#### Technical University of Denmark



#### An isotope view on ionising radiation as a source of sulphuric acid

Enghoff, Martin Andreas Bødker; Bork, Nicolai Christian; Hattori, S.; Meusinger, C.; Nakagawa, M.; Pedersen, Jens Olaf Pepke; Danielache, S.; Ueno, Y.; Johnson, M. S.; Yoshida, N.; Svensmark, Henrik *Published in:* Atmospheric Chemistry and Physics Discussions

Link to article, DOI: 10.5194/acpd-12-5039-2012

Publication date: 2012

Document Version Publisher's PDF, also known as Version of record

Link back to DTU Orbit

Citation (APA):

Enghoff, M. A. B., Bork, N. C., Hattori, S., Meusinger, C., Nakagawa, M., Pedersen, J. O. P., ... Svensmark, H. (2012). An isotope view on ionising radiation as a source of sulphuric acid. Atmospheric Chemistry and Physics Discussions, 12(2), 5039-5064. DOI: 10.5194/acpd-12-5039-2012

#### DTU Library Technical Information Center of Denmark

#### **General rights**

Copyright and moral rights for the publications made accessible in the public portal are retained by the authors and/or other copyright owners and it is a condition of accessing publications that users recognise and abide by the legal requirements associated with these rights.

• Users may download and print one copy of any publication from the public portal for the purpose of private study or research.

- You may not further distribute the material or use it for any profit-making activity or commercial gain
- You may freely distribute the URL identifying the publication in the public portal

If you believe that this document breaches copyright please contact us providing details, and we will remove access to the work immediately and investigate your claim. Atmos. Chem. Phys. Discuss., 12, 5039–5064, 2012 www.atmos-chem-phys-discuss.net/12/5039/2012/ doi:10.5194/acpd-12-5039-2012 © Author(s) 2012. CC Attribution 3.0 License.



This discussion paper is/has been under review for the journal Atmospheric Chemistry and Physics (ACP). Please refer to the corresponding final paper in ACP if available.

# An isotope view on ionising radiation as a source of sulphuric acid

M. B. Enghoff<sup>1</sup>, N. Bork<sup>1,2</sup>, S. Hattori<sup>3</sup>, C. Meusinger<sup>4</sup>, M. Nakagawa<sup>5</sup>, J. O. P. Pedersen<sup>1</sup>, S. Danielache<sup>3,6</sup>, Y. Ueno<sup>6</sup>, M. S. Johnson<sup>4</sup>, N. Yoshida<sup>3,5</sup>, and H. Svensmark<sup>1</sup>

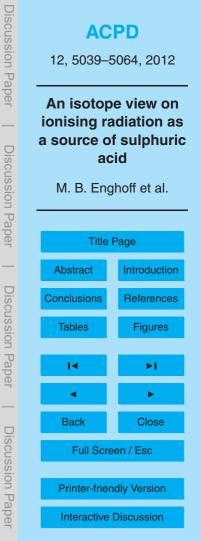
<sup>1</sup>National Space Institute, Technical University of Denmark, 2100, Copenhagen Ø, Denmark <sup>2</sup>Division of Atmospheric Science, Department of Physics, P.O. Box 64, 00014 University of Helsinki, Finland

 <sup>3</sup>Department of Environmental Science and Technology, Interdisciplinary Graduate School of Science and Engineering, Tokyo Institute of Technology, Yokohama, 226-8502, Japan
 <sup>4</sup>University of Copenhagen, Department of Chemistry, 2100, Copenhagen Ø, Denmark
 <sup>5</sup>Department of Environmental Chemistry and Engineering, Interdisciplinary Graduate School of Science and Engineering, Tokyo Institute of Technology, Yokohama, 226-8502, Japan
 <sup>6</sup>Department of Earth and Planetary Science, Tokyo Institute of Technology, Meguro-ku, Tokyo, 152-8551, Japan

Received: 17 November 2011 – Accepted: 31 January 2012 – Published: 13 February 2012

Correspondence to: M. B. Enghoff (enghoff@space.dtu.dk)

Published by Copernicus Publications on behalf of the European Geosciences Union.





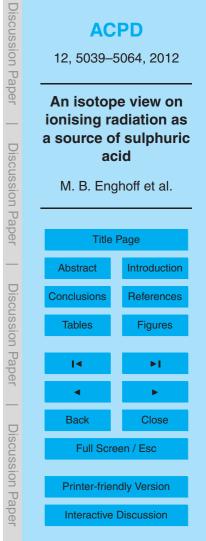
#### Abstract

Sulphuric acid is an important factor in aerosol nucleation and growth. It has been shown that ions enhance the formation of sulphuric acid aerosols, but the exact mechanism has remained undetermined. Furthermore some studies have found a deficiency

- in the sulphuric acid budget, suggesting a missing source. In this study the production of sulphuric acid from SO<sub>2</sub> through a number of different pathways is investigated. The production methods are standard gas phase oxidation by OH radicals produced by ozone photolysis with UV light, liquid phase oxidation by ozone, and gas phase oxidation initiated by gamma rays. The distributions of stable sulphur isotopes in the products and substrate were measured using isotope ratio mass spectrometry. All methods produced sulphate enriched in <sup>34</sup>S and we find a δ<sup>34</sup>S value of 8.7 ± 0.4 ‰ (1 standard deviation) for the UV-initiated OH reaction. Only UV light (Hg emission at 253.65 nm) produced a clear non-mass-dependent excess of <sup>33</sup>S. The pattern of iso-
- topic enrichment produced by gamma rays is similar, but not equal, to that produced by aqueous oxidation of SO<sub>2</sub> by ozone. This, combined with the relative yields of the experiments, suggests a mechanism in which ionising radiation may lead to hydrated ion clusters that serve as nanoreactors for S(IV) to S(VI) conversion.

#### 1 Introduction

Aerosols – small particles suspended in air – are frequently formed by condensation
 of gas molecules in Earth's atmosphere. The effect of these particles on the radiative balance of our climate system, through direct and indirect mechanisms, is the greatest uncertainty in the radiative forcing budget (Forster et al., 2007). One issue is that the mechanism behind the formation of aerosols remains unknown, even though several mechanisms have been proposed (Kulmala, 2003; Curtius, 2006; Kurtén et al., 2008) and major advances have recently been made (Kirkby et al., 2011). One of



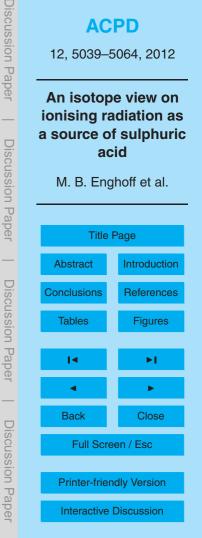


these mechanisms is that of ion-induced nucleation (Enghoff and Svensmark, 2008;

Kazil et al., 2008; Hirsikko et al., 2011) where ionising radiation promotes the early stages of the aerosol nucleation. The prevailing theory on exactly how the ions affect nucleation is that they lower the energy barrier for condensing molecules caused by the increase in energy required to create the interface between the aerosol phase and

- the gas phase the surface tension (Lovejoy et al., 2004). Other factors include the recombination of oppositely charged clusters (Yu and Turco, 2000) and the increased growth rate caused by a charge in a molecular cluster (Yu and Turco, 2001). Ionic cluster formation is estimated to account for 1–30 % of total particle formation, based on 7 European measurement sites (Manninen et al., 2010) and ionisation by galactic cosmic
- rays may increase binary (sulphuric acid and water) and ternary (sulphuric acid, water, ammonia) nucleation rates of nanometer size particles in the tropospheric boundary layer (Kirkby et al., 2011). Svensmark et al. (2007) proposed a chemical mechanism where sulphuric acid is produced by ion chemistry, but until now there has been no experimental evidence supporting this mechanism.
- <sup>15</sup> Sulphuric acid is thought to be one of the main components of aerosols nucleated in the atmosphere (Sipilä et al., 2010) and an extra source might therefore be important for understanding aerosol formation. During a measurement campaign in Hyytiälä, Finland the calculated sulphuric acid concentrations were about 10 percent lower than the measured values (Boy et al., 2005), and a study at the Antarctic also showed a
   <sup>20</sup> missing sulphuric acid source (Jefferson et al., 1998) stimulating the search for possible
- <sup>20</sup> missing sulphuric acid source (Jefferson et al., 1998) stimulating the search for possible additional sources of sulphate. An ion mechanism could additionally operate at night, providing a source of sulphuric acid not depending on sunlight.

