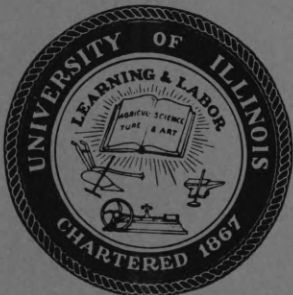




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PASSIVE DAMPING OF THE GENERAL  
RELATIVITY SATELLITE GYRO

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## Abstract

A passive damping method for aligning the instantaneous spin, symmetry and angular momentum axes of a solid, axially symmetric almost spherical satellite gyro is analyzed. Damping is accomplished by the dissipation of energy due to cyclic strains in the gyro body caused by its torque-free precession. Damping time is calculated for a particular gyro design.

PASSIVE DAMPING OF THE GENERAL  
RELATIVITY SATELLITE GYRO

1. Damping Mechanics of Precessing Body

The demands on a gyro spin axis readout system based on a preferred moment of inertia axis are simplified if the gyro symmetry axis  $\omega_3$ , angular momentum vector  $\underline{h}$  and instantaneous spin axis  $\underline{\omega}$  are colinear. In general, when a spinning gyro is suddenly released in free fall, these three axes will not be colinear as shown in Fig. 1, resulting in a torque-free motion of the gyro about its angular momentum vector.<sup>1</sup> Even if these axes were colinear, an environmental disturbance such as a micrometeorite collision could cause cratering, thereby shifting the symmetry axis with respect to the angular momentum axis,<sup>2</sup> resulting in a torque-free motion. For an axially symmetric gyro this motion is a steady precession of its instantaneous spin axis  $\underline{\omega}$  and symmetry axis  $\omega_3$  about its angular momentum axis  $\underline{h}$ , as shown in Fig. 2 from the point of view of an observer fixed in inertial space. In this figure the outer cone (body cone), whose axis is  $\omega_3$ , rolls without slipping on the inner cone (space cone), whose axis is  $\underline{h}$ , and the line of intersection is the instantaneous spin axis  $\underline{\omega}$ . This precession of the symmetry axis  $\omega_3$  about  $\underline{h}$  complicates the readout problem; consequently, a damping mechanism which aligns the three axes within a reasonable time is required. This report analyzes a passive damping scheme in which energy is dissipated by virtue of cyclic strains in the gyro body caused by its torque-free precession.<sup>3</sup> The gyro will be considered an axially symmetric

solid, spherical in shape except for two diametrically opposite flats, which give a preferred moment-of-inertia  $C$  (polar). The moment of inertia about the perpendicular axis is  $A$ . The case for a thin, spherical shell has been analyzed.<sup>4,4a</sup>

The analysis of the combined effects of gravity gradient, centrifugal distortion and the statistics of micrometeorite cratering leads to an optimum gyro diameter of about one foot.<sup>5</sup> For a gyro of this diameter the analysis of micrometeorite cratering<sup>6</sup> gives a relationship between the number of hits per year, each of which could cause an angular disturbance of 0.6 arc sec per year, versus  $(C-A)/C$ . For one hit per year the ratio  $(C-A)/C \approx .01$ . Assuming a Poisson distribution for the meteorite flux, this gives a probability of 0.92 for having one month of undisturbed data. It is important, then, that the damping time be quite smaller than one month in order to separate the effects of such cratering from the spin axis orientation data.

The quantities  $\underline{\omega}_3$ ,  $\underline{h}$ , and  $\underline{\omega}$  in Fig. 1 are coplaner, and the angles  $\theta$  and  $\alpha$  are related by<sup>1</sup>

$$\tan \theta = \frac{A}{C} \tan \alpha \quad (1-1)$$

so that if  $A = C$ , then  $\underline{h}$  and  $\underline{\omega}$  become colinear and  $\underline{\omega}_3$  loses its significance. For  $(C-A)/C = .01$  the angle  $\epsilon$  between  $\underline{h}$  and  $\underline{\omega}$  will be small and is given approximately as  $\epsilon \approx [(C-A)/C] \tan \alpha$ . For example, if  $\alpha = 0.4$  degrees,  $\epsilon \approx 14$  arc sec. During initial gyro spinup attempts will



be made to keep  $\alpha$  as small as possible, but there will be some misalignment error. The maximum allowable error is determined by the tolerance within which the satellite spin axis must lie in its orbital plane, which is of the order of 0.4 degrees for the present gyro parameters.<sup>5</sup> This means that  $\underline{\omega}_3$  must be known with respect to the gyro body to better than 0.4 degrees.

For axially symmetric bodies the rate  $\dot{\Psi}$  at which  $\underline{\omega}_3$  and  $\underline{\omega}$  precess about  $\underline{h}$  for the case of free precession (zero torque) is given by<sup>1</sup>

$$\dot{\Psi} = \frac{C}{A-C} \frac{\dot{\phi}}{\cos\theta} \quad (1-2)$$

where  $\theta$  is the angle between  $\underline{\omega}_3$  and  $\underline{h}$ . The quantity  $\dot{\phi}$ , (later referred to as the elastic vibrating frequency) is the angular rate at which the  $\underline{\omega}$  vector moves about the body as viewed by an observer stationed on the body. Therefore, if the spinning body is centrifugally distorted, an observer stationed along  $\underline{\omega}_3$  will see the body undergo periodic deformation at a fundamental rate  $\dot{\phi}$  corresponding to the rotation rate of the  $\underline{\omega}$  vector about the observer. This is shown rigorously in Section 3. Equation (1-2) can be rewritten with the aid of Fig. 3, which shows the geometrical relation of the involved quantities. Since  $\omega_o = \dot{\Psi} + \dot{\phi}/\cos\theta$  eliminating  $\dot{\Psi}$  gives

$$\dot{\phi} = \frac{A-C}{A} \omega_o \cos\theta \quad (1-3)$$

where  $\omega_o$  is the initial gyro satellite angular velocity.



