ON THE POLARIZATION OF RADIO-WAVE TRAVELLING THROUGH THE IONOSPHERE

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ABSTRACT. The limiting polarization as deduced by Baker and Green without taking the usual Appleton-Hartree formulae has been shown to be identical with the limiting polarization deduced directly from the Appleton-Hartree formula. From Bailey's formulae for the amplitude-ratio of the normal to the abnormal components of the magnetic vector of the radio-wave, the phase-difference between them at any level of the ionosphere has been deduced for $\mathbf{r} << \mathbf{r}_c$ where \mathbf{r}_c is the critical collisional frequency. It has also been shown that for $\mathbf{r} << \mathbf{r}_c$, the amplitude ratio is nearly unaffected at any level by electron collisional frequency.

 Identity of the expression for the limiting polarization as given by Baker and Green with that deduced directly from the Appleton-Hartree formulae

Without taking the usual Appleton-Hartree formulae (1927, 1929) it was shown by Baker and Green (1933) that the limiting polarization R_a is given by the roots of the following equation.

$$R_a^2 + \frac{l^2}{n} q' R_a - 1 = 0$$
(1)

where

$$R_{n} = -\frac{E_{y}}{iE_{x}} \cdot \frac{l^{2}}{n} = \frac{\sin^{2}\theta}{\cos\theta} \quad q' = \frac{p_{0}}{p - jv} \quad p_{0} = \frac{eH}{mc}$$
 (2)

 $\theta=$ angle between the direction of propagation of the radio-wave and the positive direction of earth's magnetic field.

 $E_y =$ component of the electric vector along OY (Fig. 1).

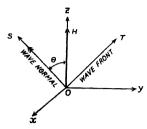
 $E_t = {
m component}$ of the electric vector along OT (Fig. 1).

H = intensity of earth's magnetic field.

e, m = electronic charge and mass.

v = electronic collisional frequency.

c = velocity of light in vacuum.



y y H

Fig 1. Co-ordinate system used by Baker and Green. OS—wavenormal OT—wavefront

H - earth's magnetic field

Fig 2 Right-handed co-ordinate system used by Appleton OX —wave normal H—curth's magnetic field

When the right-handed co-ordinate system of Fig. 2 is used, where the direction of propagation of the radio-wave has been taken along OX-axis, it can be shown that

$$R_{\sigma} = -\frac{E_{y}}{jE} = -j \left(\frac{\hbar z}{\hbar y} \right) \tag{3}$$

where

hz = component of the magnetic vector along OZ (Fig. 2).

hy = component of the magnetic vector along OY (Fig. 2).

From Eq. (3)

$$\left(\frac{hz}{hy}\right) = jR_a$$
 (3a)

Solving Eq. (1), we have

$$R_{q} = -\frac{l^{2}}{2n} q' \pm \sqrt{\frac{l^{4}}{4n^{2}}} q'^{2} + 1$$
 (1a)

Putting the value of l^2/n and q' from Eq. (2) in Eq. (1a)

$$R_{\sigma} = -\frac{eH}{2mc\cos\frac{\theta}{\theta(p-jv)}} \pm \sqrt{\frac{e^2H^2\sin^4\theta}{4m^2c^2\cos^2\theta(p-jv)^2}} + 1 \qquad \dots \quad (1b)$$

Since the critical collisional frequency is

$$v_c = \frac{p_H}{2} \frac{\sin^2 \theta}{\cos \theta}$$
 , where $p_H = \frac{eH}{mc}$

we get

$$R_a = -\frac{v_e}{p - j_V} \perp \sqrt{\frac{v_o^2}{(p - j_V)^2} + 1}$$
 (1c)

Using (le) and (3a):

$$\begin{pmatrix} hz \\ hy \end{pmatrix} = j \left[-\frac{v_e}{p-j\mathbf{v}} \pm \sqrt{\frac{v_e^2}{(p-j\mathbf{v})^2} + 1} \right] \qquad \dots \quad (1d)$$

