

The radiation impedance of orbiting conductors

J. R. Sanmartín

Escuela Técnica Superior de Ingenieros Aeronáuticos, Universidad Politécnica, Madrid

M. Martínez-Sánchez

Department of Aeronautics and Astronautics, Massachusetts Institute of Technology, Cambridge

Abstract. The dispersion relation for waves in a cold, magnetized plasma is discussed using the potential for the longitudinal part of the electric field. This clarifies wave emission from a conductor in low Earth orbit and should be useful in considering the far field and both hot plasma and nonlinear, near-field effects. General formulas for radiation impedance are directly obtained. For tethers a fundamental dependence on contactor size is discussed. Spherical and ellipsoidal contactors and an (anodeless) bare tether are considered. Simple arguments on nonlinear contactor effects lead to a surprisingly simple result for impedances off the Alfvén branch.

1. Introduction

In a seminal paper, *Drell et al.* [1965] showed how an orbiting conductor, if in electrical contact with the ionosphere, would excite Alfvén waves. Since the emf induced by the geomagnetic field is proportional to the perpendicular conductor length, using a long space tether, with contactor ends, would enhance radiation [*Banks et al.*, 1981]. A quantity of interest for both power generation and Alfvén (and higher frequency) signal propagation is the wave impedance [*Rasmussen et al.*, 1985; *Dobrowolny and Veltri*, 1986]. Fundamental results were obtained by *Barnett and Olbert* [1986] and *Estes* [1988].

Further work on the impedance was carried out by *Donohue et al.* [1991] and by *Hastings and Wang* [1987] and *Hastings et al.* [1988], who extended the analysis to conductors carrying ac currents. *Estes* [1988] and *vom Steirn and Neubauer* [1992] studied the field near a tether, *Rasmussen et al.* [1990] and *McKenzie* [1991] studied the far field, and *Hastings and Wang* [1989] and *vom Steirn and Neubauer* [1992] studied two-ion effects. Tether radiation has been modeled in the laboratory [*Urrutia and Stenzel*, 1989; *Stenzel and Urrutia*, 1990].

Here, contrary to published analyses, we do not directly study wave emission. Instead, we first collect convenient results from the well-known dispersion relation for a cold, magnetized plasma and then solve the wave equation in terms of the potential ϕ for the longitudinal part of the electric field of the wave. Only then are conditions particular to orbiting conductors considered; this corrects or clarifies results found in the literature.

The ϕ formalism, a powerful tool for studying nonlinear or hot-plasma effects, is used here to directly derive complete impedance formulas (sections 2 and 3).

For tethers the impedance for branches other than Alfvén is shown to depend heavily on the model for the cathodic and anodic contactors at the ends. Taking into account a nonvanishing contactor length along the tether proves essential for a proper model and yields an impedance varying as the inverse of contactor area. Simple results are obtained for spherical (and ellipsoidal) contactors and for an anodeless bare tether, proposed as an efficient anodic contactor [*Sanmartín et al.*, 1993]. Nonlinear contactor effects suggest an impedance varying as the inverse of current (sections 4 and 5).

2. Wave Emission

We first recall three well-known results from the usual wave equation for the Fourier-transformed electric field $\mathbf{E}(\mathbf{k}, \omega)$ in a two-component, cold plasma [*Akhiezer et al.*, 1975],

$$-\frac{\mathbf{k} \wedge (\mathbf{k} \wedge \mathbf{E})}{k^2} - \frac{\epsilon_c \cdot \mathbf{E}}{n^2} = \frac{4\pi \mathbf{j}_s}{\omega n^2}. \quad (1)$$

Here \mathbf{j}_s is the source current density, \mathbf{k} is the wave vector, $n \equiv ck/\omega$ is the refractive index, and $\epsilon_c(\omega)$ is the dielectric tensor (with z axis along the field \mathbf{B}_0),

$$\epsilon_c = \begin{pmatrix} \epsilon_1 & i\epsilon_2 & 0 \\ -i\epsilon_2 & \epsilon_1 & 0 \\ 0 & 0 & \epsilon_3 \end{pmatrix}, \quad \epsilon_3 \simeq 1 - \frac{\omega_{pe}^2}{\omega^2},$$

$$\epsilon_1 = \frac{\omega^2 - \omega_{UH}^2}{\omega^2 - \Omega_e^2} \frac{\omega^2 - \omega_{LH}^2}{\omega^2 - \Omega_i^2}, \quad \epsilon_2 \simeq \frac{\omega_{pe}^2}{\omega^2 - \Omega_e^2} \frac{\Omega_e \omega}{\omega^2 - \Omega_i^2},$$

with $\Omega_e > 0$, $\omega_{UH}^2 \simeq \omega_{pe}^2 + \Omega_e^2$, $\omega_{LH}^2 \simeq \Omega_e^2(\Omega_e^2 + \omega_{pi}^2)/(\Omega_e^2 + \omega_{pe}^2)$. Here ω_{pe} , ω_{pi} and Ω_e , Ω_i are plasma

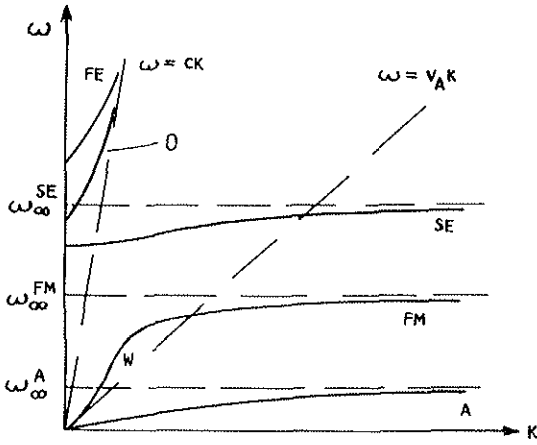


Figure 1. Schematics of branches (fast extraordinary (FE), ordinary (O), slow extraordinary (SE), fast magnetosonic (FM), and Alfvén (A)) for the dispersion relation of a cold, magnetized plasma; W is the whistler region in branch FM. The figure is dependent on the angle between \mathbf{k} and the ambient field.

frequencies and gyrofrequencies for electrons and ions; ω_{UH} , ω_{LH} are upper and lower hybrid frequencies. (1) Equation (1) yields the Astrom dispersion relation, here written as

$$\left(\epsilon_1 \sin^2 \theta + \epsilon_3 \cos^2 \theta - \frac{c_3 c_1}{n^2} \right) \left(1 - \frac{c_1}{n^2} \right) + \left(\sin^2 \theta - \frac{c_3}{n^2} \right) \frac{c_2^2}{n^2} \equiv D(k, \theta, \omega) = 0, \quad (2)$$

θ being the angle between \mathbf{k} and \mathbf{B}_0 . There are two branches for $n^2(\theta, \omega)$ but there are five branches for $\omega^2(k, \theta)$, shown schematically in Figure 1, as follows: fast extraordinary (FE), ordinary (O), slow extraordinary (SE), fast magnetosonic or compressional Alfvén (FM), and Alfvén or shear Alfvén (A). (2) FE and O branches have $n < 1$. (3) SE, FM, and A branches have asymptotes (resonances) $\omega_{\infty}^{SE}(\theta)$, $\omega_{\infty}^{FM}(\theta)$, and $\omega_{\infty}^A(\theta)$, obtained from (2) in the limit $k \rightarrow \infty$ (that is, $n \rightarrow \infty$ or $|\epsilon_j|/n^2 \rightarrow 0$, $j = 1-3$), yielding

$$c_1 \sin^2 \theta + c_3 \cos^2 \theta \equiv D(\infty, \theta, \omega) = 0. \quad (2')$$

We add here two more points. (1) On the A branch the limit $n \rightarrow \infty$ may be obtained from condition $\omega \rightarrow 0$, instead of condition $k \rightarrow \infty$. Then, one would also have $|\epsilon_3| \rightarrow \infty$, and the ratio c_3/n^2 could take any value. If c_1/n^2 , ϵ_2/n^2 are small but c_3/n^2 is large, (2) becomes

$$\epsilon_3 \cos^2 \theta - \frac{c_3 c_1}{n^2} \equiv D_A(k, \theta, \omega) \simeq 0, \quad \cos^2 \theta \ll 1. \quad (2'')$$

(2) Introducing longitudinal l and transverse t parts of the field,

$$\mathbf{E}_l \equiv \mathbf{k} \mathbf{k} \cdot \mathbf{E} / k^2, \quad \mathbf{E}_t = \mathbf{E} - \mathbf{E}_l,$$

