

### <u>F. Kwabia Tchana</u><sup>1</sup>, M. Ndao<sup>1</sup>, L. Manceron<sup>2</sup>, A. Perrin<sup>1</sup>, J. M. Flaud<sup>1</sup>, W.J. Lafferty<sup>3</sup>

<sup>1</sup>Laboratoire Interuniversitaire des Systèmes Atmosphériques, UMR CNRS 7583, Université Paris Diderot, France <sup>2</sup>Ligne AILES, Synchrotron SOLEIL and MONARIS, CNRS UMR 8233, Université Pierre et Marie Curie, France <sup>3</sup>Sensor Science Division, National Institute of Standards and Technology, USA



### Modes of vibration of phosgene <sup>35</sup>Cl<sub>2</sub>CO



Phosgene has 6 IR modes, 2 in the mid-infrared ( $v_1$  and  $v_5$ ) and 4 low energy modes:  $v_3$ ,  $v_6$ ,  $v_2$  and  $v_4$ 

Lots of strong hot bands in the 11.8 μm spectral region where the phosgene is detected and modelled

These hot bands complicate the analysis of spectra !!

### Detection of atmospheric Cl<sub>2</sub>CO: Use of the atmospheric window around 11 μm

- Phosgene is relatively more abundant in the stratosphere, where it has a lifetime of several years, but is also present in the troposphere in spite of a shorter lifetime (seventy days)
- Strong infrared absorption of phosgene (v<sub>5</sub>) occur in the same spectral region (850 cm<sup>-1</sup>/11.8 μm) as <sup>\$</sup>Freon-11 (CCl<sub>3</sub>F), which leads to an overestimation of the concentration of Freon-11, if one does not take into account the absorption of phosgene as can be seen on the residual of this figure

\$ Toon et al., Geophys. Res. Lett., 28 (2001) 2835

### Detection of atmospheric Cl<sub>2</sub>CO: Use of the atmospheric window around 11 μm

- Phosgene is relatively more abundant in the stratosphere, where it has a lifetime of several years, but is also present in the troposphere in spite of a shorter lifetime (seventy days)
- Strong infrared absorption of phosgene (v<sub>5</sub>) occur in the same spectral region (850 cm<sup>-1</sup>/11.8 μm) as <sup>\$</sup>Freon-11 (CCl<sub>3</sub>F), which leads to an overestimation of the concentration of Freon-11, if one does not take into account the absorption of phosgene as can be seen on the residual of this figure



Spectral fit to a MkIV limb transmittance spectrum at 15.36 km tangent altitude **performed** without absorption of the  $v_5$ band of atmospheric Cl<sub>2</sub>CO

<sup>\$</sup>Toon et al., Geophys. Res. Lett., 28 (2001) 2835

### Modelling of atmospheric Cl<sub>2</sub>CO: Performed with a linelist simulating the absorption of the v<sub>5</sub> band of atmospheric Cl<sub>2</sub>CO



Precise modelling of phosgene absorptions in this infrared atmosoheric windows requires the study of the  $v_5$  band but also the low vibrational energy such as:  $v_3$  (285 cm<sup>-1</sup>) and  $v_6$  (440 cm<sup>-1</sup>) bands.

#### Toon et al., Geophys. Res. Lett., 28 (2001) 2835

## Modelling of atmospheric $Cl_2CO$ : Performed with a linelist simulating the absorption of the $v_5$ band of atmospheric $Cl_2CO$



Precise modelling of phosgene absorptions in this infrared atmosoheric windows requires the study of the  $v_5$  band but also the low vibrational energy such as:  $v_3$  (285 cm<sup>-1</sup>) and  $v_6$  (440 cm<sup>-1</sup>) bands.

The omission of these hot bands leads to a systematic error of more than 20% in the retrieved profiles.

Toon et al., Geophys. Res. Lett., 28 (2001) 2835

Indeed these far infrared fundamentals are responsible for hot bands ( $v_5 + v_3 - v_3$  and  $v_5 + v_6 - v_6$ ,...), not analysed but of great importance for the correct retrieval of Freon-11.

## Previous study

 Microwave spectra: determination of phosgene structure and ground state parameters

Wilse and Robinson, *J. Chem. Phys.* **21** (1953) 1741 Mirri et *al., Spectro. Chem. Acta A* **27** (1971) 937 Carpenter and Rimmer, *J. Chem. Soc. Faraday Trans.* 2 **74** (1978) 466 Nakata et *al., J. Mol. Spectrosc.* **83** (1980) 105 Nakata et *al., J. Mol. Spectrosc.* **83** (1980) 118

• Tunable diode laser spectrum: very partial study of the strong  $v_1$  and  $v_5$  bands

