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# applied spectroscopy

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## Effects of Metallic, Semiconducting, and Insulating Substrates on the Coupling Involving Radiative Polaritons in Thin Oxide Films

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Through simulations, this work explores the effects of conducting, semiconducting, and insulating substrates on the absorption of infrared radiation by radiative polaritons in oxide layers with thicknesses that range from 30 nm to 9  $\mu$ m. Using atomic layer deposition, oxide layers can be formed in the nanometer scale. Our results suggest that the chemistry and conductivity of the substrate determine the amount of absorption by radiative polaritons in oxide layers thinner than the skin depth. The effects of the chemistry and conductivity of the substrate are especially effective for oxide films thinner than about 250 nm, which we label as the *substrate sensitive thickness* of the oxide film.

Index Headings: Infrared spectroscopy; Radiative polaritons; Atomic layer deposition; Thin oxide films.

#### **INTRODUCTION**

Electromagnetic radiation interacts with matter in different ways depending upon frequency. High-frequency  $\gamma$  and X-ray photons interact through the Compton effect, or Compton scattering, while ultraviolet (UV) and visible photons interact through the photoelectric effect. Visible light, in addition, participates in many chemical reactions of biological importance, such as photosynthesis. Infrared (IR) radiation, on the other hand, interacts with matter via coupling to dynamic dipoles. In particular, in inorganic solid and ionic crystals IR photons can interact and couple with the collective lattice vibrations called phonons, thus exciting polaritons.<sup>1,2</sup> Understanding the interaction of radiation with matter is crucial from a fundamental and a practical point of view. Indeed, radiationmatter interactions are at the basis of energy conversion devices such as solar photovoltaics. Presently, photovoltaic cells are used to convert absorbed visible photons into usable energy. There is also emerging interest in capturing IR radiation for energy conversion,<sup>3</sup> thermal photovoltaics,<sup>3,4</sup> and enhanced thermal transport at interfaces.<sup>3,5–7</sup> Such applications have led to the study of polaritons,<sup>1,2</sup> which form when IR photons couple with phonons present in dielectric materials featuring ionic bonds. Multiple types of polaritons exist, including surface phonon polaritons (SPPs)<sup>3-5</sup> and radiative polaritons (RPs).<sup>1,2</sup>

This work focuses on RPs in thin oxide films. The effects of the chemistry and conductivity of the oxide film substrate on the interaction between IR radiation and planar thin oxide films have been discussed in the literature only to a limited extent. Authors have studied films on metal substrates,<sup>8-13</sup> semiconducting substrates,<sup>14–19</sup> and insulating substrates.<sup>20–22</sup> A comparison of the effects of the conductivity of different metal substrates can also be found.<sup>9,10</sup> In addition, there is experimental and simulated data available for oxide layers less than 1 µm thick, where the IR optical properties are sensitive to the chemistry and conductivity of the substrate.<sup>19,22</sup> However, there has been no comprehensive study that examines how variations in substrate chemistry and conductivity impact IR radiation absorption in thin oxide films. Oxide films less than 1  $\mu m$  thick are of technological relevance (e.g., in microelectronics)<sup>23</sup> and can be formed with high quality by utilizing, e.g., atomic layer deposition (ALD).<sup>24</sup> Therefore, knowing the effects of the chemistry and conductivity of the substrate on the oxide film's ability to absorb IR radiation might become of compelling importance in future applications. It is important to point out that oxide films deposited by ALD, especially Al<sub>2</sub>O<sub>3</sub>, are very uniform in thickness and have low roughness (less than 10% of the film thickness).<sup>24</sup>

This work investigates numerically the absorption of IR radiation in thin planar oxide films on metallic, semiconducting, and insulating substrates. The absorption of the IR radiation is related to the excitation of RPs, as in the theory by Kliewer et al.<sup>2,25</sup> This theory, which originally considers simple cubic oxides with one couple of longitudinal ( $\omega_{LO}$ ) and transverse ( $\omega_{TO}$ ) optical phonon frequencies, will be extended here to oxides characterized by more than one couple of  $\omega_{LO}$  and  $\omega_{TO}$  frequencies. Such oxides, despite being very common, are neglected in the present literature on RPs. Their complex crystalline structure is reflected in their dielectric function ( $\varepsilon_{\text{oxide}}$ ), as will be discussed.