(1) becomes, in all generality,

$$\mathbf{E}_t - \frac{c_c}{n^2} \cdot (\mathbf{E}_t + \mathbf{E}_l) = \frac{4\pi i \mathbf{j}_s}{\omega n^2}, \quad (1')$$

yielding the \mathbf{E}_t components along z and perpendicular \perp to \mathbf{B}_0 ,

$$E_{tz} = \frac{c_3 E_{lz}}{n^2 - c_3} + \frac{4\pi i j_{sz}}{\omega (n^2 - c_3)}, \quad (3)$$

$$\mathbf{E}_{t\perp} = \frac{H^{-1} \cdot \mathbf{c}_{\perp} \cdot \mathbf{E}_{l\perp}}{n^2} + \frac{4\pi i}{\omega n^2} H^{-1} \cdot \mathbf{j}_{s\perp},$$

$$\mathbf{c}_{\perp} = \begin{pmatrix} c_1 & i c_2 \\ -i c_2 & c_1 \end{pmatrix}, \quad H = I - \frac{\mathbf{c}_{\perp}}{n^2}$$

with I the two-dimensional unit tensor. Use of

$$\mathbf{k} \cdot \mathbf{E}_t = 0, \quad \mathbf{E}_t \equiv -i \mathbf{k} \phi$$

gives an explicit scalar equation for ϕ ,

$$k^2 D(k, \theta, \omega) \phi = \left[\left(1 - \frac{c_1}{n^2} \right)^2 - \frac{c_2^2}{n^4} \right] \frac{4\pi}{\omega} \times \left[k_z j_{sz} + \left(1 - \frac{c_3}{n^2} \right) \mathbf{k}_{\perp} \cdot H^{-1} \cdot \mathbf{j}_{s\perp} \right]. \quad (4)$$

Equation (4) then becomes

$$k^2 D(\infty, \theta, \omega) \phi = \frac{4\pi}{\omega} \mathbf{k} \cdot \mathbf{j}_s, \quad (4')$$

for conditions leading to (2'), and

$$k^2 D_A(k, \theta, \omega) \phi = -\frac{4\pi}{\omega} \frac{c_3}{n^2} \mathbf{k}_{\perp} \cdot \mathbf{j}_{s\perp} \simeq \frac{\omega_{pe}^2}{c^2 k_{\perp}^2} \frac{4\pi}{\omega} \mathbf{k} \cdot \mathbf{j}_s \quad (4'')$$

for conditions leading to (2''); we used $\cos^2 \theta \ll 1$ to write $\mathbf{k}_{\perp} \cdot \mathbf{j}_{s\perp} \simeq \mathbf{k} \cdot \mathbf{j}_s$.

We are now in a position to discuss wave emission from an orbiting conductor sustaining a steady current. In the terrestrial reference frame the Doppler relation reads

$$\omega = k_r V; \quad (5)$$

as usual, we take \mathbf{B}_0 and orbital velocity \mathbf{V} horizontal and perpendicular to each other, with x axis along \mathbf{V} . In the ionosphere, the Alfvén velocity $V_A \equiv c \Omega_e / \omega_{pe}$ satisfies the conditions [Barnett and Olbert, 1986] (hereinafter referred to as BO)

$$V \ll V_A \ll c, \quad (6a)$$

$$\Omega_e^2 \ll \omega_{pe}^2 \quad \text{or} \quad V_A^2 \ll c^2 m_e / m_i, \quad (6b)$$

($V \simeq 7$ km/s, $V_A \sim 300$ km/s, $m_i/m_e \simeq 30,000$). It also satisfies the condition

$$V_A^2 \ll V^2 m_i / m_e; \quad (6c)$$

m_i/m_e is the ion-electron mass ratio. The lower-hybrid frequency is now $\omega_{LH} \simeq (\Omega_e \Omega_i)^{1/2}$.

Note from (5) and (6a) that n is here very large,

$$1/n \equiv \omega/c k = k_x V/kc \ll 1. \quad (7)$$

This means, first, that no FE or O emission is possible. It suggests, secondly, that (2') might apply to the SE and FM branches, which is indeed the case. For SE, use of $\omega_{\infty}^{SE}(\theta)$ [Akhiezer et al., 1975],

$$\omega_{pe} \leq \omega_{\infty}^{SE} \simeq (\omega_{pe}^2 + \Omega_e^2 \sin^2 \theta)^{1/2} \leq \omega_{UH}, \quad (8a)$$

gives $\epsilon_2 \simeq \Omega_e/\omega_{pe}$, $\epsilon_1 \simeq -c_2^2 \cos^2 \theta$, $\epsilon_3 \simeq \epsilon_2^2 \sin^2 \theta$; condition (7) (and (6b)) then shows all ϵ_j/n^2 to be small. Alternatively, one can find $(\omega_{\infty}^{SE} - \omega)/\omega \ll 1$, directly from (2), for $\Omega_e/c k$ small as in the present case ($\Omega_e/c k \simeq \epsilon_2/n$). For FM, use of $\omega_{\infty}^{FM}(\theta)$ [Akhiezer et al., 1975],

$$\omega_{LH} \leq \omega_{\infty}^{FM} \simeq (\Omega_e^2 \cos^2 \theta + \omega_{LH}^2 \sin^2 \theta)^{1/2} \leq \Omega_e, \quad (8b)$$

similarly proves that $|\epsilon_j|/n^2$ ($j = 1-3$) and $(\omega_{\infty}^{FM} - \omega)/\omega$ are small for $\omega_{LH}/V_A k_{\perp} \ll 1$. Here $\omega_{LH}/V_A k_{\perp} \equiv \omega_{LH} V k_x/\omega V_A k_{\perp}$ is indeed small because of condition $V/V_A \ll 1$.

The large refractive index finally suggests that (2') or (2''), or an intermediate dispersion relation, might apply to the A branch. Here we have

$$\frac{|\epsilon_3|}{n^2} \simeq \frac{\omega_{pe}^2}{c^2 k^2} = \frac{\Omega_e^2 k_x^2 m_i V^2}{\omega^2 k^2 m_e V_A^2}.$$

Since this branch always has $\omega \leq \Omega_i$, condition (6c) makes $|\epsilon_j|/n^2$ large and (2'') valid, except for k_y/k_x large enough. The equation $D_A = 0$ has the root

$$\omega_A(k_z) \equiv \Omega_i V_A k_z / (\Omega_i^2 + V_A^2 k_z^2)^{1/2}. \quad (9)$$

Note that $\cos \theta \equiv k_z/k = (k_x V/k V_A) V_A k_z/\omega_A$ is small as required, unless ω is very close to Ω_i . We now make the ansatz that negligible energy is carried by such frequencies and by large k_y/k_x values, which may thus be ignored. Note also that $c_1/n^2 (= \cos^2 \theta)$ and $c_2/n^2 (\simeq \epsilon_1 \omega/n^2 \Omega_i)$ are small as assumed.

For both the SE and FM modes, (2'), (3), and (4') give

$$\phi = \frac{4\pi \mathbf{k} \cdot \mathbf{j}_s}{\omega (\epsilon_1 k_{\perp}^2 + \epsilon_3 k_z^2)}, \quad \mathbf{E} \simeq \mathbf{E}_l = -i \mathbf{k} \phi, \quad (10a)$$

with $\mathbf{E}_t = 0(n^{-2})$. For the Alfvén mode, (2''), (3), and (4'') give

$$\phi = \frac{4\pi \mathbf{k} \cdot \mathbf{j}_s}{\omega (\epsilon_1 k_{\perp}^2 - n^2 k_z^2)}, \quad \mathbf{E} \simeq \mathbf{E}_{l\perp} = -i \mathbf{k}_{\perp} \phi, \quad (10b)$$

with $\mathbf{E}_{tz} \simeq -\mathbf{E}_{tz}$, $\mathbf{E}_{t\perp} = 0(n^{-2})$. Since ω_{∞} is a function of just θ , or k_z/k , SE and FM waves have a group velocity perpendicular to \mathbf{k} in the \mathbf{B}_0, \mathbf{k} plane, a fact not entirely clear in the literature [BO; Hastings and Wang, 1987]. Since ω_A depends only on k_z , A waves have group velocity along \mathbf{B}_0 , with \mathbf{k} nearly perpendicular. Note that certain results in BO's analysis (figures 2b and 2c, for bands II and III, of that paper) represent just the well-known formulas (8a) and (8b) (figure 2a, for band I, of BO's paper represents $\omega_A(k_x/k_z)$ from (9), with ω_A written as $V k_x$ where necessary).

