

# *Monsoons: global energetics and local physics as drivers of past, present and future monsoons*

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# 1 **Monsoons: global energetics and local physics**

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26 **Global constraints on momentum and energy govern the structure of the zonal mean trop-**  
27 **ical circulation and rainfall. The continental-scale monsoon systems are also facets of a**  
28 **momentum- and energy-constrained global circulation, but their modern and paleo vari-**  
29 **ability deviates substantially from that of the longitudinal mean through mechanisms neither**  
30 **fully understood nor well simulated. A framework grounded in global constraints yet encom-**  
31 **passing the complexities of monsoon dynamics is needed to identify the causes of mismatch**  
32 **between theory, models, and observations and, ultimately, improve regional climate projec-**  
33 **tion. In a first step towards this goal, disparate regional processes must be distilled into gross**  
34 **measures of energy flow in and out of continents and from the surface to the tropopause, so**  
35 **that monsoon dynamics may be coherently diagnosed across modern and paleo observations**  
36 **and across idealized and comprehensive simulations. Accounting for zonal asymmetries in**  
37 **the circulation, land/ocean differences in surface fluxes, and the character of convective sys-**  
38 **tems, such a monsoon framework would integrate our understanding at all relevant scales:**  
39 **from the fine details of how moisture and energy are lifted in the updrafts of thunderclouds,**  
40 **up to the global circulations.**

41 Most tropical precipitation, whether steady rain or intense showers, falls from cloud clus-  
42 ters where individual, small-scale updrafts are organized over a few hundred kilometers in dis-  
43 crete weather systems. These, in turn, are orchestrated by planetary-scale circulation features: the  
44 monsoons and the Inter-Tropical Convergence Zone (ITCZ, Figure 1a and b). The clustering of  
45 individual clouds is the visible signature of an otherwise invisible global stirring. The notion that  
46 a link exists between smaller and larger scales underpins our understanding of tropical rain belt  
47 dynamics and is the basis of their representation in global climate models. Our understanding,  
48 however, remains incomplete, as evidenced by our inability to achieve reliable predictions of how  
49 the ITCZ and monsoons respond to external forcings.

50 Without the benefit of outdoor controlled experiments, or the possibility of validating pre-

51 ditions of long-term future changes, the reliability of our prediction tools must be tested against  
52 past records. Confidence in projected climate responses increases when dynamic theories built on  
53 contemporary observations also explain past conditions and when simulations skillfully reproduce  
54 the paleo record.

55 Good tests of our theories and models are the Holocene waxing and waning of monsoons  
56 in response to orbitally-driven changes in incoming solar radiation<sup>1,2</sup> (Figure 2). During the early  
57 to mid Holocene (11,000-5,000 years before present), the Northern Hemisphere received more  
58 insolation during its summer than today, while the Southern Hemisphere received less during its  
59 summer. In the modern climate, monsoons export energy from regions where the sun delivers  
60 most—so we would expect mid-Holocene monsoons to be stronger in the Northern Hemisphere  
61 and weaker in the Southern Hemisphere, compared to today. Paleoenvironmental data from that  
62 period indeed indicates increased rainfall in North and Central America<sup>3,4</sup>, a stronger Indian mon-  
63 soon and increased inland penetration of the monsoon into China<sup>5,6</sup>, and a spectacular rainfall  
64 expansion in northern Africa<sup>7,8</sup>: abundant lake, pollen and archaeological evidence documents  
65 wetting and vegetation increase over much of the Sahara.

66 While qualitatively consistent with our expectations, the major expansion of the Northern  
67 Hemisphere monsoons is widely underestimated in climate model simulations<sup>1,2</sup>, in particular  
68 over the Sahara<sup>9</sup> (Figure 2b). This bias can be reduced by modeling earth-system feedbacks<sup>10</sup>  
69 or by imposing the observed changes in vegetation, surface water storage, wetlands, soils, and  
70 mineral dust as boundary conditions<sup>11-13</sup>. Yet, simulations that organically produce both the rain-  
71 fall distribution and the vegetation types that are consistent with the records still elude us<sup>2</sup>. Moist  
72 atmospheric dynamics and its coupling with other aspects of the Earth system, including vegetation  
73 cover and soil properties, remain prime suspects for the failures. Moreover, the sign of observed  
74 Southern-Hemisphere changes is not fully consistent with the predictions of reduced rainfall from

75 both arguments based on the hemispheric summertime insolation forcing and complex model sim-  
76 ulations. Although many more quantitative estimates of rainfall anomalies from the Southern  
77 Hemisphere are needed to paint a full picture, many palaeoenvironmental records from South-  
78 ern Africa imply *increased* precipitation<sup>14</sup> and palaeoenvironmental records from South America  
79 and Australia seem to show (their interpretation being somewhat controversial) mixed wetting and  
80 drying signals<sup>15,16</sup>. Thus, neither complex models nor theoretical intuition are sufficient to explain  
81 past monsoon records.

82 In a traditional dry paradigm, monsoon circulations are akin to continental-scale land-sea  
83 breezes driven by surface *temperature* contrast, their strength increasing with the contrast. But this  
84 is too simplistic. Monsoon lands are hottest before the start of the monsoon, but the circulation is  
85 strongest in late summer (when increased rain and cloudiness have cooled the land and reduced  
86 the contrast). In future climate projections, land-sea temperature contrasts universally strengthen,  
87 but monsoon circulations generally weaken<sup>17</sup> as does, in some instances, regional early-season  
88 rainfall<sup>18,19</sup>.

89 A better paradigm views monsoons not as giant heat lows for which rainfall is a side effect,  
90 or as circulations “driven by” the latent heating of rainfall, but as moist energetically-direct circu-  
91 lations tightly coupled to precipitating convection; a facet of the general overturning of the tropical  
92 atmosphere inextricably linked to the Hadley circulation and the zonal mean ITCZ<sup>20</sup>.

93 At seasonal time scales, convection acts to release any column instability and to bring the  
94 free-tropospheric temperature in line with the moist static energy (MSE) of the boundary layer be-  
95 low it<sup>21</sup>. Horizontal atmospheric motions homogenize the free-tropospheric tropical temperatures  
96 to one vertical profile which reflects the conditions of the major convective centers. In this view,  
97 peak rainfall should coincide with peak low-level MSE; cooler and dryer surfaces can support  
98 only shallower or suppressed convection. Land and ocean observations confirm this theoretical

99 prediction<sup>22</sup>. Meanwhile, the upper-level divergence over these convective regions (maxima of  
100 low-level MSE) implies that the circulations associated with peak rainfall export total energy away  
101 from the centers of deep convection. In other words, monsoons and the ITCZ are components of  
102 a planetary energetically-direct circulation that links land and ocean rainfall (similar to a “global  
103 monsoon”<sup>23</sup>) and shows coherent variability from seasonal<sup>24,25</sup> to geological time scales<sup>20,26,27</sup>.  
104 Such an encompassing view is an important theoretical advance. Its formulation for the zonal  
105 mean tropical circulation—known in the literature as the energy-budget framework—underpins our  
106 understanding of why the ITCZ moves meridionally in response to forcings that originate well  
107 outside the tropics<sup>20,28,29</sup> (Section 1).