The initial steps of a mechanism involving ion chemistry have been investigated by e.g. Fehsenfeld and Ferguson (1974) and more recently by Bork et al. (2011a,b) who used ab inito calculations to show that the gas-phase oxidation of SO<sub>2</sub> by ozone is facilitated in ionic clusters, but until now there has been no direct experimental investigation. One of the difficulties of such a study is that nucleation requires a certain amount of condensable gas and thus either a liquid or photochemical source of sulphuric acid is normally used (Berndt et al., 2006). Any effect from an ionising source then occurs in



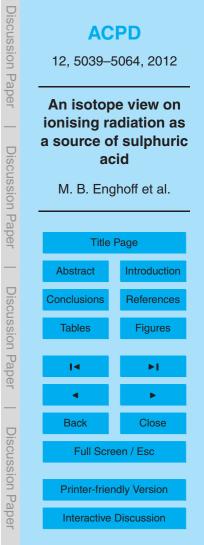


parallel with this other source and is thus hard to distinguish. An early study by Raes et al. (1985) showed an increase in particle number when a gamma source was used in addition to photochemical sulphuric acid production and they speculated that both UV light and gamma rays could produce sulphuric acid from SO<sub>2</sub>.

- An alternative technique to distinguish the processes an airmass undergoes is that of isotopic analysis. Fractionating processes alter the isotopic composition of chemical species and are used to constrain atmospheric budgets or interpret geological records and ice cores, e.g. Johnson et al. (2002); Brenninkmeijer et al. (2003). Traditional physical processes and most chemical reactions follow a so-called mass-dependent rela-
- tionship (MDF, mass-dependent fractionation), where enrichments of different isotopes follow a linear relationship according to the ratio of their masses. A deviation from the linear behavior is sometimes observed, but the origin(s) of non-mass-dependent fractionation, abbreviated NMD, is not fully understood. NMD is, however, known to be produced by a small number of mechanisms occuring in the gas phase including pho-
- tolysis and recombination reactions, as well as some nuclear processes (Guo et al., 2010; Romero and Thiemens, 2003; Thiemens, 2006). Photolytic reactions can result in NMD due to differences in absorption cross sections caused by isotopic substitution (Danielache et al., 2008; Schmidt et al., 2011). Sulphur has four stable isotopes (<sup>32</sup>S, <sup>33</sup>S, <sup>34</sup>S, and <sup>36</sup>S) and is widely used for iso-

Sulphur has four stable isotopes (<sup>32</sup>S, <sup>33</sup>S, <sup>34</sup>S, and <sup>36</sup>S) and is widely used for isotopic analysis. It is known that the photolytic fractionation of sulphur compounds is wavelength dependent (Farquhar et al., 2000b) and a study by Farquhar et al. (2001) showed that NMD signals can be produced by photolysing SO<sub>2</sub> at low wavelenghts (<220 nm) and that the type of UV lamp had significant impact on the isotopic composition. Another possible source for NMD is the oxidation channel of excited SO<sup>\*</sup><sub>2</sub>, a

<sup>25</sup> process very likely to happen in the atmosphere. NMD signatures have been found in atmospheric sulphate aerosols (Baroni et al., 2007; Pavlov et al., 2008) and Archean rock samples (Farquhar et al., 2000a; Ono et al., 2003; Ueno et al., 2008; Bao et al., 2007). In a previous study, Harris et al. (2012) investigated different gas and liquid phase oxidation pathways of SO<sub>2</sub> in laboratory experiments, but did not observe clear





NMD with respect to <sup>33</sup>S, while the sulphate production processes under consideration were generally distinguishable by their <sup>34</sup>S fractionation factors.

While the isotopic fractionation of ionising radiation is not known, NMD is not expected. From the Born-Oppenheimer approximation it is evident that molecular polarizability is not changed by isotopic substitution. In particular the S in SO<sub>2</sub> is at a symmetric site and it has previously been shown, for ozone, that NMD effects are primarily associated with assymetric sites (e.g. the two end atoms in ozone) but not at symmetric sites, such as the central atom in ozone (Gao and Marcus, 2001).

In this study we generate sulphuric acid using UV light, gamma sources, and via ozone oxidation in the aqueous phase. Using multiple sulphur isotopic measurements allows us to distinguish between these reactions and evaluate their relative importance.

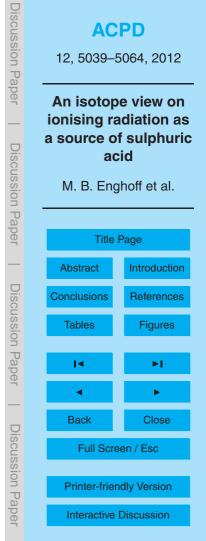
#### 2 Experimental setup and procedures

The experiments were performed using two separate systems. Sulphuric acid was generated using a chemical reaction chamber at the National Space Institute in Copenhagen. Samples were collected as BaSO<sub>4</sub> by flowing the air leaving the reactor through a series of bubblers containing BaCl<sub>2</sub>. The isotopic composition of the samples were then analysed with an isotope ratio mass spectrometer (IRMS) at the Tokyo Institute of Technology. The following sections describe each experimental system and the procedures employed.

#### 20 2.1 Chamber setup (Copenhagen)

25

A schematic of the setup can be seen in Fig. 1. The reaction volume was contained in a cylindrical 501 reaction chamber (100 cm long with a diameter of 25 cm) made of electropolished stainless steel, previously described by Enghoff et al. (2008). A mixture of pure humidified air (Labline 5.0, Strandmøllen, further cleaned with a Thermo Scientific model 1160 Zero Air Supply) was continuously flowed through the vessel at a





rate of  $1.01 \text{ min}^{-1}$ . Previous to entering the chamber  $0.41 \text{ min}^{-1}$  of the purified air went through an ozone generator (using photolysis at 184 nm) resulting in about ~400 ppb ozone in the chamber. The rest of the air ( $0.61 \text{ min}^{-1}$ ) was humidified by circulating purified (milli-Q) water through a GoreTex tube inserted into the air stream, giving a relative humidity of 40–45%. In addition  $100 \,\mu \text{lmin}^{-1}$  of pure SO<sub>2</sub>(I) evaporated from a container (Gerling Holz+Co, ≥99.95% wt.) was sent through the chamber. All air entering the chamber was additionally cleaned with a PALL Kleenpack 3 nm particle filter.

The temperature and absolute pressure was that of the laboratory, typically 22°C (with UV lamps off) and 30°C (UV on) and 1 bar, with a 1 mbar overpressure in the reaction vessel. Temperature, absolute and differential pressure, and relative humidity were monitored.

An array of UV (Philips TUV PL-L 18 W mercury discharge) lamps were used to illuminate the chamber with UV-radiation primarily at 253.65 nm (the lamp spectrum is shown in the Supplement). The lamps were placed about 20 cm from the chamber, separated by 2 layers of Teflon foil to maintain an air-tight system but transmit UV. The UV light initiates photochemistry in the gas phase, where ozone is photolysed producing excited oxygen atoms that form OH via reaction with H<sub>2</sub>O, initiating the oxidation of SO<sub>2</sub> into sulphuric acid (see Appendix C).

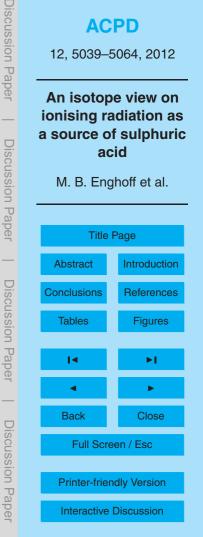
20

25

15

Alternatively the chamber could be exposed to gamma radiation from two 35 MBq Cs-137 sources, placed 27 cm from the side of the chamber.

The outlet air from the chamber was connected to a series of glass gas washing bottles (typically 3) used to collect the sulphuric acid. These bottles contained 400 ml of milli-Q water adjusted to pH 2 using hydrochloric acid with an excess (1.0 g) of barium chloride (BaCl<sub>2</sub>(H<sub>2</sub>O)<sub>2</sub>) compared to the amounts of sulphate formed in the experiments (milligrams). The bottles were magnetically stirred to maximize the airwater interaction. When the sulphuric acid entered the bottles it reacted with the barium ions to form BaSO<sub>4</sub> which has a very low solubility in water. The pH of 2 was chosen





to prevent SO<sub>2</sub> from the gas phase from dissolving in water forming the sulphite ion

 $(SO_3^{2-})$ , which could potentially precipitate as BaSO<sub>3</sub>. Acidic conditions at pH 2 shift the equilibrium away from  $SO_3^{2-}$ , instead favoring aqueous SO<sub>2</sub> and HSO<sub>3</sub>.