Note also that there is no whistler emission (part W on the FM branch of Figure 1); Stenzel and Urrutia's experiments showing whistlers do not apply to an ionospheric tether; they either fail to reproduce the steady condition $\omega = k_x V$ [Urrutia and Stenzel, 1989] or correspond to the opposite regime $V \simeq 2 \times 10^7 \text{ cm/s} \gg V_A \simeq 4 \times 10^5 \text{ cm/s}$ [Stenzel and Urrutia, 1990]. Radiation occurs at the contactors, where $\nabla \cdot \mathbf{j}_s \neq 0$, because it depends on the source divergence $\mathbf{k} \cdot \mathbf{j}_s$, a fact first noticed by Estes [1988] and clearly arising from the quasi-electrostatic character of the field, $\mathbf{E} \simeq -i \mathbf{k} \phi$ or $-i \mathbf{k}_{\perp} \phi$. For FM waves this followed from condition $V \ll V_A$; FM emission along a tether, as found by Stenzel and Urrutia [1990], corresponds to the opposite, whistler regime, $V \gg V_A$. Note, however, that condition $V \ll V_A$, being dependent on planetary parameters (surface gravity, radius, ionospheric density, and ambient field \mathbf{B}_0), is not an intrinsic property of tethers; tether radiation might thus depend, in general, on current source features other than its divergence [Donohue et al., 1991].

In the following sections we shall use the present ϕ formalism to determine radiation impedance formulas. The formalism, however, should also be useful in studying (1) how to match linear results for the field to non-linear results near the tether, where \mathbf{E} should be nearly electrostatic; (2) the radiation pattern in the transverse far field of the waves, given clearly as

$$\mathbf{E}_t \simeq -i \frac{\epsilon_c \cdot \mathbf{k}}{n^2} \phi, \quad (\text{SE, FM}) \quad \text{or} \quad \mathbf{E}_t \simeq i k_z \phi \mathbf{1}_z; \quad (\text{A})$$

(3) hot plasma effects [(10a) would still hold, with the hot dielectric tensor ϵ_h replacing ϵ_c in the expression $\epsilon_1 k_{\perp}^2 + \epsilon_3 k_z^2 \equiv \mathbf{k} \cdot \epsilon_c \cdot \mathbf{k}$].

3. The Impedance Formulas

To determine the impedance, there is no need to compute the near field [Estes, 1988] or the far field [Barnett and Olbert, 1986; Hastings et al., 1988]. For SE or FM waves we have $\mathbf{E}(\mathbf{r}, t) \simeq -\nabla \phi(\mathbf{r}, t)$ in space-time coordinates \mathbf{r}, t . Since \mathbf{j}_s vanishes outside certain volume, we find a radiated power

$$\begin{aligned}
 P_{rad} &= - \int \mathbf{j}_s \cdot \mathbf{E} d\mathbf{r} = - \int \phi \nabla \cdot \mathbf{j}_s d\mathbf{r} \\
 &= - \int d\mathbf{r} \int \frac{d\mathbf{k}_1 d\omega_1}{4\pi^2} e^{i\mathbf{k}_1 \cdot \mathbf{r} - i\omega_1 t} \phi(\mathbf{k}_1, \omega_1) \\
 &\quad \times \int \frac{d\mathbf{k} d\omega}{4\pi^2} e^{i\mathbf{k} \cdot \mathbf{r} - i\omega t} i\mathbf{k} \cdot \mathbf{j}_s(\mathbf{k}, \omega)
 \end{aligned} \tag{11}$$

with $\phi(\mathbf{k}, \omega)$ given by (10a). For A waves we have $\mathbf{E} \simeq -\nabla_{\perp} \phi$ and, again setting $\mathbf{k}_{\perp} \cdot \mathbf{j}_{s\perp} \simeq \mathbf{k} \cdot \mathbf{j}_s$, we arrive back at (11) with $\phi(\mathbf{k}, \omega)$ given by (10b). The \mathbf{r} integral is immediate, $\int d\mathbf{r} \exp[i(\mathbf{k}_1 + \mathbf{k}) \cdot \mathbf{r}] = 8\pi^3 \delta(\mathbf{k}_1 + \mathbf{k})$. This makes the \mathbf{k}_1 integral immediate too.

The source divergence can be written as

$$\begin{aligned}
 i\mathbf{k} \cdot \mathbf{j}_s &= \int \frac{d\mathbf{r} dt}{4\pi^2} e^{-i\mathbf{k} \cdot \mathbf{r} + i\omega t} \nabla \cdot \mathbf{j}_s(x - Vt, y, z) \\
 &\equiv iI_s \delta(\omega - k_x V) g(\mathbf{k}).
 \end{aligned}$$

We have set $x - Vt \rightarrow x$ and introduced the current I_s in the conductor and a dimensionless function $g(\mathbf{k})$ defined by

$$g(\mathbf{k}) \equiv -i \int d\mathbf{r} \nabla \cdot \mathbf{j}_s(\mathbf{r}) \exp(-i\mathbf{k} \cdot \mathbf{r}) / 2\pi I_s. \tag{12}$$

The ω_1 and ω integrals in (11) are now straightforward, yielding the impedance, $Z \equiv \text{Power}/I_s^2$, as

$$\begin{aligned}
 Z &= \int \frac{2i |g(\mathbf{k})|^2 d\mathbf{k}}{\omega k_{\perp}^2} [\epsilon_1(\omega) \\
 &\quad + k_z^2 \left\{ \frac{-c^2/\omega^2}{\epsilon_3(\omega)/k_{\perp}^2} \right\}^{-1}, \left\{ \begin{matrix} \text{A} \\ \text{SE, FM} \end{matrix} \right\}
 \end{aligned} \tag{13}$$

with $\omega = k_x V$; we used the relation $g(-\mathbf{k}) = -g^*(\mathbf{k})$.

Alfvén Mode

For the A branch we have $\epsilon_1 \simeq \omega_{pi}^2 / (\Omega_i - \omega^2)$ and

$$\begin{aligned}
 \epsilon_1 - \frac{c^2 k_z^2}{\omega^2} &= \frac{c^2 \Omega_i^2 + V_A^2 k_z^2}{V_A^2 \Omega_i^2 - V^2 k_x^2} \\
 \times \frac{(k_x + i\Delta)^2 - [\omega_A(k_z)/V]^2}{k_x^2}, \quad \Delta \rightarrow 0^+,
 \end{aligned}$$

where we set $\omega \rightarrow \omega + i\nu$ ($\nu \rightarrow 0^+$), which is the usual rule for wave poles, arising from considerations of weak collisions, adiabatic switching in the remote past, or the radiation condition at infinity. Note that the integrand in (13) is even in k_y, k_z , and odd in k_x (for Δ strictly zero). Then we find

$$\begin{aligned}
 Z_A &= \int_0^{\infty} dk_y \int_0^{\infty} dk_z \int_0^{\infty} \frac{dk_x (\Omega_i^2 - V^2 k_x^2) |g|^2}{c^2 V k_{\perp}^2 (\Omega_i^2 + V_A^2 k_z^2)} \\
 &\quad \times \left[\frac{8iV_A^2 k_x}{(k_x + i\Delta)^2 - \omega_A^2/V^2} - \frac{8iV_A^2 k_x}{(k_x - i\Delta)^2 - \omega_A^2/V^2} \right] \\
 &= \frac{8\pi V_A^2}{c^2 V} \int_0^{\infty} \int_0^{\infty} \frac{dk_y dk_z \Omega_i^4}{(\Omega_i^2 + V_A^2 k_z^2)^2} \frac{|g(\omega_A/V, k_y, k_z)|^2}{k_y^2 + \omega_A^2/V^2}
 \end{aligned} \tag{14}$$

where we used

$$\frac{\Delta}{(k_x - \omega_A/V)^2 + \Delta^2} \rightarrow \pi \delta\left(k_x - \frac{\omega_A}{V}\right) \text{ as } \Delta \rightarrow 0^+.$$