An RP is a mixed excitation resulting from the strong coupling between the TO phonons and the IR photon field in dielectric (oxide or semiconductor) layers.<sup>2</sup> Unlike nonradiative polaritons,<sup>1</sup> RPs have a phase velocity such that  $(\omega/|\mathbf{k}|)^2 > (c)^{2,2}$  where  $\omega$  is the angular frequency,  $|\mathbf{k}|$  is the modulus of the wave-vector, and c is the speed of light. A consequence of this physical paradox is that the RPs leak energy to the surrounding space.<sup>2</sup> The RP frequency is therefore complex:  $\omega = \operatorname{Re}(\omega) + i\operatorname{Im}(\omega)$ , where  $\operatorname{Re}(\omega)$  is the central frequency, and  $Im(\omega)$  gives the spread in frequency of the RP.<sup>2,19</sup> The value of  $2|Im(\omega)|$  is the inverse of the lifetime (or damping)<sup>2,19</sup> of the RP, which is of the order of picoseconds. Since the wave vector  $|\mathbf{k}|$  is also complex,<sup>2,19</sup> the dispersion relations of RPs are represented by plotting  $Re(\omega)$  and  $Im(\omega)$  versus the angle of incidence  $\theta_0$  of the IR radiation, not versus k. This procedure is possible because the

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relationship between **k** and  $\theta_0$  is, e.g., in the *x*-axis,  $k_x = |\mathbf{k}| \sin \theta_0$ .<sup>2,19</sup> Radiative polaritons appear at frequencies above or below the  $\omega_{\rm LO}$  and  $\omega_{\rm TO}$  frequencies. The frequency and number of RPs that the substrate can support depend on the thickness *d* of the dielectric layer.<sup>2</sup> For dielectric layers on a metallic substrate, only half of all possible RPs exist because the boundary conditions eliminate those RPs that are accompanied by electric fields parallel to the metal surface.<sup>2,25</sup>

In this work, the simulations are performed utilizing a wellestablished method based on Fresnel's equations extended to multilayer systems through a system of transfer matrices. These matrices can efficiently describe the refraction and reflection of IR radiation incident on multilayer systems<sup>26</sup> at various values of incidence angle  $\theta_0^{14}$  and oxide film thickness d,<sup>15</sup> and for various geometries.<sup>20</sup> Here, we examine the absorption of IR radiation by  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and antase-TiO<sub>2</sub> with dielectric functions  $\epsilon_{Al_2O_3}$  and  $\epsilon_{TiO_2}$ , respectively.<sup>27</sup> The dielectric functions depend not only on oxide chemical composition, but also on its crystalline structure; e.g.,  $\epsilon_{TiO_2}$  is different for anatase and rutile TiO<sub>2</sub>.<sup>20,28</sup> The angular dependence for absorption is investigated in the 0 to 81° range.<sup>11,25</sup> The dispersion relations of  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> films will be discussed in a separate report.

#### SIMULATION PROCEDURE

This work considers thin oxide layers with  $d < \lambda_{IR}$ , where d is the oxide film thickness and  $\lambda_{IR}$  is the wavelength of the IR radiation (15–30  $\mu$ m). According to the description of IR absorption by oxide films in terms of RP coupling given by Kliewer et al.,<sup>2,25</sup> in thin oxide films, for  $\text{Re}(\omega) \ge \omega_{\text{LO}}$ , only the 0TH type-RP exists. This RP is a solution to a tangenttype (T) of dispersion relation  $-i\varepsilon_{\text{oxide}}(\beta_0/\beta) = \tan\beta d$ . This kind of solution generates polarization vectors that always have a non-zero component normal at the oxide layer. In the above equation,  $\beta_0 = \sqrt{(\omega^2/c^2) - k_x^2}$ , and  $\beta = \sqrt{\epsilon_{\text{oxide}}(\omega^2/c^2) - k_x^2}$ , where  $\epsilon_{\text{oxide}}$  is the dielectric function of the oxide layer. The 0TH type-RP has  $Re(\omega) \ge \omega_{LO}$ , and it is labeled with H to emphasize that it is a high-frequency RP.<sup>2</sup> On the other hand, the 2TL and 3CL type-RPs are the solutions of a tangent- and a cotangent-type (C) of dispersion relation, respectively. The cotangent-type dispersion relation is  $i\varepsilon_{\text{oxide}}(\beta_0/\beta) = \cot\beta d$ , and its solutions generate polarization vectors whose component normal to the oxide layer surface is always zero. The 2TL and 3CL type-RPs have  $\text{Re}(\omega) < \omega_{\text{TO}}$ . These RPs are labeled with L, to emphasize that they are lowfrequency RPs.<sup>2</sup> In summary, three types of RPs are excited in oxide films with thicknesses  $d \le 0.1 \lambda_{IR}$ ,<sup>2,25</sup> and these RPs are assumed to be responsible for the absorption of IR radiation: the 0TH,<sup>2</sup> appearing at Re( $\omega$ )  $\geq \omega_{LO}$ , and the 2TL/ 3CL,<sup>2</sup> appearing at Re( $\omega$ )  $\leq \omega_{TO}$ . The optical properties of oxide films have been discussed in terms of RPs by various authors.8,12,13,18,29-31