The periodic deformation of an anelastic body gives rise to a rate of energy dissipation which, among other things, depends upon the fraction of the elastic energy which is dissipated in each deformation or strain cycle.<sup>3</sup> This fraction,  $\gamma$ , is called the hysteretic damping factor, and is a measure of the internal friction of the anelastic body. Metallurgists who measure internal friction usually state results in terms of logarithmic decrement  $D$ , quality factor  $Q$ , or angle  $\delta$  by which strain lags stress.<sup>7</sup> The logarithmic decrement is the logarithm to the base  $e$  of two successive amplitudes of a freely oscillating body. The various factors which measure internal friction are related as follows, for  $Q > 10$ :

$$Q = \frac{2\pi}{\gamma} = \frac{\pi}{D} = \tan\delta . \quad (1-4)$$

The effect of internal energy dissipation is to decrease the angle  $\theta$  between the symmetry axis  $\underline{\omega}_3$  and  $\underline{h}$ , which is shown as follows. The kinetic energy  $T$  of the axially symmetric body in Fig. 1 can be written as

$$T = \frac{1}{2} A(\omega_1^2 + \omega_2^2) + \frac{1}{2} C\omega_3^2 . \quad (1-5)$$

Also,

$$\begin{aligned} \underline{h} &= \hat{i}A\omega_1 + \hat{j}A\omega_2 + \hat{k}C\omega_3 \\ \underline{h} \cdot \underline{h} &= h^2 = A^2(\omega_1^2 + \omega_2^2) + C^2\omega_3^2 . \end{aligned} \quad (1-6)$$

Multiply Eq. (1-5) by  $2A$  and subtract from Eq. (1-6) to get

$$h^2 - 2AT = C(C-A)\omega_3^2.$$

Since  $C\omega_3 = C\omega_0 \cos\theta = h\cos\theta$ , solving for  $T$  gives

$$T = \frac{h^2}{2A} \left[ 1 - \left( \frac{C-A}{C} \right) \cos^2\theta \right].$$

For a finite dissipation, with constant  $\underline{h}$ , the time rate of change of kinetic energy is

$$\dot{T} = \frac{h^2}{A} \left( \frac{C-A}{C} \right) \cos\theta \sin\theta \dot{\theta}. \quad (1-7)$$

For the case  $C > A$ , and for a negative value of  $\dot{T}$  (energy dissipation),  $d\theta/dt$  is negative; therefore,  $\theta$  decreases.

It is shown later that the total gyro elastic strain energy can be classified into two parts. The first part is independent of  $\dot{\phi}$  and is represented as a dc or constant term. Strictly speaking, it is dependent upon  $\theta$  and  $\dot{\theta}$  and hence slowly changing with time, but this change is negligible compared with the second part. The second part of the elastic strain energy varies with time at a rate  $\dot{\phi}$ , and all higher harmonics of  $\dot{\phi}$  up to the fourth. It is this time varying part which is responsible for the hysteretic damping of precession. If we call  $W$  that portion of the gyro elastic strain energy per cycle of stress (whose fundamental frequency is  $\dot{\phi}$ ), then the fraction of this energy which is dissipated per cycle of stress is  $\gamma W$  and the rate of dissipation is  $\gamma W \dot{\phi} / 2\pi$ . This must be equal to the rate of

decrease in kinetic energy  $\dot{T}$  as given by Eq. (1-7). We then have

$$\dot{W} = \frac{\gamma W \dot{\phi}}{2\pi} = -\dot{T} .$$

Substituting  $\dot{\phi}$  from Eq. (1-3), letting  $h = C\omega_0$  and solving for  $\dot{\theta}$ , one obtains

$$\dot{\theta} = \frac{\gamma W}{2\pi C\omega_0 \sin\theta} . \quad (1-8)$$

Sections 2 and 3 describe the method of determining  $W$  for a solid, spherical body with preferred moment-of-inertia axis  $C$ , such that  $C/A \approx 1.01$ . There it will be shown that  $W$  is a function of gyro radius  $a$ , gyro material, spin speed  $\omega_0$  and angle  $\theta$  as in the following equation for small values of  $\theta$ .

$$W = \frac{4\pi^2 \rho^2 \omega_0^4 a^7 \theta^2}{E} \Gamma \quad (1-9)$$

where  $\Gamma$  is a dimensionless quantity which is a function of gyro geometry and material. Substituting this into Eq. (1-8), with  $\sin\theta \approx \theta$  and  $C \approx (2/5)Ma^2 = (8/15)\pi a^5 \rho$  gives

$$\dot{\theta} = \frac{15}{4E} \gamma \rho \omega_0^3 a^2 \Gamma \theta . \quad (1-10)$$

The solution of (1-10) for the damping time  $t$  for an initial angle  $\theta_i$  and a final angle  $\theta_f$  is

$$t = \frac{4E}{15 \gamma \rho \omega_0^3 a^2 \Gamma} \text{Ln} \left( \frac{\theta_f}{\theta_i} \right) . \quad (1-11)$$



It remains to determine  $\Gamma$ , which is obtained from the strain energy  $W$  for a rotating solid sphere in torque-free precession.

## 2. Inertia Force Field

Consider a solid sphere rotating about an instantaneous axis  $\underline{\omega}$  as in Fig. 3. The instantaneous axis of rotation  $\underline{\omega}$  is misaligned with the angular momentum vector  $\underline{h}$  due to some disturbance which has also shifted the symmetry axis  $\underline{\omega}_3$  from the momentum axis  $\underline{h}$  by a small angle  $\theta$  according to Eq. (1-1). If  $\dot{\theta}$  is assumed to be small in comparison with the spin velocity  $\dot{\phi}$  and the precession  $\dot{\psi}$ , then the angular velocity  $\underline{\omega}$  can be written as

$$\underline{\omega} = \hat{i}\dot{\psi} \sin\theta \cos\phi + \hat{j}\dot{\psi} \sin\theta \sin\phi + \hat{k}(\dot{\phi} + \dot{\psi} \cos\theta) \quad (2-1)$$

where  $\theta, \psi, \phi$  are Euler angles defining the orientation of the body axes  $x, y, z$  with respect to the space axes  $X, Y, Z$  as shown in Fig. 4 and  $\hat{i}, \hat{j}, \hat{k}$  are unit vectors along the axes  $x, y, z$  respectively.