A formula somewhat simpler than (14) is obtained by changing variables in the double integral from (k_y, k_z) to $(k_y, \omega_A/V)$, i.e., effectively changing to (k_y, k_x) . The result can thus be directly obtained from (13) by carrying out the integration over k_z . We find

$$\epsilon_1 - \frac{c^2 k_z^2}{\omega^2} = -\frac{c^2 k_z^2 - [k_A^*(k_x + i\Delta)]^2}{V^2 k_x^2},$$

$$k_A^*(k_x) \equiv \frac{\Omega_i k_x V / V_A}{(\Omega_i^2 - k_x^2 V^2)^{1/2}},$$

$$Z_A = \int_0^{\infty} dk_y \int_0^{\infty} \frac{V dk_x}{c^2 k_{\perp}^2} \int_{-\infty}^{\infty} dk_z |g|^2$$

$$\times \left\{ \frac{4ik_x}{k_z^2 - [k_A^*(k_x - i\Delta)]^2} - \frac{4ik_x}{k_z^2 - [k_A^*(k_x + i\Delta)]^2} \right\}.$$

Since we have $k_A^*(k_x \pm i\Delta) = k_A^*(k_x) \pm i\Delta dk_A^*/dk_x$, with dk_A^*/dk_x positive (for k_x positive), the k_z integral yields $8\pi k_x |g(k_x, k_y, k_A^*)|^2 / k_A^*$ for k_x such that $k_A^*(k_x)$ is real ($k_x < \Omega_i/V$), and zero otherwise. Defining $k_i \equiv \Omega_i/V$ we arrive at

$$\begin{aligned}
 Z_A &= \frac{8\pi V_A}{c^2} \int_0^{k_i} \frac{dk_x}{k_i} (k_i^2 - k_x^2)^{1/2} \\
 &\quad \times \int_0^{\infty} \frac{dk_y |g[k_x, k_y, k_A^*(k_x)]|^2}{k_y^2 + k_x^2}.
 \end{aligned} \tag{14'}$$

Slow Extraordinary and Fast Magnetosonic Modes

For the SE branch we have $\epsilon_3 = (\omega^2 - \omega_{pe}^2) / \omega^2$, $\epsilon_1 \simeq -(\omega_{UH}^2 - \omega^2) / \omega^2$. Proceeding as with Alfvén waves we get

$$\epsilon_1 + \frac{k_z^2}{k_{\perp}^2} \epsilon_3 = \frac{k_x^2 + k_{2s}^2}{k_x^2 + k_y^2} \frac{(k_x + i\Delta)^2 - k_{1s}^2}{k_x^2}, \quad \Delta \rightarrow 0^+,$$

$$Z_{SE} = \frac{8\pi}{V} \int_0^{\infty} dk_y \int_0^{\infty} dk_z \frac{|g(k_{1s}, k_y, k_z)|^2}{k_{2s}^2 + k_{1s}^2} \tag{15}$$

where k_{1s}, k_{2s} are known functions of k_y, k_z which we will not pause to give. Changing variables from k_y, k_z to k_y, k_{1s} (effectively k_y, k_x) or directly integrating over k_z in (13), we finally find

$$Z_{SE} = \frac{8\pi}{V} \int_{k_{pe}}^{k_{UH}} \frac{k_x dk_x}{(k_{UH}^2 - k_x^2)^{1/2} (k_x^2 - k_{pe}^2)^{1/2}} \times \int_0^\infty \frac{dk_y |g(k_x, k_y, k_{SE}^*)|^2}{(k_x^2 + k_y^2)^{1/2}}, \tag{15'}$$

$$k_{SE}^* \equiv (k_x^2 + k_y^2)^{1/2} (k_{UH}^2 - k_x^2)^{1/2} / (k_x^2 - k_{pe}^2)^{1/2},$$

$$k_{pe,UH} \equiv \omega_{pe,UH} / V.$$

For the FM branch we have $\epsilon_3 \simeq -\omega_{pe}^2 / \omega^2$, $\epsilon_1 \equiv -\epsilon_3(\omega^2 - \omega_{LH}^2) / (\Omega_e^2 - \omega^2)$,

$$\epsilon_1 + \frac{k_z^2}{k_x^2} \epsilon_3 = \frac{k_{pe}^2 k_x^2 + k_{2F}^2 (k_x + i\Delta)^2 - k_{1F}^2}{k_x^2 k_x^2 + k_y^2} \frac{k_e^2 - k_x^2}{k_e^2 - k_x^2}, \quad \Delta \rightarrow 0^+,$$

$$Z_{FM} = \frac{8\pi}{V} \int_0^\infty \int_0^\infty \frac{dk_y dk_z}{k_{pe}^2} \frac{k_e^2 - k_{1F}^2}{k_{2F}^2 + k_{1F}^2} |g(k_{1F}, k_y, k_z)|^2 \tag{16}$$

with $k_e \equiv \Omega_e / V$ and k_{1F}, k_{2F} , some given functions of k_y, k_z . Also,

$$Z_{FM} = \frac{8\pi}{V} \int_{k_{LH}}^{k_e} \frac{k_x dk_x}{k_{pe}^2} \left(\frac{k_e^2 - k_x^2}{k_x^2 - k_{LH}^2} \right)^{1/2} \times \int_0^\infty dk_y \frac{|g(k_x, k_y, k_{FM}^*)|^2}{(k_x^2 + k_y^2)^{1/2}}, \tag{16'}$$

$$k_{FM}^* \equiv (k_x^2 + k_y^2)^{1/2} (k_x^2 - k_{LH}^2)^{1/2} / (k_e^2 - k_x^2)^{1/2},$$

$$k_{LH} \equiv \omega_{LH} / V.$$

BO derived formulas (14')-(16') for a particular source current by determining the far field; formulas (14)-(16) are new.

4. The Source Divergence of a Tether

As usual, we will take the y axis along the vertical tether, with velocity \mathbf{V} and field \mathbf{B}_0 along axes x and z , as previously noted (Figure 2). Most current divergences discussed in the literature may then be written as

$$\nabla \cdot \mathbf{j}_s(\mathbf{r}) = f(x, z)j[\delta(y + L/2) - \delta(y - L/2)] \tag{17}$$

for an upward current and a tether length L ; $f(x, z)$ is a step function equal to 1(0)inside (outside) certain cross sections representing equal bottom and top contactors. BO considered a circle of radius b ($j = I_s / \pi b^2$), *Hastings*

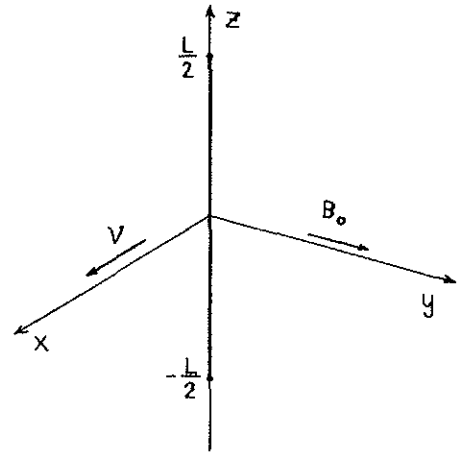


Figure 2. Coordinate system used to compute tether impedance.

and Wang [1987] and *Hastings et al.* [1988] considered rectangular surfaces of sides L_x, L_z ($j = I_s / L_x L_z$). *Estes* [1988] source divergence is that of a ribbon of width L_x , given by (17) with a segment L_x along x as cross section, and $j = I_s \delta(z) / L_x$. An additional case of interest is a ring of radius R , with $j = I_s \delta(r_{per} - R) / 2\pi R$, $r_{per} = (x^2 + z^2)^{1/2}$ (perpendicular to tether axis).

Using (17) in (12) we find

$$g(\mathbf{k}) = \frac{1}{\pi} \sin\left(k_y \frac{L}{2}\right) g_{per}(k_x, k_z) \tag{18}$$

with $(k_x, k_z) \equiv \mathbf{k}_{per}$, and $g_{per} \equiv \langle \exp(-i\mathbf{k}_{per} \cdot \mathbf{r}_{per}) \rangle$ being the average of the exponential in the cross section. For the above models we obtain

$$g_{per}^{ring} = J_0(k_{per} R), \quad g_{per}^{Est} = \frac{\sin(k_x L_x / 2)}{k_x L_x / 2},$$

$$g_{per}^{B0} = \frac{J_1(k_{per} b)}{k_{per} b / 2}, \quad g_{per}^{Hast} = \frac{\sin(k_x L_x / 2) \sin(k_z L_z / 2)}{k_x L_x / 2 \quad k_z L_z / 2}$$

with Est denoting *Estes* [1988] and Hast denoting *Hastings and Wang* [1987] and *Hastings et al.* [1988]. For models deriving from (17), contactors have vanishing characteristic length along the tether (as different from distance L between contactors).