Yamamoto et al., J. Mol. Spectrosc. 106 (1984) 376

 High resolution Fourier transform spectra: detailed and extensive analysis of the strong v<sub>1</sub> and v<sub>5</sub> bands
Kwabia Tchana et al., Molecular Physics 113 (2015) 3241

## Present study

- We recorded the high resolution (0.00102 cm<sup>-1</sup>) FT far-infrared spectra of phosgene: <sup>35</sup>Cl<sub>2</sub>CO (57%), <sup>35</sup>Cl<sup>37</sup>ClCO (37%) and <sup>37</sup>Cl<sub>2</sub>CO (6%)
- First line position analysis of the v<sub>3</sub> (285 cm<sup>-1</sup>) and v<sub>6</sub> (440 cm<sup>-1</sup>) bands of the most abundant isotopomers <sup>35</sup>Cl<sub>2</sub>CO and <sup>35</sup>Cl<sup>37</sup>CICO



Experimental setup: Synchrotron source coupled to the high resolution Fourier transform spectrometer and long-path cryogenic cell



# **Experimental setup**: Synchrotron source coupled to the high resolution Fourier transform spectrometer and **long-path cryogenic cell**



Four compartment chamber

Sample gas

- Convecteur-cooling gas (He)
- Liquid Nitrogen reservoir
- Insulating vacuum

For this experiment, the spectrum was recorded at a regulated temperature of 197 K, an optical path of 93.14 m and phosgene pressure of 1.06 hPa



Watson-type Hamiltonian used to calculate the upper states ro-vibrational 3<sup>1</sup> and 6<sup>1</sup> of phosgene

The upper state rotational constants were obtained using a Watson A-type Hamiltonian written in the I<sup>r</sup> representation (x = b, y = c and z = a)

$$\begin{split} H_{W} &= E_{v} + \left[A^{v} - \frac{1}{2} \left(B^{v} + C^{v}\right)\right] J_{z}^{2} + \frac{1}{2} \left(B^{v} + C^{v}\right) J^{2} + \frac{1}{2} \left(B^{v} - C^{v}\right) J_{xy}^{2} \\ &- \Delta_{K}^{v} J_{z}^{4} - \Delta_{JK}^{v} J_{z}^{2} J^{2} - \Delta_{J}^{v} \left(J^{2}\right)^{2} - \delta_{K}^{v} \left\{J_{z}^{2}, J_{xy}^{2}\right\} - 2\delta_{J}^{v} J_{xy}^{2} J^{2} \\ &+ H_{K}^{v} J_{z}^{6} + H_{KJ}^{v} J_{z}^{4} J^{2} + H_{JK}^{v} J_{z}^{2} \left(J^{2}\right)^{2} + H_{J}^{v} \left(J^{2}\right)^{3} \\ &+ h_{K}^{v} \left\{J_{z}^{4}, J_{xy}^{2}\right\} + h_{KJ}^{v} \left\{J_{z}^{2}, J_{xy}^{2}\right\} J^{2} + 2h_{J}^{v} J_{xy}^{2} \left(J^{2}\right)^{2} + \dots \end{split}$$

Energy level is defined by the quantum numbers J, K<sub>a</sub> and K<sub>c</sub>

$$0 \le K_a \le J, 0 \le K_c \le J, K_a + K_c = J \text{ or } J + 1$$

## Results of the fit for the v<sub>3</sub> bands of <sup>35</sup>Cl<sub>2</sub>CO and <sup>35</sup>Cl<sup>37</sup>ClCO

Range of quantum numbers observed for experimental energy levels, number of lines and RMS of the vibrational state  $v_3 = 1$ 

Isotopomers	J range	K <sub>a</sub> range	No. of lines	RMS (cm <sup>-1</sup> )
<sup>35</sup> Cl <sub>2</sub> CO	2-78	0-41	5032	0.00021
2				
<sup>35</sup> Cl <sup>37</sup> ClCO	2-75	0-41	4634	0.00038

For the two isotopomers we have reproduced more than 9000 transitions ( $J_{max}$  = 78,  $K_{amax}$  = 41), with an RMS better than 0.0004 cm<sup>-1</sup>

## Results of the fit for the v<sub>6</sub> bands of <sup>35</sup>Cl<sub>2</sub>CO and <sup>35</sup>Cl<sup>37</sup>ClCO

Range of quantum numbers observed for experimental energy levels, number of lines and RMS of the vibrational state  $v_6 = 1$ 

(cm <sup>-1</sup> )
0030
0043

For the two isotopomers we have reproduced more than 9000 transitions ( $J_{max}$  = 77,  $K_{amax}$  = 34), with an RMS better than 0.00045 cm<sup>-1</sup>

# Comparaison of spectra: observed and simulated spectra, $v_3$ band of $Cl_2CO$



# Portion of the R-branches of the v<sub>3</sub> bands of <sup>35</sup>Cl<sub>2</sub>CO and <sup>35</sup>Cl<sup>37</sup>ClCO