The simulations apply to a four-layer system represented in Fig. 1, where it is assumed that the surfaces are perfectly smooth and parallel, consisting of air (layer 1), oxide film (layer 2), substrate (layer 3), and air (layer 4). The IR radiation illuminates the sample at an angle of incidence  $\theta_0$  and changes its direction of propagation according to Snell's law, as schematically described in Fig. 1. The method by Hansen,<sup>26</sup> used to derive the IR transmittance (*T*) and reflectance (*R*) through the four layers, characterizes each layer *j* by a transfer



Fig. 1. Schematic of the four-layer system whose transmittance *T*, reflectance *R*, and absorptance *A* are studied using Fresnel's equation with the transfer matrix method formulated by Hansen.<sup>26</sup> Layer 1 is air, layer 2 is the oxide film, layer 3 is the substrate, and layer 4 is again air. The angle  $\theta_0$  is the incidence angle of the IR radiation on the planar surface, and the path of the IR radiation is traced by the arrowed straight lines. The dashed-dotted arrowed lines represent the reflected IR radiation.

matrix:

$$M_{j} = \begin{vmatrix} \cos\beta_{j} & -\frac{i}{p_{j}}\sin\beta_{j} \\ -ip_{j}\sin\beta_{j} & \cos\beta_{j} \end{vmatrix}$$
(1)

with  $\beta_j = (2\pi/\lambda)d_jN_j\cos\theta_j$ , where  $d_j$ ,  $N_j$ , and  $\theta_j$  are, respectively, layer thickness, refractive index, and orientation with respect to the IR radiation. The term *i* is the imaginary unit,  $p_{/TE} = (N_j/\sqrt{\mu_j})\cos\theta_j$ , and  $p_{JTM} = (\sqrt{\mu_j}/N_j)\cos\theta_j$ , where TE (transverse electric) and TM (transverse magnetic) denote the polarization of the IR radiation, and  $\mu_j$  is the magnetic permeability of the layer. Transmittance (*T*), reflectance (*R*), and absorptance (A = 100% - T - R) of thin  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub> films are all obtained as a percentage of the total IR radiation incident onto the planar oxide films. The IR radiation of interest for the present investigation is located in the low part of the middle-IR region (400–1800 cm<sup>-1</sup>), where the relevant 0TH and 2TL/3CL type-RPs can experimentally be detected.

The dielectric constant for air (layers 1 and 4) is  $\epsilon_{air} = 1.00059 + i0.00004,^{32}$  where the imaginary part causes the absorption of IR radiation in air. The dielectric function  $\epsilon(\omega)_{Al_2O_3}$  for  $\gamma$ -Al\_2O\_3 (layer 2) is given by Eq. 2 and the parameters are listed in Table I:^{33}

$$\varepsilon(\omega)_{Al_2O_3} = \varepsilon^{\infty} \prod_{m=1}^n \frac{\omega_{LOm}^2 - i\omega\gamma_{LOm} - \omega^2}{\omega_{TOm}^2 - i\omega\gamma_{TOm} - \omega^2}$$
(2)

Here, *n* is the total number of oscillators,  $\gamma_{\text{LO}m}$  and  $\gamma_{\text{TO}m}$  are their damping factors, and  $\varepsilon_{\infty} = Lim_{\omega \to \infty}\varepsilon(\omega)_{\text{oxide}}$ . For anatase-TiO<sub>2</sub> (also layer 2), the dielectric function  $\varepsilon(\omega)_{\text{TiO}_2}$  is given by the linear combination of the A<sub>2u</sub> and the E<sub>u</sub> symmetry contributions<sup>34,35</sup> as follows:

$$\varepsilon(\omega)_{\text{TiO}_2} = \frac{\varepsilon(\omega)_{A_{2u}} + 2\varepsilon(\omega)_{E_u}}{3}$$
(3)

The functions  $\epsilon(\omega)_{A_{2u}}$  and  $\epsilon(\omega)_{E_u}$  appear as Eq. 2, with the parameters reported in Table II.<sup>35</sup> As seen in Tables I and II,  $\gamma$ -

TABLE I. Type of RP, and frequencies (in cm<sup>-1</sup>) plus damping factors of the phonons related to the RPs in  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>.<sup>33</sup> These parameters are used in the IR spectra simulations. The value of  $\epsilon_{\infty}$  is 2.91.<sup>33</sup>

Radiative polariton	Frequency ( $\omega$ ) [cm <sup>-1</sup> ] and related phonon	Damping factor ( $\gamma$ ) [cm <sup>-1</sup> ]
_	357, TO1	72
2TL	536, TO2	101
3CL		
-	744, TO3	56
-	807, TO4	43
-	403, LO1	79
-	669LO2	68
-	783, LO3	53
0TH	917, LO4	49

TABLE II. Type of RP, and frequencies (in cm<sup>-1</sup>) plus damping factors of the phonons related to the RPs in anatase-TiO<sub>2</sub>.<sup>35</sup> These parameters are used in the IR spectra simulations. The value of  $\epsilon_{\infty}$  for the A<sub>2u</sub> symmetry is 5.41, whereas the one for the E<sub>u</sub> symmetry is 5.82.<sup>35</sup>