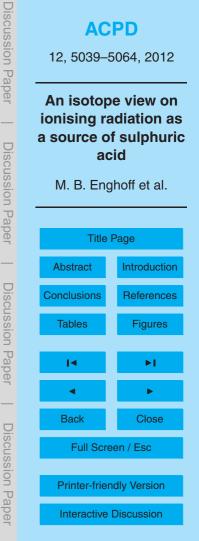
All experiments ran for one week. The sample was then collected by allowing the BaSO<sub>4</sub> to settle, decanting the excess liquid and thoroughly resuspending the remain-<sup>5</sup> ing solids in milli-Q water. This procedure was repeated three times for each sample to remove any remaining BaCl<sub>2</sub>. The remaining solution was transferred to a smaller container and the water was evaporated in an oven, leaving the BaSO<sub>4</sub> as a powder.

The first column of Table 1 shows the experiments used for further analysis. Five different types of experiments were done, to distinguish oxidation pathways for  $SO_2$ .

- <sup>10</sup> UV light and the gamma source were used separately and together. In addition, a series of experiments were run without UV and gamma source to quantify any uptake and oxidation of SO<sub>2</sub> in the liquid phase. Finally, an experiment was run with no UV, no gamma source, and no ozone, which did not yield any detectable amount of sulphuric acid.
- The isotopic composition of the liquid SO<sub>2</sub> had to be determined, as a reference point for the isotopic data. About 5 ml of SO<sub>2</sub> (0.2 mmol) was added to a solution consisting of 500 ml milli-Q water with 15 mmol H<sub>2</sub>O<sub>2</sub>, as oxidizing agent. The solution was adjusted to pH 11 before SO<sub>2</sub> was added to ensure dissolution. The solution was stirred overnight and BaCl<sub>2</sub>(H<sub>2</sub>O)<sub>2</sub> was added. The resulting BaSO<sub>4</sub> was collected
  using the method described above. A second reference was prepared in the same fashion, except pH was not adjusted, in order to test for any pH dependence, even though the reaction is supposedly almost pH-independent (Seinfeld and Pandis, 2006, chap. 7).

#### 2.2 Isotope fractionation setup (Tokyo)

<sup>25</sup> The collected BaSO<sub>4</sub> was converted to SF<sub>6</sub> for isotopic analysis. The first step was conversion into Ag<sub>2</sub>S: the BaSO<sub>4</sub> was held at 300 °C and reduced to H<sub>2</sub>S with the modified Kiba reagent (concentrated, dehydrated phosphoric acid containing Sn<sup>2+</sup>, Sakai





et al., 1984). The generated H<sub>2</sub>S was then carried by N<sub>2</sub>(g), washed through a water trap, and reacted with a 0.1 M AgNO<sub>3</sub> solution to convert it into Ag<sub>2</sub>S. The Ag<sub>2</sub>S was then converted into SF<sub>6</sub> by reaction with excess F<sub>2</sub> at 200 °C in a nickel reaction tube and then purified by gas chromatography using a technique similar to that of Ono et al. (2006). Additionally a separate setup with an improved purification system was used to analyze the samples for <sup>36</sup>S with better precision. The isotopic composition of SF<sub>6</sub> was determined using a Finnigan MAT 253 mass spectrometer (Thermo Fishier Scientific) equipped with a dual inlet system. The SF<sub>5</sub><sup>+</sup> ions with mass 127, 128, 129, and 131 amu were measured, corresponding to <sup>32</sup>S, <sup>33</sup>S, <sup>34</sup>S, and <sup>36</sup>S.

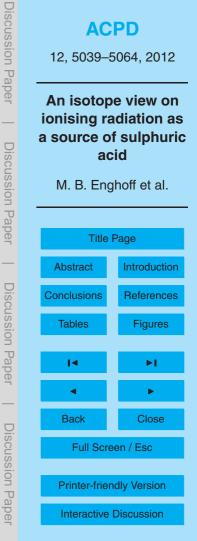
#### 10 3 Results and discussion

The results of the analysis are presented in Table 1. Definitions for the isotopic designations are given in Appendix B. Columns three and four of the table show  $\delta^{34}$ S and  $\Delta^{33}$ S for all samples. Data for  $\Delta^{36}$ S was also measured and is presented in Appendix A. The fractionation data of  $\delta^{34}$ S and  $\Delta^{33}$ S is presented graphically in Fig. 2.

The two references yielded 41.8 and 36.7 mg of BaSO<sub>4</sub>, respectively, from the ~5 ml of SO<sub>2</sub>. This corresponds to 0.18 and 0.16 mmol. The collected amounts suggest full conversion of the SO<sub>2</sub>, considering the uncertainty of determining the volume of SO<sub>2</sub> transferred and losses of BaSO<sub>4</sub> from sample collection. We attempted to dissolve a small part of the first reference with 4 M HCl. It did not dissolve, demonstrating that
 the product was not BaSO<sub>3</sub>. Samples from other experiments were also subjected to 4 M HCl, yielding similar results.

#### 3.1 Sulphate production

The sulphate mass obtained from each bottle is listed in the second column of Table 1. Sulphate is generated in the setup in all different operation modes and yields vary widely – for instance more sulphuric acid is formed with UV light present than for



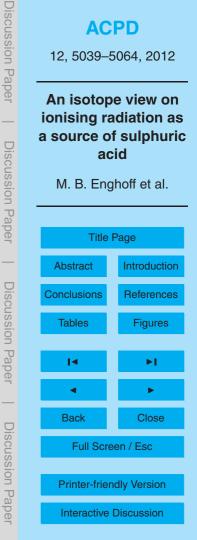


the gamma source but the sources combined yield less than each individual source. Seeing sulphate in the samples where UV and gamma radition was not used indicates a process occurring in the trapping bottles themselves. The experiment without ozone, which showed no precipitation, strongly indicates that this process is liquid phase ox-

- <sup>5</sup> idation of SO<sub>2</sub> by ozone where the reactants become dissolved in the liquid phase depending on their Henry's Law constants, leading to the conversion of SO<sub>2</sub> to  $H_2SO_4$  and not simply dissolved SO<sub>2</sub> precipitating as BaSO<sub>3</sub>. Note that the yields in the experiments without UV and gamma radiation increase for each bottle counting from the chamber.
- <sup>10</sup> We conclude that while the liquid phase ozone oxidation process takes place in all experiments, the amount of sulphate produced via this process should be small in the first bottle, compared to the amount produced in experiments that had other formation processes present – any effect this would have on the  $\delta^{34}$ S values is less than one standard deviation. Additionally the concentration of ozone is substantially smaller (about half) in the UV experiments, due to photolysis.

### 3.2 <sup>34</sup>S

The  $\delta^{34}$ S values of our sample O<sub>3</sub>-UV-1 (8.7 ± 0.4 ‰) can be compared to the recently published values by Harris et al. (2012) for the OH reaction. In the Harris et al. study and this work, UV photolysis was used to generate OH via ozone photolysis, in the first case prior to the reaction chamber but in the second case within the reaction chamber – see Sect. 3.3. Their <sup>34</sup>  $\alpha$  value of 1.0077 ± 0.0022 at 30 °C is easily converted to our  $\delta^{34}$ S notation by subtracting 1, giving 7.7 ± 2.2 ‰, which appears to agree with our result. The  $\delta^{34}$ S of the OH reaction has also been estimated by Leung et al. (2001), using RRKM transition-state theory and by Tanaka et al. (1994) using ab initio calcula-tions. They do, however, arrive at very different values: 140 and -9.0 ‰, respectively. While the Leung et al. result shows enrichment in <sup>34</sup>S, as in this study, the effect is much larger than what we measure.





Harris et al. (2012) also measured liquid phase oxidation, but it is not obvious that we can make a direct comparison to our experiments without UV and gamma radiation. It is clear that not all sulphuric acid is collected in the first bottle in the line as we get more material in the second and third bottles and thus it is not possible to distinguish

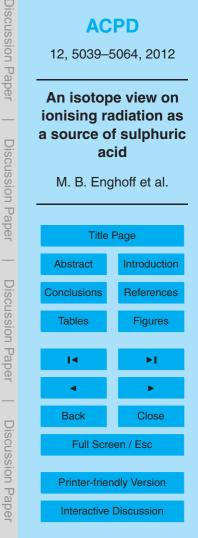
the fractionation for the liquid phase process. However we note that the δ<sup>34</sup>S value for our sample O<sub>3</sub>-1 (16.0 ± 0.4 ‰) agrees with the value at 22 °C from the Harris et al. (2012) study of 14.8 ± 2.7 ‰. The Harris et al. (2012) results are also shown in Fig. 2. Note that Harris et al. (2012) find a Δ<sup>33</sup>S value of ~8‰ for the SO<sub>2</sub> + O<sub>3</sub> reaction, but this "may be a measurement artefact as only two samples were measured" (Harris et al., 2012). Also, their reported uncertainties for the <sup>33</sup>S values of the SO<sub>2</sub> + OH reaction are quite large due to counting statistics. Therefore we report their data in Fig. 2 with errors on the Δ<sup>33</sup>S scale exceeding the depicted y-axis.