This equation can be deduced directly from the Appleton-Hartree formula. Appleton used the right-handed co-ordinate system of Fig. 2 and the polarization is given by

$$\left(\frac{hz}{hy}\right) - R = -\frac{J}{\gamma_L} \left[-\frac{\gamma_T^2}{2(1+\alpha+j\beta)} \pm \sqrt{\frac{\gamma_T^4}{4(1+\alpha+j\beta)^2} + \gamma} \right] \quad \dots \quad (4)$$

Using $\alpha = p^2/p_0^2$, $\beta = p\nu/p_0^2$, $p_0^2 = 4\pi Ne^2/m$, γ , $\tau = pp_{L,T}/p_0^2$ Eq. (4) can be written as

$$R = -\frac{J}{p_L} \left[-\frac{pp_T^2}{2(p_0^2 - p^2 + jp_V)} \right] \sqrt{\frac{p^2 \bar{p}_T^3}{4(p_0^2 - p^2 + jp_V)^2} + p_L^2} \right] \qquad \dots (4a)$$

Hence the limiting value is obtained by putting $p_0^2 = 0$ (i.e. N = 0) in Eq. (4a)

$$R = \begin{pmatrix} hz \\ hy \end{pmatrix} = j \left[-\frac{\frac{v_e}{p - j_v} \pm \sqrt{\frac{v_e^2}{(p - j_v)^2}} + 1} \right] \qquad \dots \quad (4b)$$

It is seen that Eqs. (1d) and (4b) are identical.

Evaluation of the phase-difference between the normal and the abnormal Components of the magnetic vector of a radio-wave at any level of the ionosphere for v << v_e

It was shown by Bailey (1934) that the amplitude-ratios of the normal to the abnormal components of magnetic vector are given by

$$\rho_0 = a \left[1 - \frac{d_1}{d_2} \cot \phi_0 \right] \text{ for the O-mode} (5a)$$

and

$$\rho_z = a \Big[\, 1 + \frac{d_1}{d_2} \cot \phi_0 \, \Big] \text{ for the X-mode .} \qquad \qquad \dots \quad (5b)$$

where

$$a = \frac{1}{2} [\sqrt{1 + Y} + \sqrt{1 + Y'}]$$

$$Y = 2 \frac{d_1 d_2^2}{d_1^2 + d_2^2} + \frac{d_1^2 d_2^2}{d_1^2 + d_2^2} \dots (50)$$

$$\begin{split} Y' &= -2 \ \frac{d_1 d_2{}^2}{d_1{}^2 + d_2{}^2} + \frac{d_1{}^2 d_2{}^2}{d_1{}^2 + d_2{}^2} \\ d_1 &= \mathrm{v_c/v}, \ d_2 = \mathrm{v_c/p'}, \ p' = p \ \left(1 - \frac{p_0{}^2}{p^2}\right), \ p_0{}^2 = \frac{4\pi N e^2}{m} \end{split}$$

 ϕ_0 = phase-difference for the O-mode.

Using Eq. (5e), it can be shown :

$$Y = \frac{v_c}{v^2 + a'^2} (2v + v_c) \qquad ... \quad (6a)$$

$$Y' = \frac{\mathsf{v}_{\mathfrak{c}}}{\mathsf{v}^{2} + p'^{2}} (\mathsf{v}_{\mathfrak{c}} - 2\mathsf{v}) \qquad \qquad \dots \quad (6b)$$

Using Eqs. (5a) and (5b) and the relation, $\rho_0 \rho_x = 1$,

$$\cot \phi_0 = \pm \frac{d_2}{d_1} \sqrt{1 - \frac{1}{a^2}}$$
 ... (7a)

When $v << v_c$, we get from (6a), (6b) and (5c),

$$Y = Y' \simeq \frac{v_e^2}{v^2 + p'^2}$$
 ... (6c)

$$a^2 \simeq \frac{p'^2 + v_c^2}{p'^2 + v^2}$$
 ... (6d)

Using (6d) and (7a)

$$\tan \phi_0 \simeq \pm \frac{p'}{\nu} \sqrt{1 + \frac{p'^2}{\nu_e^2}}$$
 ... (7b)