Radiative polariton	Frequency ( $\omega$ ) [cm <sup>-1</sup> ] and related phonon	Damping factor $(\gamma) \ [cm^{-1}]$
-	367. TO1 - A <sub>20</sub>	68
-	755, LO1 - A <sub>2u</sub>	79
-	262, TO2 - E <sub>u</sub>	36
-	366, LO2 - E <sub>u</sub>	4.1
2TL	435, TO3 - E <sub>u</sub>	32
3CL		
0TH	876, LO3 - E <sub>u</sub>	33

Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub> have four and three TO and LO frequency couples (related by the Lyddane–Sachs–Teller relation), respectively. Previous reports<sup>14,34</sup> show that the strongly absorbing 0TH type-RPs, which we study in this work, are near the  $\omega_{LO4}$  and  $\omega_{LO3}$  frequencies for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub>, respectively. The 2TL/3CL type-RPs, on the other hand, appear near the  $\omega_{TO2}$  and  $\omega_{TO3}$  frequencies for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub>, respectively.<sup>14,34</sup>

Aluminum (Al), silicon (Si), and silicon dioxide (SiO<sub>2</sub>) are selected as substrate materials (layer 3) of different conductivities and defined thicknesses. For a 10 µm thick Al substrate, the dielectric function is  $\varepsilon_{A1} = 1 + (i2\sigma/\omega)$ , with  $\sigma = (\sigma_0/1 - 2i\pi\tau\omega c)$ . Here,  $\sigma_0 = 3.67 \times 10^{17} \text{ s}^{-1}$  is the direct current electrical conductivity, and  $\tau = 0.8 \times 10^{-14} \text{ s}$  is the relaxation time. The Si substrate has the dielectric function  $\varepsilon(\omega)_{Si} = 12.25 - i0.021$ ,<sup>36</sup> and a 0.5 µm thickness. A thicker Si substrate would introduce interference fringes that limit the observation of the features in the IR spectra.<sup>37</sup> For a 10 µm thick SiO<sub>2</sub> substrate, the dielectric function is also given by Eq. 2, with parameters from Kirk<sup>39</sup> summarized in Table III. The simulated spectra obtained in this work are the result of the interaction between oxide films and IR radiation that can be characterized as consisting of 50% TE and 50% TM polarized radiation.<sup>39</sup>

The classical skin depth is calculated as follows:<sup>40,41</sup>

$$\delta = \frac{1}{\alpha} = \frac{\lambda_{\rm IR}}{2\pi} \frac{1}{{\rm Im}N(\omega)_{\rm oxide}} \tag{4}$$

TABLE III. Type of RP, and frequencies (in cm<sup>-1</sup>) plus damping factors of the phonons related to the RPs in SiO<sub>2</sub>.<sup>38</sup> These parameters are used in the IR spectra simulations. The value of  $\varepsilon_{\infty}$  is 2.356.<sup>38</sup>

Radiative polariton	Frequency ( $\omega$ ) [cm <sup>-1</sup> ] and related phonon	Damping factor $(\gamma) \ [cm^{-1}]$
2TL 3CL 0TH	1072, TO1	30
	1235, LO1	35

where  $\alpha$  is the film absorption coefficient, and  $\alpha = (c/\omega)\text{Im}N(\omega)_{\text{oxide}}$ . The thin oxide film refractive index is  $N(\omega)_{\text{oxide}} = \sqrt{\epsilon(\omega)}_{\text{oxide}}$ , where  $\epsilon(\omega)_{\text{oxide}}$  is given by Eq. 2 for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and by Eq. 3 for anatase-TiO<sub>2</sub>. Figure 2 plots  $\delta$  versus frequency, which shows in Fig. 2a that  $\delta = 4 \ \mu\text{m}$  at  $\omega_{\text{LO4}}$  and  $\delta = 1.1 \ \mu\text{m}$  at  $\omega_{\text{TO2}}$  for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>, and in Fig. 2b that  $\delta = 4.5 \ \mu\text{m}$  at  $\omega_{\text{LO3}}$  and  $\delta = 0.6 \ \mu\text{m}$  at  $\omega_{\text{TO3}}$  for anatase-TiO<sub>2</sub>.