108 Further progress in understanding and simulating regional rainfall anomalies requires us to  
109 extend the energy-budget framework developed for the Hadley circulation to describe how the full  
110 tropical circulation redistributes energy vertically and horizontally, from its poleward boundaries  
111 and across the tropics. The extension of the energy-budget framework from a purely zonal for-  
112 mulation to one that includes zonal asymmetries would parallel the progress that has been made  
113 in understanding how the momentum budget constrains tropical circulations. The momentum-  
114 budget framework has led us from a starting view of the Hadley circulation<sup>30</sup> and—to a degree—  
115 the monsoons<sup>31</sup> as axisymmetric angular-momentum-conserving circulations, to the recognition  
116 that momentum transport by eddies is crucial to maintaining the zonal mean circulation in most  
117 cases<sup>32,33</sup>, to a deepening understanding of the role of the vorticity transport by stationary ed-  
118 dies in both localizing the monsoon circulation and regulating its intensity<sup>34,35</sup>. An analogous  
119 energy-budget framework would invoke zonal asymmetries in surface properties, which introduce  
120 horizontal gradients in the distribution of energy, amplify the importance of both stationary and  
121 transient eddy transports, and can generate shallow circulations which can be energetically in-  
122 direct (Section 2). These asymmetries are expressed in the observable differences in convective  
123 weather characteristics between land and ocean (Figure 1c and d), which aggregate into different

124 ascent profiles and vertical energy transport (Section 3).

125       The eventual goal is to combine the momentum-based and energy-based theories of tropical  
126 circulations and rainfall into a self-contained model for the tropical climate, one that describes how  
127 the interplay between energy and momentum fluxes in a moist atmosphere—where a profusion of  
128 weather phenomena, cloud types, and scales of motion is organized in mean and eddy effects—  
129 drives the seasonal evolution of oceanic and continental rainfall and controls monsoon diversity,  
130 variability, and response to external forcing. Understanding how the clouds within such circula-  
131 tions modify the energy input into the atmosphere through changes in both radiative fluxes and  
132 turbulent surface fluxes would be the next challenge<sup>36–38</sup>. In the rest of this paper, we propose a  
133 first step towards our goal: an energy-budget framework suitable for monsoons.

## 134 **1 The explanatory power of energetic constraints.**

135 Energetic constraints provide a parsimonious explanation of how the zonal-mean tropical rainfall  
136 shifts its position in response to internal variability and external forcings<sup>20,28,29</sup>. The ascending  
137 branches of the Hadley cells are in the deep tropics, where the solar radiation absorbed by the  
138 earth most greatly exceeds the terrestrial radiation emitted to space. Energy is transferred from  
139 the surface to the atmosphere by fluxes of radiation, sensible heat and moisture (latent heat) and a  
140 planetary circulation moves this excess energy towards high latitudes (Fig. 3a). Moist surface air  
141 converges and rises in the ITCZ, thus cooling adiabatically, condensing moisture, and forming pre-  
142 cipitating clouds. As a result, maximum rainfall is broadly co-located with the boundary between  
143 the Hadley cells (Fig. 3c).

144       These ideas can be formalized by focusing on the atmospheric energy budget in the annual  
145 and zonal means<sup>20,28</sup>. Averaging in time and longitude disposes of tendency terms and zonal fluxes,  
146 so that, under the assumption that eddy terms are unimportant, net energy input into the atmosphere



147 is balanced by the divergence of the vertically integrated energy flux by the Hadley circulation  $\mathcal{F}$   
148 and the ITCZ coincides with the “energy-flux equator,” i.e., the latitude at which the energy flux  
149  $\mathcal{F} = 0$  and changes direction.

150 The Coriolis force is weak in the tropics and atmospheric waves are effective in smooth-  
151 ing out pressure gradients. Because the tropical troposphere cannot maintain strong temperature  
152 gradients, an extra-tropical eddy heat flux that reaches the tropics will be carried into the tropics  
153 and through—to the mid latitudes of the other hemisphere. This argument appears to apply more  
154 broadly: any asymmetry in the energy flux across the northern and southern edges of the trop-  
155 ics is felt throughout the tropical band<sup>39</sup>, making the position of the energy-flux equator sensitive  
156 to extra-tropical forcings<sup>20</sup>. The same argument explains how the inter-hemispheric asymmetry  
157 in seasonal insolation drives the north-south annual migration of the ITCZ<sup>24</sup>. However, because  
158 the cross-equatorial Hadley cell is always the strongest throughout the annual cycle, and because  
159 maximum ascent and rainfall are concentrated within the cross-equatorial cell and equatorward of  
160 the energy-flux equator, seasonal rainfall shifts are less pronounced than those of the energy-flux  
161 equator<sup>24</sup> (Fig. 3a). In comprehensive models, changes in the annual-mean position of the ITCZ,  
162 their inter-model spread, and the seasonal migration of the rain belt have been shown to follow  
163 the same quantitative relationship with energy transports, albeit with some scatter and uncertainty  
164 (a 1PW change in the cross-equatorial energy flux leads to about a 3° shift in the position of the  
165 ITCZ, Fig 3b). This correspondence has been used to suggest that the same dynamics control both  
166 the meridional shifts characteristic of the annual cycle and the variability of mean rainfall at paleo  
167 time scales<sup>20,24,40</sup>.

168 This theory for the ITCZ position sees the tropical rain belt as the expression of the conser-  
169 vation law that governs energy in the climate system; it links the occurrence of convection not just  
170 to the local environment, but to the planetary adjustments that bring the global atmosphere towards

171 equilibrium; and it subsumes arguments that link the ITCZ position to gradients in tropical SST. It  
172 has been used to explain the location of the modern ITCZ north of the equator as a consequence  
173 of energy transport by the thermohaline ocean circulation<sup>41</sup>, the southward shift of tropical rainfall  
174 during the last glacial maximum and Dansgaard-Oeschger events as a consequence of northern  
175 high-latitude cooling<sup>29,40</sup>, the effect of Eurasian afforestation on monsoons in the mid Holocene<sup>42</sup>,  
176 the Sahel drought in the 1970's and 1980's as part of a global-scale southward ITCZ shift due  
177 to sulfate aerosols<sup>43</sup>, and the role of Southern Ocean heat uptake in setting up the hemispheric  
178 asymmetry in future tropical rainfall changes<sup>44</sup>.