If contactor lengths in the plane (x, z) perpendicular to the tether are small enough, the above models collapse into a common form, with $g_{per} \rightarrow 1$, $g \rightarrow \sin(k_y L / 2) / \pi$. One may also consider negligible perpendicular lengths for a general source divergence not given by (17); we may then write $g(\mathbf{k}) \rightarrow g(k_y)$. This will later be of interest for the Alfvén mode of a bare tether.

Alfvén Mode

In (14') we have $k_z = k_A^*(k_x) \sim k_x V / V_A \ll k_x \ll \Omega_i / V$ ($\simeq 1/36$ m). Thus we may certainly neglect the

characteristic z length and set $k_z = k_A^* \simeq 0$ in g_{per} . Using (18) in (14') we can carry out the k_y integration to obtain

$$Z_A \simeq \frac{2V_A}{c^2} \int_0^{k_i} \frac{dk_x}{k_x} \left(1 - \frac{k_x^2}{k_i^2}\right)^{1/2} \times (1 - e^{-k_x L}) |g_{per}(k_x, 0)|^2. \quad (19)$$

BO gave (19) with g_{per}^{B0} ; Estes [1988] gave a close approximation for his model, using g_{per}^{Est} in place of $|g_{per}^{Est}|^2$.

If the characteristic length along x may also be neglected, one sets $|g_{per}|^2 = 1$ in (19) and, for $L \gg V/\Omega_i$, obtains

$$Z_A = (2V_A/c^2) \ln(2e^{\gamma-1} \Omega_i L/V), \quad \gamma \simeq 0.577, \quad (19')$$

a result found by BO and by Estes. The dominant, logarithmic contribution comes from frequencies $\omega \sim V/L$. Since k_x is ω/V , the single integral over k_x in (19) contains the Alfvén power spectrum; clearly, frequencies $\omega \simeq \Omega_i$ contribute negligibly to the impedance, proving a previous ansatz.

For an arbitrary source divergence with negligible perpendicular lengths, a new, general formula results from using $g(k_y)$ in (14') and carrying out the k_x integration

$$Z_A = \frac{4\pi^2 V_A}{c^2} \int_0^\infty \frac{dk_y}{k_y} \frac{(k_i^2 + k_y^2)^{1/2} - k_y}{k_i} |g(k_y)|^2. \quad (19'')$$

For $g = \sin(k_y L/2)/\pi$, and $Lk_i \gg 1$, one recovers (19'). When deriving (19)-(19'') from (14'), one can verify that large values of k_y/k_x may be ignored, again proving a previous ansatz.

Slow Extraordinary and Fast Magnetosonic Modes

As later detailed, perpendicular contactor lengths are large compared with $1/k_{per}$ for the FM and SE modes. This makes impedances quite dependent on the size of contactors, a case opposite that of A waves, as first observed by BO for FM waves, and repeatedly discussed afterward. Here we note that the effect is heavily dependent on contactor model. At large sizes the factor $|g|^2$ in the (15')-(16') integrals, and therefore the impedance itself, scales differently with contactor size for different contactor shapes. We have

$$\begin{aligned} Z^{ring} &\propto 1/R, & Z^{Est} &\propto 1/L_x^2, \\ Z^{B0} &\propto 1/b^3, & Z^{Hast} &\propto 1/L_x^2 L_z^2, \end{aligned} \quad \text{SE, FM} \quad (20)$$

a square sine or cosine averaging to 1/2 inside the integrals (see Appendix).

This sensitivity of impedance to the source divergence model raises the fundamental question of the proper model. We now make two points. First, there is a

simple rule applying to the above results; for both the ring and Estes sources, the divergence occurs at a finite, one-dimensional set (line contactors); if the set closes on itself, with no boundary (ring), one has $Z \propto (\text{length})^{-1}$; otherwise [Estes, 1988], one has $Z \propto (\text{length})^{-2}$. Now the BO and Hastings and Wang [1987] and Hastings et al. [1988] sources use two-dimensional sets (surface contactors), the first having one boundary (radially), the Hastings set having boundaries on both dimensions; one would thus expect scalings $Z^{B0} \propto (\text{length})^{-1} \times (\text{length})^{-2}$ and $Z^{Hast} \propto (\text{length})^{-2} \times (\text{length})^{-2}$, in agreement with the above results (see the Appendix).

Secondly, this sensitivity questions the validity of neglecting the contactor characteristic length along the tether, as in all those models. We now show that taking into account that length actually determines the proper model, i.e., the correct dependence of Z on contactor size. Nonvanishing lengths along the tether were considered in studying contactor planar surfaces of different orientations [Hastings et al., 1988], the Alfvén far field [Rasmussen et al., 1990], and Debye sheath effects [Donohue et al., 1991].

In a proper model the current divergence should clearly occur at a surface topologically equivalent to a sphere, e.g., a rotational ellipsoid with axis along \mathbf{B}_0 . This is a finite two-dimensional set as in the models of Hastings and Wang [1987] and Hastings et al. [1988] and BO, but it has no boundaries at all; we would thus expect finding $Z \propto (\text{length})^{-2}$, and this is indeed the case. For instance, for (symmetrical) spherical contactors of radius R the source divergence is

$$\begin{aligned} \nabla \cdot \mathbf{j}_s &= \frac{I_s}{4\pi R^2} \left[\delta \left(\left| \mathbf{r} + \frac{L}{2} \mathbf{1}_y \right| - R \right) \right. \\ &\quad \left. - \delta \left(\left| \mathbf{r} - \frac{L}{2} \mathbf{1}_y \right| - R \right) \right] \end{aligned}$$

and one arrives at

$$g = \frac{\sin(k_y L/2)}{\pi} \frac{\sin(kR)}{kR}, \quad \rightarrow \quad Z \propto \frac{1}{R^2}. \quad (21)$$

Note that (1) among sources collapsed along the tether, Estes [1988] model has, coincidentally, the right length dependence and (2) the scaling may be read as $Z \propto A^{-1}$ ($A \equiv$ contactor area).

Using (21) in (15') with $k^2 = k_{SE}^2 + k_\perp^2 = k_\perp^2 k_e^2 / (k_x^2 - k_{pe}^2)$, setting $\sin^2(k_y L/2) \simeq \sin^2(kR) \simeq 1/2$, $k_{UH} \simeq k_{pe} + k_e^2/2k_{pe}$ and calling $k_y/k_x \equiv u$, we find

$$\begin{aligned} Z_{SE} &= \frac{2V}{\pi R^2 \Omega_e^2} \int_{k_{pe}}^{k_{UH}} \frac{dk_x}{k_x} \left(\frac{k_x^2 - k_{pe}^2}{k_{UH}^2 - k_x^2} \right)^{1/2} \\ &\times \int_0^\infty \frac{du}{(1+u^2)^{3/2}} \simeq \frac{V}{2\omega_{pe}^2 R^2} \\ &= \frac{\pi V}{\omega_{pe}^2} \left(\frac{1}{A_{cat}} + \frac{1}{A_{an}} \right), \end{aligned} \quad (22)$$

with $A_{\text{cat}} = A_{\text{an}} = 4\pi R^2$. The only previous result ($Z_{\text{SE}} = 2V^2/b^3\omega_{pe}^3$) corresponds to BO's model, which underestimates Z_{SE} , for $b \sim R$, by a typically small factor, $4V/R\omega_{pe} \sim 1/R(\text{mm})$. When (21) is used in (16'), only values $k_x \sim k_{\text{LH}}$ contribute to the integral. We may then set $k_x/k_e \rightarrow 0$, integrate throughout the range $k_{\text{LH}} \leq k_x < \infty$, and write $k^2 = k_{\text{FM}}^2 + k_{\perp}^2 \simeq k_{\perp}^2$ to find

$$Z_{\text{FM}} \simeq \frac{2V}{\pi R^2 \omega_{pe}^2} \int_{k_{\text{LH}}}^{\infty} \frac{dk_x}{k_x} \frac{k_e}{(k_x^2 - k_{\text{LH}}^2)^{1/2}} \times \int_0^{\infty} \frac{du}{(1+u^2)^{3/2}} = \frac{V}{\omega_{pe} \omega_{pi} R^2} \quad (23)$$

$$= \frac{2\pi V}{\omega_{pe} \omega_{pi}} \left(\frac{1}{A_{\text{cat}}} + \frac{1}{A_{\text{an}}} \right).$$

No previous analytical formula for Z_{FM} was available. Note the simple results (22) and (23).

