

1       **Mercury anomalies associated with three extinction events**  
2       **(Capitanian Crisis, Latest Permian Extinction, and the**  
3       **Smithian/Spathian Extinction) in NW Pangea**  
4

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15

## 16 **Abstract**

17 Permian through Early Triassic strata that include a record of three major extinction events  
18 (Capitanian Crisis, Latest Permian Extinction, and the Smithian/Spathian Extinction) were  
19 examined at the Festningen section, Spitsbergen. Over the ~ 12 Ma record examined, mercury in  
20 the sediments shows relatively constant background values of 0.005 to 0.010  $\mu\text{g/g}$ . However,  
21 there are notable spikes in Hg concentration, over an order of magnitude above background,  
22 associated with the three extinctions. The Hg/TOC ratio shows similar large spikes, indicating  
23 that they represent true increase in Hg loading to the environment. We argue that these represent  
24 Hg loading events associated with enhanced Hg emissions from large igneous province (LIP)  
25 events that are synchronous with the extinctions. The Hg anomalies are consistent across the NW  
26 margin of Pangea, indicating widespread mercury loading occurred. While this provides utility as  
27 a chemostratigraphic marker, the Hg spikes may also indicate loading of toxic metals to the  
28 environment that would be a contributing cause to the mass extinction events.

29 **Keywords:** mercury, Early Triassic, Latest Permian Extinction, chemostratigraphy

30

## 31 **1. Introduction**

32 Mercury (Hg) emissions associated with the emplacement of Large Igneous Provinces (LIPs)  
33 were first recognised by (Sanei et al., 2012), who showed a large Hg spike associated with the  
34 Siberian Traps eruptions. This event was coincident with the Latest Permian Extinction (LPE),  
35 the largest extinction in Earth History that had a devastating impact on both terrestrial and  
36 marine ecosystems (Erwin, 2006). High Hg loading associated with the Siberian Traps has been  
37 supported by a similar Hg spike at the LPE boundary in Spitsbergen (Grasby et al., 2015).  
38 Recurrent Siberian Trap volcanism may have also influenced Hg loading during the  
39 Smithian/Spathian Extinction in the Sverdrup Basin (Grasby et al., 2013a). While these initial  
40 studies are focused on the sedimentary records from NW Pangea, the global extent of Hg loading  
41 related to Siberian Trap volcanism has yet to be demonstrated. However, subsequent work has  
42 shown similar Hg anomalies associated with other LIP events such as the Cretaceous–Paleogene  
43 transition (Sial et al., 2014; Sial et al., 2013; Silva et al., 2013) related to Deccan Trap volcanism.

44 Hg is extremely toxic to life. This combined with the ease of transport over long distances  
45 and persistence of Hg in the environment makes modern anthropogenic mercury emissions the  
46 subject of significant global concern (AMAP, 2011). The two largest natural source of Hg to the  
47 environment are volcanic emissions and natural coal combustion (Pirrone et al., 2010). These  
48 sources release Hg to the atmosphere where it can be globally transported prior to deposition in  
49 terrestrial and marine environments. In the marine environment organic matter and clay minerals  
50 scavenge Hg and transports it to the sea floor to become fixed in bottom sediments (Gehrke et  
51 al., 2009; Turner et al., 2004; Andren and Harriss, 1975; Lindberg et al., 1975; Cranston and  
52 Buckley, 1972). The control of primary productivity on Hg sequestration is shown by the close

53 relationship between sedimentary organic matter and Hg in modern (Sanei et al., 2014; Stern et  
54 al., 2009; Outridge et al., 2007), as well as ancient sediments (Grasby et al., 2013a).

55 Modern volcanic eruptions have a significant Hg flux that produce global Hg anomalies  
56 (Pyle and Mather, 2003; Schuster et al., 2002; Slemr et al., 1995; Slemr and Scheel, 1998). Hg is  
57 sourced from both volcanic gases as well as from rock units intruded by magma (Grasby et al.,  
58 2015; Sanei et al., 2015; Sanei et al., 2012). During LIP events emission rates would greatly  
59 exceed normal Hg release (Sanei et al., 2012; Grasby et al., 2015), such that the normal marine  
60 buffering control on Hg can be potentially overwhelmed (Sanei et al., 2012; Grasby et al., 2013a)  
61 and be recorded as a Hg spike in the sediment record. For instance Grasby et al. (2015) estimated  
62 that the Siberian Traps may have released up to 9.98 Gg/a Hg, or 400% above modern natural  
63 emissions. In comparison, the modern anthropogenic Hg emissions that roughly equal natural  
64 are subject of global concern over impact on marine and terrestrial ecosystems. Therefore, Hg  
65 anomalies formed at time of LIP eruptions suggest that: 1) Hg could be an effective marker in  
66 the geologic record of periods of enhanced volcanic activity, and 2) LIP eruptions could  
67 potentially release toxic amounts of Hg to the environment. Such a scenario would add to the  
68 variety of kill mechanisms already associated with major volcanic eruptions (Keller and Kerr,  
69 2014), as well as providing a direct link between terrestrial and marine extinction.

70 To further the assessment of Hg in the geologic record, we examined the Permian  
71 through Early Triassic record at the Festningen section in Spitsbergen (Fig. 1). This well-known  
72 location records the LPE (Grasby et al., 2015; Wignall et al., this volume) as well as evidence for  
73 the Earlier Capitanian crisis (Bond et al., 2015) and Smithian/Spathian extinction (Wignall et al.,  
74 this volume), all of which have been linked to periods of major volcanic activity. Of these recent  
75 studies, the Hg record at Festningen has only been examined in a narrow (40 m) zone straddling

76 the LPE boundary (Grasby et al., 2015). In this present paper we examined the Hg record at  
77 Festningen over 600 m of section through the Middle Permian to Early Triassic, representing  
78 ~12 million years of Earth history.

79

## 80 **2. Study Area.**

81 The Festningen section is located at Kapp Starostin, Spitsbergen (Fig. 1). The ~45° eastward  
82 dipping beds occur along a ~7 km low sea-cliff. The section provides a near continuous exposure  
83 of Carboniferous to Cenozoic strata, from Kapp Starostin to Festningsdodden, deposited in a  
84 distal broad epicontinental shelf setting on the northwestern margin of Pangea (Wignall et al.,  
85 1998; Blomeier et al., 2013; Stemmerik and Worsley, 2005). During Permian time Spitsbergen  
86 was at a paleolatitude of ~40-45° N (Scotese, 2004; Golonka and Ford, 2000).