The samples  $O_3$ -1 and  $O_3$ -3 show a difference in  $\delta^{34}$ S which can be explained by the isotopic composition of the substrate of the bottles which differs when operated in series, i.e. the first bottle leaves depleted substrate to be trapped in later bottles, so that  $\delta^{34}$ S values are shifted to lower numbers.

As indicated by the "Isotopic Mixing line" in Fig. 2, we note that the values of both  $\delta^{34}$ S and  $\Delta^{33}$ S of sample O<sub>3</sub>-UV- $\gamma$ -1 lie between the values for samples O<sub>3</sub>-UV-1 and O<sub>3</sub>- $\gamma$ -1, indicating that in this case both processes are happening at the same time. Assuming that sample O<sub>3</sub>-UV- $\gamma$ -1 is a combination of the two processes, we find, using a linear combination of the  $\delta^{34}$ S values, that 79 % of sample O<sub>3</sub>-UV- $\gamma$ -1 originates from the UV process and 21 % from the  $\gamma$  process, ignoring ozone oxidation.

#### 3.3 Non-mass dependent fractionation

NMD was observed in our samples, especially for the UV experiment (O<sub>3</sub>-UV-1). This stands in contrast to the mass dependent signal reported by Harris et al. (2012), however their sensitivity to <sup>33</sup>S is not very high and their experiment was not directly irradiated with UV. NMD signals in SO<sub>2</sub> mainly arise from differences in the absorption spectra of isotopologues (Danielache et al., 2008) and in this experiment the reaction





volume, which contains a high  $SO_2$  concentration (0.01 %), is directly exposed to light at 253.65 nm, which was not the case in the Harris et al. experiment. These conditions make a host of new reactions possible (Ueno et al., 2009), the most important ones being

$$_{5}$$
 SO<sub>2</sub> +  $hv \rightarrow$  SO<sub>2</sub><sup>\*</sup> (R1

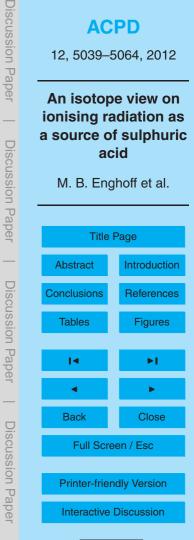
$$SO_2^* + O_2 \rightarrow SO_3 + O, \quad k = 1 \times 10^{-16} \text{ cm}^3 \text{ s}^{-1}$$
 (R2)

$$SO_2^* + SO_2 \rightarrow SO_3 + SO$$
,  $k = 4 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$ 

$$SO_2^* \to SO_2, \quad k=2.2 \times 10^4 \, \text{s}^{-1}$$
 (R4)

$$SO_2^* + M \rightarrow SO_2 + M$$
,  $k = 1 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$ 

Note that SO<sub>2</sub><sup>\*</sup> can denote both a singlet and triplet state, where the singlet can 10 be relaxed into the triplet state. The shown reaction rates are for the singlet state but similar relations occur for the triplet. The absorption cross section for SO2 at 253.65 nm is  $\sim 1 \times 10^{-19}$  cm<sup>2</sup> (Danielache et al., 2008), and the concentration of SO<sub>2</sub> is 0.01%. The ozone cross section is  $\sim 1 \times 10^{-17}$  cm<sup>2</sup> at 253.65 nm and we have ~400 ppb ozone in the reaction volume. At 40–45 % RH, 0.2 of each O(<sup>1</sup>D) from Eq. (R7) (Appendix C) becomes OH (Seinfeld and Pandis, 2006, chap. 6). The photon flux and pathlength for the light is the same in both cases so relative to each other we have  $1 \times 10^{-19} \text{ cm}^2 \times 1 \times 10^{-4} = 1 \times 10^{-23} \text{ cm}^2$  for the SO<sub>2</sub> excitation pathway and  $1 \times 10^{-17}$  cm<sup>2</sup> × 4 × 10<sup>-7</sup> × 0.2 = 8 × 10<sup>-25</sup> cm<sup>2</sup> for ozone. This slightly favours the excitation pathway, starting with Reaction (R1). However, the excited SO<sub>2</sub><sup>\*</sup> can easily be 20 relaxed back to the ground state before it is oxidised to SO<sub>3</sub>. With 0.01 % SO<sub>2</sub> Reaction (R2) is the quickest pathway to SO<sub>3</sub> of Reactions (R2) and (R3), but the quenching reaction, Reaction (R5), is faster still, meaning that most of the excited SO<sub>2</sub> should be relaxed back to the ground state before reacting further, even at these highly elevated concentrations of SO<sub>2</sub>. 25



(R3)

(R5)

() BY Using the high resolution spectra obtained by Danielache et al. (2008) we can make an estimate of the expected NMD signal due to photo-excitation. Using the cross section at 253.65 nm, we predict a  $\Delta^{33}$ S of 18.3 to 24.7 ‰ for SO<sup>\*</sup><sub>2</sub> compared to SO<sub>2</sub> depending on the linewidth of the excitation source. Using more recent data (Danielache,

- <sup>5</sup> 2012) we get 13.5–36.0 ‰. Since we only have a small conversion of our SO<sub>2</sub> reservoir this can be compared directly to our measured  $\Delta^{33}$ S of 0.30 ‰. This shows that the measured NMD can be explained by photo-excitation. Our  $\delta^{34}$ S result is thus obtained from a combination of the OH reaction and the UV excitation reaction. However, since our measured  $\Delta^{33}$ S is so small compared to the value expected from the excitation pathway most of the sulphuric acid production in the UV experiments most likely pro
  - ceeds by the OH pathway and thus justifies comparison with the results of Harris et al. (2012). Furthermore the lack of NMD in sample  $O_3$ -UV-3 is due to the liquid phase oxidation dominating.
- A different study by Farquhar et al. (2001) focused on NMD signals in SO<sub>2</sub> photolysis reactions and described systems with different light sources. Comparing our data to theirs shows the same trends (our lamp compares best, but not perfectly, to their KrF excimer laser), but one has to keep in mind that minute changes in the irradiance may shift isotopic effects drastically (Danielache et al., 2008). Nevertheless, the  $\Delta^{36}$ S values reported here (see Appendix A) are pointing towards the negative, while the  $\Delta^{33}$ S values show a positive trend, just as in the Farquhar et al. (2001) study.

A small non-zero  $\Delta^{33}$ S is seen in samples O<sub>3</sub>-3, O<sub>3</sub>-UV-3, and O<sub>3</sub>- $\gamma$ -1, while sample O<sub>3</sub>-1 has a stronger signal. They are all significant with respect to the error, but we do not have a full understanding of the underlying process. In order to understand this effect a model of the heterogeneous reaction could be constructed similar to the evaluated reduction model of Page (1072) (and place laboration 2011).

<sup>25</sup> sulphate reduction model of Rees (1973) (see also Johnston, 2011).

	ACPD 12, 5039–5064, 2012 An isotope view on ionising radiation as a source of sulphuric acid M. B. Enghoff et al.			
	Title	Title Page		
-	Abstract	Introduction		
	Conclusions	References		
	Tables	Figures		
	I	۶I		
5	•	•		
5	Back	Close		
	Full Screen / Esc			
2	Printer-frie	Printer-friendly Version		
202	Interactive Discussion			



#### 3.4 Gamma experiments

lonising radiation gives energy to the electron clouds in molecules often generating ion pairs:

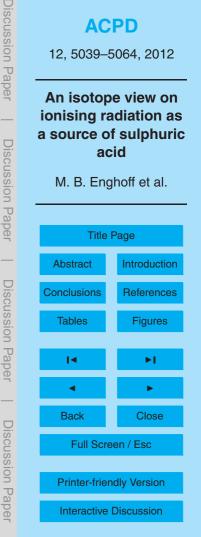
 $M + \gamma \rightarrow M^+ + e^-$ 

In our experiment, as well as the atmosphere, M is typically N<sub>2</sub>. The resulting electrons eventually attach to molecules, depending on their electron affinity and concentrations. Oxygen and nitrogen exist in the greatest abundance and since oxygen has a positive electron affinity the O<sub>2</sub><sup>-</sup> ion is spontaneously formed. The core ion then attracts water to form hydrated clusters, such as O<sub>2</sub><sup>-</sup>(H<sub>2</sub>O)<sub>n</sub> – this happens even at otherwise unsaturated conditions, due to the electric charge. There are many options for further reaction – one relevant scheme is presented by Bork et al. (2011a,b). In the troposphere ionisation is mostly caused by secondary particles (muons and electrons) formed by galactic cosmic rays with energies >1 GeV (Bazilevskaya et al., 2008). Although the energy of the gamma source (662 keV) is somewhat lower than that of cosmic rays penetrating

