Microwave Spectral Taxonomy and Astronomical Searches for Vibrationally-Excited C₂S and C₃S



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THE ROTATIONAL EMISSION-LINE SPECTRUM OF ORION A BETWEEN 247 AND 263 GHZ

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~250,000 FIDs





~250,000 FIDs





~250,000 FIDs





Microwave spectral taxonomy: A semi-automated combination of chirped-pulse and cavity Fourier-transform microwave spectroscopy

Kyle N. Crabtree,¹ Marie-Aline Martin-Drumel,² Gordon G. Brown,³ Sydney A. Gaster,³ Taylor M. Hall,³ and Michael C. McCarthy^{2,a)}





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MM-WAVE SPECTRA







Evidence for: $2v_1$, $v_1 + v_2$, v_3 in cm-wave











C₃S VIBRATIONAL TEMPERATURES



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Non-Detections

Quantitative Search PRIMOS, Belloche IRAM 30m, HEXOS Upper Limits (10¹³ cm⁻²)

Т (К)	CCS v ₂	CCS v ₁	C ₃ S v ₃	C ₃ S 2v ₄	C ₃ S 1v ₅
100	15.5	15.7	4.2	4.1	8.1
150	21.1	21.5	5.2	5.1	10.1
200	26.2	27.1	6.3	6.2	12.3

Qualitative Search

TMC-1 (9 - 50 GHz) Kaifu et al. (2004)

IRC+10216 (18 - 26 GHz) Gong et al. (2015)

IRC+10216 (215 - 285 GHz) Tenenbaum et al. (2010)

> Orion KL (HEXOS) Crockett et al. (2014)

CCS: 20 K, 8.3 x 10¹³ cm⁻² C₃S: Present, but strange

Half a Decade of ALMA: Cosmic Dawns Transformed

20-23 September 2016, Indian Wells, CA (near Palm Springs) To register and submit an abstract (now open!) visit: http:// go.nrao.edu/ALMA5years

Invited Talks & Speakers

Galaxy Formation and Evolution I: Cosmic Evolution (Caitlin Casey) Galaxy Formation and Evolution II: Gas & Star Formation Properties (Linda Tacconi) Galactic Centers: Star Formation, AGN, Black Holes & ULIRGs (Masatoshi Imanishi) Nearby Galaxies I: Normal Galaxies (Karin Sandstrom) Nearby Galaxies II: Starburst & Super Star Clusters (Kazushi Sakamoto) Massive Star Formation (Jill Rathborne) Low Mass Star Formation (Adele Plunkett) Chemical Evolution During Star and Planet Formation (Jeong-Eun Lee) Protostellar Disks & Planet Formation (Laura Perez) Debris Disks (Brenda Matthews) Stars and Stellar Evolution (Leen Decin) Solar System (Arielle Moullet) Synergy between ALMA and JWST (Klaus Pontoppidan) ALMA after 5 Years (Pierre Cox) Future ALMA (John Carpenter, Al Wootten, Neal Evans) Conference Summary (Anneila Sargent)









 $CO(v_1) + CO(v_3) \longrightarrow 2CO(v_2)$

Endothermic

Requires energy input into the system, due to anharmonicity







 $CO(v_1) + CO(v_3) - 2CO(v_2)$

Exothermic Energy *leaves* the system, due to anharmonicity

Energy



$CO(v_1) + CO(v_3) \implies 2CO(v_2)$

Exothermic Energy *leaves* the system, due to anharmonicity