179         However, the existing framework is limited in important ways. First, it is incomplete as  
180 a predictive theory because internal radiative feedbacks from clouds and water vapour can over-  
181 whelm the external forcings in setting energy gradients, are not easily predicted from gross en-  
182 ergetic constraints, and vary substantially between models<sup>36,37,45</sup>. Even when the radiative effect  
183 of clouds is carefully controlled in models<sup>46,47</sup>, changes in oceanic heat transport can oppose the  
184 inter-hemispheric difference in energy input in the atmosphere and complicate the response of  
185 the ITCZ<sup>48</sup> to external forcings. Indeed, the tight coupling between atmospheric and oceanic heat  
186 transport<sup>20,49-51</sup> suggests an expansion of atmospheric-only arguments. Second, changes in tropical  
187 rainfall are often better described as intensifications<sup>52</sup> or contractions<sup>19,53-56</sup> of the climatological  
188 net rainfall pattern, rather than shifts, and some regional anomalies are strongly driven by localized  
189 gradients in surface conditions<sup>57-59</sup>, including between ocean and land<sup>60,61</sup>. Third, the assumption  
190 that rainfall is the product of deep convection in which ascent extends throughout the troposphere  
191 and maximizes at mid levels is crucial to the portrayal of the Hadley cell as energetically direct.  
192 Quantitatively, this requires that the same large-scale circulation that converges moisture into the  
193 ITCZ also diverges enough static energy from the ITCZ at upper levels that the vertically-integrated  
194 result is a net export of MSE—a result that depends heavily on where exactly in the column air is  
195 converging and diverging. Yet, rainfall is also produced in circulations with shallow components

196 and bottom-heavy ascent profiles (see also Fig 4), and direct calculation of the energy flux out of  
197 such regions indicate a net import of energy<sup>62,63</sup>. Fourth, the assumption in the ITCZ energy-budget  
198 framework that eddies are not important is debatable. Even well within tropical latitudes, transient  
199 eddies transport latent heat poleward<sup>64</sup> and thus might be as important as the zonal circulation in  
200 transporting energy, especially where the latter is weak. While the bulk effect of transient eddies  
201 could be subsumed into the framework developed for the ITCZ in an aquaplanet<sup>28,65</sup>, the presence  
202 of land introduces stationary eddies<sup>34,66</sup> and localizes transient eddies<sup>31</sup> in ways that preclude the  
203 straightforward application of the zonal mean framework to limited longitude bands (Fig 2c). This  
204 is particularly true in a moist atmosphere, where feedbacks between surface fluxes, clouds, and  
205 latent and radiative heating can amplify and extend the asymmetries in the forcing<sup>45</sup>. These limi-  
206 tations hamper our ability to use the existing energy-budget framework for the ITCZ to accurately  
207 predict the response of regional rainfall to past or future forcings. Yet the fundamental insight that  
208 regional changes are expressions of global conservation laws should not be abandoned.

## 209 **2 The need for an energy-budget framework for monsoon systems.**

210 Zonal asymmetries and contrasts between land and ocean are fundamental to the energy and the  
211 momentum budget of monsoon circulations<sup>13,34,67,68</sup>. To close, the energy budget must include  
212 the effect of zonal transports and of the complex vertical structure of the meridional circulation,  
213 both of which are the result of inhomogeneities in surface properties. Inhomogeneities do oc-  
214 cur over oceans (such as between warm pools and the equatorial cold tongues), but gradients in  
215 surface properties are especially strong at coastlines and over continents due to orography, vari-  
216 ability in the characteristics of soils and vegetation, and—in a positive feedback—to the response  
217 of the land system to differences in precipitation. The conceptual model of the Hadley cell as a  
218 simple deep overturning meridional circulation (Fig. 3c) can broadly capture the main features of  
219 the ITCZ, but is insufficient to describe the more complex monsoon circulations. In many regional

220 monsoons<sup>22,69</sup>, gradients in surface temperature and sensible heat fluxes (e.g., between the hot sub-  
221 tropical desert and the cooler oceans and equatorial rainforests) force a shallow circulation with  
222 dry ascent poleward of the monsoon rainfall. This circulation includes the low-level monsoon flow  
223 that fuels moisture to the rainfall band and the dry return flow above that can cap deep convection,  
224 affecting the frequency of rainfall and the occurrence of severe convection (Figure 1c,d and 3c).  
225 Re-evaporation of rainfall in a dryer lower troposphere (Fig. 3c, cloud types) can also affect the  
226 vertical distribution of latent heating and, concurrently, ascent (see also the next section). The  
227 presence of land, thus, changes the vertical transports in the meridional divergent circulation. The  
228 strong horizontal gradients in temperature and moisture associated with surface type also induce  
229 strong non-divergent, rotational flows that transport energy and moisture horizontally (Fig. 3c,  
230 broad arrows). Past seminal work<sup>67</sup> has shown that ventilation, the transport of low-MSE oceanic  
231 air by the rotational flow, is key to setting the poleward extent of the monsoon rainfall. Recent  
232 studies with more comprehensive models have confirmed the important role of the rotational flow  
233 in balancing the energy budget of monsoon regions<sup>70</sup> and in driving the onset of off-equatorial  
234 rainfall<sup>34</sup>. The annual changes in surface properties that take place during the progression of the  
235 rainy season are also reflected in changing flows of energy in the atmosphere. The most obvious  
236 change is soil water content, which affects evaporative fluxes and albedo and, because of its high  
237 spatial variability, can introduce sharp gradients in surface properties at small spatial scales<sup>71</sup>.

238         There is a vast literature that focuses on the individual regional monsoons and emphasizes the  
239 zonally asymmetric regional flow and its interaction with mountains and regional oceans. There is  
240 value in formulating such regional analysis within a MSE framework. For example, the extent and  
241 intensity of the South Asian monsoon depends on the presence of the Himalayan mountains be-  
242 cause the latter shield the MSE maximum over India from low-MSE extratropical air<sup>72</sup>. In another  
243 example, the African monsoon appears to be particularly susceptible to moisture anomalies enter-  
244 ing from the North Atlantic<sup>73-76</sup> and the Mediterranean<sup>77</sup> (even though the main moisture source

245 for the monsoonal rains is in the tropical ocean) because they regulate the depth of convection,<sup>78</sup>  
246 which is in turn linked to energetic requirements of the large-scale flow<sup>70</sup>. How the profile of con-  
247 vective ascent evolves across the monsoon season modulates such sensitivity, by making changes  
248 in either the boundary layer MSE (such as the horizontal advection or recycling of moisture) or the  
249 free-tropospheric dynamics (such as changes in upper-level vertical stability) more or less relevant  
250 (a similar dependence of rainfall sensitivity on low-level and upper-level processes is also seen  
251 across models<sup>70,73</sup>).

252         The broad-brush framework built around the MSE budget has been successful in explain-  
253 ing why and how processes as disparate as aerosol microphysics and the oceanic thermohaline  
254 circulation can affect the zonal mean tropical rainfall. Here, we advocate extending the energy-  
255 budget framework from its current form, appropriate to the ITCZ/Hadley cell, to one appropriate  
256 for monsoons: one that quantifies how the presence of zonally-confined land masses modifies  
257 both the input and the flow of energy in the atmosphere and thus changes the leading terms of  
258 the vertically-integrated atmospheric energy budget. In this section we have emphasized (i) how  
259 land/ocean differences in surface properties change the input of moist and dry energy to the atmo-  
260 sphere, (ii) how the continental-scale stationary circulations that develop in response advect MSE  
261 gradients horizontally through the geostrophic flow, and vertically through shallow circulations,  
262 and (iii) how changes in the dominant balance shape both the diversity of regional monsoons and  
263 monsoon evolution through the development and decay of the rainy season. In the next section we  
264 focus on the role of vertical motion at the scale of clouds, and how it shapes the vertical fluxes of  
265 moist energy and its input in the atmosphere.