Note also that Z_{FM} is not influenced by hot-plasma effects, despite concerns to that effect raised earlier (see BO, *Hastings et al.* [1988], *Donohoue et al.* [1991]). This is a pleasant side effect of the fact that a contactor will have a nonvanishing characteristic length along the tether, making the contribution from values, say, $k_x, k_y < 10k_{\text{LH}}$ (approximately inverse electron gyro-radius), dominant in the cold-plasma double integral in (16'). Hot effects on SE waves suggest the energy in such waves should experience heavy collisionless damping [*Hastings et al.*, 1988]; the small value found here for the ratio $Z_{\text{SE}}/Z_{\text{FM}} = m_e/2m_i \sim 10^{-5}$ further suggests that such energy is negligible. Note also the ratio

$$\frac{Z_{\text{A}}}{Z_{\text{FM}}} = 2 \left(\frac{m_i}{m_e} \right)^{1/2} \frac{R^2 \Omega_i^2}{V_A V} \ln \left(2e^{\gamma-1} \frac{\Omega_i L}{V} \right) \quad (24)$$

$$\simeq \left[\frac{n_e (\text{cm}^{-3})}{3.6 \times 10^5} \right]^{1/2} \left[\frac{R(\text{m})}{13} \right]^2 \ln \frac{L(\text{km})}{0.028},$$

taken from (19') and (23); n_e is electron density.

We now check our having assumed k_x, k_z to be large compared with inverse perpendicular contactor lengths, for the integrals in (15') and (16'). This is clear for the SE branch, where we have $k_z = k_{\text{SE}}^* \sim k_x \simeq \omega_{pe}/V \sim 1/0.2 \text{ mm}$. The FM case is more involved. Although we do have $k_x \geq \omega_{\text{LH}}/V \sim 1/20 \text{ cm}$, which is comparatively large, $k_z \equiv k_{\text{FM}}^*$ is smaller by a factor of order $k_{\text{LH}}/k_e = (m_e/m_i)^{1/2}$ for $k_x \sim \omega_{\text{LH}}/V$. However, in a proper model of g such as given by (21), k_z appears only in $k_{\text{per}} = (k_x^2 + k_z^2)^{1/2} > k_x$, and our assumption holds.

Nonetheless, consider uniform current divergence on the surface of a rotational, prolate ellipsoid of major and minor semi-axes a and b along and perpendicular to \mathbf{B}_0 respectively. A straightforward development gives the result

$$g = \frac{\sin(k_y L/2)}{\pi} \frac{2 \int_0^1 d\eta (1 - \eta^2 e^2)^{1/2}}{(1 - e^2)^{1/2} + (\arcsin e)/e} \quad (25)$$

$$\times \cos(k_z a \eta) J_0 \left[k_{\perp} b (1 - \eta^2)^{1/2} \right],$$

where $e = (1 - b^2/a^2)^{1/2}$. For $b = a \equiv R$ ($e = 0$) one recovers (21). For an opposite case, $V/\omega_{\text{LH}} \ll b \ll a \ll V/\Omega_i$, in the FM branch, one gets

$$g = \frac{\sin(k_y L/2)}{\pi} \left[J_0^2 \left(\frac{k_{\perp} b}{2} \right) - J_1^2 \left(\frac{k_{\perp} b}{2} \right) \right]$$

$$\simeq \frac{\sin(k_y L/2)}{\pi} \frac{4 \sin(k_{\perp} b)}{\pi k_{\perp} b}.$$

Clearly, the impedance then reads

$$Z_{\text{FM}} \simeq \frac{16}{\pi^2} \frac{V}{b^2 \omega_{pe} \omega_{pi}} = \frac{a}{b} \frac{8V}{\omega_{pe} \omega_{pi}} \left(\frac{1}{A_{\text{cat}}} + \frac{1}{A_{\text{an}}} \right), \quad (26)$$

which is much larger than the value for the spherical case as given by (23). Note, however, that the actual divergence might be highly nonuniform.

5. Bare Tether Impedance

It has been shown recently that a tether could work efficiently without an anodic contactor, by collecting electrons along certain length of its anodic end, if bare [*Sanmartín et al.*, 1993]. If actually bare along its entire length, a (generator) tether was found to be optimal if positively biased over a length $l_a \simeq \frac{1}{7}L$. The current $\tilde{I}(y)$ would vanish at $y = L/2$ and reach a nearly constant value I_s at $y = L/2 - l_a$, with $d\tilde{I}/dy \propto -(y - \frac{1}{2}L + l_a)^{1/2}$ (Figure 3). If b is the collecting radius in the x, z plane and the cathodic contactor is a sphere of radius R , we then have

$$\nabla \cdot \mathbf{j}_s = \frac{I_s \delta(|\mathbf{r} + \frac{1}{2}L\mathbf{1}_y| - R)}{4\pi R^2} - I_s \frac{(y - \frac{1}{2}L + l_a)^{1/2}}{\frac{2}{3}l_a^{3/2}}$$

$$\times \frac{\delta(r_{\text{per}} - b)}{2\pi b} \left[h\left(y - \frac{L}{2} + l_a\right) - h\left(y - \frac{L}{2}\right) \right]$$

where h is the unit step function. For $k_y > 0$ we find

$$g = \frac{\sin kR}{2\pi i k R} e^{ik_y L/2} - \frac{3J_0(k_{\text{per}} b)}{4\pi k_y l_a} e^{-ik_y L/2}$$

$$\times \left[1 - e^{ik_y l_a} \frac{C_2(k_y l_a) - iS_2(k_y l_a)}{(2k_y l_a/\pi)^{1/2}} \right],$$

with Fresnel integrals $[C_2(s), S_2(s)] \equiv \int_0^s \{\cos u, \sin u\} \times du/(2\pi u)^{1/2}$, and

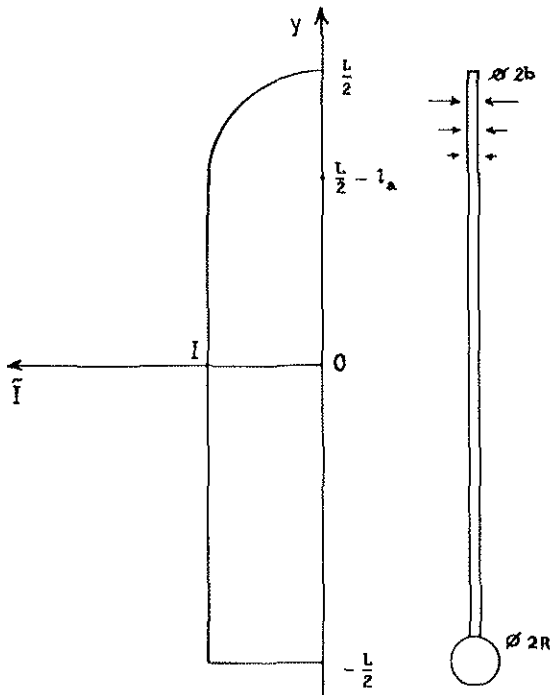


Figure 3. Anodeless, bare tether collecting electrons at its upper (anodic) end, over a positively biased length $l_a (\simeq L/7)$; the bias vanishes at height $L/2 - l_a$.