Based on the difference in  $\delta^{34}$ S, the mechanism for sulphate formation in the gamma reaction is clearly different from the OH oxidation pathway. The amount produced is also too high to be explained by this mechanism. The estimated OH production per ionisation is about 2 (Müller and Crutzen, 1993) and the ionisation in our chamber is estimated to be 1000 cm<sup>-3</sup> s<sup>-1</sup> (based on a simple model using the activity of the source

- <sup>20</sup> timated to be 1000 cm<sup>-3</sup> s<sup>-1</sup> (based on a simple model using the activity of the source and distance from the chamber to calculate the energy deposited in the volume), resulting in  $6 \times 10^{13}$  OH molecules formed during a run of one week, which would give 23 ng of BaSO<sub>4</sub>, orders of magnitude less than the 19.4 mg collected, even if a small part of this comes from the liquid phase oxidation process. The fractionation pattern
- resembles that of the experiments without UV and gamma radiation more even if they are some standard deviations apart, but in the gamma experiment we see much more material in the first bottle than in the following, the reverse pattern of what is seen in the experiments wihout UV and gamma radiation. This suggests that the mechanism

(R6)





of the gamma pathway is gas-phase or heterogenous chemistry in the chamber and by a mechanism that somehow enhances the yield far beyond one sulphate per ionisation. That we see a small NMD effect also suggests a chemical pathway. One possible explanation is the mechanism under investigation in Bork et al. (2011a,b) where hydrated

- ion clusters (Zatula et al., 2011), such as described above, serve as nanoreactors for 5 sulphur oxidation by a process similar to the oxidation of SO<sub>2</sub> by ozone observed in the experiments without UV and gamma radiation. We note that the SO<sub>2</sub> concentration in this experiment is many orders of magnitude above atmospheric levels which probably decreases the impact of this mechanism drastically. No enhanced sulphuric acid production is, for instance, reported in the recent experiment by Kirkby et al. (2011) even
- 10

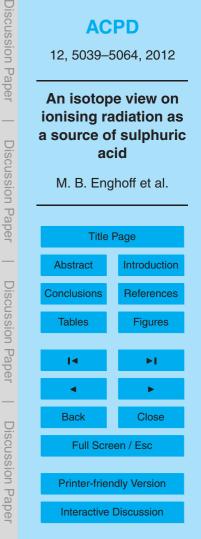
#### Conclusions 4

though they radiated an SO<sub>2</sub> containing volume.

In this study we investigated the production of sulphuric acid from SO<sub>2</sub> through a number of different pathways. The sulphuric acid was produced using standard gas phase oxidation by OH radicals from ozone photolysis, liquid phase oxidation by ozone, and 15 oxidation initiated by gamma rays. The distributions of stable sulphur isotopes in the products and substrate were measured using isotope ratio mass spectrometry. All methods produced sulphate enriched in <sup>34</sup>S and we find a  $\delta^{34}$ S value of 8.7 ± 0.4 ‰ (1 standard deviation) for the UV-initiated OH reaction. Only the experiment with UV light (Hg emission at 253.65 nm) produced a clear non-mass-dependent excess of <sup>33</sup>S 20 of around 0.3%, resulting from a small contribution from photo-excitation. Our results for UV photolysis and possibly aqueous oxidation by ozone, matched isotope patterns seen in earlier measurements. However, these UV photolysis result do not agree with theoretical studies (Tanaka et al., 1994; Leung et al., 2001), suggesting that additional

theoretical work is needed. 25

> The pattern of isotopic enrichment produced by gamma rays was distinctly different from that caused by the UV light and similar, but not equal, to that produced by the





aqueous oxidation of  $SO_2$ . This, combined with the relative yields of the experiments, suggests a mechanism by which ionising radiation may lead to hydrated ion clusters that serve as nanoreactors for S(IV) to S(VI) conversion.

#### Appendix A

## Δ<sup>36</sup>S

5

10

In addition to  $\delta^{34}$ S and  $\Delta^{33}$ S we also measured  $\Delta^{36}$ S, using an improved purification system for most of the samples. The values can be seen in Table 2. We refrain from going into a detailed analysis of this data but note that the linear relation between  $\Delta^{33}$ S and  $\Delta^{36}$ S is consistent with the isotope mixing shown in Fig. 2. Additionally, Table 2 displays further information on the amount of material used and recovered during the isotopic analysis.

#### Appendix B

#### 15 **Definitions**

Sulphur isotopic compositions are reported as  $\delta^{34}$ S and  $\Delta^{33/36}$ S values normalized against the mean value of our two SO<sub>2</sub> references:

 $\delta^{34}S = \frac{{}^{i}R_{sample}}{{}^{i}R_{reference}} - 1$ 

where *R* denotes the isotope ratios  $\left(\frac{{}^{34}S}{{}^{32}S}\right)$  of samples and reference.

<sup>20</sup> 
$$\Delta^{33}S = \delta^{33}S - [(\delta^{34}S + 1)^{0.515} - 1]$$
  
 $\Delta^{36}S = \delta^{36}S - [(\delta^{34}S + 1)^{1.90} - 1]$ 

Discussion Paper ACPD 12, 5039-5064, 2012 An isotope view on ionising radiation as a source of sulphuric **Discussion** Paper acid M. B. Enghoff et al. **Title Page** Introduction Abstract **Discussion** Paper Conclusions References **Tables Figures** 14 Back Close **Discussion** Paper Full Screen / Esc **Printer-friendly Version** Interactive Discussion



The  $\Delta$  notation describes the excess or deficiency of <sup>33</sup>S and <sup>36</sup>S relative to the reference mass-dependent fractionation array (Farquhar et al., 2000b).

The precision based on replicated analysis of Ag<sub>2</sub>S standards (IAEA-S1 and our house Ag<sub>2</sub>S) is typically ±0.3 and ±0.006‰ (1 standard deviation) for  $\delta^{34}$ S and  $\Delta^{33}$ S.

<sup>5</sup> For  $\Delta^{36}$ S the standard deviation is 0.7% for the method used to measure the other isotopes and 0.1% for the improved method. Measurements of the two SO<sub>2</sub> reference samples are close to the stated precisions, suggesting that errors originating from sample handling are minor.

#### Appendix C

#### 10

#### Gas-phase oxidation of $SO_2$ by OH

Here we list the reactions relevant to the oxidation of  $SO_2$  by OH in the chamber.

$O_3 + hv \rightarrow O(^1D) + O_2$	(R7)
-------------------------------------	------

 $O(^{1}D) + H_{2}O \rightarrow 2OH$ 

$$^{5}$$
 SO<sub>2</sub> + OH + M  $\rightarrow$  HSO<sub>3</sub> + M

 $HSO_3 + O_2 \rightarrow SO_3 + HO_2$ 

 $SO_3 + H_2O \rightarrow H_2SO_4$ 

Supplementary material related to this article is available online at: http://www.atmos-chem-phys-discuss.net/12/5039/2012/

<sup>20</sup> acpd-12-5039-2012-supplement.pdf.

**Discussion** Paper ACPD 12, 5039-5064, 2012 An isotope view on ionising radiation as a source of sulphuric **Discussion** Paper acid M. B. Enghoff et al. **Title Page** Introduction Abstract **Discussion** Paper Conclusions References **Tables Figures** 14 Back Close **Discussion** Paper Full Screen / Esc **Printer-friendly Version** Interactive Discussion

(R8)

(R9)

(R10)

(R11)



Acknowledgements. This work was supported by the Global Environment Research Fund (A-0904) of the Ministry of the Environment, and Grant-in-Aid for Scientific Research (S-23224013), The Ministry of Education, Culture, Sports, Science and Technology (MEXT), Japan. We thank the IntraMIF network sponsored by the European Communitys Seventh

Framework Programme (FP7/2007-2013) under grant agreement number 237890 for support. Martin B. Enghoff acknowledges the Carslberg Foundation and The Danish Council for Independent Research | Natural Sciences for financial support. Shohei Hattori is supported by Grant in Aid for JSPS Research Fellows (DC1, No. 22-7563) and Global COE program Earth to Earths of MEXT, Japan. Yuichiro Ueno is supported by the NEXT program of MEXT, Japan.