The angular acceleration  $\underline{\dot{\omega}}$  is

$$\underline{\dot{\omega}} = \frac{\partial \underline{\omega}}{\partial t} + \underline{\omega} \times \underline{\omega} .$$

Assuming that  $\theta, \dot{\phi}$  and  $\dot{\psi}$  are constant,

$$\underline{\dot{\omega}} = \dot{\phi}\dot{\psi} \sin\theta (-\hat{i} \sin\phi + \hat{j} \cos\phi) .$$

Substituting (2-1) and (2-2) into the equation for linear acceleration



$$\underline{\underline{a}} = \underline{\underline{a}}_0 + \underline{\underline{a}}' + \underline{\underline{\omega}} \times (\underline{\underline{\omega}} \times \underline{\underline{r}}) + \dot{\underline{\underline{\omega}}} \times \underline{\underline{r}} + 2\underline{\underline{\omega}} \times \underline{\underline{v}}' \quad (2-3)$$

and noting the following approximations<sup>1,3</sup>

$$\underline{\underline{a}}_0 = \underline{\underline{a}}' = \underline{\underline{v}}' = 0 \quad (2-4)$$

$$\begin{aligned} \underline{\underline{a}} = \omega_0^2 \hat{i} & \left[ -x \left( \cos^2 \theta + \frac{C^2}{A^2} \sin^2 \theta \sin^2 \phi \right) + y \left( \frac{C}{A} \right)^2 \sin^2 \theta \sin \phi \cos \phi \right. \\ & \left. + z \left( \frac{C}{A} \right) \sin \theta \cos \theta \cos \phi + z \left( 1 - \frac{C}{A} \right) \frac{C}{A} \sin \theta \cos \theta \cos \phi \right] \\ & + \omega_0^2 \hat{j} \left[ x \frac{C^2}{A^2} \sin^2 \theta \sin \phi \cos \phi - y \left( \cos^2 \theta + \frac{C^2}{A^2} \sin^2 \theta \cos^2 \phi \right) \right. \\ & \left. + z \frac{C}{A} \sin \theta \cos \theta \sin \phi + z \left( 1 - \frac{C}{A} \right) \frac{C}{A} \sin \theta \cos \theta \sin \phi \right] \\ & + \omega_0^2 \hat{k} \left[ x \frac{C}{A} \sin \theta \cos \theta \cos \phi + y \frac{C}{A} \sin \theta \cos \theta \sin \phi - z \frac{C^2}{A^2} \sin^2 \theta \right. \\ & \left. - x \left( 1 - \frac{C}{A} \right) \frac{C}{A} \sin \theta \cos \theta \cos \phi + y \left( 1 - \frac{C}{A} \right) \frac{C}{A} \sin \theta \cos \theta \sin \phi \right]. \end{aligned} \quad (2-5)$$

Under the assumption that  $\dot{\theta}$  is negligible compared with  $\dot{\phi}$  and  $\dot{\psi}$ , the only time-varying quantity in (2-5) is  $\phi = \dot{\phi}t$  and the inertia force varies harmonically at a rate  $\dot{\phi}$  and  $2\dot{\phi}$  as can be seen from the above equation.

If it is assumed that the body can be approximated by a homogeneous sphere so that the ratio  $\frac{C}{A}$  is equal to unity, then the above acceleration becomes

$$\begin{aligned}
\underline{a} = & \hat{i}\omega_0^2 \left[ -x(\cos^2\theta + \sin^2\theta \sin^2\phi) + y \sin^2\theta \sin\phi \cos\phi + z \sin\theta \cos\theta \cos\phi \right] \\
& + \hat{j}\omega_0^2 \left[ x \sin^2\theta \sin\phi \cos\phi - y(\cos^2\theta + \sin^2\theta \cos^2\phi) + z \sin\theta \cos\theta \sin\phi \right] \\
& + \hat{k}\omega_0^2 \left[ x \sin\theta \cos\theta \cos\phi + y \sin\theta \cos\theta \sin\phi - z \sin^2\theta \right].
\end{aligned} \tag{2-6}$$

The problem is now reduced to finding the displacement field in a sphere subjected to an inertia force  $\underline{F} = -\rho \underline{a}$ .

A general method of solution for a sphere subjected to body force was given by Chree<sup>8</sup> who, as an example, has worked out the displacement field of a sphere rotating about a diametric axis. In that case the problem becomes axisymmetric but such a symmetry is lost when the body force field is that due to acceleration (2-6) which takes into account the influence of precession of the spin axis. In the following, a brief account of Chree's method and its application to the present non-axisymmetric case will be given.

### 3. Displacement Field in a Sphere Subjected to Inertia Force

Chree's method mentioned above is essentially based on the existence of a body force potential which is expanded in spherical harmonics. Consider Navier's equation of equilibrium

$$(\lambda + \mu)\nabla\nabla \cdot \underline{u} + \mu\nabla^2 \underline{u} + \rho \underline{F} = 0$$

where  $\lambda$  and  $\mu$  are Lamé's constants,  $\underline{u}$  is the displacement and  $\underline{F}$  is the inertia force per unit mass due to the acceleration  $\underline{a}$  ( $= -\underline{F}/\rho$ )

expressed by (2-6). This may be written in indicial notation as

$$(\lambda+\mu)\Delta_{,i} + \mu u_{i,kk} + \rho F_i = 0 \quad (3-1)$$

$\Delta \equiv \nabla \cdot \underline{u}$  is the dilatation,  $P_{,i} \equiv \frac{\partial P}{\partial x_i}$ ,  $i \equiv 1,2,3$  for any function  $P$  and the triad  $x,y,z$  becomes  $x_1, x_2, x_3$ . If  $F_i$  and  $u_i$  are derivable from potentials  $V_n$  (an  $n^{\text{th}}$  order spherical harmonic) and  $\Phi$ , respectively, such that

