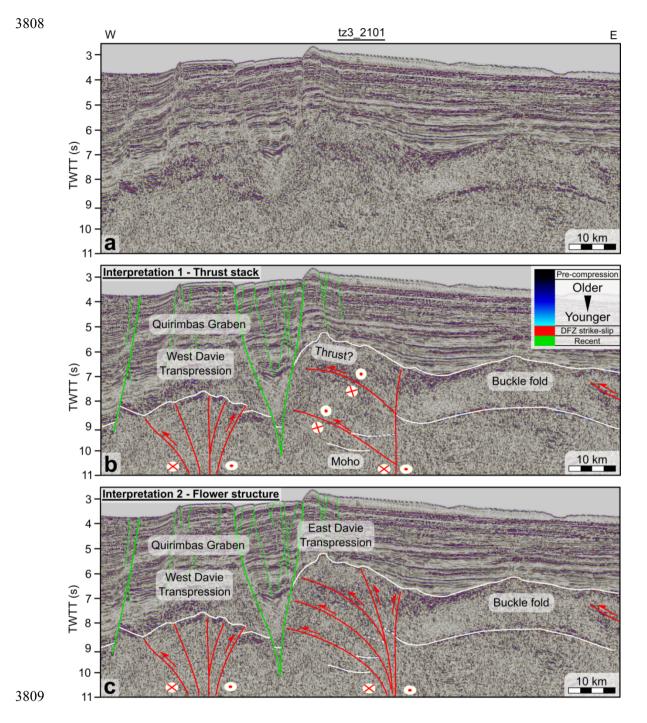
3760 In the western part of Figure 4.11, the sudden change in the dip of the top basement at the 3761 eastern end of the fold suggests a decoupling of the fold from crust to the east. This may 3762 have occurred along the sub vertical boundary between zones of different crustal reflectivity, 3763 and/or east dipping thrust faults, which may have also offset and folded the top basement 3764 interface and mid crustal reflections. These thrust faults lie along the trend of the TCBtb and 3765 likely form part of this structure. Here, however, at this southern limit of the TCBtb, very 3766 little crustal thickening has occurred. This is consistent with an earlier cessation of 3767 compression in the south, as inferred from the deformation of sedimentary packages 3768 overlying the TCBtb (i.e. Section 4.4.2.5), which would result in less crustal thickening in 3769 this southern location. The presence of a buckle fold to the west of the TCBtb, is consistent 3770 with earlier observations of a swap from the development of buckles on the east of the 3771 TCBtb in the north to on the west in the south, likely related to the age of the crust before 3772 deformation.

3773 Slightly to the east, the folded and offset segments of crust that form part of the DFZ (west) 3774 may be bounded by discontinuities that dip towards the centre of the lens-shaped area of 3775 thickened crust, defining a positive flower structure. In the central region of the DFZ (west), thrust faults, which stack and thicken the crust, may also present. In this interpretation, these 3776 3777 gently dipping faults have been cut by later, steeper, strike-slip faults, consistent with the 3778 onset of strike-slip deformation after thrusting, as inferred elsewhere along the DFZ (e.g. 3779 tz3 3600 (east)). The DFZ (east), which likely also forms a positive flower structure, is 3780 again offset from the DFZ (west) by ~20 km, suggesting that these features run parallel to 3781 each other. This is confirmed by observations made at several other locations as summarised 3782 in Figure 4.1.

### 3783 **4.4.3.5. Line tz3\_2101**

3784 Line tz3 2101 intersects the DFZ some 215 km to the south of the previous line. It is located 3785 to the south of the extinct spreading centre of the WSB (Chapter 2), and shows elements of 3786 the Quirimbas Graben (Franke et al., 2015). This recent extension is related to the East 3787 African Rift System and overprints deformation associated with the DFZ (Figure 4.12). 3788 Nonetheless, at the western edge of the section beneath the Quirimabas Graben, top 3789 basement reflections are identifiable and step down to the west across normal faults. 3790 Approximately 20 km from the western edge of Figure 4.12, normal faulting becomes 3791 reduced and a horst structure forms a relatively uplifted area of seafloor. Beneath the horst, 3792 segmented, but high amplitude, basement reflections define a generally domed shape of the 3793 top basement, which sits nearly along the trend of the DFZ (west). To the east, a subsidiary 3794 basin separates this basement dome from a second region of uplifted basement, which sits at 3795 depths as shallow as 5.3 s TWTT, as much as 1.7 s TWTT above the surrounding basement. 3796 Beneath this basement uplift, which sits nearly along the trend of the DFZ (east), a diffuse 3797 band of east dipping, mid-high amplitude, reflections sits at a depth of  $\sim 6.9$  s. Below this, at 3798 depths of 9.3 s and 9.8 s TWTT, other reflections also dip gently to the east or lie sub-3799 horizontal. These may define an offset in the Moho, which has been upthrown on the east. 3800 Inspection of crosslines to tz3 2101 along the western edge of this uplifted crust reveals that 3801 the faults, with a generally E-W strike, bend northwards as they approach the western edge 3802 of the uplift.

To the east of this belt of crustal thickening, top basement and Moho reflections have been folded, and define a buckle fold with a wavelength of  $\sim$ 50 km that has developed in the oceanic crust. Near the eastern edge of Figure 4.12, a step up in the top of the basement is aligned with a high amplitude east dipping reflector in the mid crust.