266 **3 The diversity of convection in tropical precipitation systems: interaction with the large-**  
267 **scale climate.**

268 Recent observations have highlighted the rich diversity of convective systems (Figs. 1 and 4), and  
269 have allowed fundamental insights into the processes that govern them. Convective precipitation  
270 appears controlled by both low-level and deep-column moisture, which set the buoyancy of en-  
271 training ascending parcels. Thus, while early theories predict the position of the rainfall maximum  
272 from just the boundary-layer MSE (Sec. 1), full tropospheric water vapour is key to fully account  
273 for the spatial and temporal variability of rainfall intensity. In turn, processes such as detrainment  
274 from precipitating clouds and re-evaporation of rain make the tropospheric water vapour depend  
275 on the occurrence of rainfall. This two-way coupling underpins the correspondence between rain-  
276 fall amounts and total humidity found over land and ocean (see also Fig 1a) and is encapsulated  
277 in an exponential relationship of daily precipitation intensity on the integrated humidity<sup>79</sup> (with  
278 some variations across convective system<sup>80</sup> and their drivers<sup>81</sup>). However, rainfall characteristics,  
279 such as intensity, organization, and duration, and the vertical and temporal distribution of clouds  
280 depend on factors other than column humidity, including wind shear, the larger-scale flow, and the  
281 properties of the surface boundary<sup>71,82,83</sup>. Continental updrafts are often deeper and more intense  
282 than oceanic updrafts (Fig. 4a,b), as evidenced by the preferential occurrence of lightning over  
283 land (Fig. 1c), but land convection varies greatly through the day (cf. Figs. 4 c and d) and the  
284 season in depth, organization, and lifetime—affected by surface inhomogeneity and by stronger  
285 triggering and inhibition processes.

286 The growing appreciation of the diversity of convective cloud systems has yet to mature into  
287 enough understanding of the interplay between clouds and large-scale dynamics to create much-  
288 needed convective parameterizations able to describe such diversity. Parameterizations are still  
289 overly reliant on so-called quasi-equilibrium formulations: the gross effects of convection are taken

290 to be nearly instantaneous and deterministic functions of large-scale forcings, thermodynamics  
291 drivers are overemphasized over dynamics, and convective organization is typically ignored<sup>84,85</sup>.  
292 The approximations lead to a preference for tall, disorganized convection and the lack of both  
293 shallower, developing convection<sup>86</sup> and more persistent cloud systems and is likely the source of  
294 climate models' systematic errors in rainfall timing (for example, the too-early peak in the diurnal  
295 cycle of land rainfall). It affects the climate at longer time scale as well: model biases in seasonal  
296 rainfall are typically established within few days after initialization from observations, and only  
297 later amplified in the coupled system, pointing to the dominance of fast atmospheric processes<sup>87</sup>  
298 in setting the bias. Such long-ranging effects are not surprising: the aggregate effect of convection  
299 is reflected in the profiles of horizontal convergence and divergence and in cloud and moisture  
300 radiative effects and will thus affect both the net vertically-integrated energy flux and the total  
301 energy input into the atmosphere. The extent and vertical distribution of clouds modify the net  
302 energy in the atmosphere via changes in radiative fluxes and, by modifying the structure of the  
303 boundary layer, turbulent fluxes. The vertical profile of ascent modulates the energy transport by  
304 the circulation, to the point that a predominance of bottom-heavy or top-heavy convection can  
305 determine whether the circulation imports or exports MSE to or from a convective region<sup>62,63</sup>.  
306 Thus, shortcomings in the mix of deep convective and stratiform rain production, of warm and  
307 cold cloud microphysics, and of mixing and re-evaporation, can translate into global scale biases<sup>88</sup>  
308 and introduce another source of uncertainty in future projections<sup>70</sup>.

309 High-resolution cloud-resolving dynamical models can portray the full evolution and organi-  
310 zation of cloud systems<sup>88,89</sup> and—when coupled to parameterized large-scale fields<sup>90</sup> that describe  
311 a broad suite of boundary conditions and environments— are being used to investigate how the  
312 environment shapes the rich diversity of cloud characteristics and how cloud processes feed back  
313 on the environment. These experiments have shown that the specific way in which land convection  
314 evolves during the day (such as through morning fog or land-sea breezes) is a key determinant of

315 the large-scale mean and the seasonal evolution of rainfall and other environmental variables<sup>91,92</sup>.  
316 More research on tropical land convection is needed to elucidate whether this cross-scale link is  
317 achieved because land processes and clouds modify the energy input into the atmosphere, or be-  
318 cause the daily evolution of convection changes the profile of ascent and this effect is rectified into  
319 changes in the vertical energy flux.

## 320 **4 Synthesis**

321 Common biases in the simulation of the diurnal and seasonal cycle of rainfall highlight structural  
322 deficiencies across global climate models. Moreover, common biases in the simulation of rainfall  
323 in past climate states indicate that common structural deficiencies also affect the modeled responses  
324 to changes in forcing. This diminishes the value of consensus in projections of climate response to  
325 anthropogenic forcings as an indication of reality. Convective parameterizations, which produce  
326 most tropical rainfall within current climate models, have inherent limitations that are a likely cause  
327 of bias, but whether cloud-resolving global models will be able to adequately capture observed  
328 rainfall variations in modern or past climate states remains to be seen. Oceanic processes that  
329 amplify atmospheric biases in the seasonal cycle<sup>93</sup> are likely to play a role in setting long-term  
330 trends<sup>94</sup> and paleo simulations highlight the additional importance of earth-system feedbacks, such  
331 as between precipitating atmospheric dynamics, vegetation cover, and soil composition. Having a  
332 small set of metrics by which to characterize simulations of a climate phenomenon would make  
333 the task of model development less daunting; a modern energy-budget framework, one modeled on  
334 past efforts<sup>20,21,32,95,96</sup> but aware of the advances and challenges outlined in this review, can supply  
335 such metrics for tropical rainfall.

336         Developing the framework we are advocating will require the combination of observational,  
337 modeling and theoretical approaches. Both modern and paleo observations must be the ultimate



338 tests of our theories. Fine-resolution, high-frequency observations of rainfall and of the vertical  
339 atmospheric structure, together with cloud-resolving simulations, have the potential to shed much-  
340 needed light on how convection becomes organized and how monsoon rainfall interacts with its  
341 environment from sub-daily to monthly timescales<sup>84</sup>. To be truly useful for benchmarking met-  
342 rics and developing parameterizations, current efforts must be expanded to encompass a broader  
343 cross section of tropical environments and climate regimes—including a variety of chronically  
344 undersampled land regions. Collection of new paleo evidence and quantitative reconstructions of  
345 precipitation from the Southern Hemisphere will refine our picture of paleo monsoons in the mid  
346 Holocene and provide a well defined target for testing model simulations and their interpretation  
347 within this common framework. The opposite is also true: forward proxy models and climate  
348 models are needed to guide the interpretation of data, morph sparse environmental observations  
349 into a coherent portrait of past climates, and help prioritize new data acquisition. Multi-model  
350 simulations<sup>97</sup>, simulations with a broad set of forcings<sup>48</sup>, and simulations with models of different  
351 vintage are needed together to provide a robust test for a new monsoon framework.