#### **RESULTS AND DISCUSSION**

Figure 3 shows the simulated percent (%) absorptance spectra (A = 100% - T - R) obtained for various film thicknesses from 50 nm to 3  $\mu$ m at  $\theta_0 = 30^\circ$  and  $60^\circ$  for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al,  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si, and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub>. The  $\omega_{LO4}$  and  $\omega_{TO2}$ frequencies are indicated by the vertical lines. The 0TH and 2TL/3CL type-RPs appear near the  $\omega_{LO4}$  and  $\omega_{TO2}$  frequencies, respectively. In panels (c) and (f) of Fig. 3, additional vertical lines point out the  $\omega_{LO}$  and  $\omega_{TO}$  frequencies of SiO<sub>2</sub> (labeled LO<sub>s</sub> and TO<sub>s</sub>, respectively).<sup>38</sup> The 0TH and 2TL/3CL type-RPs for SiO<sub>2</sub> appear near these frequencies. The same sequence of spectra is illustrated in Fig. 4 for anatase-TiO<sub>2</sub> in a thickness range from 50 nm to 7  $\mu$ m. The spectra in Figs. 3 and 4 indicate that the absorptance peaks at the  $\omega_{LO4}$  and  $\omega_{LO3}$ frequencies, corresponding to the 0TH type-RP, increase significantly by increasing the oxide film thickness and the IR radiation incidence angle  $\theta_0$ . The angular dependence of the absorptance of the 0TH type-RP near the  $\omega_{LO4}$  and  $\omega_{LO3}$ frequencies is known as the Berreman effect.<sup>11</sup> Its thicknessdependent increase and the blue-shift in frequency is in agreement with the RPs behavior described by Kliewer et al.<sup>2,25</sup> Similar trends are also revealed in the reflectance and transmittance spectra, also in agreement with Kliewer et al.<sup>25</sup> (spectra not shown).

From Figs. 3 and 4 we infer that the thickness- and angledependence of the absorption of the 0TH and 2TL/3CL type-RPs is related to the chemistry and conductivity of the substrate (specifically Al, Si, and SiO<sub>2</sub>). In addition, the results on the absorption by oxide films on the insulating SiO<sub>2</sub> substrate (Figs. 3c and 3f, and Figs. 4c and 4f) are largely affected by the 2TL/3CL and 0TH type-RPs of SiO<sub>2</sub>, as already noted by Ahn et al.<sup>21</sup>

The 0TH and 2TL/3CL type-RPs absorptance dependence on thickness for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub> films is illustrated in Figs. 5 through 8. Figure 5 shows the absorptance spectra of the 2TL/3CL type-RPs versus oxide film thicknesses near  $\omega_{TO2}$ for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al (a),  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si (b), and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> (c) in 30 nm to 5  $\mu$ m thick films, and near  $\omega_{TO3}$  for anatase-TiO<sub>2</sub>/Al (d), anatase-TiO<sub>2</sub>/Si (e), and anatase-TiO<sub>2</sub>/SiO<sub>2</sub> (f) in 30 nm to 9  $\mu$ m thick films. The simulations were performed fixing the incidence angle  $\theta_0$  at 0°, 30°, 60°, and 81°. The data indicate



FIG. 2. The skin depth  $\delta$  as a function of the angular frequency  $\omega$  for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> (**a**), and anatase-TiO<sub>2</sub> (**b**). The vertical lines indicate the location of the  $\omega_{LO}$  and  $\omega_{TO}$  frequencies of the oxides in the middle-IR region from 400 to 1800 cm<sup>-1</sup>.



FIG. 3. Simulated absorptance spectra (100% - T - R) obtained for various film thicknesses at  $\theta_0 = 30^\circ$  for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al (**a**),  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si (**b**), and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> (**c**), and at  $\theta_0 = 60^\circ$  for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al (**d**),  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si (**e**), and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> (**f**). The vertical lines indicate the  $\omega_{LO4}$  and  $\omega_{TO2}$  frequencies. The 0TH and 2TL/3CL type-RPs are the peaks that appear near the  $\omega_{LO4}$  and  $\omega_{TO2}$  frequencies, respectively. In panels (**c**) and (**f**) additional vertical lines indicate the LO<sub>s</sub> and TO<sub>s</sub> frequencies of SiO<sub>2</sub>.<sup>38</sup> The 0TH and 2TL/3CL type-RPs of SiO<sub>2</sub> are the peaks near the LO<sub>s</sub> and TO<sub>s</sub> frequencies, respectively. The oxide film thicknesses are labeled in panel (**a**), and the sequence is the same in all panels.



FIG. 4. Simulated absorptance spectra (100% - T - R) obtained for various film thicknesses at  $\theta_0 = 30^\circ$  for anatase-TiO<sub>2</sub>/Al (**a**), anatase-TiO<sub>2</sub>/Si (**b**), and anatase-TiO<sub>2</sub>/SiO<sub>2</sub> (**c**), and at  $\theta_0 = 60^\circ$  for anatase-TiO<sub>2</sub>/Al (**d**), anatase-TiO<sub>2</sub>/Si (**e**), and anatase-TiO<sub>2</sub>/SiO<sub>2</sub> (**f**). The vertical lines indicate the  $\omega_{LO4}$  and  $\omega_{TO2}$  frequencies. The 0TH and the 2TL/3CL type-RPs are the peaks that appear near the  $\omega_{LO4}$  and  $\omega_{TO2}$  frequencies, respectively. In panels (**c**) and (**f**) additional vertical lines point out the LO<sub>s</sub> and TO<sub>s</sub> frequencies of SiO<sub>2</sub>.<sup>38</sup> The 0TH and 2TL/3CL type-RPs of SiO<sub>2</sub> are the peaks near the LO<sub>s</sub> and TO<sub>s</sub> frequencies, respectively. The oxide film thicknesses are labeled in panel (**a**), and the sequence is the same in all panels.