**Figure 4.12.** Folding, thrusting, and strike-slip deformation along seismic line tz3\_2101 in response to compression and development of the DFZ, and graben formation in response to extension across the Quirimbas Graben. (a) Uninterpreted section. (b) Interpreted section based on the transportation of thrust sheets along the strike of the DFZ (east). (c) Interpreted section based on flower structure development along the DFZ (east). Line descriptions are as for Figure 4.4. The relative ages of the faulting are shown in the key. Location shown in Figure 4.1.

The uplift of the top basement beneath the horst block within the Quirimbas graben is not wholly accounted for by the reduced extension across this structure, and may therefore indicate a relative thickening of the crust in this region. To the east, top basement and Moho reflections record a crustal thickening to 4.7 s TWTT, more than twice the original crustal thickness. Allowing for 5 - 12 km of recent extension across the DFZ (Franke et al., 2015), the two regions of thickened crust in Figure 4.12 are closely aligned with the trends of the large offset DFZ (east) and DFZ (west), suggesting that they are the southern continuations of these structures.

3825 The geometry of the top of the basement around the DFZ (west) is consistent with a 3826 compressional flower structure, as seen elsewhere along this fracture zone. The amount of 3827 crustal thickening recorded at the DFZ (east), however, is greater than anywhere else along 3828 this structure. Eastward dipping reflections within the crust and an offset in the Moho imply 3829 a general tilting of the crust to the east due to east dipping compressional faults. The detailed 3830 nature of the crustal structure is, however, not well constrained by the data, and in Figure 3831 4.12 b and c, we therefore present two scenarios that are consistent with the main 3832 observations. Interpretation 1 (Figure 4.12b) employs a sequence of stacked thrusts sliding 3833 along each other, into the plane of Figure 4.12, to achieve greater than twice the original 3834 crustal thickness in this location. Interpretation 2 (Figure 4.12c) employs steeper faults, 3835 within a large flower structure, to achieve a similar effect.

3836 The cause of this possible change in style and amount of crustal thickening in this location 3837 may be revealed by the presence of E-W trending faults that bend northwards as they 3838 approach the western boundary of the DFZ (east). Ordinarily, normal faults generated at an 3839 oceanic spreading centre near to a MOR transform offset bend in the direction of the 3840 adjacent spreading centre (Davies et al., 2005). This is caused by a rotation of  $\sigma_1$  on 3841 approach to the transform fault due to the local influence of opposing plate motions across 3842 the fault. Within the WSB, MOR segments have a left lateral offset (e.g. Cochran, 1988; 3843 Chapter 2), and the adjacent MOR segment across the DFZ lies in the Mozambique basin. 3844 Ordinarily, therefore, faults within the WSB adjacent to the DFZ should bend to the south, 3845 opposite to the observed trend. Faults that develop along the borders of transpressional 3846 stepover basins, however, can show the opposite trend (e.g. (McClay and Bonora, 2001)). A 3847 step to the east, when moving south along the DFZ (east), would result in such a 3848 transpressional push-up structure, with faults within the WSB bending to the north adjacent 3849 to the DFZ (east). The large offset across the DFZ would likely supply a large amount of 3850 crust into this transpressional zone, allowing a large amount of observed crustal thickening. 3851 Both low-angled thrusts and steep faulted flower structures have been shown to develop 3852 under such transpressional conditions from analogue modelling (e.g. McClay and Bonora,

2001), and it is likely that a combination of these structures, from both Figure 4.12b and c,
contributed to the crustal thickening of the DFZ (east).

Recent extension across the Quirimbas Graben is localised along the flanks of the DFZ (east and west), and likely reactivates these structures, resulting in the development of a horst above the DFZ (west). The northward termination of extension associated with the Quirimbas Graben coincides with the extinct MOR system of the WSB. It is possible, therefore, that coupling across the DFZ to the north of the Quirimbas Graben due to the passing of a spreading centre and associated magmatism stopped the extensional reactivation of the DFZ farther north.

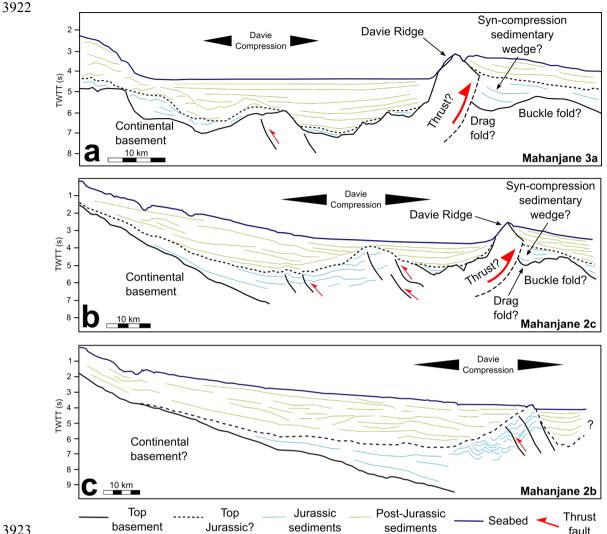
3862 East of the DFZ, a buckle fold and thrust fault, which results in a step in the top of the 3863 basement on the east flank of the fold, have developed. These observations oppose those 3864 from farther north along the DFZ, where to the east of the DFZ (east), the oceanic crust 3865 shows little sign of compressional deformation. Figure 4.12, however, is located to the south 3866 of the MOR system of the WSB, which may provide an explaination for the observed 3867 compressional deformation in this location. To the north of the MOR, no relative motion 3868 occurrs across the DFZ subsequent to the formation of adjacent oceanic crust. To the south 3869 of the MOR, however, this is no longer true. Oceanic crust accreted to the south of the 3870 MOR, and adjacent to the DFZ, formed part of the East Gondwana plate. This crust will 3871 have, therefore, subsequently undergone a protracted history of transportation along the 3872 transpressional DFZ, until the final cessation of spreading in the WSB at  $\sim 125$  Ma. This 3873 transpression may have resulted in the observed buckle folding and compressional 3874 deformation of Figure 4.12. Transpression of this crust would have begun immediately after 3875 its formation at the MOR, and is consistent with the short 50 km wavelength of the buckle 3876 fold, which must have formed in extremely young oceanic lithosphere.