352 Drawing our lessons from the development of budget-based theories of the ITCZ, we pro-  
353 pose that a systematic understanding of the monsoons will require a comprehensive hierarchy of  
354 model simulations and a common set of energy and momentum diagnostics to compare results  
355 across the model hierarchy and with observations. A key element of such a hierarchy are idealized  
356 model setups in which energy is conserved through fluxes at the atmospheric boundaries and that  
357 are designed to highlight the effect of different land characteristics—themselves isolated through  
358 ad-hoc idealizations—on the response of the tropical rain belt to a range of external forcings<sup>98–100</sup>.  
359 Accurate diagnostics of how energy flows into and within the tropical atmosphere, and via differ-  
360 ent elements of the atmospheric circulation (from the zonal mean, to stationary eddies, to synoptic  
361 eddies, to much faster variability at the scale of convection) are often difficult to calculate from  
362 available model output<sup>54</sup>. Yet they must underpin any metrics used to elucidate which small- and

363 large-scale processes are essential to the monsoon systems (and by which interactions); to test the  
364 coherence of our dynamical theories; and to benchmark model development. A theoretical model  
365 that can deduce the behavior of tropical rainfall from simple indicators based on conservation laws  
366 of the physical system with minimal empirical or ad-hoc assumptions would cap the hierarchy and  
367 provide the grounds for interpreting both idealized and realistic simulations and for understand-  
368 ing the observed behavior of the earth system. To derive a unified theory of the ITCZ and the  
369 regional monsoons from basic conservation laws is a formidable challenge, but we must meet it to  
370 confidently link past to present to future.

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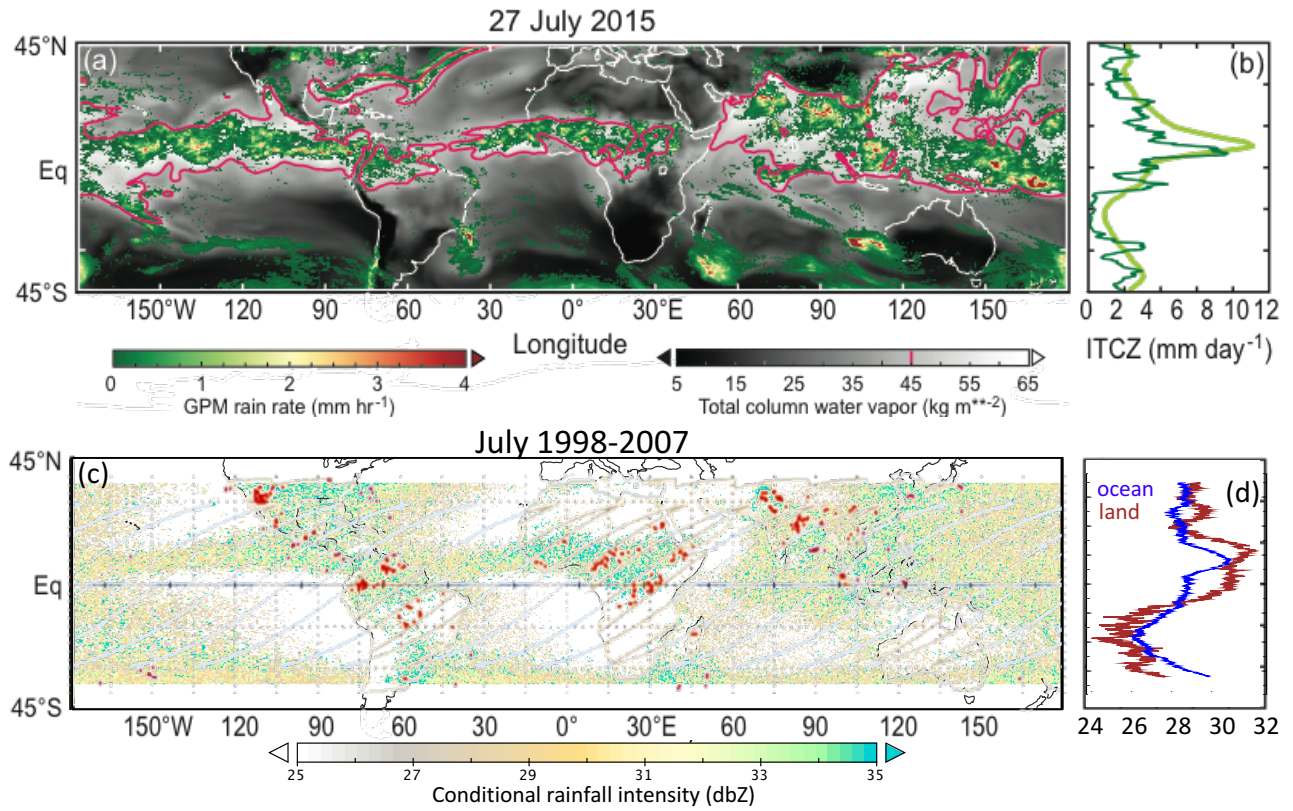


Figure 1: (a) Rainfall (color) on July 27th, 2015 and the high atmospheric moisture enveloping it (indicated by the 45mm contour of column-integrated water vapour, the full field is in gray); (b) zonal-mean rainfall for the same day (dark green) and climatological values for the same period (light green). (c) July climatological mean intensity of instantaneous near-surface rainfall (from TRMM precipitation radar, units of reflectivity) and occurrence of lightning (red dots) on July 27th, 2014 (ascending passes of the Lightning Imaging Sensor on TRMM). (d) zonal mean rainfall intensity for land and ocean regions. See on-line method section for further details.