#### 10 References

20

Bao, H., Rumble, D., and Lowe, D. R.: The five stable isotope compositions of Fig Tree barites: Implications on sulfur cycle in ca. 3.2 Ga oceans, Geochim. Cosmochim. Ac., 71, 4868–4879, doi:10.1016/j.gca.2007.05.032, 2007. 5042

Baroni, M., Thiemens, M., Delmas, R., and Savarino, J.: Mass-Independent Sul-

- <sup>15</sup> fur Isotopic Compositions in Stratospheric Volcanic Eruptions, Science, 315, 84–87, doi:10.1126/science.1131754, 2007. 5042
  - Bazilevskaya, G. A., Usoskin, I. G., Flueckiger, E. O., Harrison, R. G., Desorgher, L., Buetikofer, R., Krainev, M. B., Makhmutov, V. S., Stozhkov, Y. I., Svirzhevskaya, A. K., Svirzhevsky, N. S., and Kovaltsov, G. A.: Cosmic ray induced ion production in the atmosphere, Space Sci. Rev.,
  - 137, 149–173, doi:10.1007/s11214-008-9339-y, 2008. 5051 Berndt, T., Böge, O., and Stratmann, F.: Formation of atmospheric H<sub>2</sub>SO<sub>4</sub>/H<sub>2</sub>O particles in the absence of organics: A laboratory study, Geophys. Res. Lett., 33, L15817,

doi:10.1029/2006GL026660, 2006. 5041 Bork, N., Kurtén, T., Enghoff, M. B., Pedersen, J. O. P., Mikkelsen, K. V., and Svensmark, H.: Ab

- initio studies of  $O_2^-(H_2O)n$  and  $O_3^-(H_2O)n$  anionic molecular clusters,  $n \le 12$ , Atmos. Chem. Phys., 11, 7133–7142, doi:10.5194/acp-11-7133-2011, 2011a. 5041, 5051, 5052
  - Bork, N., Kurtén, T., Enghoff, M. B., Pedersen, J. O. P., Mikkelsen, K. V., and Svensmark, H.: Structures and reaction rates of the gaseous oxidation of SO<sub>2</sub> by an O<sub>3</sub>-(H<sub>2</sub>O)<sub>05</sub> cluster – a density functional theory investigation, Atmos. Chem. Phys. Discuss., 11, 29647–29679, doi:10.5104/cond.11.00047.00114.00144.5011.5051
- <sup>30</sup> doi:10.5194/acpd-11-29647-2011, 2011b. 5041, 5051, 5052

Discussion Paper ACPD 12, 5039–5064, 2012 An isotope view on ionising radiation as a source of sulphuric **Discussion** Paper acid M. B. Enghoff et al. **Title Page** Introduction Abstract **Discussion** Paper Conclusions References **Tables Figures** Back Close **Discussion** Paper Full Screen / Esc **Printer-friendly Version** Interactive Discussion



863–878, doi:10.5194/acp-5-863-2005, 2005. 5041
Brenninkmeijer, C. A. M., Janssen, C., Kaiser, J., Röckmann, T., Rhee, T. S., and Assonov, S. S.: Isotope Effects in the Chemistry of Atmospheric Trace Compounds, Chem. Rev., 103, 5125–5162, doi:10.1021/cr020644k, 2003. 5042

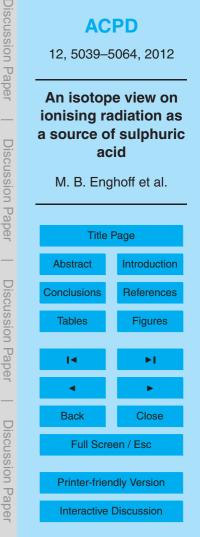
Boy, M., Kulmala, M., Ruuskanen, T. M., Pihlatie, M., Reissell, A., Aalto, P. P., Keronen, P., Dal Maso, M., Hellen, H., Hakola, H., Jansson, R., Hanke, M., and Arnold, F.: Sulphuric

acid closure and contribution to nucleation mode particle growth, Atmos. Chem. Phys., 5,

- Curtius, J.: Nucleation of atmospheric aerosol particles, CR Phys., 7, 1027–1045, doi:10.1016/j.crhy.2006.10.018, 2006. 5040
- <sup>10</sup> Danielache, S. O.: manuscript in preparation, 2012. 5050
- Danielache, S. O., Eskebjerg, C., Johnson, M. S., Ueno, Y., and Yoshida, N.: High-precision spectroscopy of <sup>32</sup>S, <sup>33</sup>S, and <sup>34</sup>S sulfur dioxide: Ultraviolet absorption cross sections and isotope effects, J. Geophys. Res., 113, D17314, doi:10.1029/2007JD009695, 2008. 5042, 5048, 5049, 5050
- Enghoff, M. B. and Svensmark, H.: The role of atmospheric ions in aerosol nucleation a review, Atmos. Chem. Phys., 8, 4911–4923, doi:10.5194/acp-8-4911-2008, 2008. 5040
  - Enghoff, M. B., Pedersen, J. O. P., Bondo, T., Johnson, M. S., Paling, S., and Svensmark, H.: Evidence for the Role of Ions in Aerosol Nucleation, J. Phys. Chem. A, 112, 10305–10309, doi:10.1021/jp806852d, 2008. 5043
- <sup>20</sup> Farquhar, J., Bao, H., and Thiemens, M.: Atmospheric Influence of Earth's Earliest Sulfur Cycle, Science, 289, 756–758, doi:10.1126/science.289.5480.756, 2000a. 5042
  - Farquhar, J., Jackson, T. L., and Thiemens, M. H.: A <sup>33</sup>S enrichment in ureilite meteorites: Evidence for a nebular sulfur component, Geochim. Cosmochim. Ac., 64, 1819–1825, doi:10.1016/S0016-7037(00)00356-2, 2000b. 5042, 5054
- Farquhar, J., Savarino, J., Airieau, S., and Thiemens, M. H.: Observation of wavelengthsensitive mass-independent sulfur isotope effects during SO<sub>2</sub> photolysis: Implications for the early atmosphere, J. Geophys. Res., 106, 32829–32839, doi:10.1029/2000JE001437, 2001. 5042, 5050

Fehsenfeld, F. C. and Ferguson, E. E.: Laboratory studies of negative ion reactions with atmo-

- <sup>30</sup> spheric trace constituents, J. Chem. Phys., 61, 3181–3193, doi:10.1063/1.1682474, 1974. 5041
  - Forster, P., Ramaswamy, V., Artaxo, P., Berntsen, T., Betts, R., Fahey, D. W., Haywood, J., Lean, J., Lowe, D. C., Myhre, G., Nganga, J., Prinn, R., Raga, G., Schulz, M., and Van Dorland, R.:

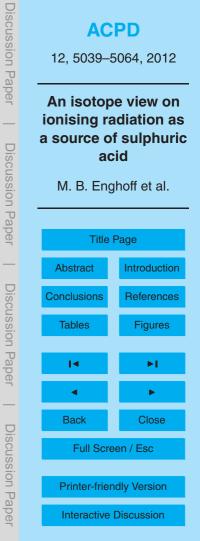




Changes in Atmospheric Constituents and in Radiative Forcing, in: Climate Change 2007: The Physical Science Basis. Contribution of Working Group 1 to the Fourth Assessment Report of the Intergovenrmental Panel on Climate Change, edited by: Solomon, S., Qin, D., Manning, M., Chen, Z., Marquis, M., Averyt, K. B., Tignor, M., and Miller, H. L., Cambridge University Press, 2007. 5040

#### University Press, 2007. 5040 Gao, Y. Q. and Marcus, R. A.: Strange and Unconventional Isotope Effects in Ozone Formation, Science, 293, 259–263, doi:10.1126/science.1058528, 2001. 5043