$$\begin{aligned} F_i &= V_{n,i} \\ u_i &= \Phi_{,i} \end{aligned} \quad (3-2)$$

then (3-1) becomes

$$(\lambda+\mu)\Phi_{,kk} + \rho V_n = 0 \quad (3-3)$$

Because of the identity

$$(r^m V_n)_{,kk} = m(m+2n+1)r^{m-2} V_n \quad (3-4)$$

where  $m$  and  $n$  are positive integers, Eq. (3-4) is satisfied by

$$\Phi = - \frac{\rho}{2(\lambda+2\mu)(2n+3)} r^2 V_n \quad (3-5)$$

The displacement field corresponding to this  $\Phi$  is, from (3-6) and (3-3),

$$u_i = - \frac{\rho}{2(\lambda+2\mu)(2n+3)} (r^2 V_n)_{,i} \quad (3-6)$$



which is accompanied by a surface traction

$$T_i = - \frac{\rho\omega}{r(\lambda+2\mu)} \left[ \frac{n+1}{2n+3} r^2 V_{n,i} + \left\{ \frac{\lambda}{\mu} + \frac{2(n+1)}{2n+3} \right\} x_i V_n \right] \quad (3-7)$$

across any spherical surface  $r = \text{constant}$ . If a body is bounded by the surface  $r = a$  where the surface traction is zero, the displacement  $u_i$  is determined by adding to (3-6) the displacements corresponding to  $\omega$ - and  $\Phi$ -type solutions.<sup>9</sup> This yields surface tractions

$$\frac{rT_i}{\mu} = (2n+\alpha_n) r^2 \omega_{n,i} + \left[ 2n \left( \frac{\lambda}{\mu} + 1 \right) + \alpha_n \left\{ (n+3) \frac{\lambda}{\mu} + (n+2) \right\} \right] x_i \omega_n \quad (3-8)$$

$$\alpha_n = -2 \frac{n\lambda + (3n+1)\mu}{(n+3)\lambda + (n+5)\mu}$$

and

$$\frac{rT_i}{\mu} = 2(n-1)\Phi_{n,i} \quad (3-9)$$

respectively, where, in this case,  $\omega_n$  and  $\Phi_n$  are to be taken as

$$\begin{aligned} \omega_n &= \rho B_1 V_n \\ \Phi_n &= \rho B_2 V_n \end{aligned} \quad (3-10)$$



For the case  $T_i = 0$  on  $r = a$ , these constants  $B_1$  and  $B_2$  are given by

$$B_1 = \frac{\{(2n+3)\lambda + (2n+2)\mu\} \{(n+3)\lambda + (n+5)\mu\}}{2(2n+3)(\lambda+2\mu)\mu \{(2n^2+4n+3)\lambda + 2(n^2+n+1)\mu\}} \quad (3-11)$$

$$B_2 = \frac{n\{(n+2)\lambda + (n+1)\mu\} a^2}{2(n-1)\mu \{(2n^2+4n+3)\lambda + 2(n^2+n+1)\mu\}}$$

where  $n$  is the order of the spherical harmonic,  $V_n$ .

It can be shown easily that the inertia force field due to (2-6) is derivable from the potential

$$V = r^2 V_0 + V_2 \quad (3-12)$$

where  $V_0$  and  $V_2$  are spherical harmonics of zeroth and second order given by

$$V_0 = \frac{\omega_0^2}{3} \quad \text{and} \quad (3-13a)$$

$$\begin{aligned} V_2 = & \frac{\omega_0^2}{6} (3\cos^2\theta + 3\sin^2\theta\sin^2\phi - 2)x^2 \\ & + \frac{\omega_0^2}{6} (3\cos^2\theta + 3\sin^2\theta\cos^2\phi - 2)y^2 \\ & + \frac{\omega_0^2}{6} (3\sin^2\theta - 2)z^2 + \\ & - \omega_0^2 \sin^2\theta \sin\phi \cos\phi xy \\ & - \omega_0^2 \sin\theta \cos\theta \sin\phi yz \\ & - \omega_0^2 \sin\theta \cos\theta \cos\phi xz . \end{aligned} \quad (3-13b)$$

It is interesting to note that the potential given by (3-12) and (3-13) reduces, as it should, to that of a spinning sphere given by Chree<sup>8</sup> when the misalignment  $\theta$  is set to zero.

By substituting (3-12) into (3-6) and superposing the displacements due to  $\omega_n$ - and  $\Phi_n$ -type potentials given by (3-10) and (3-11) to ensure the satisfaction of the boundary condition on  $r = a$ , after considerable amount of algebra, one obtains the following expressions for the components of the displacement.

$$\begin{aligned}
 u_1 = & \left[ \frac{\rho \omega_o^2 a^2 (5\lambda + 6\mu)}{15(\lambda + 2\mu)(3\lambda + 2\mu)} - \frac{2B_2}{3} \omega_o^2 + B_2 \omega_o^2 \cos^2 \theta + B_2 \omega_o^2 \sin^2 \theta \sin^2 \phi \right] x \\
 & - \left[ B_2 \omega_o^2 \sin^2 \theta \sin \phi \cos \phi \right] y - \left[ B_2 \omega_o^2 \sin \theta \cos \theta \cos \phi \right] z \\
 & + \left[ \left\{ -\frac{2}{3} B_1 + \frac{1}{35} L + \frac{B_1 M}{3} \right\} \omega_o^2 \right. \\
 & + \left\{ B_1 - \frac{1}{7} L - \frac{B_1}{2} M \right\} \omega_o^2 \cos^2 \theta \\
 & \left. + \left\{ B_1 - \frac{1}{7} L - \frac{B_1}{2} M \right\} \omega_o^2 \sin^2 \theta \sin^2 \phi \right] x^3 \\
 & + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin^2 \theta \sin \phi \cos \phi y^3 \\
 & + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin \theta \cos \theta \cos \phi z^3 \\
 & + \left[ -B_1 + \frac{3}{14} L + B_1 M \right] \omega_o^2 \sin^2 \theta \sin \phi \cos \phi x^2 y
 \end{aligned}$$