# 3877 4.4.3.6. South of the Quirimbas Graben

3878 To the south of the Quirimbas Graben, gravity modelling and investigation of seismic 3879 reflection lines mz1 8100, mz1 8000, and mz1 7500, has been performed in Chapter 3 (Figures 3.5 and 3.11). Along line mz1 8100, The DFZ is obscured by the St Lazare 3880 3881 volcanic edifice, but lines mz1 8000 and mz1 7500 are clear of volcanics, and allow for an 3882 investigation of basement structures associated with the DFZ. Along the trend of the DFZ 3883 (east), line mz1 8000 displays a thickening of the crust from 2 s TWTT to 2.5 s TWTT. To 3884 the west of the DFZ (east), a zone of possible exhumed mantle exists, which may result in 3885 the lack of observed crustal thickening along the trend of the DFZ (west). The amount of 3886 crustal thickening along the DFZ (east) is similar to observations along the northern DFZ, 3887 and the large amount of crustal thickening observed along line tz3 2101, possibly associated

3888 with a transpressional stepover basin along the DFZ (east), does not continue to this point. 3889 380 km to the south of this, line mz1\_7500 displays a well-developed positive flower 3890 structure along the approximate trend of the DFZ (west), and a 20 km wide zone of 3.3 s 3891 TWTT thick crust has developed. However, 20 km to the east no evidence for strike-slip 3892 tectonics along the trend of the DFZ (east) is present and instead a thrust fault, which offsets 3893 the Moho, thickens the crust in this location.

- 3894 The Rovuma Transform Margin (Figure 4.1, dark green symbols) has also been identified 3895 along these three seismic and gravity profiles (Chapter 3), and runs in a SSE direction. It is 3896 aligned with the Davie Compression (Figure 4.1, red circles), a SSE trending fold and thrust 3897 belt composed predominantly of Jurassic and basement rocks (Mahanjane, 2014; Figures 3898 4.13a, b, and c). The natural prolongation of the large offset Rovuma Transform Margin into 3899 the trend of the Davie Compression (Figure 4.1), suggests a genetic relationship between 3900 these two features, and we propose that the Davie Compression forms the transform 3901 continent-ocean boundary along the eastern edge of the Angoche basin.
- 3902 To the east of the SSE trending Davie Compression, the Davie Ridge, an ~N-S trending 3903 prominent basement high which generally protrudes 2-3 s TWTT above the surrounding 3904 basement (Figure 4.13a and b), forms the natural extension of the DFZ farther north (Figure 3905 4.1). This uplifted basement structure has previously been interpreted as a rift shoulder uplift 3906 (e.g. Mahanjane, 2014), similar to the uplifted flanks of the Quirimbas Graben farther north. 3907 However, dredge samples recovered from the DFZ show a metamorphic P-T path associated 3908 with compressional tectonics, (Bassias, 1992). The geometry of the Davie Ridge, visible in 3909 published seismic sections (e.g. Mahanjane, 2014), shares characteristics with thrusts 3910 observed along the TCBtb. To the east of the Davie Ridge, flexure of the basement has led to 3911 the development of a sedimentary basin and, approaching the ridge, the top basement is bent 3912 upwards towards the uplifted basement. An erosion surface has developed along the top of 3913 the highly rotated basement of the Davie Ridge, and possibly bounds syn-compressional 3914 sediments deposited in the adjacent basin (Figure 4.13.a and b). This is similar to the 3915 geometries that developed along the TCBtb in response to compressional flexure, thrust 3916 faulting, and associated fault drag folding. We therefore tentatively suggest that the Davie 3917 Ridge may, in fact, be composed of oceanic crust and meta-sediments, compressed and 3918 uplifted by thrusting. If the Davie Ridge is such a thrust structure, it likely developed in 3919 response to the compressional stresses that led to the development of the Davie 3920 Compression, or during the compression near the end of the Jurassic, when this crust may 3921 have been attached to East Gondwana, conjugate to the compressed crust within the TCB.



3923

3924 Figure 4.13. Line drawings of seismic reflection data from Mahanjane (2014), which have 3925 been re-interpreted in this study. (a) Figure 3a of Mahanjane (2014). (b) Figure 2c of 3926 Mahanjane (2014). (c) Figure 2b of Mahanjane (2014). Location shown in Figure 4.1.