that for  $d > \delta$  the dependence of absorption by the 2TL/3CL type-RPs on oxide film thickness is not affected by the chemistry and conductivity of the substrates. For  $d < \delta$ , however, the absorption by the 2TL/3CL type-RPs differs depending upon substrate chemistry and conductivity but does not depend upon  $\theta_0$  (except when  $\theta_0 > 80^\circ$ ). These phenomena can readily be observed in Fig. 6, which displays the same data as in Fig. 5 for d < 250 nm. In Fig. 6 we observe that the 2TL type-RP does not absorb IR radiation for d < 250 nm in  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub> films on Al. This phenomenon derives directly from the model described by Kliewer et al.,<sup>2,25</sup> which predicts that ultra-thin dielectric layers on a metallic substrate only support 2TL type-RPs with very long lifetime, thus ultranarrow band width, and therefore cannot be observed. Surprisingly, this effect of the substrate on the absorption is more pronounced for oxide films less than 250 nm thick, and this result was not predicted in the original theory by Kliewer et al.<sup>2</sup> These conclusions are only slightly affected by the different oxide chemistry and crystalline structure of  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub>.

Figure 7 shows the simulated percent (%) absorptance spectra versus film thicknesses of the 0TH type-RP near  $\omega_{LO4}$ for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al (a),  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si (b), and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> (c) in films from 30 nm to 5 µm thick, and near  $\omega_{LO3}$  for anatase-TiO<sub>2</sub>/Al (d), anatase-TiO<sub>2</sub>/Si (e), and anatase-TiO<sub>2</sub>/SiO<sub>2</sub> (f) in films from 30 nm to 9 µm thick. Also in this case the simulations were performed by fixing the angle of incidence  $\theta_0$ at 0°, 30°, 60°, and 81°. The data suggest that for  $d > \delta$  the dependence of the absorption by the 0TH type-RP on the oxide film thickness is not affected by the chemistry and conductivity of the substrate. Instead, for  $d < \delta$ , the absorption of the 0TH type-RPs always increases with  $\theta_0$  (Berreman effect)<sup>11</sup> but in a different way depending upon the chemistry and conductivity of the substrate. The absorption increase with  $\theta_0$  is highly



FIG. 5. Percent reflectance (*R*, empty symbols) and absorptance (*A*, filled symbols) versus oxide film thickness at  $\theta_0 = 0^\circ$  (squares), 30° (circles), 60° (triangles), and 81° (stars). For the 2TL/3CL type-RPs, the peaks near  $\omega_{TO2}$  were considered for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al (**a**),  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si (**b**), and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> (**c**), whereas those near  $\omega_{TO3}$  were considered for anatase (**a**)-TiO<sub>2</sub>/Al (**d**), anatase-TiO<sub>2</sub>/Si (**e**), and anatase-TiO<sub>2</sub>/SiO<sub>2</sub> (**f**). The thick lines on the left of the graphs labeled "sst" represent the *substrate sensitive thickness*.

pronounced on the metallic substrate for all oxide thicknesses, especially below 250 nm, as can clearly be observed in Figs. 8a and 8d. For oxide films on semiconducting (Figs. 7b and 7e, and Figs. 8b and 8e) and insulating substrates (Figs. 7c and 7f, and Figs. 8c and 8f) the increased absorption as  $\theta_0$  increases, rises up to about 70°, and drops thereafter. Additionally, when dealing with insulating substrates, there is little variation of absorption with oxide film thickness, even for very thin films with  $d \ll 250$  nm. In particular, our simulated spectra indicate that the Berreman effect appears only in oxide films with d <250 nm, especially for the films on a metallic substrate. This result was also not predicted in the theory by Kliewer et al.<sup>2</sup> However, in agreement with the Berreman effect<sup>11</sup> and the model described by Kliewer et al.,<sup>2,14</sup> we observe that the 0TH type-RPs do not absorb IR radiation at  $\theta_0 = 0^\circ$ . These conclusions hold regardless of the oxide film chemistry

considered. The reflection (R) data in Figs. 5 through 8 support the conclusions derived from the absorptance (A) data (spectra not shown).

Because ALD films are very uniform in thickness and smooth,<sup>24</sup> we do not expect film thickness deviations and interface roughness to interfere with the reported results.