87 The Kapp Starostin Formation was deposited during a period of passive subsidence that  
88 followed a major relative sea level drop at the Lower/Middle Permian boundary (Blomeier et al.,  
89 2013). Deposition of widespread heterozoan carbonate (Vøringen Member) occurred in Roadian  
90 time, followed by the progradation of heterozoan carbonates and cherts over much of the Barents  
91 Shelf and Svalbard (Blomeier et al., 2013). Spitsbergen shares a similar depositional history to  
92 the palaeogeographically adjoining Sverdrup Basin (van Hauen, Degerbøls and Trold Fiord  
93 formations; Beauchamp et al., 2009) (Fig. 2). Approximately 40 m below the contact with the  
94 overlying uppermost Permian–Lower Triassic Vardebukta Formation, fossiliferous carbonates of  
95 the Kapp Starostin Formation transition to Late Permian spiculitic chert (Blomeier et al., 2013).  
96 These Later Permian chert beds are considered equivalent to the Black Stripe and Lindström  
97 formations of the Sverdrup Basin (Beauchamp et al., 2009). The Vardebukta and overlying  
98 Tvillingodden formations are dominated by shale, siltstone and minor sandstone of Early to

99 Middle Triassic age (Mørk et al., 1982), that are equivalent to the Blind Fiord Formation of the  
100 Sverdrup Basin (Embry, 1989). The basal ~6-7 m of the Vardebukta Formation is latest Permian  
101 in age (Grasby et al., 2015; Wignall et al., 1998).

102

### 103 **3. Methods**

104 A detailed sample suite was collected from 100 m below the Kapp Starostin/Vardebukta contact  
105 to 500 m above. Sample spacing varied from 1 to 2 m away from the contact and higher density  
106 of 20 to 50 cm spacing from ~ 40 m below to 19 m above the top of the Kapp Starostin  
107 Formation. Samples were recorded in metres above (positive) and below (negative) the last chert  
108 bed that defines the top of the Kapp Starostin Formation. Weathered surfaces were removed and  
109 then samples were collected from an isolated layer no greater than 2 cm thick. In the laboratory  
110 any remaining weathered surfaces were removed and fresh samples were powdered by agate  
111 mortar and pestle.

112 Total organic carbon (TOC) was measured using Rock-Eval 6<sup>®</sup>, with  $\pm 5\%$  analytical error  
113 of reported value, based on repeats and reproducibility of standards run after every 5<sup>th</sup> sample  
114 (Lafargue et al., 1998). Elemental determinations were conducted on powdered samples digested  
115 in a 2:2:1:1 acid solution of H<sub>2</sub>O-HF-HClO<sub>4</sub>-HNO<sub>3</sub>, and subsequently analysed using a  
116 PerkinElmer Elan 9000 mass spectrometer, with  $\pm 2\%$  analytical error. Hg was measured at  
117 GSC-Atlantic by LECO<sup>®</sup> AMA254 mercury analyzer (Hall and Pelchat, 1997) ( $\pm 10\%$ ).  $\delta^{13}\text{C}_{\text{org}}$   
118 was determined on samples washed with hydrochloric acid, and rinsed with hot distilled water to  
119 remove any carbonate.  $\delta^{13}\text{C}$  was measured using Continuous Flow-Elemental Analysis-Isotope  
120 Ratio Mass Spectrometry, with a Finnigan Mat Delta+XL mass spectrometer interfaced with a  
121 Costech 4010 elemental analyzer with combined analytical and sampling error of  $\pm 0.2\%$ .

122 Analytical results from a total of 341 samples used in this study are provided in Table 1.

123

## 124 **4. Results**

125

### 126 **4.a. Organic Carbon**

127 The organic carbon record, expressed as percent TOC, shows low values (< 1% TOC) overall  
128 (Fig. 3a). There is an increase in TOC at -40 m. Bond et al. (2015) argue, based on brachiopod,  
129 Sr isotopes records and magnetostratigraphy data, that the -40 m level is around the Middle  
130 Permian extinction level at Festningen. The TOC values decline again leading up to the LPE  
131 boundary, where there is a second brief increase in TOC just above the LPE. For the remainder  
132 of the section, TOC values in the Vardebukta Formation are < 0.2%. At these low values the  
133 accuracy of Rockeval results decreases, however data are plotted in Figure 3a to illustrate the  
134 overall low TOC values through this interval. The TOC values increase again in the basal  
135 Tvillingodden Formation, before dropping to low values in the upper 100 m of section.

136 Rock-Eval 6<sup>®</sup> results also provided information on thermal maturation and indicated that  
137 organic matter in the shales have never been heated past the upper end of the oil window, such  
138 that stable isotope values of organic carbon would not be thermally altered (Hayes et al., 1983).  
139 The only exception is a local thermal anomaly associated with a sill emplaced at ~19 m that does  
140 not extend beyond 5 m of the sill boundary (Grasby et al., 2015).

141

### 142 **4.b. Carbon isotope record**

143 Portions of the organic carbon isotope curve ( $\delta^{13}\text{C}_{\text{org}}$ ) at the Festningen section have been  
144 reported in previous studies, including Middle to Late Permian sediments (Bond et al., 2015),

145 across the Latest Permian extinction event (Grasby et al., 2015) and early Triassic sediments  
146 (Wignall et al., this volume). The combined  $\delta^{13}\text{C}_{\text{org}}$  curve is provided here (Fig. 3b) as reference  
147 for comparison with both organic carbon isotope data from NW Pangea and the global inorganic  
148 carbon isotope record. The  $\delta^{13}\text{C}_{\text{org}}$  data from Festningen shows initially relatively stable values  
149 of  $\sim -24$  ‰ through the Middle to Late Permian, with a relatively minor negative shift at  $\sim -40$  m.  
150 At the top of the Kapp Starostin Formation, carbon isotope values show the onset of a  
151 pronounced 10‰ negative shift through the basal Vardebukta Formation sediments (Grasby et  
152 al., 2015). After this point there is a progressive recovery through the next 200 m followed by  
153 another progressive negative shift to a minimum at  $\sim +300$  m which is followed by an unstable  
154 period through the remainder of the measured section (Wignall et al., this volume).

155

#### 156 **4.c. Molybdenum**

157 Similar to the carbon isotope record, trace element data has been shown as a proxy for anoxia for  
158 the 3 separate portions of the section previously studied (Grasby et al., 2015; Bond et al., 2015;  
159 Wignall et al., this volume). For reference the composite curve of Mo normalised to Al is shown  
160 through the entire interval (Fig. 3c). Results show relatively low values at the base of the section  
161 with a significant spike in Mo/Al at  $\sim -40$  m, which coincides with the shift to higher TOC  
162 values and minor negative shift in  $\delta^{13}\text{C}_{\text{org}}$ . Above the LPE boundary there is a second spike in  
163 Mo/Al that initiates  $\sim 5$  m after the main LPE event. This is followed by a gradual decline in  
164 Mo/Al until  $\sim 200$  m, coincident with the reversal in the trend of carbon isotopes. The Mo/Al  
165 values progressively increase again through the next 100 m until  $+300$  m, where peak Mo/Al  
166 ratios are coincident with the minimum in  $\delta^{13}\text{C}_{\text{org}}$  values. After this Mo/Al values become highly  
167 variable for the rest of the measured section.



