Fission of Highly Excited Nuclei: a Finite Temperature Description

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Outline



Introduction

- Fission at Finite Temperature
- Theory Overview

2 Results

- The General Picture
- Collective Masses
- Fission half-lives

3 Conclusions

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Fission at Finite Temperature Theory Overview

What are we going to do?

- An attempt to study the temperature dependence of fission in heavy nuclei using a microscopic theory. Application to two test nuclei.
 - Gogny force.
 - Reasonably big configuration space
 - Axially deformed base.
 - Breaking of reflection symmetry allowed: octupole deformation.

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Fission at Finite Temperature Theory Overview

Specifics: Gogny Force

- D1S set.
- Pairing automatically included.
- D1S set has not been adjusted to finite temperature, although it has adjusted the surface energy term, hence it has been successful with fission barriers at zero temperature and high angular momentum.

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Fission at Finite Temperature Theory Overview

Specifics: Calculation Basis

- Configuration space with 15 major shells. Checked for convergence with 17 shells in selected cases.
- Axially deformed basis: we are going to put a constraint on $\langle \hat{Q}_2 \rangle = q_2$.
- Standard truncation condition:

$$N_{\perp} + rac{n_z}{q} < N_0$$

where N_{\perp} , n_z are the HO quantum numbers and q is the nuclear axis ratio. A value q = 1.5 is used.

• i.e: *N*₀ shells in the perpendicular direction and up to 1.5*N*₀ in the axial (*z*) direction.

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Fission at Finite Temperature Theory Overview

Finite Temperature HFB

• For a system at constant *T* and average number of particles *N*, the equilibrium state is obtained by minimizing the grand canonical potential

$$\Omega = F - \mu N$$
 with $F = E - TS$

By solving

$$\left(\begin{array}{cc}h&\Delta\\-\Delta^*&-h^*\end{array}\right)\left(\begin{array}{c}U_k\\V_k\end{array}\right)=\left(\begin{array}{c}U_k\\V_k\end{array}\right)E_k$$

with *h* the HF hamiltonian and Δ the pairing potential, *U*, *V* and *E_i* are obtained.

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Fission at Finite Temperature Theory Overview



 This in turn allows for the calculation of the density matrix and pairing tensor:

$$\rho = UfU^+ + V^*(1-f)V^t$$

$$\kappa = UfV^+ + V^*(1-f)U^t$$

with
$$f_i = \frac{1}{1+e^{\beta E_i}}$$
 and $\beta = 1/kT$

 From this, expected values of an observable Ô are obtained as thermal averages

$$O = \operatorname{Tr}(\hat{D}\hat{O})$$
 where

$$\hat{D} = Z^{-1} \exp(-eta(\hat{H} - \mu \hat{N}))$$
 and $Z = \operatorname{Tr}\left[\exp(-eta(\hat{H} - \mu \hat{N}))
ight]$

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Fission at Finite Temperature Theory Overview

An additional constraint

 In order to study the barriers as the nucleus elongates, we add a constraint on the quadrupole deformation. Minimize:

$$\Omega = \boldsymbol{E} - \boldsymbol{T}\boldsymbol{S} - \mu \boldsymbol{N} - \lambda_{\boldsymbol{Q}_{20}} \boldsymbol{q}_2$$

• Once done this, thermal fluctuations come for free. The probability of obtaining a given value *q* of the constrained magnitude is given by the Free energy as

$$P(q) \propto e^{-F(q)/T}$$

Ensemble averages of the observable \hat{O} are obtained by means of its thermal expectation value O(q):

$$\left\langle \hat{O} \right\rangle = rac{\int O(q) e^{-F(q)/T} \mathrm{d}q}{\int e^{-F(q)/T} \mathrm{d}q}$$

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Fission at Finite Temperature Theory Overview

Collective Masses

- To include dynamical effects, collective masses in the quadrupole degree of freedom are calculated using the ATDHFB framework at finite temperature.
- As usual, the residual interaction is neglected and the masses are:

$$M(q_{20}) = rac{1}{2}\sum_{\mu
u} rac{\mathbb{Q}_{\mu
u}\mathbb{Q}_{\mu
u}}{|\mathbb{E}_{\mu}-\mathbb{E}_{
u}|^3}|\mathbb{F}_{\mu}-\mathbb{F}_{
u}|^3$$

- This approximation gives trouble when two levels cross or their quasiparticle energies are almost equal. An extra, constant, value added to the denominator avoids numerical divergence and simulates the residual interaction.
- Values in the 0.5-1.0 range were tested, and the results do not show a qualitative difference being quantitatively close.

Fission at Finite Temperature Theory Overview

Fission Half-life

• Spontaneous fission half-life is computed as:

$$T_{sf} = \frac{\ln 2}{\nu} \frac{1}{P}$$

• The assault frequency is set to 1MeV/ \hbar and the penetration probability is calculated using the WKB approximation as:

$$P = \frac{1}{1 + \exp(2G)}$$

• and the action: $G = \int_{q_2 min}^{q_2 max} \sqrt{2M(q_2)\Delta F} dq_2$ where $\Delta F = F(q_2) - E_0$ is the free energy above the ground state. The q_2 limits are set to span the barrier of interest. $F(q_2) = E(q_2) - TS(q_2)$, where $E(q_2)$ is obtained by substracting the zero-point energy correction for q_2 to the corrected (kinetic energy and rotational energy) HFB energy.