3927 4.4.3.7. Summary of the DFZ formation

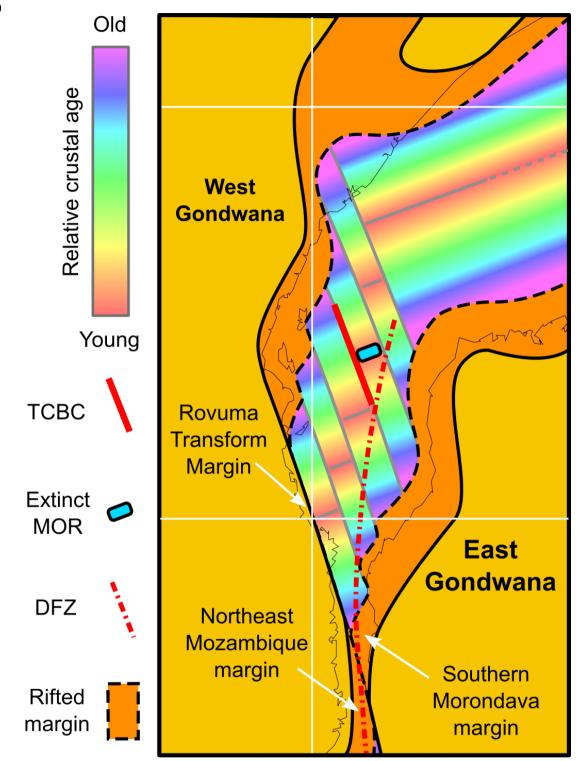
3928 Bounding the eastern edge of the TCB and the Mozambique Basin, the DFZ appears as a 2000 km arcuate free-air gravity low between -5°S and -25°S (Figure 4.1), following a 3929 3930 single small circle with the pole of rotation at 10.15 S, 74.30 E. Our interpretation of seismic 3931 reflection data supports the presence of steep strike-slip faults and transpressional flower 3932 structures along the DFZ, and different tectonic crustal histories on either side of the DFZ. 3933 This combined evidence supports the interpretation of this structure as a large-offset fracture 3934 zone along which East Gondwana was transported southwards. Along the majority of the 3935 DFZ, crustal thickening, generally to ~2.8 s TWTT, but possibly of up to 4.7 s TWTT, 3936 results from thrusting and the development of transpressional flower structures. Faults 3937 associated with strike-slip motion and flower development generally cut older thrust faults,

possibly reflecting a shift from compression to transpression over time. In general,
deformation along the DFZ has been dominated by transpression throughout its history, and
little evidence for transtension is present.

3941 Along most of the length of the WSB, the DFZ is comprised of a pair of transpressional 3942 fracture zones, which run parallel to each other from 7°S to just below the Comoros Islands 3943 c. 13°S. North and south of these latitudes, evidence for strike-slip motion only exists along 3944 the trends of the DFZ (east) and DFZ (west), respectively, with a termination of strike-slip 3945 deformation along the trend of the other DFZ branch. The northern termination of the DFZ 3946 (west), however, lies at the intersection of a SSE trending fold and thrust belt within the 3947 TCB (Figure 4.1). Given that SSE trending thrust belts within the TCB seem to have 3948 developed along pre-existing fracture zones, the initiation of the DFZ (west) at an 3949 intersection with such a thrust belt may indicate a genetic link between these two structures. 3950 In the south, along the trend of the DFZ (east) on line mz1 7500, the presence of a thrust, 3951 but no indication of strike-slip motion, may suggest a similar termination of the DFZ (east) 3952 into a pre-existing SSE trending FZ. Whether such a join between the DFZ and SSE trending 3953 fracture zones occurred gradually, through the formation of arcuate fracture zones, or 3954 suddenly, through the development of crosscutting faults, is not constrained by the available 3955 data. However, the lack of bending of the southern TCBtb into the DFZ suggests that any 3956 previous fracture zone in this location was cut by the DFZ, rather than merging with this 3957 structure. This is supported by the extinction of the MOR to the NE of the TCBtb (Figure 3958 4.3), suggesting a sudden abandonment of the spreading configuration, as opposed to a 3959 gradual change in spreading directions, which might have allowed this spreading centre to 3960 remain active.

3961 Formation of the DFZ, whether gradual or sudden, resulted from a change in plate motion 3962 near the end of the Jurassic (e.g. Reeves et al., 2016; Chapter 3), when an alignment between 3963 the obliquely rifted northeast Mozambique margin and southern Morondava margin 3964 occurred (Figure 4.14). To the north of these juxtaposed margins, plate tectonic modelling 3965 (e.g. Phethean et al., 2016) predicts a series of left-stepping spreading centres running north-3966 south, close to the eastern edge of the present day TCB. The presence of an extinct MOR 3967 segment near to the DFZ in the north of the TCB, and the occurrence of short wavelength 3968 buckle folds, which are associated with young ( $\leq 3$  Myrs) oceanic crust, close to the DFZ on 3969 the SW side of the TCBtb, is consistent with this prediction. This alignment of weak rifted 3970 margins and young oceanic lithosphere (Figure 4.14) may have influenced the timing of the 3971 plate rotation, and the location of the DFZ development. The plate rotation, near the end of 3972 the Jurassic, likely resulted in transpression along incompatible fracture zones in the TCB,

3973 and presents a possible driving mechanism for the compression seen there. Development of 3974 the DFZ, which was compatible with the new plate motion, however, would remove this 3975 driving mechanism of compression within the TCB. As the sedimentary record along the 3976 TCBtb suggests an earlier cessation of compression along the south of this structure, 3977 compared to north, it may also be the case that the DFZ developed from the south, 3978 propagating northwards to form the edge of the TCB.