168

169 **4.d. Mercury**

170 Hg values, reported here for the first time, are plotted in Figure 3d. Hg values are 0.005 to 0.010  
171  $\mu\text{g/g}$  in the basal 60 m of section and then show a shift above  $-40$  m to values up to  $0.050 \mu\text{g/g}$ ,  
172 followed by a subsequent decline to values of  $\sim 0.020 \mu\text{g/g}$  in the 10 m below just the LPE  
173 boundary. At the LPE there is a significant spike in Hg to the highest values observed in the  
174 section ( $> 0.130 \mu\text{g/g}$ ). Hg concentrations drop rapidly to background values of  $0.005$  to  $0.010$   
175  $\mu\text{g/g}$  through the remainder of the section. The only exception is a brief spike at  $\sim +300$  m, where  
176 Hg concentrations up to  $0.065 \mu\text{g/g}$  are observed.

177 In both marine and freshwater environments dissolved Hg has been shown to have strong  
178 affinity for organic matter (OM) (Gehrke et al., 2009; Mason et al., 1996; Han et al., 2006;  
179 Gagnon et al., 1997; Benoit et al., 2001). Grasby et al. (2013a) have also shown that OM  
180 strongly controls Hg sequestration over geologic time. Therefore, along with absolute  
181 concentrations, the Hg/TOC ratio is plotted in Figure 3e. However, TOC values are too low to be  
182 considered accurate for parts of the section, where Rockeval analyses cannot accurately resolve  
183 concentrations  $< 0.2\%$  TOC. Therefore only samples with TOC  $> 0.2\%$  are plotted as lower  
184 accuracy can greatly affect the calculated ratio. These data show a general trend whereby the  
185 Hg/TOC ratio has constant low values through the section. The exceptions to this are the large  
186 spikes in Hg/TOC values that occur at the LPE boundary as well as smaller shifts that occur at  $-$   
187  $40$  m as well as at  $\sim +300$  m. These increases in Hg/TOC ratio are consistent with zones where  
188 there are large spikes in absolute Hg concentration. Outside of these three levels Hg values are  
189 low.

190

## 191 **5. Discussion**

### 192 **5.a. Carbon isotope chemostratigraphy**

193 The  $\delta^{13}\text{C}$  record through Permian to Early Triassic sediments shows significant shifts that are  
194 comparable to those observed in the Sverdrup Basin (Grasby et al., 2013b) as well as with  
195 inorganic  $\delta^{13}\text{C}$  trends from records in the Panthalassa (Horacek et al., 2009) and the Tethys  
196 (Horacek et al., 2007; Payne et al., 2004). This demonstrates that Festningen records global  
197 variation in biogeochemical cycles through this time interval. These can be used as a  
198 chemostratigraphic tool to support both regional and global correlation.

199 Bond et al. (2015) argued that the minor negative carbon anomaly at – 40 m correlates with  
200 the Capitanian crises. The LPE event, as marked by the loss of chert forming siliceous sponges is  
201 characterised by the onset of a large negative shift in  $\delta^{13}\text{C}_{\text{org}}$  that reaches a minimum in the basal  
202 Vardebukta Formation (Grasby et al., 2015; Wignall et al., 1998). Above the Kapp Starostin  
203 Formation Wignall et al. (this volume) show that the next negative low point at ~300 m is  
204 correlative with the end Smithian Substage. This is also consistent with comparison to the  
205  $\delta^{13}\text{C}_{\text{org}}$  record from the Smithian stratotype (Grasby et al., 2013b), as illustrated in Figure 4. This  
206 makes the low in carbon isotope values equivalent to the Smithian/Spathian Extinction event that  
207 was associated with renewed Siberian Trap volcanism and rapid global temperature increase  
208 (Brayard et al., 2006; Orchard, 2007; Sun et al., 2012; Xie et al., 2010).

### 210 **5.b. Mercury deposition**

211 The Hg record at Festningen shows a relatively constant background value through the majority  
212 of the succession analysed ( $< 0.020 \mu\text{g/g}$ ). However notable spikes in Hg concentration occur  
213 coincident with the three extinction levels represented in this section (Capitanian Crisis, LPE,

214 and Smithian/Spathian Extinction). Of these three spikes the most significant occurs at the LPE  
215 boundary. Aside from the three most prominent Hg spikes there is also a slight increase in late  
216 Dienerian portion of the section. It is interesting to note that these prominent Hg spikes are all  
217 associated with shifts to lower  $\delta^{13}\text{C}$  values. In addition there is a general association of higher Hg  
218 concentrations associated with high Mo/Al values (Fig. 2). While this may suggest increased Hg  
219 sequestration associated with changes to more anoxic environments, a plot of Mo/Al versus Hg  
220 reveals that there is no correlation (Fig. 5). So despite the general relationship, the data suggest  
221 that anoxia has no direct influence on Hg sequestration in sediments.

222         Given the low organic matter content, throughout much of the Early Triassic portion of the  
223 studied section, reliable Hg/TOC ratios are not always possible to obtain. For values  $<0.2\%$  TOC  
224 inaccuracies in measurement can lead to magnified errors and highly variable Hg/TOC values  
225 that are not reflective of natural conditions, and therefore not shown. However, for TOC values  $>$   
226  $0.2\%$ , the Hg/TOC ratios for the Festningen section have relatively constant background values  
227 (vertical dashed line in Fig. 2e). These results are consistent with Hg/TOC through the Early  
228 Triassic at the Smithian stratotype, Sverdrup Basin (Grasby et al., 2013a), indicating a general  
229 background level of Hg sequestration into sediment that is largely controlled by organic matter  
230 deposition.

231         The notable exceptions to this background Hg deposition are spikes in both total Hg  
232 concentrations as well as Hg/TOC levels at the 3 extinction levels (Fig. 2). Previously Sanei et  
233 al. (2012) and (Grasby et al., 2013a) have argued that Hg anomalies in the geologic record are  
234 related to massive Hg emissions associated with periods of major volcanic eruptions. Our results  
235 from Festningen provide support for this hypothesis, showing over the  $\sim 12$  ma record that there  
236 are three prominent Hg anomalies superimposed on background Hg concentrations. In each case

237 these Hg spikes are associated with global extinction events that have been previously tied to LIP  
238 eruptions: Capitanian Crisis - Emeishan eruptions, LPE – Siberian Traps, Smithian/Spathian –  
239 renewed Siberian Traps (Bond and Wignall, 2014; Xie et al., 2010; Paton et al., 2010).