#### SUMMARY AND CONCLUSIONS

We presented a study of the absorption of IR radiation by selected RPs in oxide films of variable thicknesses and orientations. The observed RPs are the 0TH-type, appearing near the  $\omega_{LO4}$  frequency for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> films and near the  $\omega_{LO3}$  frequency for anatase-TiO<sub>2</sub>, and the 2TL and 3CL-types, appearing near the  $\omega_{TO2}$  frequency for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> films and near the  $\omega_{TO3}$  frequency for anatase-TiO<sub>2</sub>. The Hansen method was



FIG. 6. Same as Fig. 5, zoomed into the oxide film thickness range between 30 and 250 nm.

used to simulate transmittance (*T*), reflectance (*R*), and absorptance (100% - T - R) of thin  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub> films on Al, Si, and SiO<sub>2</sub>.

The results presented in Figs. 5 through 8 show that the chemistry and the conductivity of the substrate determine the amount of absorption by the 0TH and 2TL/3CL type-RPs for oxide thicknesses less than the skin depth ( $d < \delta$ ). In oxide layers with thickness  $d > \delta$ , there is no significant dependence of the RP absorption on the chemistry and conductivity of the substrate. This conclusion holds especially for the 2TL/3CL type-RPs.

Furthermore, the data in Figs. 7 and 8 indicate that the Berreman effect (i.e., the dependence of the absorption of the 0TH type-RP on the IR radiation incidence angle  $\theta_0$ ) is apparent in oxide films below about 250 nm, especially for oxide films on metallic substrates. The strength of absorption by the 0TH type-RP shows a dependence on film substrate

previously unnoted, and not discussed in the original theory by Kliewer et al.<sup>2,25</sup> This phenomenon is, however, built into the Kliewer et al. formalisms describing the optical properties of the RPs.<sup>14,25</sup>

Finally, the data in Figs. 7 and 8 indicate that, in general, oxide chemistry, such as that of  $\gamma$ -Al<sub>2</sub>O<sub>3</sub> and anatase-TiO<sub>2</sub> films on various substrates, only slightly affects the absorption of IR radiation by RPs, but substrate chemistry and conductivity can be noticeable in very thin oxide films with d < 250 nm. This finding was obtained by combining the Kliewer et al.<sup>2,25</sup> theory with the Fresnel's equations through the Hansen method<sup>26</sup> for simulating IR spectra of oxide films on substrates of any type of chemistry and conductivity. The upper limit thickness where the IR radiation interaction with the thin oxide film significantly depends on substrate chemistry and conductivity is labeled in Figs. 5



Fig. 7. Percent reflectance (*R*, empty symbols) and absorptance (*A*, filled symbols) versus oxide film thickness at  $\theta_0 = 0^\circ$  (square), 30° (circle), 60° (triangle), and 81° (star). For the 0TH type-RP, the peaks near  $\omega_{LO4}$  were considered for  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Al (**a**),  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/Si (**b**), and  $\gamma$ -Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> (**c**), whereas that near  $\omega_{LO3}$  was considered for anatase (**a**)-TiO<sub>2</sub>/Al (**d**), anatase-TiO<sub>2</sub>/Si (**e**), and anatase-TiO<sub>2</sub>/SiO<sub>2</sub> (**f**). The thick lines on the left of the graphs labeled "sst" represent the *substrate sensitive thickness*.

and 7 as the substrate sensitive thickness (sst), or d = 250 nm.

In conclusion, our data show that (1) the chemistry and the conductivity of the substrate affect the IR radiation absorption by RPs in oxide films thinner than their IR skin depth  $\delta$ , (2) the Berreman effect (the dependence of the absorption of the 0TH type-RP on the IR radiation incidence angle  $\theta_0$ ) occurs in oxide films on a metallic substrate, and (3) the chemistry of the oxide on various substrates only slightly affects the absorption by RPs, especially for oxide films thinner than 250 nm thick. Point (1) is supported by the behavior of the 0TH type-RP, where the increase in absorptance is higher in oxide

films on Al, followed by those on Si and, lastly, on SiO<sub>2</sub>. These results suggest that the chemistry and the conductivity of the substrate determine the amount of absorption by RPs in oxide layers thinner than the skin depth. In addition, the ability of an oxide film to interact with IR radiation is effective below the *substrate sensitive thickness*, which is about 250 nm regardless of oxide film chemistry.

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FIG. 8. Same as Fig. 7, zoomed into the oxide film thickness range between 30 and 250 nm.