- Guo, Z., Li, Z., Farquhar, J., Kaufman, A. J., Wu, N., Li, C., Dickerson, R. R., and Wang, P.: Identification of sources and formation processes of atmospheric sulfate by sulfur iso-
- tope and scanning electron microscope measurements, J. Geophys. Res., 115, D00K07, doi:10.1029/2009JD012893, 2010. 5042
  - Harris, E., Sinha, B., Hoppe, P., Crowley, J. N., Ono, S., and Foley, S.: Sulfur isotope fractionation during oxidation of sulfur dioxide: gas-phase oxidation by OH radicals and aqueous oxidation by H<sub>2</sub>O<sub>2</sub>, O<sub>3</sub> and iron catalysis, Atmos. Chem. Phys., 12, 407–427, doi:10.5194/acp-12-407-2012, 2012. 5042, 5047, 5048, 5049, 5050, 5064
- 12-407-2012, 2012. 5042, 5047, 5048, 5049, 5050, 5064
   Hirsikko, A., Nieminen, T., Gagne, S., Lehtipalo, K., Manninen, H. E., Ehn, M., Horrak, U., Kerminen, V.-M., Laakso, L., McMurry, P. H., Mirme, A., Mirme, S., Petaja, T., Tammet, H., Vakkari, V., Vana, M., and Kulmala, M.: Atmospheric ions and nucleation: a review of observations, Atmos. Chem. Phys., 11, 767–798, doi:10.5194/acp-11-767-2011, 2011. 5041
- Jefferson, A., Tanner, D. J., Eisele, F. L., and Berresheim, H.: Sources and sinks of H<sub>2</sub>SO<sub>4</sub> in the remote Antarctic marine, J. Geophys. Res., 103, 1639–1645, doi:10.1029/97JD01212, 1998. 5041
  - Johnson, M. S., Feilberg, K. L., von Hessberg, P., and Nielsen, O. J.: Isotopic processes in atmospheric chemistry, Chem. Soc. Rev., 31, 313–323, doi:10.1039/b108011n, 2002. 5042
- Johnston, D. T.: Multiple sulfur isotopes and the evolution of Earth's surface sulfur cycle, Earth-Sci. Rev., 106, 161–183, doi:10.1016/j.earscirev.2011.02.003, 2011. 5050
  - Kazil, J., Harrison, R. G., and Lovejoy, E. R.: Tropospheric New Particle Formation and the Role of Ions, Space Sci. Rev., 137, 241–255, doi:10.1007/s11214-008-9388-2, 2008. 5041
    Kirkby, J., Curtius, J., Almeida, J., Dunne, E., Duplissy, J., Ehrhart, S., Franchin, A., Gagne,
- S., Ickes, L., Kuerten, A., Kupc, A., Metzger, A., Riccobono, F., Rondo, L., Schobesberger, S., Tsagkogeorgas, G., Wimmer, D., Amorim, A., Bianchi, F., Breitenlechner, M., David, A., Dommen, J., Downard, A., Ehn, M., Flagan, R. C., Haider, S., Hansel, A., Hauser, D., Jud, W., Junninen, H., Kreissl, F., Kvashin, A., Laaksonen, A., Lehtipalo, K., Lima, J., Lovejoy, E. R.,





Makhmutov, V., Mathot, S., Mikkila, J., Minginette, P., Mogo, S., Nieminen, T., Onnela, A., Pereira, P., Petaja, T., Schnitzhofer, R., Seinfeld, J. H., Sipila, M., Stozhkov, Y., Stratmann, F., Tome, A., Vanhanen, J., Viisanen, Y., Vrtala, A., Wagner, P. E., Walther, H., Weingartner, E., Wex, H., Winkler, P. M., Carslaw, K. S., Worsnop, D. R., Baltensperger, U., and Kulmala, M.: Role of sulphuric acid, ammonia and galactic cosmic rays in atmospheric aerosol nucleation, Nature, 476, 429–433, doi:10.1038/nature10343, 2011. 5040, 5041, 5052

Kulmala. M.: How Particles Nucleate and Grow, Science, 202, 1000-1001, doi:10.1126/science.1090848, 2003. 5040

Kurtén, T., Loukonen, V., Vehkamäki, H., and Kulmala, M.: Amines are likely to enhance neu-

tral and ion-induced sulfuric acid-water nucleation in the atmosphere more effectively than 10 ammonia, Atmos. Chem. Phys., 8, 4095–4103, doi:10.5194/acp-8-4095-2008, 2008. 5040 Leung, F. Y., Colussi, A. J., and Hoffmann, M. R.: Sulfur isotopic fractionation in the gas-phase oxidation of sulfur dioxide initiated by hydroxyl radicals, J. Phys. Chem. A, 105, 8073-8076,

doi:10.1021/ip011014+.2001.5047.5052

5

- 15 Loveiov, E. R., Curtius, J., and Frovd, K. D.: Atmospheric ion-induced nucleation of sulfuric acid and water, J. Geophys. Res., 109, 1–11, doi:10.1029/2003JD004460, 2004. 5041
  - Manninen, H. E., Nieminen, T., Asmi, E., Gagne, S., Hakkinen, S., Lehtipalo, K., Aalto, P., Vana, M., Mirme, A., Mirme, S., Horrak, U., Plass-Duelmer, C., Stange, G., Kiss, G., Hoffer, A., Toeroe, N., Moerman, M., Henzing, B., de Leeuw, G., Brinkenberg, M., Kouvarakis,
- G. N., Bougiatioti, A., Mihalopoulos, N., O'Dowd, C., Ceburnis, D., Arneth, A., Svenningsson, 20 B., Swietlicki, E., Tarozzi, L., Decesari, S., Facchini, M. C., Birmili, W., Sonntag, A., Wiedensohler, A., Boulon, J., Sellegri, K., Laj, P., Gysel, M., Bukowiecki, N., Weingartner, E., Wehrle, G., Laaksonen, A., Hamed, A., Joutsensaari, J., Petaja, T., Kerminen, V. M., and Kulmala, M.: EUCAARI ion spectrometer measurements at 12 European sites – analysis of new parti-
- cle formation events, Atmos. Chem. Phys., 10, 7907–7927, doi:10.5194/acp-10-7907-2010, 25 2010. 5041
  - Müller, R. and Crutzen, P. J.: A Possible Role of Galactic Cosmic Rays in Chlorine Activation During Polar Night, J. Geophys. Res., 98, 20483–20490, doi:10.1029/93JD02455, 1993. 5051
- Ono, S., Eigenbrode, J. L., Pavlov, A. A., Kharecha, P., Rumble III, D., Kasting, J. F., and Freeman, K. H.: New insights into Archean sulfur cycle from mass-independent sulfur isotope records from the Hamersley Basin, Australia, Earth Planet. Sc. Lett., 213, 15-30, doi:10.1016/S0012-821X(03)00295-4, 2003. 5042

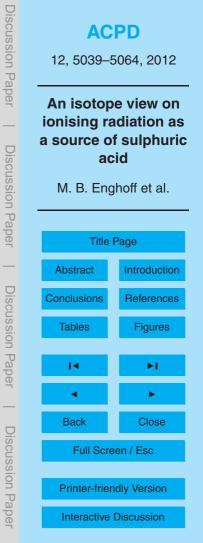
Discussion Paper	ACPD 12, 5039–5064, 2012 An isotope view on ionising radiation as a source of sulphuric acid M. B. Enghoff et al.				
per   Discussion Pape					
Paper	Title	Title Page			
	Abstract	Introduction			
Disc	Conclusions	References			
<b>Discussion</b> Paper	Tables	Figures			
Pap	I.4	►I.			
θŗ	•	•			
	Back	Close			
<b>Discussion Paper</b>	Full Screen / Esc				
on P	Printer-friendly Version				
aper	Interactive Discussion				



- Ono, S., Wing, B., Johnston, D., Farquhar, J., and Rumble, D.: Mass-dependent fractionation of quadruple stable sulfur isotope system as a new tracer of sulfur biogeochemical cycles, Geochim. Cosmochim. Ac., 70, 2238–2252, doi:10.1016/j.gca.2006.01.022, 2006. 5046
- Pavlov, A. A., Mills, M. J., and Toon, O. B.: Mystery of the volcanic mass-independent sulfur isotope fractionation signature in the Antarctic ice core, Geophys. Res. Lett., 32, L12816,