240 On a regional perspective the Hg spikes observed at Festningen closely correspond to those  
241 in the Sverdrup Basin (Fig. 4). This indicates that these periods of anomalous Hg deposition are  
242 regional in scope and suggest periods of enhanced Hg deposition over broad areas.

243

## 244 **6. Conclusions**

245 The organic carbon isotope records from the Festningen section show trends through the Late  
246 Permian and Early Triassic that closely correspond to those of the Sverdrup Basin, as well as the  
247 global inorganic carbon record. These results illustrate that NW Pangea records perturbations to  
248 the global carbon cycle. Previously it has been demonstrated that Hg is an excellent proxy for  
249 periods of major volcanic activity in the geologic record (Sanei et al., 2012; Grasby et al., 2013a;  
250 Sial et al., 2013). The Festningen section records a general constant background of Hg deposition  
251 through time. However, there are notable spikes in Hg concentration as well as in the Hg/TOC  
252 ratio that correspond to periods of mass extinction (Capitanian Crisis, the LPE event, and the  
253 Smithian/Spathian extinction), all of which have been associated with LIP events. Our results  
254 thus support the use of Hg as a marker for LIP eruptions. What remains uncertain is what role  
255 such Hg release could have on ecosystems. Certainly Hg is one of the most toxic elements for  
256 life, and enhanced Hg flux from large igneous events would likely have global impact in both the  
257 terrestrial and marine realm. Our results show that there are consistent records of Hg spikes  
258 associated with LIP events and mass extinctions across NW Pangea. The global nature of these  
259 records remains to be demonstrated. If they are more widely distributed then Hg release could  
260 have played a role as a significant extinction mechanism throughout Earth history.

261

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269

## 270 **Figure Legends**

271 Figure 1. Map showing location of the Festningen section on Spitsbergen (base of section is  
272 located at 78.0950°N, 13.8240°E (WGS84 datum).

273 Figure 2. Paleogeographic map after (Embry, 1992) showing relative location of Sverdrup Basin  
274 and Spitsbergen. Inset map (after Scotese) showing location on NW Pangea.

275 Figure 3. Plots of geochemical data from Festningen, showing: a) percent total organic carbon  
276 (TOC), b) carbon isotope values for organic carbon, c) Mo normalised by Al, d) Hg  
277 values, and e) Hg normalised by TOC (the Hg/TOC values are only shown for values  
278 of TOC > 0.2% as below that value Rockeval analyses provide less accurate results  
279 that are magnified in calculated ratios).

280 Figure 4. Correlation of carbon isotope and Hg records from the Smithian stratotype (Grasby et  
281 al., 2013b) with composite record from Festningen (this study).

282 Figure 5. Plot of the Mo/Al ratio versus Hg concentration.

283

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Table 1. Geochemical data from the Festningen section. Depths are measured relative to the Latest Permian Extinction Boundary (LPE) marked by the top of the Kapp Starostin Formation.

Sample ID	depth relative to LPE metres	$\delta^{13}\text{C}$ org ‰	Hg ppb	TOC %	Mo PPM	Al %
C-556979	-97.7	-24.5	9	0.02	0.45	0.54
C-556980	-97.2	-21.0	5	0.07	0.48	0.31
C-556981	-95.7	-22.2	4	0.05	0	0.59
C-556982	-94.2	n.d.	5	0.04	0	0.58
C-556983	-92.7	-22.4	6	0.11	0.09	1.02
C-556984	-91.2	-23.0	2	0.01	0.05	0.42
C-556985	-89.7	-23.0	2	0.02	0.07	0.3
C-556986	-86.7	-25.0	11	0.22	0.31	1.59
C-556987	-83.7	-24.0	7	0.11	0.2	1.15
C-556988	-80.7	-23.7	5	0.02	0.08	0.6
C-556989	-77.7	-25.0	6	0.03	0.22	0.61
C-556990	-74.7	-23.8	9	0.12	0.1	1.55
C-556991	-71.7	-24.3	6	0.02	0	0.53
C-556992	-68.7	-24.5	4	0.02	0	0.71
C-556993	-65.7	-25.5	4	0.05	0	0.64
C-556994	-62.7	-24.4	4	0.01	0	0.85
C-556995	-59.7	-24.8	4	0.03	0.08	0.59
C-556996	-56.7	-24.6	4	0.07	0.09	0.75
C-556997	-53.7	-25.2	4	0.07	0.21	1.21
C-556998	-50.7	-25.3	3	0.10	0	1.01
C-556999	-47.7	-24.5	4	0.01	0	0.7
C-557000	-46.2	-24.9	9	0.09	0.13	1
C-557001	-44.7	-24.5	6	0.05	0.13	1.28
C-557002	-43.2	-24.2	5	0.03	0.08	0.9
C-557003	-41.7	-23.6	6	0.08	0.06	0.45
C-557004	-40.2	-23.3	7	0.04	0.2	1.55
C-556572	-27.2	-24.3	28	0.25	0.53	2.65
C-556573	-26.7	-24.6	34	0.44	0.82	4.18
C-556574	-26.2	-24.6	36	0.39	0.99	3.85
C-556575	-25.7	-24.4	22	0.33	0.57	3.63
C-556576	-25.2	-24.7	22	0.29	2.66	2.84
C-556577	-24.7	-24.5	17	0.22	1.05	2.6
C-556578	-24.2	-24.4	35	0.34	1.83	3.3
C-556579	-23.7	-24.0	41	0.43	0	
C-556580	-23.2	-23.8	34	0.49	0.34	3.68
C-556581	-22.2	-23.74	48	0.65	0.25	4.36
C-556582	-21.7	-23.6	42	0.69	0.39	4.43