**Figure 4.14.** Plate configuration around the end of the Jurassic. Transformation of East Gondwana along the Rovuma Transform Margin has led to the alignment of the southern Morondava and north-east Mozambique oblique rifted margins, and release of East Gondwana from the constraints of strong lithosphere along this transform boundary. A contemporaneous alignment of young oceanic crust to the north may have influenced the location at which the DFZ developed.

# 3988 **4.5. Regional tectonic interpretation**

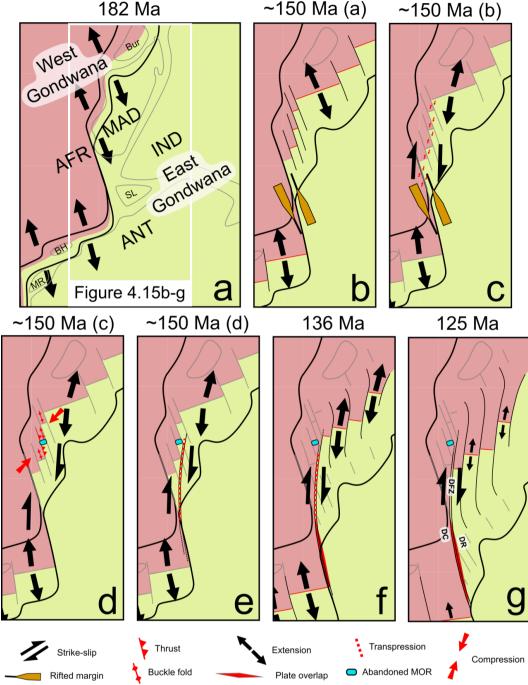
3989 Observations of compression within the TCB, followed by strike-slip and transpression 3990 along the DFZ, likely result from the change in plate motions of East Gondwana relative to 3991 West Gondwana during plate separation. These observations are, therefore, tied to this 3992 broader regional tectonic context below:

- Following a tight initial fit of Madagascar within Africa (Section 4.3), rifting
   between East and West Gondwana began at ~182 Ma (Figure 4.15a; Geiger et al.,
   2004), coincident with the eruption of the Karoo large igneous province in
   Mozambique (e.g. Riley and Knight, 2001). This resulted in oblique margin
   formation along northeastern Mozambique and the Southern Morondava Basin in
   Madagascar (Chapter 3).
- 3999 • Following continental breakup at ~170 Ma (e.g. Chapter 2), an initial phase of SSE 4000 plate separation translated Southern Madagascar (attached to East Gondwana) along 4001 the Rovuma Transform Margin of Northern Mozambique (Figure 4.15a-b; Chapter 4002 3). Due to the large offset of the Angoche Basin and Tanzania Coastal Basin across the Rovuma Transform Margin (Chapter 3), these basins are unlikely to have 4003 4004 overlapped and were, therefore, likely connected by transform faults at the onset of 4005 rifting (Section 1.2.2.1). This means no rifting of the lithosphere along the Rovuma 4006 Transform Margin would have occurred prior to strike-slip motions, maintaining the 4007 thickness and strength of this lithosphere.
- 4008 At ~150 Ma, an alignment of the obliquely rifted margins of northeastern ٠ 4009 Mozambique and the Southern Morondava Basin occurred across the Rovuma 4010 Transform Margin (Figure 4.15b). Around this time, MOR segments to the north 4011 within the TCB were also aligned with these opposed rifted margins (Figure 4.14). 4012 The timing of this alignment, near the end of the Jurassic, was coincident with a 4013 change in plate motion from SSE to N-S, (Figure 4.15b-c) and it is possible that the 4014 alignment of weaker lithosphere, along the rifted margins and MOR segments, 4015 influenced this change.
- This new southerly plate motion was incompatible with SSE trending fracture zones within the TCB, and resulted in the build-up of transpressional deformation along these structures (Figure 4.15c), leading to thrusting of oceanic crust along pre-existing fracture zones and short wavelength buckle folding of young (<3 Ma) oceanic crust (Figure 4.15d). The extinction of spreading centres within the TCB at the same time as the onset of compression was probably ultimately caused by this change in spreading direction. A prediction of this model, which may be tested in</li>

4023future as more widespread data becomes available, is that this compressional event4024will have also affected the Jurassic oceanic crust to the west and northwest of4025Madagascar, which at the time was attached to East Gondwana.

This compression, which continued for longer in the north of the basin, was most
 likely ended by the development of the DFZ, which propagated northwards and
 accommodated the new plate motion (Figure 4.15e). This fracture zone was,
 nonetheless, dominated by transpression throughout its history.