doi:10.1029/2005GL022784, 2008. 5042

- Raes, F., Janssens, A., and Eggermont, G.: A synergism between ultraviolet and gamma radiation in producing aerosol particles from SO<sub>2</sub>-H<sub>2</sub>SO<sub>4</sub> laden atmospheres, Atmos. Environ., 19, 1069–1073, doi:10.1016/0004-6981(85)90191-X, 1985. 5042
- Rees, C. E.: Steady-state model for sulfur isotope fractionation in bacterial reduction processes, Geochim. Cosmochim. Ac., 37, 1141–1162, doi:10.1016/0016-7037(73)90052-5, 1973. 5050
  - Romero, A. B. and Thiemens, M. H.: Mass-independent sulfur isotopic compositions in presentday sulfate aerosols, J. Geophys. Res., 108, 4524, doi:10.1029/2003JD003660, 2003. 5042
- <sup>15</sup> Sakai, H., Des Marais, D. J., Ueda, A., and Moore, J. G.: Concentrations and isotope ratios of carbon, nitrogen and sulfur in ocean-floor basalts, Geochim. Cosmochim. Ac., 48, 2433– 2441, doi:10.1016/0016-7037(84)90295-3, 1984. 5045
  - Schmidt, J. A., Johnson, M. S., and Schinke, R.: Isotope effects in N<sub>2</sub>O photolysis from first principles, Atmos. Chem. Phys., 11, 8965–8975, doi:10.5194/acp-11-8965-2011, 2011. 5042
- <sup>20</sup> Seinfeld, J. H. and Pandis, S. N.: Atmospheric Chemistry and Physics: From Air Pollution to Climate Change, Second Edn., John Wiley and Sons, 2006. 5045, 5049
  - Sipilä, M., Berndt, T., Petaja, T., Brus, D., Vanhanen, J., Stratmann, F., Patokoski, J., Mauldin, III, R. L., Hyvarinen, A.-P., Lihavainen, H., and Kulmala, M.: The Role of Sulfuric Acid in Atmospheric Nucleation, Science, 327, 1243–1246, doi:10.1126/science.1180315, 2010. 5041
- Svensmark, H., Pedersen, J. O. P., Marsh, N. D., Enghoff, M. B., and Uggerhøj, U. I.: Experimental evidence for the role of ions in particle nucleation under atmospheric conditions, P. Roy. Soc. Lond. A Mat., 463, 385–396, doi:10.1098/rspa.2006.1773, 2007. 5041
  - Tanaka, N., Rye, D. M., Xiao, Y., and Lasag, A. C.: Use of stable sulfur isotope systematics for evaluating oxidation reaction pathways and in-cloud-scavenging of sulfur-dioxide in the
- <sup>30</sup> atmosphere, Geophys. Res. Lett., 21, 1519–1522, doi:10.1029/94GL00893, 1994. 5047, 5052
  - Thiemens, M. H.: History and Applications of Mass-Independent Isotope Effects, Annu. Rev. Earth. Pl. Sc., 34, 217–262, doi:10.1146/annurev.earth.34.031405.125026, 2006. 5042





- Ueno, Y., Ono, S., Rumble, D., and Maruyama, S.: Quadruple sulfur isotope analysis of ca.
  3.5 Ga Dresser Formation: New evidence for microbial sulfate reduction in the Early Archean, Geochim. Cosmochim. Ac., 72, 5675–5691, doi:10.1016/j.gca.2008.08.026, 2008. 5042
- Ueno, Y., Johnson, M. S., Danielache, S. O., Eskebjerg, C., Pandey, A., and Yoshida, N.: Geological sulfur isotopes indicate elevated OCS in the Archean atmosphere, solving faint young sun paradox, P. Natl. Acad. Sci. USA, 106, 14784–14789, doi:10.1073/pnas.0903518106, 2009. 5049
  - Yu, F. and Turco, R. P.: Ultrafine aerosol formation via ion-mediated nucleation, Geophys. Res. Lett., 27, 883–886, doi:10.1029/1999GL011151, 2000. 5041
- Yu, F. and Turco, R. P.: From molecular clusters to nanoparticles: Role of ambient ionization in tropospheric aerosol formation, J. Geophys. Res., 106, 4797–4814, doi:10.1029/2000JD900539, 2001. 5041
  - Zatula, A. S., Andersson, P. U., Ryding, M. J., and Uggerud, E.: Proton mobility and stability of water clusters containing the bisulfate anion, HSO<sub>4</sub><sup>-</sup>(H<sub>2</sub>O)<sub>n</sub>, Phys. Chem. Chem. Phys., 13, 12287, 12204, doi:10.1020/c1cp21070i.2011.5052
- 13287–13294, doi:10.1039/c1cp21070j, 2011. 5052

]					
	ACPD				
	12, 5039–5064, 2012				
-	An isotope view on ionising radiation as a source of sulphuric acid				
-	M. B. Enghoff et al.				
J					
	Title Page				
_	Abstract Introduction				
	Conclusions References				
-	Tables Figures				
	I4 FI				
	▲ ▶				
,	Back Close				
	Full Screen / Esc				
5	Printer-friendly Version				
	Interactive Discussion				

Jiscussion Pape

Jiscussion Paper

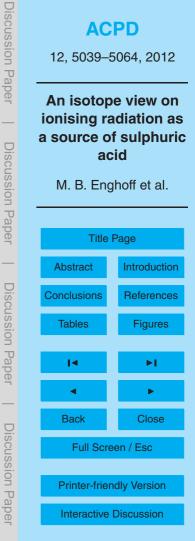
Discussion Paper

Jiscussion Paper



**Table 1.** Samples and results. The table displays a sample ID, the mass of  $BaSO_4$  recovered, and the isotopic fractionation data (normalized against the references as defined in Appendix B). The sample ID refers to the method of sulphate production:  $O_3$  is the liquid phase oxidation by ozone, UV is when the UV lamps were turned on, and  $\gamma$  is when the gamma source was used. The number at the end of the ID refers to the position of the bottle in the sampling line (see Fig. 1).

$\begin{array}{c ccccccccccccccccccccccccccccccccccc$				
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Sample ID	BaSO <sub>4</sub> Rec	$\delta^{34}$ S	$\Delta^{33}S$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	_	mg	‰	‰
$\begin{array}{cccccccc} O_3^-UV-1 & 26.1 & 8.7 & 0.296 \\ O_3^-UV-3 & 63.7 & 12.6 & -0.063 \\ O_3^-\gamma-1 & 19.4 & 13.9 & -0.041 \end{array}$	O <sub>3</sub> -1	2.8	16.0	-0.170
$O_3^{"}$ -UV-363.712.6 $-0.063$ $O_3^{"}$ - $\gamma$ -119.413.9 $-0.041$	O <sub>3</sub> -3	392.8	10.4	-0.072
$O_3 - \gamma - 1$ 19.4 13.9 -0.041	O <sub>3</sub> -UV-1	26.1	8.7	0.296
	O <sub>3</sub> -UV-3	63.7	12.6	-0.063
O <sub>3</sub> -UV-γ-1 16.8 9.7 0.232	Ο <sub>3</sub> -γ-1	19.4	13.9	-0.041
	Ο <sub>3</sub> -UV-γ-1	16.8	9.7	0.232





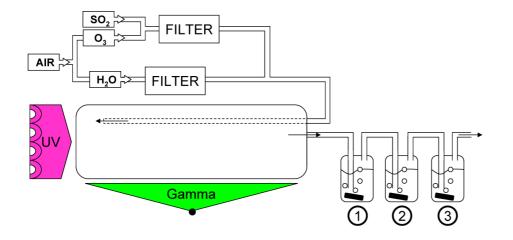
<b>Table 2.</b> Isotopic composition of samples. The table displays the sample ID, $\Delta^{36}$ S (normalized
against the references as defined in Appendix B), the amount of BaSO <sub>4</sub> used, the amount of
$Ag_2S$ recovered, the amount of $Ag_2S$ used, and the amount of $SF_6$ recovered.

Sample ID –	∆ <sup>36</sup> S ‰	BaSO <sub>4</sub> Used mg	Ag <sub>2</sub> S Rec mg	Ag <sub>2</sub> S Used mg	SF <sub>6</sub> µmol
O <sub>3</sub> -1	0.5	_	0.689	0.689	2.85
O <sub>3</sub> -3	0.4	8.99	8.84	2.07	7.91
O <sub>3</sub> -UV-1	-1.3	-	_	-	6.98
0 <sub>3</sub> -UV-3	0.2	9.39	9.48	2.00	8.87
Ο <sub>3</sub> -γ-1	0.1*	5.51	5.09	2.21	9.45
Ο <sub>3</sub> -UV-γ-1	-1.0	7.90	7.34	2.13	9.80

\* This sample was not measured with the improved purification system.

Discussion Par	<b>ACPD</b> 12, 5039–5064, 2012				
per   Discussion	ionising ra a source o ac	An isotope view on ionising radiation as a source of sulphuric acid M. B. Enghoff et al.			
Paper	Title	Page			
	Abstract	Introduction			
Disc	Conclusions	References			
ussion	Tables	Figures			
Pap	I	►I			
		•			
	Back	Close			
iscussi	Full Screen / Esc				
on P	Printer-friendly Version				
aper	Interactive Discussion				





**Fig. 1.** Schematic of the chamber setup in Copenhagen. The purified airstream is split in two, one passing through the humidifier and the other through the ozone generator before rejoining along with the  $SO_2$  flow and entering the chamber (see text). The airstream leaving the reaction volume is passed through 3 numbered gas washing bottles, in series, with magnetic stirrers. Collimated UV light enters through a Teflon foil at the end of the chamber, while a gamma source is place at the side.

	<b>ACPD</b> 12, 5039–5064, 2012			
	An isotope view on ionising radiation as a source of sulphuric acid M. B. Enghoff et al.			
Title Page				
-	Abstract	Introduction		
	Conclusions	References		
	Tables	Figures		
	I	۶I		
	•	•		
_	Back	Close		
	Full Screen / Esc			
2	Printer-friendly Version			
	Interactive Discussion			