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- 1. K. L. Kliewer and R. Fuchs, Phys. Rev. 144, 495 (1966).
- 2. K. L. Kliewer and R. Fuchs, Phys. Rev. 150, 573 (1966).
- 3. S. Ghassemi and J. S. Hammonds, Jr., Phys. Lett. A 374, 4945 (2010).
- M. Francoeur, R. Vaillon, and M. P. Mengüç, IEEE T. Energy Conver. 26, 686 (2011).
- M. Francoeur, M. Pinar Mengüç, and R. Vaillon, Appl. Phys. Lett. 93, 043109 (2008).
- 6. S. Shen, A. Narayanaswamy, and G. Chen, Nano Lett. 9, 2909 (2009).
- E. Rousseau, A. Siria, G. Jourdan, S. Volz, F. Comin, J. Chevrier, and J.-J. Greffet, Nat. Photonics 3, 514 (2009).
- 8. H. Tanaka and K. Hisano, J. Phys.: Condens. Matter 1, 9539 (1989).
- 9. E. A. Vinogradov, Phys. Reports 217, 159 (1992).
- 10. E. A. Vinogradov, Physics-Uspekhi 45, 1213 (2002).
- 11. D. W. Berreman, Phys. Rev. 130, 2193 (1963).
- 12. R. Kälin and F. Kneubühl, Infrared Phys. 16, 491 (1976).
- 13. S. Mochizuki, Phys. Stat. Sol. B 126, 105 (1984).
- G. Scarel, J.-S. Na, and G. N. Parsons, J. Phys. Condens. Matter 22, 155401 (2010).

- G. Scarel, J.-S. Na, B. Gong, and G. N. Parsons, Appl. Spectrosc. 64, 120 (2010).
- M. Klevenz, S. Wetzel, M. Trieloff, H.-P. Gail, and A. Pucci, Phys. Stat. Sol. B 247, 2179 (2010).
- B. C. Trasferetti, C. U. Davanzo, R. A. Zoppi, N. C. da Cruz, and M. A. B. de Moraes, Phys. Rev. B 64, 125404 (2001).
- J. Rivnay, L. Jimison, M. F. Toney, M. Preiner, N. A. Melosh, and A. Salleo, J. Vac. Sci. Technol. B 26, 1454 (2008).
- 19. F. Comas, H. Calas, and C. Trallero-Giner, Phys. Rev. B 60, 8238 (1999).
- G. Scarel, G. K. Hyde, D. Hojo, and G. N. Parsons, J. Appl. Phys. 104, 094314 (2008).
- 21. J. S. Ahn, H. S. Choi, and T. W. Noh, Phys. Rev. B 53, 10310 (1996).
- V. A. Yakovlv, N. N. Novikova, E. A. Vinogradov, G. Rossetto, A. Sartori, and M. Bolzan, J. Nanopart. Res., February 17 (2011), doi: 10. 1007/s1105-011-0280-8
- G. D. Wilk, R. M. Wallace, and J. M. Anthony, J. Appl. Phys. 89, 5243 (2001).
- M. Ritala and M. Leskelä, in *Handbook of Thin Film Materials*, H. S. Nalwa, Ed. (Academic Press, San Diego, CA, 2002), 1, pp. 103–159.
- 25. R. Fuchs, K. L. Kliewer, and W. J. Pardee, Phys. Rev. 150, 589 (1966).
- 26. W. N. Hansen, J. Opt, Soc. Am. 58, 380 (1968).
- 27. D. W. Berreman and F. C. Unterwald, Phys. Rev. 174, 791 (1968).

- 28. G. Scarel, C. J. Hirschmugl, V. V. Yakovlev, R. S. Sorbello, C. R. Aita, H. Tanaka, and K. Hisano, J. Appl. Phys. 91, 1118 (2002).
- 29. L. Wendler and T. Kraft, Physica B 271, 33 (1999).
- 30. H. Welsch and J. Lafait, Opt. Commun. 116, 369 (1995).
- 31. M. Załuźny and W. Zietkowski, Physica E 13, 370 (2002).
- 32. L. G. Hector and H. L. Schulz, J. Appl. Phys. 7, 133 (1936).
- Y. T. Chu, J. B. Bates, C. W. White, and G. C. Farlow, J. Appl. Phys. 64, 3727 (1988).
- G. Scarel, C. R. Aita, H. Tanaka, and K. Hisano, J. Non-Cryst. Solids 303, 50 (2002).
- 35. R. J. Gonzalez, R. Zallen, and H. Berger, Phys. Rev. B 55, 7014 (1997).
- R. D. Kekatpure, E. S. Barnard, W. Cai, and M. L. Brongersma, Phys. Rev. Lett. 104, 243901 (2010).
- 37. K. Yamamoto and H. Ishida, Appl. Opt. 34, 4177 (1995).
- 38. C. T. Kirk, Phys. Rev. B 38, 1255 (1988).
- J. Ibáňez, E. Tarhan, A. K. Ramdas, S. Hernández, R. Cuscó, L. Artús, M. R. Melloch, and M. Hopkinson, Phys. Rev. B 69, 075314 (2004).
- G. R. Fowles, *Introduction to Modern Optics* (New York, Dover Publications, 1975), pp. 157.
- 41. J. D. Jackson, *Classical Electrodynamics* (New York, Wiley, 1999), pp. 310.