# Portion of the R-branches of the v<sub>6</sub> bands of <sup>35</sup>Cl<sub>2</sub>CO and <sup>35</sup>Cl<sup>37</sup>ClCO



## The good agreement between observation and simulation demonstrates the quality of the analysis and the fitting

### Conclusions and Outlook

- First high-resolution infrared absorption spectra of Cl<sub>2</sub>CO isotopomers: <sup>35</sup>Cl<sub>2</sub>CO and <sup>35</sup>Cl<sup>37</sup>ClCO between 250 to 480 cm<sup>-1</sup>.
- First determination of the upper-state rotationnal and distortion constants for  $v_3$  and  $v_6$  bands.
- We also recorded the  $v_2$  (567 cm<sup>-1</sup>) and  $v_4$  (580 cm<sup>-1</sup>) bands of  ${}^{35}Cl_2CO$  and  ${}^{35}Cl^{37}CICO$ . The analysis of these two bands is in progress.
- Outlook: Study the hots bands, measuring the intensities and the widths of the mid-infrared band of Cl<sub>2</sub>CO as a function of temperature. This is important for solving the 40 % discrepancies between the existing room temperature measurements, and to clarify the contribution of hot bands.
- **Objective:** Provide full prediction including intensities and linewidths for remote sensing in the 11.8  $\mu$ m spectral region.

### END

### Thanks you for your attention!!

### Specifications of the LISA-SOLEIL long-path cryogenic cell



Convecteur-cooling gas, He (3)

Liquid Nitrogen reservoir (2)

Insulating vacuum (1)

- Variable optical path :
- 3 to more than 141 m
- Spectral range 10 to 4000 cm<sup>-1</sup>
- Temperature adjustable in the **80 to 400 K** range, within  $\pm$  2 K
- Cryogenic gauge 0.2 to 40 mb, measured with  $\pm 1\%$

#### Four compartment chamber

Sample gas (4)

• Convecteur-cooling gas, He (3)

• Liquid Nitrogen reservoir (2)

• Insulating vacuum (1)

We chose a concept with a completely static configuration (no forced circulation pump, no closed cycle cooler) including a large cryostat around the cell body and additional gas convection cooling. The cooling power originates from the heat of vaporization of liquid nitrogen, giving off approximately 69 W of cooling power per liter of liquid nitrogen evaporated per hour, even at the lowest temperature, and thus keeping chamber 2 at a constant 77 K temperature. The heat is transmitted through radiative cooling to the inner envelopes or, more efficiently, through convection by filling the chamber 3 with helium gas. The helium pressure can be varied to adapt the cooling power to the desired end temperature.

Atmospheric Cl<sub>2</sub>CO

Atmospheric  $COCl_2$  is formed from the breakdown of chlorinated hydrocarbons such as CCl<sub>4</sub>, CH<sub>3</sub>CCl<sub>3</sub>, CHCl<sub>3</sub>, C<sub>2</sub>Cl<sub>4</sub>, and C<sub>2</sub>HCl<sub>3</sub> [e.g. Helas and Wilson, 1992; Kindler et al., 1995]. It is believed not to have any significant source from chlorofluorocarbons, which break down to COF<sub>2</sub> or COCIF due to the greater strength of the C-F bond as compared with the C-Cl bond [Sen et al., 1996]. COCl<sub>2</sub> has both a tropospheric source (OH-initiated oxidation of these parent compounds) and a stratospheric source (mainly UV photolysis of CCl<sub>4</sub>).

Toon et al., Geophys. Res. Lett., 28 (2001) 2835



## Phosgene production: Triphosgene was heated to generate phosgene

 $C_3CI_6O_3(s) \xrightarrow{\Delta_{sub}H} 3 \times (CI_2CO)(g)$ 



Phosgene production begins at a temperature of 80°C, below which the gas is not produced. The reaction proceeds cleanly up to 110°C at a steady rate

> The gas is trapped in the liquid nitrogen

> Pumping of the not condensable impurities (CO,  $CO_2$ ,...)

## Homogeneity of the temperature along the entire optical path length



FIG. 5. Temperature (°C) profile vs. position along the cell body for the SOLEIL-LISA Cryocell at nine different cooling temperatures near 293, 274, 242, 224, 209, 170, 127, 106, and 84 K (average temperatures). The temperature sensor indexes are in parenthesis. Sensor 5 is slightly outside the optical path. The numbers on the right indicate average temperatures and  $1\sigma$  deviations.

Along the cell body five pairs of flexible heaters (500 W, Kapton encapsulated NiCr) were glued with Stycast 2850FT and connected to five separate temperature controllers. This allows a fine adjustment of the temperature and compensations of the small differences between center and extremities of the cell body. The regulation is based on five Pt100 class A temperature sensors, with stated accuracy of  $\pm 0.1^{\circ}C$