$$\begin{aligned}
& + \left[ -B_1 + \frac{3}{14}L + B_1 M \right] \omega_o^2 \sin\theta \cos\theta \cos\phi \ x^2 z \\
& + \left[ \left\{ -\frac{2}{3}B_1 + \frac{1}{35}L + \frac{B_1}{3}M \right\} \omega_o^2 \right. \\
& + \left\{ B_1 - \frac{1}{7}L - \frac{B_1}{2}M \right\} \omega_o^2 \cos^2\theta - \left\{ \frac{1}{14}L \right\} \omega_o^2 \sin^2\theta \\
& \left. + \left\{ B_1 - \frac{B_1}{2}M \right\} \omega_o^2 \sin^2\theta \cos^2\phi \right] y^2 x \\
& + \left[ -B_1 + \frac{1}{14}L \right] \omega_o^2 \sin^2\theta \cos\theta \cos\phi \ y^2 z \\
& + \left[ \left\{ -\frac{2}{3}B_1 - \frac{3}{70}L + \frac{B_1}{3}M \right\} \omega_o^2 + B_1 \omega_o^2 \cos^2\theta - \frac{B_1}{2}M \omega_o^2 \sin^2\theta \right. \\
& \left. + \left\{ B_1 - \frac{1}{14}L \right\} \omega_o^2 \sin^2\theta \sin^2\phi \right] z^2 x \\
& + \left[ -B_1 + \frac{1}{14}L \right] \omega_o^2 \sin^2\theta \sin\phi \cos\phi \ z^2 y \\
& + \left[ \frac{1}{7}L + B_1 M \right] \omega_o^2 \sin\theta \cos\theta \sin\phi \ xyz . \tag{3-14a}
\end{aligned}$$

$$u_2 = -B_2 \omega_o^2 \sin^2\theta \sin\phi \cos\phi \ x .$$

$$+ \left[ \left\{ -\frac{2}{3}B_2 + \frac{\rho a^2 (5\lambda + 6\mu)}{15(\lambda + 2\mu)(3\lambda + 2\mu)} \right\} \omega_o^2 + B_2 \omega_o^2 \cos^2\theta \right]$$



$$\begin{aligned}
& + B_2 \omega_o^2 \sin^2 \theta \cos^2 \phi \Big] y - B_2 \omega_o^2 \sin \theta \cos \theta \sin \phi z \\
& + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin^2 \theta \sin \phi \cos \phi x^3 \\
& + \left[ \left\{ -\frac{2}{3} B_1 + \frac{1}{35} L + \frac{B_1}{3} M \right\} \omega_o^2 + \left\{ B_1 - \frac{1}{7} L - \frac{B_1}{2} M \right\} \omega_o^2 \cos^2 \theta \right. \\
& + \left. \left\{ B_1 - \frac{1}{7} L - \frac{B_1}{2} M \right\} \omega_o^2 \sin^2 \theta \cos^2 \phi \right] y^3 + \left[ -B_1 + \frac{1}{14} L \right] \times \\
& \quad \omega_o^2 \sin \theta \cos \theta \sin \phi z^3 \\
& + \left[ \left\{ -\frac{2}{3} B_1 + \frac{1}{35} L + \frac{B_1}{3} M \right\} \omega_o^2 + \left\{ B_1 - \frac{1}{7} L - \frac{B_1}{2} M \right\} \omega_o^2 \cos^2 \theta \right. \\
& - \left. \frac{1}{14} L \omega_o^2 \sin^2 \theta \right. \\
& + \left. \left\{ B_1 - \frac{B_1}{2} M \right\} \omega_o^2 \sin^2 \theta \sin^2 \phi \right] x^2 y \\
& + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin \theta \cos \theta \sin \phi x^2 z \\
& + \left[ -B_1 + \frac{3}{14} L + B_1 M \right] \omega_o^2 \sin^2 \theta \sin \phi \cos \phi y^2 x \\
& + \left[ -B_1 + \frac{3}{14} L + B_1 M \right] \omega_o^2 \sin \theta \cos \theta \sin \phi y^2 z \\
& + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin^2 \theta \sin \phi \cos \phi z^2 x
\end{aligned}$$



$$\begin{aligned}
& + \left[ \left\{ -\frac{2}{3} B_1 - \frac{3}{70} L + \frac{B_1}{3} \right\} \omega_o^2 + B_1 \omega_o^2 \cos^2 \theta \right. \\
& + \left. \left\{ B_1 - \frac{1}{14} L \right\} \omega_o^2 \sin^2 \theta \cos^2 \phi \right] z^2 y \\
& + \left[ \frac{1}{7} L + B_1 M \right] \omega_o^2 \sin \theta \cos \theta \cos \phi \quad xyz . \tag{3-14b}
\end{aligned}$$

$$\begin{aligned}
u_3 = & - B_2 \omega_o^2 \sin \theta \cos \theta \cos \phi \quad x - B_2 \omega_o^2 \sin \theta \cos \theta \sin \phi \quad y \\
& + \left[ \left\{ \frac{B_2}{3} + \frac{\rho a^2 (5\lambda + 6\mu)}{15(\lambda + 2\mu)(3\lambda + 2\mu)} \right\} \omega_o^2 - B_2 \omega_o^2 \cos^2 \theta \right] z \\
& + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin \theta \cos \theta \cos \phi \quad x^3 + \left[ -B_1 + \frac{1}{14} L \right] \times \\
& \quad \omega_o^2 \sin \theta \cos \theta \sin \phi \quad y^3 \\
& + \left[ \left\{ -\frac{12}{15} L + \frac{B_1}{3} M \right\} \omega_o^2 + \left\{ \frac{1}{7} L - B_1 M \right\} \omega_o^2 \cos^2 \theta \right] z^3 \\
& + \left[ -B_1 + \frac{1}{14} L \right] \omega_o^2 \sin \theta \cos \theta \sin \phi \quad x^2 y \\
& + \left[ \left\{ \frac{B_1}{3} - \frac{3}{70} L + \frac{B_1}{3} M \right\} \omega_o^2 - \left\{ B_1 + \frac{B_1}{2} M \right\} \omega_o^2 \cos^2 \theta \right. \\
& \left. - \left\{ \frac{1}{14} L + \frac{B_1}{2} M \right\} \omega_o^2 \sin^2 \theta \sin^2 \phi \right] x^2 z
\end{aligned}$$

