LASER SPECTROSCOPIC STUDY OF CaH IN THE B $^{2}\Sigma^{+}$ AND D $^{2}\Sigma^{+}$ STATES

<u>KYOHEI WATANABE</u>, KANAKO UCHIDA, KAORI KOBAYASHI, FUSAKAZU MATSUSHIMA, YOSHIKI MORIWAKI, *Department of Physics, University of Toyama, Toyama, Japan*.

Calcium hydride is one of the abundant molecules in the stellar environment, and is considered as a probe of stellar analysis^{*a*}. Ab initio calculations have shown that the electronic excited states of CaH have complex potential curves. It is suggested that the $B^{2}\Sigma^{+}$ state has an interesting double minimum potential due to the avoided crossing^{*b*}. Such a potential leads to drastic change of the rotational constants when the vibrational energy level goes across the potential barrier. Spectroscopic studies on CaH began in the 1920's^{*c*}, and many studies have been carried out since then. Bell et al. extensively assigned the $D^{2}\Sigma^{+}-X^{2}\Sigma^{+}$ bands in the UV region^{*d*}. Bernath's group has observed transitions in the IR and visible regions and identified their upper states as the $A^{2}\Sigma^{+}$, $B^{2}\Sigma^{+}$ and $E^{2}\Sigma^{+}$ states^{*efgh*}. We have carried out a laser induced fluorescence (LIF) study in the UV region between 360 and 430 nm. We have produced CaH by using laser ablation of a calcium target in a hydrogen gas environment, then molecules have been excited by a second harmonic pulse of dye laser and the fluorescence from molecules have been detected through a monochromator. Detection of the $D^{2}\Sigma^{+}$ - $X^{2}\Sigma^{+}$ bands already identified by Bell et al. indicates the production of CaH. In addition, many other bands have been also found and a few bands have been assigned by using the combination differences, the lower state of these bands have been confirmed to the vibrational ground state of $X^{2}\Sigma^{+}$ state. We have tentatively assigned these bands as the $B^{2}\Sigma^{+}-X^{2}\Sigma^{+}$ transition. We will discuss the assignment of these bands, together with the rotational constants comparing with those calculated from the ab initio potential.

- ^fR. S. Ram, K. Tereszchuk, I. E. Gordon, K. A. Walker, and P. F. Bernath, J. Mol. Spec. 266, 86 (2011).
- ^gG. Li, J. J. Harrison, R. S. Ram, C. M. Western, and P. F. Bernath Quant. Spectrosc. Rad. Transfer. 113, 67 (2012).
- ^hA. Shayesteh, R. S. Ram, and P. F. Bernath, J. Mol. Spec. 288, 46 (2013).

^aB. Barbuy, R. P. Schiavon, J. Gregorio-Hetem, P. D. Singh C. Batalha, Astron. Astrophys. Sippl. Ser. 101, 409 (1993).

^bP. F. Weck and P. C. Stabcil, J. Chem. Phys. 118, 9997 (2003).

^cR. S. Mulliken, Phys. Rev. 25, 509 (1925).

^dG. D. Bell, M, Herman, J. W. C. Johns, and E. R. Peck, *Physica Scripta* **20**, 609 (1979).

^eA. Shayesteh, K. A. Walker, I. Gordon, D. R. T. Appadoo, and P. F. Bernath, J. Mol. Struct. 695-696, 23 (2004).