C-556583	-21.2	-23.63	38	0.59	0.26	4.21
C-556584	-20.7	-23.87	30	0.60	0.48	4.26
C-556585	-20.2	-23.62	37	0.58	0.46	4.54
C-556586	-19.7	-23.6	33	0.55	0.46	4.67
C-556587	-18.2	-23.86	28	0.51	0.8	4.19
C-556588	-17.7	-23.88	30	0.50	0.97	4.18
C-556589	-17.2	-23.9	33	0.57	1.26	3.97
C-556590	-16.7	-23.7	28	0.45	0.52	3.69
C-556591	-16.2	-24.0	29	0.49	0.7	4.2
C-556592	-15.7	-23.3	31	0.45	0.66	4.62
C-556593	-15.2	-24.0	30	0.53	0.74	4.08
C-556594	-14.7	-23.6	30	0.41	0.34	4.35
C-556595	-14.2	-24.1	32	0.44	0.33	4.07
C-556596	-13.7	-23.9	38	0.65	0.22	5.28
C-556597	-13.2	-23.7	34	0.50	0.31	4.7
C-556598	-12.7	-24.0	33	0.48	0.63	4.7
C-556599	-12.2	-24.3	37	0.56	0.57	5.4
C-556600	-11.7	-24.2	38	0.64	0.59	5.58
C-556601	-11.2	-25.2	14	0.44	0.44	3.85
C-556602	-10.7	-25.1	13	0.36	0.69	4.19
C-556603	-10.2	-24.9	13	0.35	1.13	3.6
C-556604	-9.7	-24.9	13	0.21	0.73	2.5
C-556605	-9.2	-25.5	17	0.35	0.83	4.2
C-556606	-8.7	-25.4	14	0.45	1.01	3.92
C-556607	-8.2	-24.7	16	0.35	0.68	3.69
C-556608	-7.7	-25.6	13	0.32	0.73	3
C-556609	-7.2	-25.0	15	0.37	0.48	3.32
C-556610	-6.7	-24.6	16	0.37	0.63	3.2
C-556611	-6.2	-24.5	13	0.40	0.48	3.42
C-556612	-5.7	-24.3	13	0.29	0.29	3.37
C-556613	-5.2	-24.3	16	0.29	0.45	3.11
C-556614	-4.7	-24.7	9	0.45	0.38	4
C-556615	-4.2	-25.2	14	0.52	0.76	3.22
C-556616	-3.7	-25.2	14	0.24	0.79	2.89
C-556617	-3.2	-25.7	17	0.24	0.84	2.67
C-556618	-3	-25.5	16	0.22	0.72	2.77
C-556619	-3.45	-25.1	15	0.19	2.92	1.66
C-556620	-2.8	-25.3	22	0.40	0.62	4.48
C-556621	-2.6	-25.7	19	0.39	0.82	4.04
C-556622	-2.4	-25.9	18	0.33	0.96	3.84
C-556623	-2.2	-26.1	24	0.39	0.92	3.83
C-556624	-2	-25.6	19	0.41	0.6	3.84
C-556625	-1.8	-25.7	15	0.34	0.8	3.57

C-556626	-1.6	-25.6	17	0.46	0.89	3.9
C-556627	-1.4	-24.8	18	0.35	0.36	4.17
C-556628	-1.2	-25.6	24	0.30	1.14	3.65
C-556629	-1	-25.5	15	0.22	1.31	3.5
C-556630	-0.8	-25.1	18	0.30	0.75	3.82
C-556631	-0.6	-24.5	20	0.29	0.41	3.43
C-556632	-0.4	-24.2	21	0.39	0.77	4.33
C-556633	-0.2	-25.3	16	0.23	0.57	3.42
C-556634	0	-25.5	20	0.33	1.37	3.33
C-556635	1	-29.0	29	0.28	0.48	7.98
C-556636	1.2	-29.6	43	0.24	1.07	8.21
C-556637	1.4	-29.1	48	0.19	1.34	8.69
C-556638	1.6	-28.8	33	0.13	1.21	8.45
C-556639	1.8	-28.6	27	0.14	0.28	8.25
C-556640	2	-28.7	26	0.11	0.52	8.52
C-556641	2.2	-28.8	16	0.10	0.14	8.57
C-556642	2.4	-28.9	20	0.09	0.16	8.57
C-556643	2.6	-30.1	30	0.13	0.46	9.98
C-556644	2.8	-30.4	17	0.10	0.14	7.85
C-556645	2.9	-30.3	25	0.08	0.46	9.5
C-556646	3.1	-30.8	25	0.12	0.17	7.87
C-556647	3.3	-31.0	27	0.22	0.16	8.2
C-556648	3.5	-32.8	41	0.40	0.55	7.96
C-556649	3.7	-32.6	43	0.45	0.96	8.09
C-556650	3.9	-32.5	42	0.43	0.57	8.21
C-556651	4.1	-33.5	77	0.31	4.4	9.75
C-556652	4.3	-32.8	60	0.48	1.59	7.87
C-556653	4.4	-32.6	74	0.27	0.63	8.05
C-556654	5.4	-32.5	62	0.50	1.43	7.99
C-556655	5.6	-32.2	53	0.33	1.23	8.19
C-556656	5.8	-33.3	86	0.42	2.39	9.91
C-556657	6	-32.6	69	0.44	0.76	9.66
C-556658	6.2	-33.2	69	0.31	0.57	9.77
C-556659	6.4	-33.6	60	0.39	1.04	8.47
C-556660	6.6	-33.3	53	0.28	0.48	10.3
C-556661	6.8	-33.2	46	0.26	0.35	10.97
C-556662	7	-32.8	53	0.34	0.44	11.28
C-556663	7.2	-32.6	51	0.34	0.52	10.92
C-556664	7.4	-33.0	50	0.31	0.39	10.76
C-556665	7.6	-32.9	81	0.44	2.76	10.02
C-556666	7.8	-32.9	77	0.56	5.03	8.99
C-556667	8	-32.3	82	0.71	3.91	9.56
C-556668	8.15	-32.7	89	0.51	5.2	7.92



C-556669	8.2	-33.7	134	1.03	20.65	8.74
C-556670	8.4	-33.0	88	0.63	10.16	7.12
C-556671	8.6	-33.5	100	0.61	26.37	7.02
C-556672	8.8	-33.6	57	0.38	4.51	9.62
C-556673	9	-33.8	57	0.41	3.67	9.2
C-556674	9.2	-32.1	25	0.08	1.58	6.99
C-556675	9.4	-33.3	50	0.46	2.77	9.55
C-556676	9.6	-32.8	38	0.25	0.59	10.06
C-556677	9.8	-33.1	26	0.30	0.96	9.54
C-556678	10	-32.9	47	0.23	0.62	10.87
C-556679	10.2	-33.2	35	0.30	1.02	10.63
C-556680	10.4	-33.0	30	0.29	1.92	10.99
C-556681	10.6	-32.8	40	0.49	1.98	9.53
C-556682	10.8	-31.6	19	0.14	1.42	6.76
C-556683	11	-31.8	23	0.17	1.4	9.02
C-556684	11.2	-27.4	74	1.02	1.41	9.1
C-556685	11.4	-31.8	34	0.17	2.14	8.21
C-556686	11.6	-32.1	23	0.21	0.83	9.45
C-556687	11.8	-32.2	31	0.25	1.51	9.31
C-556688	12	-32.6	48	0.39	3.08	9.09
C-556689	12.2	-32.6	56	0.55	2.9	9.23
C-556690	12.4	-31.3	19	0.09	1.28	6.7
C-556691	12.6	-33.1	34	0.29	2.55	9.55
C-556692	12.8	-32.7	35	0.15	3.04	6.67
C-556693	13	-33.1	45	0.33	7.18	7.38
C-556694	13.2	-32.8	47	0.29	4	7.1
C-556695	13.05	-33.9	41	0.48	12.34	9.75
C-556696	13.4	-32.7	34	0.30	4.13	7.19
C-556697	13.6	-33.6	42	0.50	4.96	7.1
C-556698	13.8	-33.3	43	0.48	3.89	7.31
C-556699	14	-32.4	35	0.17	2.55	7.44
C-556700	14.2	-33.4	27	0.31	3.45	4.54
C-556701	14.4	-33.5	24	0.26	3.5	4.57
C-556702	14.8	-32.8	33	0.45	2.31	7.7
C-556703	15.3	-32.4	15	0.18	1.71	7.25
C-556704	15.8	-31.2	10	0.10	0.37	7.99
C-556705	16.3	-31.8	7	0.28	1.53	7.82
C-556706	16.8	-32.4	1	0.35	1.27	7.98
C-556707	17.3	-32.4	1	0.19	3.08	8.2
C-556708	17.8	-25.1	0.43	0.00	3.34	7.79
C-556709	18.8	n.d.	0.49	0.01	3.01	7.72
C-556710	20.3	-31.9	0.36	0.19	3.51	7.98
C-556711	21.8	-32.6	4	0.16	1.32	8.08