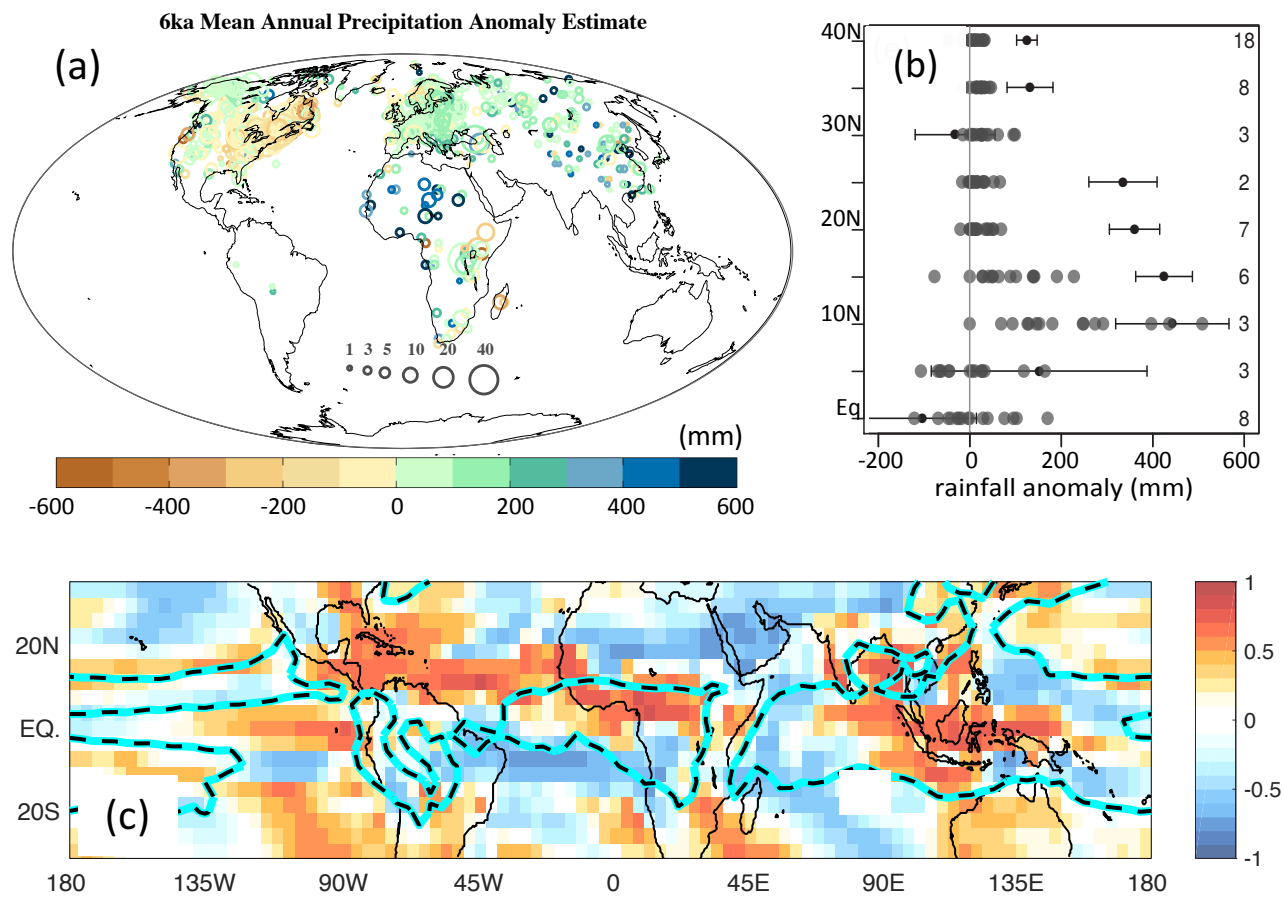
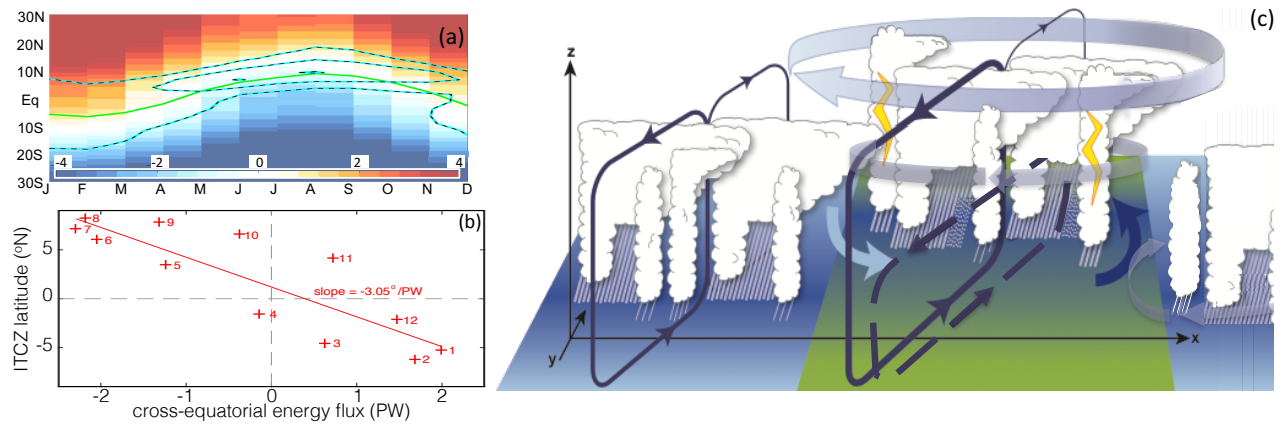
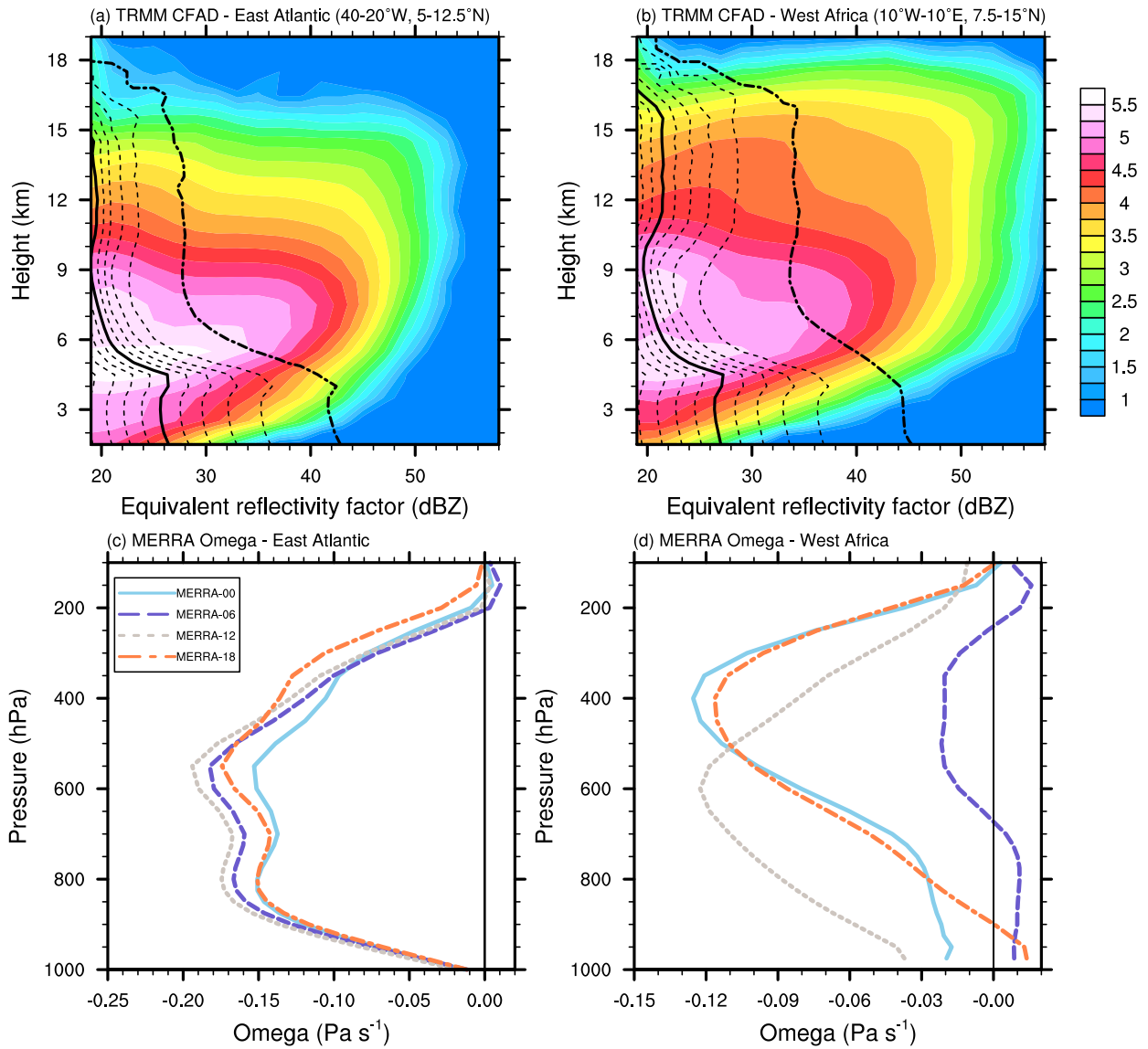