$$|g|^2 = \left(\frac{\sin kR}{2\pi kR} \right)^2 + \frac{J_0^2(k_{\text{per}}b)F(k_y l_a)}{4\pi^2} + \frac{3J_0(k_{\text{per}}b)}{4\pi^2} \times \frac{\sin kR}{kR} [g_1(k_y l_a) \cos(k_y L) + g_2(k_y l_a) \sin(k_y L)], \quad (27)$$

$$F(s) = \frac{9\pi}{8s^3} \left\{ \left[S_2(s) - \left(\frac{2s}{\pi} \right)^{1/2} \sin s \right]^2 + \left[C_2(s) - \left(\frac{2s}{\pi} \right)^{1/2} \cos s \right]^2 \right\},$$

$$g_1(s) = (\pi/2s^3)^{1/2} [S_2(s) \cos s - C_2(s) \sin s],$$

$$g_2(s) = (\pi/2s^3)^{1/2} [C_2(s) \cos s + S_2(s) \sin s] - 1/s.$$

For the Alfvén branch and $kR, k_{\text{per}}b \ll 1$, we have $g \rightarrow g(k_y)$, (27) becoming

$$4\pi^2 |g|^2 \simeq 1 + F(k_y l_a) + 3g_1(k_y l_a) \cos(k_y L) + 3g_2(k_y l_a) \sin(k_y L),$$

for use in (19''). Note that if $k_y l_a \rightarrow 0$, we recover $|g|^2 = \sin^2(k_y L/2)/\pi^2$. At large $s \equiv k_y l_a$ we have $F(s) = 0(s^{-2})$, $g_1(s) = 0(s^{-3/2})$, $g_2(s) = 0(s^{-1})$, and $4\pi^2 |g|^2 \rightarrow 1$. We can thus integrate separately ranges $k_y < k_m$ and $k_y > k_m$, with

$$7/L \simeq 1/l_a \ll k_m \ll \Omega_i/V.$$

Equation(19'') can then be written as

$$\frac{c^2 Z_A}{V_A} \simeq \int_{k_m}^{\infty} \frac{dk_y}{k_y} \frac{(k_i^2 + k_y^2)^{1/2} - k_y}{k_i} + \int_0^{k_m l_a} ds \frac{1 + F(s) + 3g_1(s) \cos 7s + 3g_2(s) \sin 7s}{s},$$

leading to an impedance

$$Z_A = \frac{V_A}{c^2} \ln \left(2e^{\alpha-1} \Omega_i \frac{L}{V} \right), \quad (28)$$

$$\alpha \equiv \int_0^1 [1 + F(s) + 3g_1(s) \cos 7s] \frac{ds}{s}$$

$$+ \int_1^{\infty} [F(s) + 3g_1(s) \cos 7s] \frac{ds}{s}$$

$$+ \int_0^{\infty} 3g_2(s) \sin 7s \frac{ds}{s} - \ln 7 \simeq 4.00.$$

For the FM branch we may take kR large. Then the last term in (27) makes a negligible contribution to the impedance Z_{FM} as shown in (16'), while the first term yields one half the value given by (23), in agreement with the general result that anodic and cathodic contactors make independent contributions to the FM impedance [Donohue et al., 1991]. For the middle term we can set $k_y \sim 1/l_a \ll \omega_{\text{LH}}/V < k_x$ and $k_z = k_{\text{FM}}^* \ll \omega_{\text{LH}}/V \sim k_x$, thus writing $k_{\perp} \simeq k_x$, $k_{\text{per}} \sim k_x$. If, additionally, $k_{\text{per}}b$ is large, we arrive at

$$Z_{\text{FM}} = \frac{V/2}{R^2 \omega_{pe} \omega_{pi}} + \frac{2V}{\pi \omega_{pe}^2 b l_a} \int_{k_{\text{LH}}}^{\infty} \frac{dk_x}{k_x} \frac{k_e \beta}{(k_x^2 - k_{\text{LH}}^2)^{1/2}} = \frac{2\pi V}{\omega_{pe} \omega_{pi}} \left(\frac{1}{A_{\text{cat}}} + \frac{\beta}{A_{\text{an}}} \right), \quad (29)$$

$$A_{\text{an}} \equiv 2\pi b l_a, \quad \beta \equiv \int_0^{\infty} \frac{F(s) ds}{\pi} \simeq 1.12.$$

Note that written in terms of areas, (23) and (29) read nearly the same.

Actually, condition $b \gg 1/k_{\text{per}} \sim 20$ cm will not be satisfied if b is the actual radius of the tether; for $k_{\text{per}}b$ small we would have $Z_{\text{FM}} \simeq 2\pi V/\omega_{pe} \omega_{pi} A_{\text{cat}}$, or one half the value given in (23). At this point we raise the question of what are the lengths really characterizing $\nabla \cdot \mathbf{j}_s$. Estes [1988] suggested that the plasma cloud emitted by an active contactor might extend the effective dimensions well beyond the dimensions of the contactor itself; Donohue et al. [1991] further suggested that the sheath radius should be an effective radius for passive contactors.

The basic argument underlying these suggestions is the need to account for nonlinear effects, which are essential for a self-consistent analysis. A crude recipe is to consider the entire nonlinear region around the contac-

tor as part of the contactor itself. For the anodic part of a bare tether, which collects electrons as a cylindrical Langmuir probe in the orbital-motion-limited regime, the faraway electric potential decays as the inverse of distance to tether [Laframboise, 1966]. Then, for typical electron temperature $T_e \simeq 0.1$ eV, tether radius (~ 1 mm) and anodic bias (hundreds of volts), the average b would be of order of meters, large, indeed, compared with $1/k_{\text{per}}$.

The above argument suggests that nonlinear effects would adjust contactor areas to effective values

$$A_{\text{cat}} \sim I_s/j_{th} \sim A_{\text{an}}, \quad (30)$$

where j_{th} is the (unperturbed) electron random current density, $j_{th} \equiv \frac{1}{4}en_e(8T_e/\pi m_e)^{1/2}$, with $e \equiv$ electron charge. Using (30) in our previous (23) or (29) and using $m_i V^2 \simeq 8.9$ eV then yields a very simple result

$$Z_{\text{FM}} I_s \sim \frac{4\pi V j_{th}}{\omega_{pe} \omega_{pi}} = \left(\frac{m_i V^2 T_e}{2\pi e^2} \right)^{1/2} \simeq 0.38 V. \quad (31)$$

Note that the power radiated in the FM branch, $Z_{\text{FM}} I_s^2$, will increase only linearly (rather than quadratically) with current. FM radiation will dominate Alfvén radiation, except at large currents, when the effective contactor area and effective radius R is large; this is reflected in (24), where Z_A/Z_{FM} scales as R^2 . A 0.5 km long (the so-called PMG) tether, flown in June 1993, reached a current $I_s \simeq 0.3A$ at the maximum density $n_e \simeq 10^6 \text{ cm}^{-3}$, yielding $R \simeq 1.54$ m in (30); (31) and (24) then give $Z_{\text{FM}} \simeq 1.27 \Omega$ and $Z_A \simeq 0.086 \Omega$. For the aborted tethered satellite system (TSS) 1 experiment the effective area was probably the actual physical surface of the passive anode (a sphere of radius 0.8 m), the current in (30) being of purely thermal origin, $I_s \simeq 0.066 A$ at $n_e \simeq 10^6 \text{ cm}^{-3}$. Equation (31) then yields $Z_{\text{FM}} \simeq 5.7 \Omega$, a result also directly obtained from (23). Since the deployed tether length was $L \simeq 300$ m, one finds $Z_A/Z_{\text{FM}} \simeq 0.015$ in (24).

6. Summary of Results

We have used the potential ϕ for the longitudinal part of the electric field of a wave and conditions particular to low Earth orbits in discussing the five branches $\omega(k, \theta)$ of the dispersion relation for a cold, magnetized plasma. No emission of fast extraordinary or ordinary waves is possible. Slow extraordinary (SE) and fast magnetosonic (FM) emission occurs near $k \rightarrow \infty$ asymptotes, with $\mathbf{E} \simeq -\nabla\phi$. Alfvén (A) emission occurs in a peculiar regime, arising from condition

$$V_A^2 \ll V^2 m_i/m_e,$$

with $\mathbf{E} \simeq -\nabla_{\perp}\phi$. The discussion clarifies or corrects some published results on group velocities, emission bands, and emission of whistlers or outside contactors.

The ϕ formalism appears useful for future work on a number of standing issues. The formalism has been used here to determine the impedance of an orbiting conductor.

For tethers we found that usual models, having vanishing contactor length along the tether ($L_y = 0$), yield values Z_{FM} (and Z_{SE}) heavily dependent on the model. A proper model requires $L_y \neq 0$ and this yields $Z_{\text{FM,SE}} \propto (\text{contactor area})^{-1}$. We found new, simple formulas for Z_{SE} (see (22)), so small that power radiated into this strongly damped branch should be negligible, and Z_{FM} (23) unaffected by hot-plasma effects. An anodeless, bare tether of equal area, has a very similar Z_{FM} (29). Nonlinear effects might make contactor area proportional to current I_s , and this would lead to an extremely simple result, (31). We have also found Z_{FM} for ellipsoidal contactors (see (26)) and Z_A for contactors with $L_y \neq 0$ but with vanishing lengths perpendicular to the tether (see (19')), particularized to a bare tether in (28).