According to Murty and Khastgir (1960), ϕ_0 lies in the first quadrant for the vertically downcoming wave in the northern hemisphere, hence taking the positive sign in (7b), we have

$$\tan \phi_0 = \frac{p'}{v} \sqrt{1 + \frac{p^{\frac{1}{2}}}{v_e^2}}$$
 ...(7c)

Similarly, for the vertically downcoming radio wave in the southern hemisphere:

$$\tan \phi_0 \simeq -\frac{p'}{\nu} \sqrt{1 + \frac{p'^2}{\nu_c^2}}$$
 ... (7d)

From Eq. (7e), we can draw several conclusions .

- (a) At the level $p_0^2 = p^2$ (i.e. p' = 0), the wave is plane-polarised.
- (b) The sense of rotation of the magnetic vector is reversed when the wave crosses the level $p_0^2 = p^2$.

- (c) Since at the lower boundary of the ionosphere, p'=p and at the level, $p_0^2=p^2$, p'=0, the phase-difference gradually increases from zero to a certain value depending on p, ν , and ν_e as the wave comes down from the level, $p_0^2=p^2$, to the lower boundary of the ionosphere.
 - (d) At the level $p_0^2 = p^2 pp_H$,

$$\left[\tan\phi_{0}\right]_{p_{0}^{2}=p^{2}-pp_{_{H}}}=\frac{p_{_{H}}}{\mathsf{v}}\left[\cot^{2}\theta+\csc^{2}\theta\right]$$

 Effect of electron collisional frequency on the amplitude-ratio at any level for v << v_e.

It can be shown from the Appleton-Hartree formulae (1927, 1929) that the amplitude-ratio for the ordinary mode for vertically down-coming radio-wave in the northern hemisphere for zero collisional frequency is given by

$$\rho_0 = -\frac{v_c}{p'} + \sqrt{1 + \frac{v_c^2}{p'^2}} \tag{8}$$

where

$$p'=p\Big(1-rac{{p_0}^2}{p^2}\Big)$$
 , $u_c=p_H\,\sin^2\! heta/2\,\cos heta$

From Eq. (8), it can be shown.

$$\left[\frac{1+\rho_0^2}{1-\rho_0^2}\right]_{\nu=0} = \frac{\sqrt{\nu_e^2 + p'^2}}{\nu_e} \tag{9}$$

It has been shown by Murty and Khastgir (1959) that

$$\rho_0 = \frac{a'}{v_a} \left[\frac{v}{\cos \phi_n} - \frac{p'}{\sin \phi_n} \right] \tag{10a}$$

and

$$\rho_z = \frac{a'}{v_a} \left[\frac{v}{\cos \phi_0} + \frac{p'}{\sin \phi_0} \right] \tag{10b}$$

where

$$a' = \frac{r_c}{v^2 + p'^2}$$

Using $\rho_0 \rho_x = 1$, we have from Eqs. (10a), (10b)

$$\rho_0^2 = \frac{\frac{\mathsf{v}}{\cos\phi_0} - \frac{p'}{\sin\phi_0}}{\frac{\mathsf{v}}{\cos\phi_0} + \frac{p'}{\sin\phi_0}} \qquad \dots \tag{10e}$$

Hence from Eq. (10c)

$$\left[\frac{1+\rho_0^2}{1-\rho_0^2}\right]_{\nu=0} - \frac{\nu}{p'} \tan \phi_0 \qquad ... (11)$$

Using (11), (7c),

$$\begin{bmatrix} 1+\rho_0^2\\ 1-\rho_0^{-2} \end{bmatrix} \qquad \qquad \sqrt{\mathsf{v}_c^2+} p'^2 \qquad \qquad \dots \quad (11a)$$

and from (11a) and (9).

$$[\rho_0]_{\nu \neq 0} \simeq [\rho_0]_{\nu = 0}$$

This means that at any level of the ionosphere, the amplitude-ratio for the O-mode is nearly unaffected by electron collisional frequency provided the critical collisional frequency is much larger than electron-collisional frequency. The same conclusion can be drawn for X-mode also.

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