Figure 2: (a) Mid-Holocene mean annual precipitation (MAP) anomalies (color circles, sized by the number of used reconstructions). (b) CMIP5/PMIP3 simulated (gray circles) and reconstructed (black circles with error bar) mid-Holocene MAP changes and grid cells contributing to the reconstruction and model means (numbers at right) for northern Africa latitude bands. (c) Correlation of decadal rainfall variations with the first principal component of the zonal-mean precipitation in transient Holocene simulations. See methods for data references and details.



**Figure 3:** (a) Zonal-mean rainfall (dashed), ITCZ position (green), and zonally- and vertically-integrated atmospheric energy transport (in PW, shaded). (b) Climatological ITCZ latitude as a function of the vertically-integrated atmospheric energy transport at the equator. Numbers correspond to calendar month. The ellipse is sloped at  $3^\circ$  ITCZ shift per 1 PW energy flux. (c) ITCZ (blue background) and monsoon (green background) schematics. The Hadley cells (dark solid lines) meet in the northern tropics, maximum ascent and rainfall occur in the winter cell (equatorward of the cell boundary) close to maximum low-level MSE (darker surface shading). Monsoons are distinguished by the additional shallow meridional circulation (dashed) and ventilation by the rotational horizontal flows (ribbon arrows) and by a different distribution of cloud types. See on-line methods.



**Figure 4:** Cloud characteristics (top) and ascent profiles (bottom) over the eastern Atlantic ITCZ (left) and western Africa (right) as function of height. (a-b) The Contoured Frequency with Altitude Diagrams (CFAD, filled contour) shows the frequency (logarithmic scale) of storm-top height as a function of rainfall intensity (units of reflectivity, dBZ). Contours: percentile lines (solid=median, dashed-dotted = 99th percentile) of reflectivity frequency. Frequency is normalized at each level and all rain types are included. (c-d) Daily evolution of the vertical velocity profiles (pressure coordinates; negative values indicate upward motion) from 6-hourly MERRA reanalysis. See on-line method section for details.

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642 **5 Methods**

643 Figure 1: Characteristics of observed tropical rainfall. On any given day, rain is produced by cloud  
644 systems that range in size between individual convective towers (1-10km) to mesoscale convective  
645 systems and tropical cyclones (> 100 kilometers). Despite substantial variability on short time  
646 scale and small spatial scales, most disturbances are organized within the large-scale rain belts  
647 formed by the monsoons and the inter-tropical convergence zone (> 1000 km). The map (a, color)  
648 and zonal mean (b, dark green) of daily rainfall are obtained from merged GPM satellite mea-  
649 surements calibrated against rain gauges (Huffman et al., 2015). Column integrated water vapour  
650 (a, grey shading and magenta contour) was obtained from the ERA Interim reanalysis product  
651 (Dee et al., 2011). The summer climatological zonal mean rainfall (b, light green) is calculated as  
652 the 1998-2014 average of the 11-day period centered around July 27th and is obtained from the  
653 TRMM-3B42 rainfall estimate (Huffman et al., 2007). Climatological conditional intensity of rain-  
654 fall (c, color shading, and d) is obtained from the precipitation-radar data on the TRMM satellite  
655 (Biasutti et al, 2011) and is expressed as a reflectivity in units of dbZ (decibels of Z). Reflectivity is  
656 the amount of transmitted power returned to the radar receiver; light rain is detected by the TRMM  
657 PR when the dBZ value reaches 18 (corresponding to about 0.4mm/hr). The higher the dBZ, the  
658 stronger the rainrate. Uncertainty in the methods for translating reflectivity into a quantitative  
659 precipitation estimates are detailed in Villarini and Krajewski (2009). Lightning flashes for July  
660 27th, 2014 are obtained from Lightning Imaging Sensor during ascending passes of the TRMM  
661 satellite (<https://lightning.nsstc.nasa.gov/lisib/lisbrowse.exe?which=qcyear=2014day=208>).

662 Figure 2: Changes in rainfall in the Holocene indicate a complex behavior across the mon-  
663 soon systems, not fully captured by climate models, and do not suggest that meridional displace-  
664 ments of the zonal mean ITCZ explain a large fraction of the variance of tropical continental  
665 rainfall. (a) The expectation that Southern Hemisphere monsoons would be weaker in periods of

666 weaker SH summer insolation is not fully supported by current observations. Quantitative recon-  
667 structions of changes in mean annual precipitation (MAP) between the mid Holocene (11,000-  
668 5,000 years before present) and present day in colored circles, the color indicates the size of the  
669 anomaly (mm, colorbar) while the size indicates the number of reconstructions used for the es-  
670 timate, as indicated by the legend inside the map. The data was first published by Bartlein et al  
671 (2011). (b) CMIP5/PMIP3 simulated and reconstructed changes in mean annual precipitation in  
672 the mid Holocene for 5° latitude bands across northern Africa (longitude 20°W to 40°E between  
673 0 and 45°N), where the model results are averages for the grid cells with observations and each  
674 model is represented by a different gray circle. The mean and standard error of the reconstruc-  
675 tions is shown in black and the number of grid cells contributing to the reconstruction is shown  
676 for each latitude band. CMIP5/PMIP3 data were available thanks to the World Climate Research  
677 Programme's Working Group on Coupled Modelling, which is responsible for CMIP. For CMIP  
678 the U.S. Department of Energy's Program for Climate Model Diagnosis and Intercomparison pro-  
679 vides coordinating support and led development of software infrastructure in partnership with the  
680 Global Organization for Earth System Science Portals. The figure is modified from Perez-Sanz et  
681 al., 2014; (c) Correlation maps (shading) of decadal rainfall variations at each gridpoint with the  
682 first principal component of the zonal mean precipitation in the fully-forced TrACE-21000 (Otto-  
683 Bliesner et al, 2014) simulation for the period 9.5ka to 0.5ka before present. The dashed line is a  
684 representative contour for the mean precipitation and indicates the climatological position of the  
685 rain belt.