Appendix

Consider an (x, z) surface contactor with polygonal contour, as in the rectangular model for (17) by Hastings *et al.* [1988]. We then have

$$\begin{aligned} g_{\text{per}} &= \int_{s_m}^{s_M} e^{-ik_{\text{per}}s} \frac{ds}{A_c} w_c(s) \\ &= \int_{s_m}^{s_M} e^{-ik_{\text{per}}s} \frac{ds}{A_c} \frac{dw_c/ds}{ik_{\text{per}}}, \end{aligned} \quad (A1)$$

where A_c is area and $w_c(s)$ is full width at given s ; we used a partial integration and the vanishing of $w_c(s)$ at both s_m and s_M (Figure 4a). Since dw_c/ds is piecewise constant, g_{per} is a sum of straightforward integrals, yielding $g_{\text{per}} \propto 1/A_c k_{\text{per}}^2 \propto 1/A_c$, $Z \propto 1/A_c^2$, in agreement with the result in (20).

If the contour is a smooth curve, as in the BO model, dw_c/ds diverges at both s_m and s_M (Figure 4b). A range $\Delta s \sim 1/k_{\text{per}}$ around each extreme then gives dominant contributions of order $1/k_{\text{per}}^{3/2}$ to (A1); within such range we have $dw_c/ds \sim (R/\Delta s)^{1/2}$, where $R \sim A_c^{1/2}$ is the radius of curvature of the contour. This yields $g_{\text{per}} \propto \Delta s (R/\Delta s)^{1/2} / A_c k_{\text{per}} \propto 1/A_c^{3/4} k_{\text{per}}^{3/2} \propto 1/A_c^{3/4}$, $Z \propto 1/A_c^{3/2}$, as in (20).

For models with line rather than surface contactors we would have

$$g_{\text{per}} = \int_{s_m}^{s_M} e^{-ik_{\text{per}}s} \frac{ds}{L_c} \frac{dl_c}{ds}, \quad (A2)$$

where $l_c(s)$ is the length lying between s_m and a given s , and L_c is total length (Figures 4a and 4b). For a polygonal (open or closed) curve, dl_c/ds is piecewise constant, yielding $g_{\text{per}} \propto 1/L_c k_{\text{per}} \propto 1/L_c$, $Z \propto 1/L_c^2$.

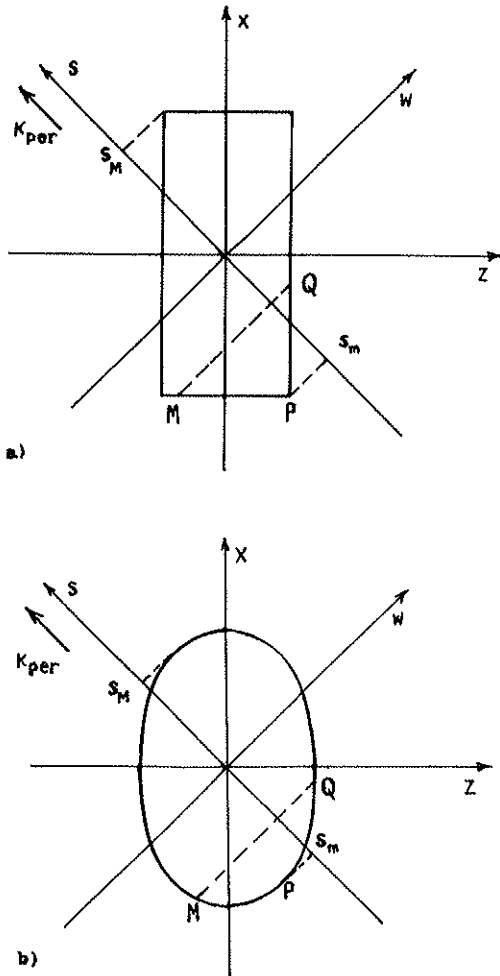


Figure 4. Cross section of contactor with vanishing length along tether and (a) polygonal, (b) smooth shape. For a surface contactor, $w_c(s)$ represents the width MQ; for a line contactor, $l_c(s)$ represents the length of the curve MPQ.

For a smooth curve, dl_c/ds diverges at s_m and s_M ; if $R \propto L_c$ is radius of curvature, (A2) gives $g_{per} \propto \Delta s(R/\Delta s)^{1/2}/L_c \propto 1/L_c^{1/2} k_{per}^{1/2} \propto 1/L_c^{1/2}$, $Z \propto 1/L_c$. This explains the results in (20) for the Estes [1988] and ring models.

Acknowledgments. This work was supported by the Comision Interministerial de Ciencia y Tecnología of Spain (project ESP92-0989-E).

The Editor thanks B.E. Gilchrist and D.J. Donohue for their assistance in evaluating this paper.

References

- Akhiezer, A.I., I.A. Akhiezer, R.V. Polovin, A.G. Sitenko, and K.N. Stepanov, *Plasma Electrodynamics*, vol.1, Pergamon, New York, 1975.
- Banks, P.M., P.R. Williamson, and K.L. Oyama, Electri-

- cal behavior of a shuttle electrodynamic tether system (SETS), *Planet. Space Sci.*, 29, 139, 1981.
- Barnett, A., and S. Olbert, Radiation of plasma waves by a conducting body moving through magnetized plasma, *J. Geophys. Res.*, 91, 10,117, 1986.
- Dobrowolny, M., and P. Veltri, MHD power radiated by a large conductor in motion through a magnetoplasma, *Nuovo Cimento Soc. Ital. Fis. C.*, 9, 27, 1986.
- Donohue, D.J., T. Neubert, and P.M. Banks, Estimating radiated power from a conducting tethered satellite system, *J. Geophys. Res.*, 96, 21,245, 1991.
- Drell, S.D., H. M. Foley, and M. A. Ruderman, Drag and propulsion of large satellites in the ionosphere: An Alfvén propulsion engine in space, *J. Geophys. Res.*, 70, 3131, 1965.
- Estes, R.D., Alfvén waves from an electrodynamic tethered satellite system, *J. Geophys. Res.*, 93, 945, 1988.
- Hastings, D.E., and J. Wang, The radiation impedance of an electrodynamic tether with end connectors, *Geophys. Res. Lett.*, 14, 519, 1987.
- Hastings, D.E., and J. Wang, Induced emission of radiation from a large space-station-like structure in the ionosphere, *AIAA J.*, 27, 438, 1989.
- Hastings, D.E., A. Barnett, and S. Olbert, Radiation from large space structures in low Earth orbit with induced alternating currents, *J. Geophys. Res.*, 93, 1945, 1988.
- Laframboise, J.G., Theory of spherical and cylindrical Langmuir probes in a collisionless, Maxwellian plasma at rest, *Rep. 100*, Univ. of Toronto, Inst. for Aerosp. Stud. Toronto, Canada, 1966.
- McKenzie, J.F., Stationary MHD waves generated by a source in a moving plasma, *J. Geophys. Res.*, 96, 9491, 1991.
- Rasmussen, C.E., P.M. Banks, and K.J. Harker, The excitation of plasma waves by a current source moving in a magnetized plasma: The MHD approximation, *J. Geophys. Res.*, 90, 505, 1985.
- Rasmussen, C.E., P.M. Banks, and K.J. Harker, The excitation of plasma waves by a current source moving in a magnetized plasma: Two-dimensional propagation, *J. Geophys. Res.*, 95, 10,459, 1990.
- Sanmartín, J.R., M. Martínez-Sánchez and E. Ahedo, Bare wire anodes for electrodynamic tethers, *J. Propul. Power*, 9, 353, 1993.
- Stenzel, R.L., and J.M. Urrutia, Currents between tethered electrodes in a magnetized laboratory plasma, *J. Geophys. Res.*, 95, 6209, 1990.
- Urrutia, J.M., and R.L. Stenzel, Transport of current by whistler waves, *Phys. Rev. Lett.*, 62, 272, 1989.
- vom Stein, R., and F.M. Neubauer, Plasma wave field generation by the tethered satellite system, *J. Geophys. Res.*, 97, 10,849, 1992.

M. Martínez-Sánchez, Department of Aeronautics and Astronautics, Massachusetts Institute of Technology, Cambridge, MA 02139. (e-mail:mmart@mit.edu)

J. R. Sanmartín, E.T.S.Ingenieros Aeronáuticos, Universidad Politécnica, Plaza Cardenal Cisneros 3, 28040-Madrid, Spain.

(Received April 20, 1994; revised October 13, 1994; accepted October 28, 1994.)