$$\begin{aligned}
& + \left[ B_1 + \frac{1}{14} L \right] \omega_o^2 \sin\theta \cos\theta \cos\phi \ y^2 x \\
& + \left[ \left\{ \frac{B_1}{3} - \frac{3}{70} L + \frac{B_1}{3} M \right\} \omega_o^2 - \left\{ B_1 + \frac{B_1}{2} M \right\} \omega_o^2 \cos^2\theta \right. \\
& \quad \left. - \left\{ \frac{1}{14} L + \frac{B_1}{2} M \right\} \omega_o^2 \sin^2\theta \cos^2\phi \right] y^2 z \\
& + \left[ -B_1 + \frac{3}{14} L + B_1 M \right] \omega_o^2 \sin\theta \cos\theta \cos\phi \ z^2 x \\
& + \left[ -B_1 + \frac{3}{14} L + B_1 M \right] \omega_o^2 \sin\theta \cos\theta \sin\phi \ z^2 y \\
& + \left[ \frac{1}{7} L + B_1 M \right] \omega_o^2 \sin\theta \sin\phi \cos\phi \ xyz . \tag{3-14c}
\end{aligned}$$

where

$$L = \frac{1}{\lambda + 2\mu}$$

$$M = \frac{4\lambda + 14\mu}{5\lambda + 7\mu} . \tag{3-15}$$

Terms in these components of displacement can be classified into three categories, the first group being the steady part which does not depend on  $\phi$  at all, the second those which vary with time at a rate  $\dot{\phi}$ , and lastly those which pulsate at a rate  $2\dot{\phi}$ . The last two categories which vary with time are responsible for the hysteretic damping of precession.

#### 4. Elastic Strain Energy of Solid Sphere

The time varying part of the elastic strain energy can be computed by first taking one half the dot product of the force field  $-\rho \underline{a}$ , where  $\underline{a}$  is given by (2-6), and the displacement field  $\underline{u}$ , given by (3-15), to give the strain energy density  $W'$  such that

$$W' = \frac{1}{2}(-\rho a_i)u_i \quad (4-1)$$

and then dropping all the steady terms which are independent of  $\phi$ . The amount of alternating strain energy  $W$  per cycle of precession is

$$W = \int_0^{2\pi} \int_{\Omega} W' d\Omega d\phi \quad (4-2)$$

where  $\Omega$  is the volume of the sphere. By substituting (2-6) and (3-14) into (4-1) and (4-2), and dropping terms of the order of  $\sin^4\theta$  in comparison with those proportional to  $\sin^2\theta$ , it is not hard, although tedious, to obtain

$$W = 4\pi^2 \rho^2 \omega_o^4 a^7 \theta^2 \left\{ \frac{B_2}{5a^2} + \frac{20}{63} B_1 - \frac{9}{980} L - \frac{19}{630} B_1 M \right\} \quad (4-3)$$

as the fundamental component of the alternating part of elastic strain energy per unit cycle of precession. Designating the quantity within brackets as  $\Gamma'$ , one has

$$\Gamma' = \frac{B_2}{5a^2} + \frac{20B_1}{63} - \frac{9}{980} L - \frac{19}{630} B_1 M . \quad (4-4)$$



Equation (4-4) may be simplified by first substituting  $\lambda$  and  $\mu$  into the expressions for L and M of (3-15) and  $B_1$  and  $B_2$  of (3-11), to get

$$L = \frac{(1+\nu)(1-2\nu)}{E(1-\nu)}$$

$$M = \frac{2(7+10\nu)}{7-4\nu},$$

and

$$B_1 = \frac{-(3+\nu)(7-4\nu)(1+\nu)}{7E(1-\nu)(7+5\nu)}$$

$$B_2 = \frac{(3+2\nu)(1+\nu)a^2}{E(7+5\nu)}$$

for  $n = 2$ .

These values of L, M,  $B_1$  and  $B_2$  are substituted into (4-4) to get

$$\begin{aligned} \Gamma' &= \frac{1+\nu}{E} \left\{ \frac{3+2\nu}{5(7+5\nu)} - \frac{9(1-2\nu)}{980(1-\nu)} + \frac{(3+\nu)(7-4\nu)}{(1-\nu)(7+5\nu)} \left[ \frac{19(7-10\nu)}{2205(7-4\nu)} - \frac{20}{441} \right] \right\} \\ &= \frac{\Gamma}{E} \end{aligned} \quad (4-5)$$

which is now only in terms of Poisson's ratio  $\nu$  and Young's modulus E. Equation (4-3) may be rewritten as

$$W = 4\pi^2 \rho^2 \omega_o^4 a^7 \theta^2 \frac{\Gamma}{E}$$

where  $\Gamma$  is a function of  $\nu$  only. This value of  $\Gamma$ , substituted in (1-11) determines the damping time  $t$  for a given gyro.