686 Figure 3: The energy-budget framework for the tropical rain belt. (a) The seasonal evo-  
687 lution of the zonal mean observed climatological rainfall (dashed contours, only the 4, 6, and  
688 8mm/day isolines are shown) and ITCZ position (defined as the centroid of zonal-mean rainfall  
689 within 20° N/S; green line) are superimposed on a reanalysis-based estimate of the zonally and  
690 vertically integrated atmospheric energy transport (shaded, warm and cool colors indicate north-

691 ward and southward transport, respectively, and the white area indicates the energy flux equator,  
692 units of PW). (b) The seasonal relationship between the vertically integrated atmospheric energy  
693 flux at the equator ( $\mathcal{F}_o$ ) and the ITCZ position. The energy flux associated with the net mass move-  
694 ment between the hemisphere is retained here, leading to a phase lag between the two fields. The  
695 slope of the relationship, given by the direction of the major axis of the ellipse, indicates a  $3^\circ$  shift  
696 for a 1 PW energy flux, consistent with calculations that omit the barotropic circulation.

697 The climatological rainfall in Figure 3a,b is calculated as the 1979-2013 average of the Cli-  
698 mate Prediction Center (CPC) Merged Analysis of Precipitation (CMAP) dataset (Adler et al.,  
699 2003). The energy fields are from ERA Interim reanalysis (Dee et al., 2011) over the same  
700 period. The cross-equatorial energy transport is calculated eliminating the mass budget resid-  
701 ual before vertically integrating the fluxes, the data was provided by the National Center for  
702 Atmospheric Research (retrieved from [https://climatedataguide.ucar.edu/climate-data/era-interim-](https://climatedataguide.ucar.edu/climate-data/era-interim-derived-components)  
703 [derived-components](https://climatedataguide.ucar.edu/climate-data/era-interim-derived-components)).

704 (c) Schematic of the ITCZ (depicted over an oceanic surface, blue half on the left) and  
705 monsoon (depicted over a continental surface, green half on the right) circulations for northern-  
706 hemisphere summer. The summer and winter Hadley cells (dark solid lines) meet in the northern  
707 tropics, close to where the low-level moist static energy (MSE) is maximum (darker shading at  
708 the surface), consistent with convective quasi-equilibrium theory. Most upward motion and thus  
709 most rainfall occurs in the ascending branch of the stronger, winter cell, so maximum rainfall is  
710 slightly equatorward of the Hadley cell boundary and the energy-flux equator. As is the case for  
711 the ITCZ, rainfall associated with the monsoonal circulation is positioned slightly equatorward of  
712 the maximum surface moist energy (dark blue shading) and is associated with large-scale ascent in  
713 local meridional overturning cells whose strength is greater when the ascent is further away from  
714 the equator. Key distinctions for the monsoons are in the complexity of the circulation and the

715 distribution of cloud types. Notice in the land portion of the diagram the presence of a shallow  
716 meridional circulation (dashed lines) with ascent poleward of the rain band and a dry return flow,  
717 the rotational circulations associated with the low-level cyclone (light and dark blue ribbon-width  
718 arrows indicate negative and positive transport of MSE), mid- and upper-level land anticyclones,  
719 and the oceanic anticyclone (anticyclonic circulations are depicted with ribbon-width grey arrows).  
720 Notice also the deeper and more intense convection over land (indicated by a distribution of clouds  
721 that include more overshoots and fewer clouds lacking an anvil), more lightning, less rain from  
722 warm cloud (fewer clouds without anvil), and more re-evaporation of rain (dotted rain from the  
723 anvil cloud).

724 This schematic highlights those aspects of an hypothetical “essential” monsoons that are  
725 addressed in this paper. It is not meant to represent any particular monsoon system, as each is  
726 highly affected by the geometry of the continent, the location and orientation of orography, the  
727 geographical distribution of surface types, including deserts, and oceanic processes unique to each  
728 ocean basin.

729 Figure 4: Cloud characteristics and ascent profiles over ocean and land during the peak of  
730 the rainy season. Left: the eastern Atlantic ITCZ (40-20°W 5-12.5°N). Right, western Africa  
731 (10°W-10°E 7.5-15°N). The Contoured Frequency with Altitude Diagrams (CFAD) in the top pan-  
732 els show that western Africa has deeper, more intense convective cells while the eastern Atlantic  
733 has more mid-level rainy cloud. The filled contour shows the log of the frequency of storm top  
734 height as a function of radar reflectivity (a measure of rainfall intensity). Both regions show that  
735 storms reaching 5km in height and measuring reflectivities of less than 30dBZ are the most fre-  
736 quent, but this peak is more pronounced over ocean than land. Conversely, the land region shows  
737 more frequent instances of storms that have intensity above 50dBZ (colors extending to the right  
738 of the diagram) and reach 18km in height (colors extending to the top of the diagram). Overlaid

739 on the color are percentile lines of reflectivity frequency at each level with the black line equal  
740 to the median at each level. The far right line is the 99th percentile line. Notice that the median  
741 surface reflectivity is slightly higher over land, while the 99th percentile reflectivity is much higher  
742 over land, indicating that land convection reaches more extreme values of intensity. The secondary  
743 reflectivity maximum in the median line is also more noticeable over land than ocean, suggest-  
744 ing again significant differences in the vertical profile of the cloud systems over land and ocean.  
745 Frequency is normalized at each level and all rain types (stratiform, convective, shallow isolated,  
746 shallow non-isolated) are included. The bottom panels show how ascent profiles in rain systems  
747 have much larger diurnal variations over western Africa and are much more top-heavy than over  
748 the adjacent ocean. Profiles of vertical velocity in pressure coordinates (negative values indicate  
749 upward motion) were obtained from the MERRA reanalysis.

750 The TRMM reflectivity data (Kummerow et al, 1998) was originally sourced from TRMM  
751 orbital files (2A23/2A25) for the 1998-2014 period (August values only). Vertical omega profiles  
752 are calculated from 6-hourly MERRA reanalysis (Rieckner et al, 2011) for August days in 1983-  
753 2007. Only samples that contributed the top 50% of rainfall by volume were included (using the  
754 MERRA surface precipitation flux data for all rain and rainfall thresholds unique to each domain  
755 and 6-hourly period). When all times are included, the eastern Atlantic ascent is even more bottom  
756 heavy than shown in Figure 4, but the qualitative comparison to land convection is unchanged.

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