Subsequent to the formation of the DFZ at  $\sim 150$  Ma, the southward drift of East 4030 • 4031 Gondwana led to the overlap of the East and West Gondwana plates across the 4032 southern Rovuma Transform Margin (Figure 4.15f), and thus the collision of south 4033 Madagascar with the obliquely rifted margin of northeast Mozambique (red area, 4034 Figure 4.15f). The Davie Compression, a thrusted and inverted sequence of Jurassic 4035 and basement rocks along the eastern edge of the Angoche Basin (Mahanjane, 4036 2014), likely resulted from this plate collision (Figure 4.15g), and may therefore 4037 represent the continent-ocean boundary in this region. The Davie Ridge, which may 4038 be a highly rotated oceanic thrust, may have also developed during this collision. 4039 This is supported by the metamorphic P-T paths from along this structure, which 4040 show a pattern generally associated with collisional settings (Bassias, 1992).



4044 Figure 4.15. Plate tectonic model showing the change in plate motion between West 4045 Gondwana (pink) and East Gondwana (green) at ~ 150 Ma, possibly influenced by the 4046 alignment of rifted margins offshore Mozambique and Madagascar, and of MOR segments 4047 in the north. This plate motion change may have resulted in compression within the TCB, 4048 extinction of MOR segments within the TCB, and development of the DFZ. Subsequently, 4049 the collision of southern Madagascar with the oblique rifted margin of northeast 4050 Mozambique during the southward drift of East Gondwana may have led to the development of the Davie Compression and Davie Ridge. AFR, Africa; ANT, Antarctica; BH, Biera 4051

4052 High; Bur, Bur High; DC, Davie Compression; DR, Davie Ridge; IND, India; MAD,
4053 Madagascar; MR, Mozambique Rise; SL, Sri Lanka.

# 4054 **4.6. Implications for plate motion controls**

4055 Plate motions are classically inferred to be controlled by a number of driving forces 4056 originating from the lithosphere and mantle, including: mantle convection (e.g. Ziegler, 4057 1993); slab-pull (Lithgow-Bertelloni, 2014); ridge-push (e.g. Mahatsente and Coblentz, 4058 2015); orogenic collapse (e.g. Rey et al., 2001); and mantle plumes (e.g. Larson, 1991). The 4059 largest change in plate driving forces within the basins surrounding Gondwana was the onset 4060 of subduction along the NeoTethys. However, this was before the Late Jurassic change in 4061 plate motions (Stampfli, 2000). The diachronous subduction of the NeoTethys spreading 4062 ridge, however, may have still been occurring at the time of plate motion changes (e.g. 4063 Stampfli and Borel, 2002). However, the sharp change in spreading direction, indicated by 4064 the cutting of the TCBtb by the DFZ, would be unlikely to have been generated by this 4065 gradual subduction process. In contrast, the alignment of the obliquely rifted northeast 4066 Mozambique margin and Southern Morondava margin with young oceanic spreading centres 4067 farther north occurred relatively rapidly, and at the moment of plate motion change. Furthermore, following plate rotation and the formation of the DFZ, this large-offset fracture 4068 zone was dominated by transpression throughout its history, suggesting it did not form in 4069 4070 perfect alignment with plate driving forces.

4071 We therefore argue that this change in plate motion was not 'driven' by alterations to plate 4072 driving forces, but was partly the consequence of a reduction in resisting forces along the 4073 Rovuma Transform Margin. This reduction was the caused by the alignment of weak, rifted 4074 and/or young, lithosphere along the N-S trend of the future DFZ, which provided a new 4075 alternate pathway for East Gondwana that was more, but not perfectly, compatible with plate 4076 driving forces. In this case, prior to the end of the Jurassic, the driving forces required for a 4077 more southerly drift of East Gondwana would have already been active, but were resisted by 4078 plate boundary forces along the strong Rovuma Transform Margin, resulting instead in the 4079 slip of East Gondwana along the trend of this margin. This would have led to transpression 4080 along the Rovuma Transform Margin, and so structural investigations along this margin 4081 could be used to test this model.

# 4082 **4.7. Conclusions**

4083 A change in the plate motion of East Gondwana near the end of the Jurassic from SSE to 4084 southwards drift led to build-up of transpressional forces along SSE trending fracture zones 4085 within the TCB. This resulted in the onset of compressional tectonics within this basin, 4086 including the development of the SSE trending TCBtb, possibly along a pre-existing fracture 4087 zone, and short wavelength buckle folds, possibly associated with young oceanic crust in 4088 proximity to MOR segments. At the same time, the extinction of spreading centres within 4089 the TCB likely also resulted from the change in spreading direction. This model for the onset 4090 of compression within the TCB predicts that crust that lay conjugate to the TCB, now 4091 situated offshore northern and western Madagascar, would also have been affected by 4092 compression, allowing it to be tested when data becomes available in these regions.

4093 The onset of this plate motion change was coincident with the passing of southern 4094 Madagascar beyond the strong lithosphere of the Rovuma Transform Margin and the 4095 alignment of weak and/or young lithosphere, associated with rifted margins and oceanic 4096 spreading centres, along the future trend of the DFZ. This reduction in lithospheric strength 4097 may have triggered the change in plate motion and would imply first-order control of 4098 resisting forces along transform plate boundaries on the motions of plates during the 4099 dispersal of Gondwana. In this case, prior to the change in plate motions, the Rovuma 4100 Transform Margin may have been under transpression; therefore, structural analyses of this 4101 margin could test this hypothesis.