C-556712	23.3	-32.5	18	0.64	3.42	7.49
C-556713	24.8	-33.0	20	0.42	2.66	7.43
C-556714	26.3	-32.7	22	0.50	2.96	7.07
C-556715	27.8	-32.7	26	0.54	2.9	7.03
C-556716	29.3	-32.8	25	0.40	1.21	7.42
C-556717	30.8	-32.3	24	0.32	1.36	7.1
C-556718	32.3	-32.8	31	0.73	3.33	7.09
C-556719	33.8	-32.1	34	0.57	2.31	7.16
C-556720	35.3	-31.3	23	0.35	1.04	7.42
C-556721	36.8	-31.8	18	0.49	3.17	7.3
C-556722	38.3	-31.7	9	0.23	0.4	6.65
C-556723	39.8	-31.3	9	0.26	0.8	7.43
C-556724	41.3	-31.3	16	0.30	0.42	7.29
C-556725	44.3	-31.5	22	0.49	0.75	7.38
C-556726	47.3	-31.4	29	0.62	0.71	6.84
C-556727	50.3	-31.1	27	0.45	0.36	7.16
C-556728	53.3	-29.2	9	0.05	0.55	4.44
C-556729	56.3	-29.9	16	0.12	0.42	6.64
C-556730	59.3	-29.7	11	0.15	0.33	6.94
C-556731	62.3	-28.7	10	0.15	0.17	7.56
C-556732	65.3	-29.2	8	0.12	0.29	6.92
C-556733	68.3	-29.8	5	0.07	0.28	6.51
C-556734	71.3	-29.6	8	0.09	0.59	6.67
C-556735	74.3	-29.6	8	0.10	0.46	6.01
C-556736	77.3	-29.5	5	0.06	0.31	6.88
C-556737	80.3	-29.4	7	0.07	0.29	6.66
C-556738	83.3	-29.2	7	0.06	0.26	6.43
C-556739	86.3	-28.6	4	0.04	0.27	6.58
C-556740	89.3	-29.1	14	0.11	0.64	6.84
C-556741	92.3	-28.5	5	0.05	0.19	6.32
C-556742	95.3	-27.9	8	0.13	0.23	6.24
C-556743	98.3	-28.5	5	0.05	0.28	6.36
C-556744	101.3	-28.0	4	0.11	0.24	5.53
C-556745	104.3	-28.3	8	0.05	0.25	5.57
C-556746	107.3	-27.6	3	0.03	0.16	6.66
C-556747	110.3	-28.3	5	0.12	0.25	6.04
C-556748	113.3	-27.6	4	0.02	0.18	5.84
C-556749	116.3	-27.8	12	0.03	0.45	4.94
C-556750	117.8	-28.7	16	0.15	0.85	6.42
C-556751	119.3	-28.3	22	0.16	0.45	6.54
C-556752	122.3	-28.9	16	0.07	0.51	4.1
C-556753	125.3	-27.6	14	0.06	0.3	5.12
C-556754	128.3	-26.7	9	0.04	0.3	5.95

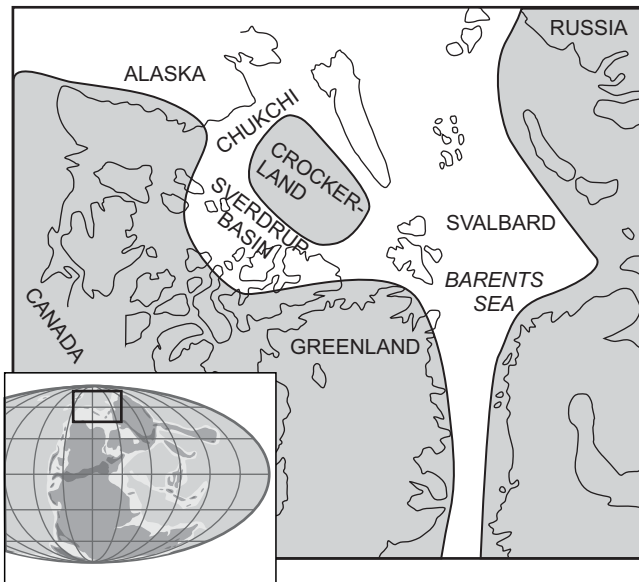
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C-556759	146.3	-27.2	21	0.09	0.27	5.97
C-556760	149.3	-26.7	10	0.08	0.29	6.92
C-556761	152.3	-26.7	16	0.10	0.19	6.55
C-556762	153.3	-26.0	7	0.03	0.25	6.03
C-556763	154.8	-26.0	22	0.11	0.38	6.86
C-556764	157.8	-26.9	10	0.07	0.18	6.32
C-556765	160.8	-27.4	17	0.12	0.27	7.81
C-556766	163.8	-26.8	7	0.02	0.41	6.26
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C-556769	182.3	-26.5	4	0.04	0.19	5.03
C-556770	196.8	-24.7	7	0.06	0.18	5.23
C-556771	199.8	-25.7	7	0.03	0.14	6.66
C-556772	202.8	-25.7	4	0.02	0.13	6.27
C-556773	205.8	-26.0	4	0.01	0.18	4.37
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C-556775	214.8	-26.0	3	0.01	0.08	5.68
C-556776	217.8	-26.6	5	0.02	0.14	6.88
C-556777	220.8	-27.1	7	0.04	0.11	6.86
C-556778	223.8	-27.4	5	0.02	0.15	6.06
C-556779	226.8	-26.8	5	0.09	0.16	5.99
C-556780	229.8	-27.1	7	0.02	0.18	5.1
C-556781	232.8	-28.7	10	0.05	0.14	5.21
C-556782	235.8	-27.4	10	0.03	0.41	4.33
C-556783	238.8	-27.4	6	0.03	0.18	6.22
C-556784	241.8	-27.7	8	0.06	0.24	6.65
C-556785	244.8	-28.1	11	0.06	0.29	5.71
C-556786	247.8	-29.2	8	0.06	0.2	5.61
C-556787	250.8	-29.4	7	0.06	0.2	6.31
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C-556789	256.8	-31.1	11	0.15	0.32	6.17
C-556790	260.8	-29.7	12	0.09	0.31	6.36
C-556791	263.8	-29.4	9	0.06	0.44	5.57
C-556792	266.8	-31.4	14	0.13	0.3	6.08
C-556793	269.8	-31.6	13	0.19	0.43	5.85
C-556794	272.8	-31.5	15	0.16	0.39	6.14
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C-556797	281.8	-30.7	5	0.05	0.18	5.13