### 5. Numerical Calculations

The expression for damping time (1-11) may be rewritten in the form

$$t = \frac{4}{15} \frac{E}{\gamma \omega_o \Gamma(\rho \omega_o^2 a^2)} \cdot \text{Ln} \left( \frac{\theta_f}{\theta_i} \right) \quad (5-1)$$

From elasticity theory, the maximum stress at the center of a solid, spinning sphere is approximately given by<sup>8</sup>

$$\sigma_{\max} = \rho \omega_o^2 a^2 \left( \frac{3 + 2\nu}{7 + 5\nu} \right) \quad (5-2)$$

where  $\nu$  is Poisson's ratio. Hence, the term  $\rho \omega_o^2 a^2$  in the denominator of (5-1) is proportional to the maximum allowable stress for the selected gyro material. Since the optimization study referred to in Section 1 fixes the value of  $a$  to be approximately six inches,  $\omega_o$  for a given material follows from equation (5-2), giving due allowance for a safety factor.

Volume electrical resistivity requirements severely restrict the choice of materials to those between the good conductors ( $\rho > 10^3$  ohm cm) and good insulators ( $\rho < 10^{10}$  ohm cm). The materials germanium, silicon and titanium dioxide, when properly doped, and certain glasses are among those which appear to satisfy the electrical resistivity requirements. At present, the hysteretic damping factor  $\gamma$  has been obtained only for certain glasses. Among other things,

$1/Q$  is a function of the elastic vibrating frequency  $\dot{\phi}$  given by equation (1-3). Present gyro parameters indicate a vibration frequency of about 1 to 3 cps. Fortunately, glass has a maximum value for  $1/Q (= \gamma/2\pi)$  of about  $4 \times 10^{-3}$  in this range of frequencies<sup>7,11,12</sup> at room temperature. This makes it a promising high damping factor material.

In order to determine experimentally the approximate value of  $\sigma_{\max}$ , thin glass disks were spun to the bursting point in a motor-driven test fixture. For such disks, the maximum stress is also at the center and is, to good approximation

$$\sigma_{\max} = \rho \omega^2 a^2 \left( \frac{3 + \nu}{8} \right). \quad (5-3)$$

For glass,  $\nu \approx 0.16$  so that the bracketed factors in (5-2) and (5-3) become 0.425 and 0.395 for the sphere and disk, respectively. Hence, the maximum stresses are nearly the same for identical materials, diameters, and spin speed, therefore justifying the use of disks for this test. These tests indicated an upper value of  $\omega_0$  of about 630 rad/sec for plate glass, with an adequate safety factor. For  $C/A = 1.01$  and  $\theta$  of the order of half degree, the elastic vibrating frequency  $\dot{\phi}$  is about one cps, as seen from equation (1-3).

Using glass as the gyro material, the following parameters have been determined:



$$\gamma = \frac{2\pi}{Q} = .025$$

$$\rho = 2.5 \times 10^3 \text{ kgm/m}^2$$

$$\omega_o = 630 \text{ rad/sec}$$

$$a = 7.5 \text{ cm} = .075 \text{ m}$$

$$\theta_f = 0.1 \text{ arc sec}$$

$$\theta_i = 0.5 \text{ degree}$$

$$E = 7 \times 10^{10} \text{ newtons/m}^2$$

$$\nu = 0.16$$

The value of  $\Gamma$  from equation (4-5) becomes  $\Gamma = -0.071$ . Substituting this and the above values into equation (1-11), the damping time  $t$  becomes 8.2 hrs, a reasonable time.

## 6. Conclusion

Using glass as a possible gyro material, the passive damping method for aligning the gyro instantaneous spin axis, angular momentum axis and symmetry axis has been shown to be feasible, requiring about 8½ hours to damp from  $\theta = 0.5$  degree to  $\theta = 0.1$  arc sec. The damping time constant  $\tau = 0.83$  hours. This value of damping time is probably required only during the initial gyro spin up. The statistics of micrometeorite collisions with this gyro show that

there will be a probability of 0.92 for no collisions within the period of one month which could cause an angular disturbance of 0.6 arc sec per year. The effect of such a collision, however, would require only  $(0.83) (1.8) = 1.5$  hours, a reasonable time.

#### 7. Acknowledgments

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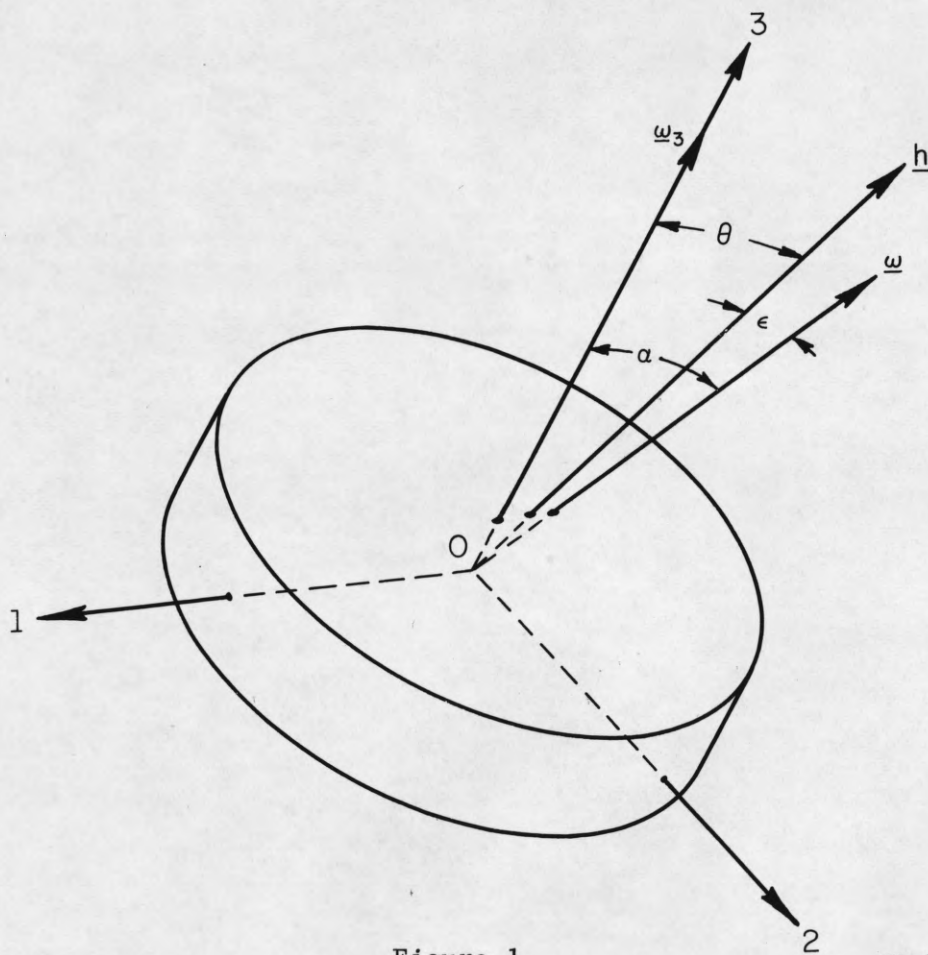