4102 The development of the DFZ likely ended compression within the TCB, and the earlier 4103 cessation of compression in the south of the basin, as compared to the north, may represent a 4104 south to north propagation of the DFZ along what was to become the eastern boundary of 4105 the TCB. The subsequent drift of East Gondwana along this 2000 km fracture zone was 4106 dominated by transpression, and led to the collision of southern Madagascar with the oblique 4107 rifted margin of northeast Mozambique, forming the Davie Compression, which may 4108 represent the continent-ocean boundary in this region. The Davie Ridge, which may be a 4109 rotated oceanic thrust, may have also developed during this collision.

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# 4235 4.9. Supplementary material

4236 <u>4.9.1. Alternative interpretation of seismic line tz4\_3300</u>

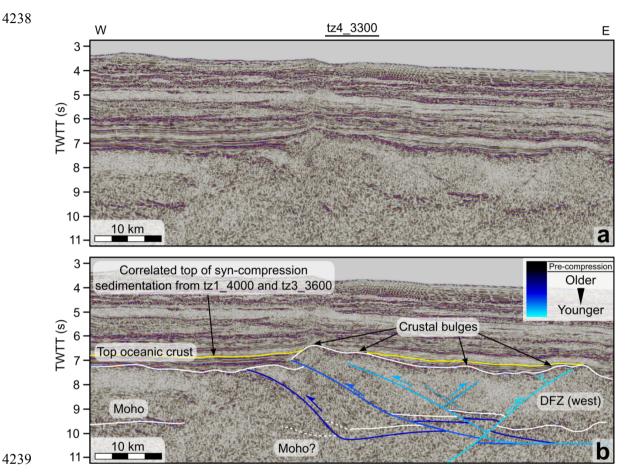


Figure S4.1. The lower of the dual bands of reflections around the level of the Moho may alternatively represent a shear boundary, which has been subsequently offset. The continuation of the structures at depth is speculative.

# 4243 5. Discussion, conclusions, and 4244 future work

4245 Plate tectonic modelling of the WSB, based on the analysis of gravity lineaments related to 4246 ocean spreading features, shows that two phases of spreading led to the development of this 4247 basin. The more recent phase, between the Late Jurassic (~150 Ma) and Aptian (~125 Ma), 4248 involved a simple north-south translation of East Gondwana along the DFZ, and has been 4249 recognised previously (e.g. Coffin and Rabinowitz, 1987; Gaina et al., 2013). The Earlier 4250 phase between Middle Jurassic (~170Ma) and Late Jurassic, however, deviates from the 4251 previous academic consensus for the region and involves a NNW-SSE motion of East 4252 Gondwana. This motion places the origin of Madagascar within the TCB, inboard of the 4253 DFZ, which is supported by the identification of oceanic crust inboard of the DFZ. The 4254 initial SSE motion also predicts the existence of a SSE trending transform margin along the 4255 Rovuma Basin (Chapter 2). Subsequent identification of the Rovuma Transform Margin 4256 using seismic and gravity methods (Chapter 3), confirms this controversial phase of NNW-4257 SSE spreading during the Jurassic. This changes the interpretation of the DFZ, previously 4258 thought to be a continent-ocean fracture zone, to an ocean-ocean fracture zone and 4259 consequently shifts the location of the continent-ocean transition landward within the TCB. 4260 This changes our understanding of the crustal affinity within the TCB, and the hydrocarbon industry should now consider that significant portions of the TCB may be oceanic in nature. 4261 4262 This may affect the likelihood of source rock presence in these regions and should also be 4263 taken into account during paleoheatflow modelling.

4264 The origin of Madagascar within the TCB also supports a tight initial fit of Gondwana 4265 fragments (Chapter 2), which, combined with the newly determined initial SSE drift of East 4266 Gondwana, has large implications for the types of margins that surround the WSB. The 4267 northern margins of the WSB likely developed roughly orthogonal to the extension direction and may therefore be interpreted as rifted margins. The western margins, however, 4268 4269 developed at an oblique angle to the extension direction and are therefore likely highly 4270 segmented and/or obliquely rifted margins (Chapter 2). Between these two margins an offset 4271 is required along the western border of the Lamu Embayment to satisfy the fit of Madagascar within this basin. The offset has a similar trend to the Royuma Transform 4272 4273 Margin and suggests the presence of another transform margin in this region. This is 4274 supported by the observation of strike-slip faults within the conjugate Bajocian deposits of 4275 Madagascar (Chapter 4). The formation of these different types of margin surrounding the 4276 WSB may have been facilitated by several mechanisms suggested to assist continental 4277 rupture. The northern margins follow branches of the Karoo rift system and pre-existing 4278 lithospheric weaknesses may therefore have influenced continental breakup (e.g. Audet and 4279 Bürgmann, 2011). The western margins of the WSB and TCB also follow the pre-existing 4280 Karoo rift system, and additionally have undergone oblique rifting, both of which may have 4281 facilitated breakup (e.g. Brune et al., 2012). The Rovuma Transform Margin, however, did 4282 not follow any apparent pre-existing rift system, but was highly oblique (Chapter 3). Rifting 4283 in the Mozambique basin to the south was contemporaneous with large amounts of 4284 volcanism and, therefore, thermal weakening due to magmatism may have played an 4285 important role in facilitating continental breakup in this basin (e.g. Buck, 2007). This 4286 suggests that a collaboration of different rifting facilitators led to the breakup of Gondwana. 4287 The prolonged and episodic phase of Karoo rifting prior to volcanism in Mozambique and 4288 subsequent Gondwana dispersal may reflect this necessity for interplay of mechanisms to 4289 achieve successful supercontinent breakup (Chapter 2). Chapter 4 shows that the change in 4290 spreading direction during the Late Jurassic may have been triggered by the alignment of 4291 lithospheric weaknesses along the future trend of the DFZ. This suggests that plate motion 4292 changes are not necessarily driven by changes in bottom-up driving forces, and, 4293 consequently, that plate motions may not purely reflect such driving forces but are also 4294 subject to top-down controls on plate motion of a first-order significance (e.g. Brune et al., 4295 2016). This questions whether observations of plate motions may be used to make inferences 4296 about dynamic processes in the mantle which are thought to drive plate tectonics, and 4297 therefore has large consequences for future academic studies in this field (e.g. Becker and 4298 Faccenna, 2011).