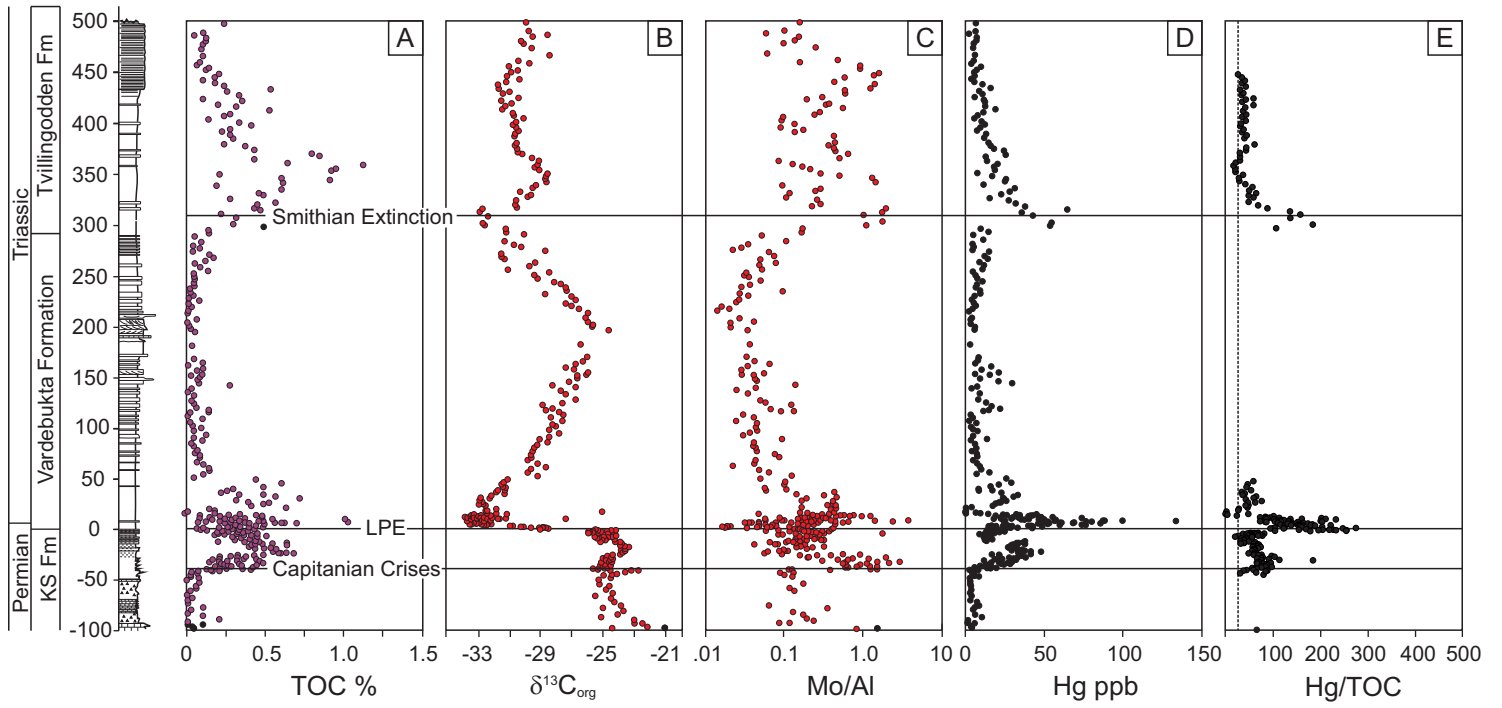
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C-556802	299.8	-32.5	54	0.50	7.52	6.65
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C-556807	318.8	-30.6	38	0.43	1.46	6.31
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C-556811	330.8	-29.6	23	0.49	0.73	6.23
C-556812	333.8	-30.3	30	0.46	1.55	6.23
C-556813	336.8	-29.6	32	0.61	1.88	6.43
C-556814	339.8	-29.3	11	0.20	0.48	5.07
C-556815	342.8	-28.6	26	0.62	7.48	5.08
C-556816	345.8	-28.7	26	0.91	8.42	6.32
C-556817	348.8	-28.6	18	0.62	3.14	6.27
C-556818	351.8	-28.6	8	0.22	1.44	4.98
C-556819	354.8	-29.1	19	0.93	1.33	6.25
C-556820	357.8	-29.4	19	0.95	0.99	5.79
C-556821	360.8	-29.3	21	1.13	0.83	6.51
C-556822	363.8	-29.1	14	0.66	0.52	6.06
C-556823	366.8	-29.6	13	0.43	2.08	4.21
C-556824	369.8	-30.2	26	0.86	4.3	6.68
C-556825	372.8	-30.5	24	0.80	3.07	6.7
C-556826	375.8	-30.6	19	0.43	2.47	5.69
C-556827	378.8	-30.7	17	0.38	2.11	5.7
C-556828	381.8	-30.5	15	0.25	2.61	5.96
C-556829	387.8	-30.7	12	0.30	2.45	5.74
C-556830	390.8	-30.6	13	0.28	0.8	5.75
C-556831	393.8	-30.4	8	0.23	0.76	4.31
C-556832	396.8	-30.6	10	0.28	0.44	4.86
C-556833	399.8	-30.8	13	0.42	0.79	5.86
C-556834	402.8	-30.5	11	0.34	0.54	5.84
C-556835	405.8	-30.1	6	0.15	0.36	3.72
C-556836	408.8	-30.8	8	0.25	1.1	4.54
C-556837	411.8	-30.8	12	0.28	1.41	4.96
C-556838	414.8	-31.5	20	0.53	3.72	6.7
C-556839	417.8	-31.3	13	0.33	1.99	6
C-556840	420.8	-30.9	13	0.21	1.96	5.37

