

## INTEGRAL detects the recurrent transients XTE J1709-267 and XTE J1739-285 in outburst

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*Published in:*  
The Astronomer's Telegram : ATel

*Publication date:*  
2012

*Document Version*  
Publisher's PDF, also known as Version of record

[Link back to DTU Orbit](#)

*Citation (APA):*  
Sanchez-Fernandez, C., Chenevez, J., Pavan, L., Bozzo, E., Cadolle Bel, M., Natalucci, L., ... Watanabe, K. (2012). INTEGRAL detects the recurrent transients XTE J1709-267 and XTE J1739-285 in outburst. The Astronomer's Telegram : ATel, [ATel #4304].

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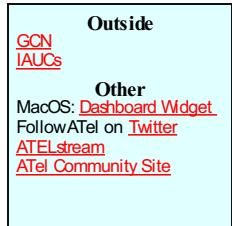
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## INTEGRAL detects the recurrent transients XTE J1709-267 and XTE J1739-285 in outburst

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on 10 Aug 2012; 21:56 UT

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Subjects: X-ray, Binary, Transient

The neutron star, recurrent X-ray transients XTE J1709-267 (=RX J1709.5-2639) and XTE J1739-285, both of which are known X-ray bursters, were detected in outburst during INTEGRAL observations of the Galactic Center region (PI, J. Wilms), carried out between (UT) 2012 August 08, 07:18 - 19:05.

XTE J1709-267 was seen by JEM-X at the 54-sigma level in the 3-10 keV band, with an average flux of 141 +/- 3 mCrab, and at the 27-sigma level in the 10-25 keV band, with an average flux of 61 +/- 7 mCrab. The effective JEM-X exposure time was 4 ks. XTE J1709-267 was only marginally detected by ISGRI in the 20-40 keV band. We estimate a rough flux of ~8 mCrab (5 sigma) for an effective ISGRI exposure time of 21 ks. A preliminary spectral fit to the JEM-X data is consistent with an absorbed disk blackbody model of temperature ~2.2 keV. The onset of this outburst was also detected by MAXI <http://maxi.riken.jp/top/index.php?cid=1&jname=J1709-266>

This new outburst from XTE J1709-267 occurs 2 years after the previous period of activity of the source, reported by MAXI/GSC (ATEL #[2729](#)), and is consistent with a outburst recurrence time of 2-3 years, claimed in ATel #[1302](#).

XTE J1739-285 was seen by JEM-X at the 9-sigma level in the 3-10 keV band, with an average flux of 16 +/- 2 mCrab, and at the 8-sigma level in the 10-25 keV band, with an average flux of 16 +/- 4 mCrab, for an effective exposure time of 10 ks. A weak (short) type-I X-ray burst was detected during these observations. The source flux in the 20-40 keV band was below the ISGRI sensitivity limit.

The last outburst from XTE J1739-285, reported by INTEGRAL (ATEL #[734](#)) was detected in 2006.

XTE J1709-267 and XTE J1739-285 will be observed regularly by INTEGRAL in the coming weeks during Galactic Center and Galactic Bulge monitoring observations.

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