Figure 1.

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General Relationship of Angular Momentum, Symmetry  
and Instantaneous Spin Axes of Axially Symmetric Gyro.

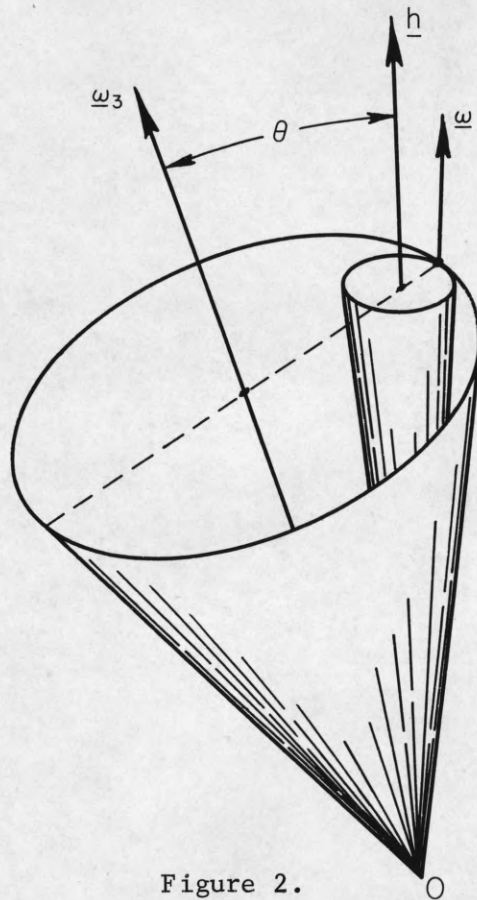


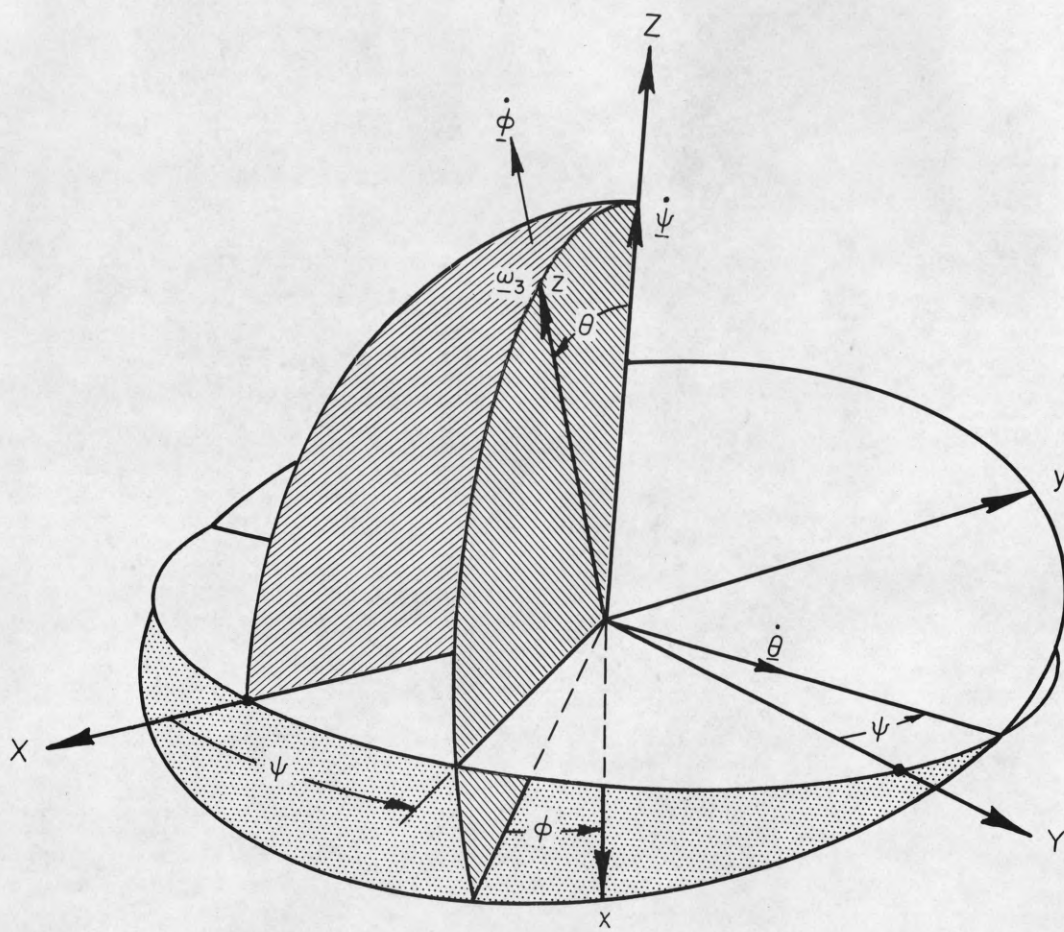
Figure 2.

Body and Space Cones of Axially Symmetric Gyro in Torque-Free Precession.

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Figure 4.  
 Satellite Gyro Coordinates  
 with Respect to an Inertial Frame.

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