4299 The change in plate motions during the Late Jurassic (Chapters 2 and 4) led to the formation 4300 of the DFZ, yet contradictory evidence exists as to whether this occurred sharply by the 4301 cutting of new faults, or gradually by the formation of curved fracture zones. Chapter 2 4302 shows that curved gravity lineaments from the northern and southern regions of the WSB are 4303 best accounted for by a gradual plate rotation. Chapter 4, however, presents evidence for the 4304 crosscutting of a SSE trending fracture zone (the TCBtb) by the DFZ suggesting a sharp 4305 plate motion change in this region. Two important factors allow for these observations to be 4306 reconciled. Firstly, in Chapter 4, a significant amount of internal plate deformation within 4307 the TCB is also observed, which, despite local manifestation as sharp crosscutting indicators 4308 of plate motion, could allow for gradual plate motion changes away from this deformation. 4309 This emphasises the importance of introducing such internal plate deformation (e.g. Eakin et 4310 al., 2015) into future generations of plate tectonic models. Secondly, Chapter 3 contains 4311 suggestions that the recent discovery of Pan African age (533 Ma) zircons within the 4312 xenoliths of Grande Comore may indicate a microcontinent cleaving event during the Late 4313 Jurassic change in plate motions (Roach et al., 2017). Such a cleaving event would allow for 4314 a period of simultaneous ocean spreading along two overlapping MOR systems within the 4315 WSB. Whilst the same total amount of extension in the WSB would result as predicted in 4316 Chapter 2, the gradual reduction in spreading and eventual extinction of the original MOR 4317 system at the expense of increased spreading along the propagating (cleaving) MOR system 4318 would lead to the development of arcuate fracture zones and the abandonment of a triangular 4319 ocean basin (Figure 5.1; e.g. Nunns, 1983). The Comoros Basin, between the Comoros 4320 island chain and the northern margin of Madagascar, has such a triangular geometry (e.g. 4321 Figure 2.4) and this scenario is thus worthy of further investigation. Microcontinent release 4322 during such plate motion changes possibly offers a new mechanism to explain the large 4323 numbers of enigmatic microntinents being discovered within oceanic domains (e.g. 4324 Whittaker et al., 2016). These continental fragments, e.g. the recently interpreted Biera High 4325 microcontinent offshore Mozambique (e.g. Mueller et al., 2016), may offer new prospective 4326 zones for hydrocarbon exploration, and understanding their development is crucial for 4327 locating potential future resource exploration opportunities.

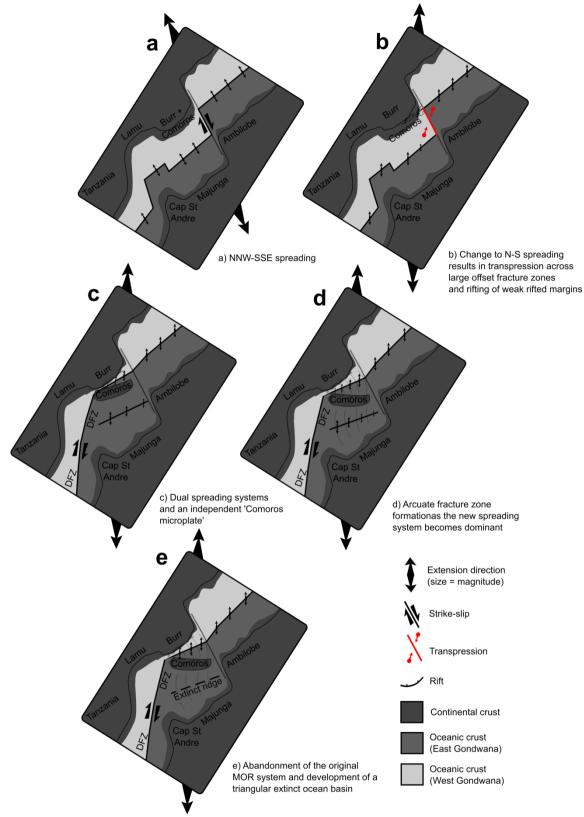


Figure 5.1. Schematic of the proposed 'Comoros Microcontinent' cleaving during the Late
Jurassic change in plate motions. The sudden change in plate motion may form arcuate
fracture zones due to the development of a second spreading axis which gradually takes over
from the first.

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