C-556841	423.8	-31.5	12	0.37	1.17	5.94
C-556842	426.8	-30.4	6	0.11	1.2	3.86
C-556843	429.8	-31.4	11	0.34	3.49	5.8
C-556844	432.8	-31.0	9	0.26	2.28	3.79
C-556845	435.8	-31.7	17	0.54	6.92	5.49
C-556846	438.8	-31.8	11	0.25	6.32	4.44
C-556847	441.8	-31.2	6	0.19	1.7	3.85
C-556848	444.8	-30.3	4	0.11	1.99	3.37
C-556849	447.8	-31.2	6	0.20	5.56	4.08
C-556850	450.8	-30.9	6	0.22	6.11	3.81
C-556851	453.8	-30.5	8	0.13	2.78	2.99
C-556852	456.8	-31.1	10	0.16	3.47	3.68
C-556853	459.8	-29.7	4	0.07	0.66	4.23
C-556854	462.8	-30.4	6	0.10	1.8	3.73
C-556855	468.3	-28.4	7	0.11	0.25	4.19
C-556856	475.3	-29.6	5	0.10	1.23	4.87
C-556857	479.8	-30.2	7	0.11	0.45	4.62
C-556858	482.8	-30.3	7	0.14	0.55	4.04
C-556859	485.8	-29.5	8	0.14	0.75	4.93
C-556860	488.8	-28.5	3	0.05	0.22	3.78
C-556861	491.8	-29.8	7	0.12	0.42	4.09
C-556862	499.8	-30.0	7	0.24	0.63	3.98



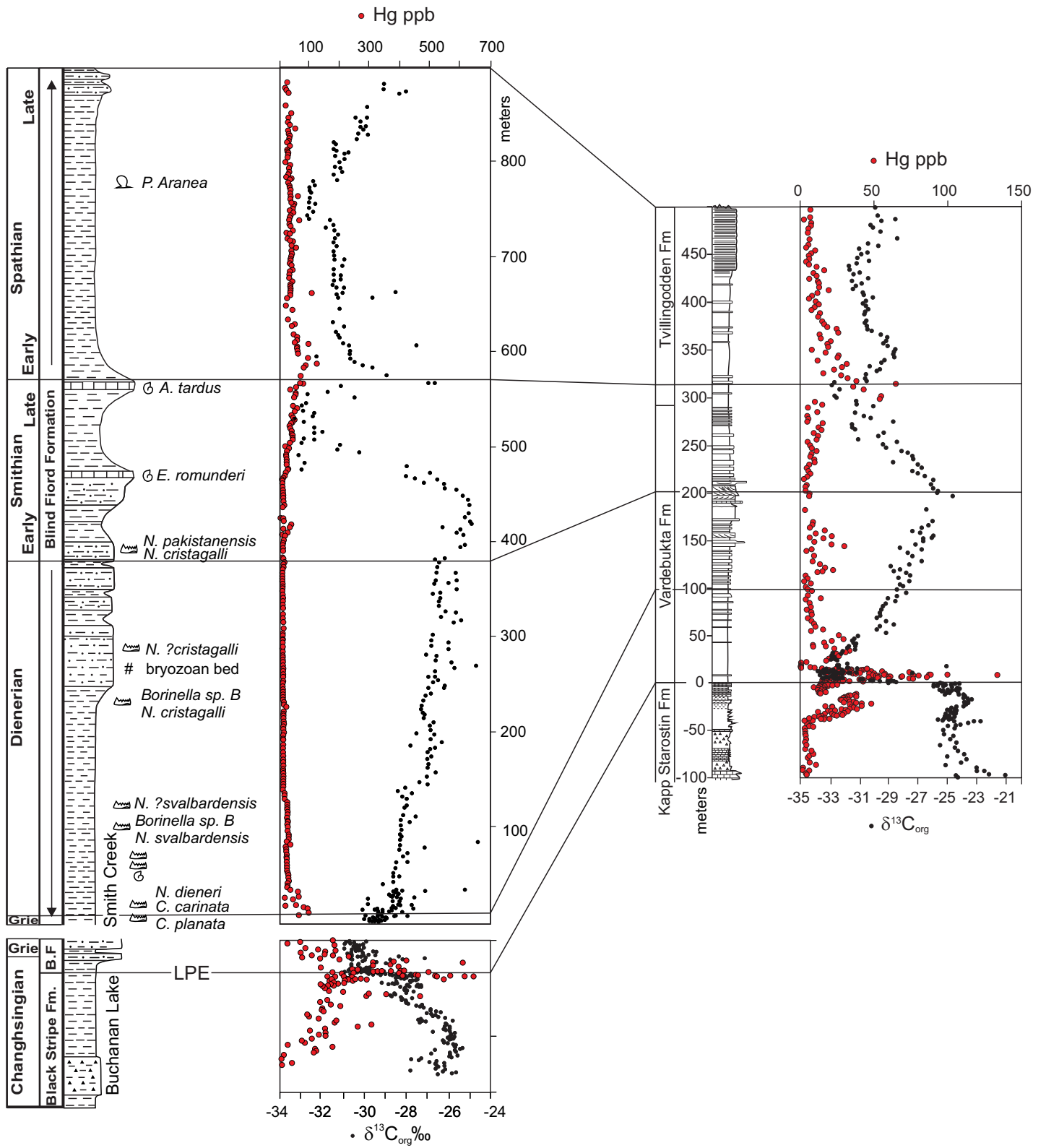
Grasby et al. Figure 1



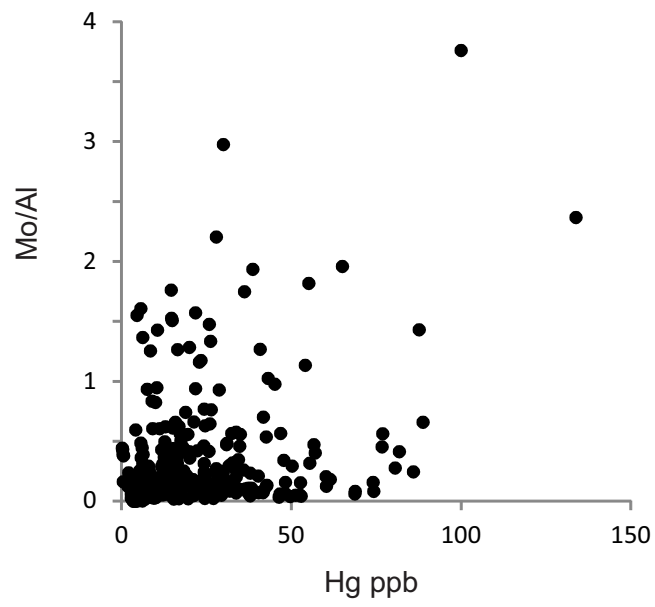


Grasby et al. Figure 3





Grasby et al. Figure 4



Grasby et al. Figure 5