Calculations have been performed on two well known nuclei:

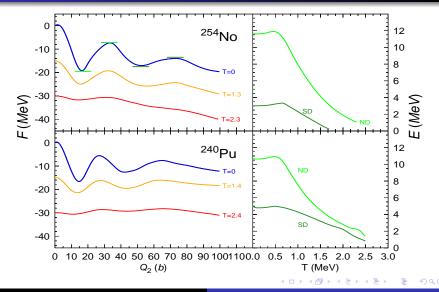
- ²⁴⁰Pu, typical benchmark case.
- ²⁵⁴No, also extensively studied. High spin D1S calculations (T=0) available. Shell stabilized ('magic' N=152).
 Appearance of octupole deformation in its path to fission.

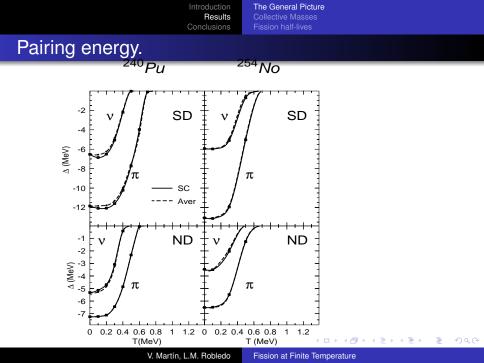
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Introduction The General Picture Results Collective Masses Conclusions Fission half-lives

Free Energy Curves and Fission Barriers.



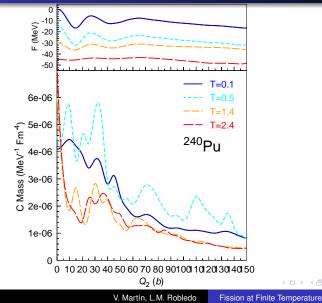


Introduction Results The General Picture Collective Masses Fission half-lives

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²⁴⁰Pu Collective Mass

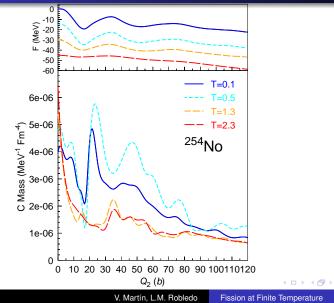


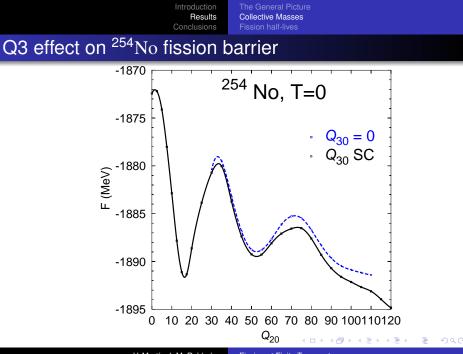
Introduction Results The General Picture Collective Masses Fission half-lives

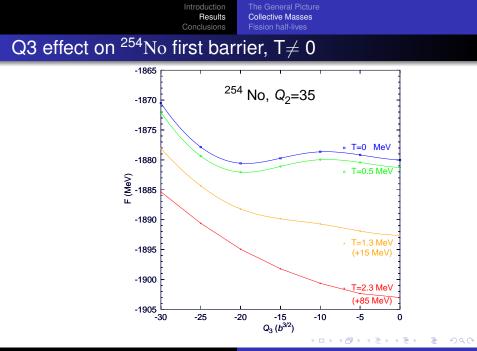
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²⁵⁴No Collective Mass





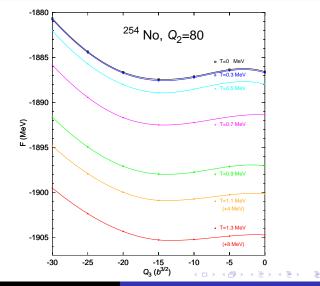


 Introduction
 The General Picture

 Results
 Collective Masses

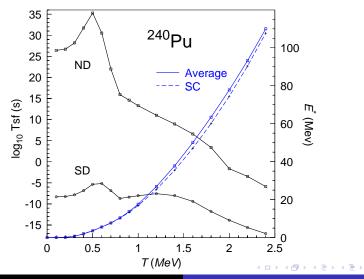
 Conclusions
 Fission half-lives

Q3 effect on ²⁵⁴No second barrier, $T \neq 0$



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²⁴⁰Pu Fission half-life



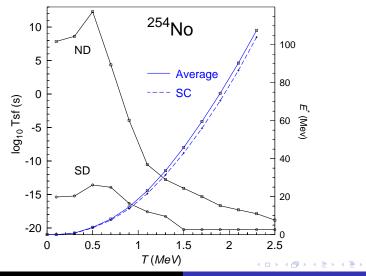
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 Introduction
 The General Pictor

 Results
 Collective Masse

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 Fission half-lives

²⁵⁴No Fission half-life



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Conclusions

- Gogny D1S force used in FTHFB fission calculations with two typical test nuclei. Axial symmetry, octupole shapes allowed.
 - Fission barriers, as expected, go to zero with temperature. First and second barriers dissappear around the same temperature in ²⁴⁰Pu.
 - A small increase in the height of the barriers is seen correlated with the pairing collapse. This is reflected in that spontaneous half-lives are bigger at around T=0.5 MeV.
 - Mass parameters behave as expected: smaller around the free energy minima and bigger near the top of the barriers. An overall increase of its value when pairing collapses.
 - Reflection assymetry also dissapears with temperature, although it is still relevant at pretty high temperatures.
- Future work.
 - Incoporate triaxial shapes. Fluctuations.