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Characterization of CZT Detectors for the ASIM Mission

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Abstract-. The National Space Institute, Technical University of Denmark is responsible for the selection and characterization of the CZT detector crystals for the X and gamma -ray instrument MXGS onboard ESA's Atmospheric Space Interaction Monitor (ASIM) mission. The Modular X- and Gamma ray Sensor (MXGS) shall, situated on the International Space Station, detect Terrestrial Gamma ray Flashes (TGF) generated in the Earth atmosphere. TGFs are believed to be produced in connection with thunderstorms and ASIM will therefore also be equipped with optical instruments which trigger on lightening.

MXGS is equipped with a sensor plane consisting of 64 detector modules each having 4 CZT crystals with dimensions of 20 mm x 20 mm x 5 mm. The crystals are divided in pixels with a pitch of 2.5 mm. The MXGS employs coded mask imaging with which the direction of the TGF with respect to Nadir can be determined. The MXGS spectral performance is important since the details of the TGF energy spectrum can be used to determine the atmospheric generation altitude of the TGFs.

Redlen has recently delivered the first CZT detector modules. Measurements at the University of Bergen demonstrate that the detectors exhibit the expected spectral performance; however it was also found that the detector modules showed unexplained pixel-to-pixel count rate variations. At The National Space Institute a series of measurements were initiated focusing on the pixel-to-pixel count rate variations. Collimated monochromatic X-ray beams were used to scan a number of Redlen pixel detectors. The main conclusion of these measurements is that the pixel-to-pixel count rate variations are likely related to the nonuniformity of the detector internal electric field. The measurements and their interpretation will be discussed in the following sections

I. INTRODUCTION

The Atmosphere-Space Interactions Monitor (ASIM) is proposed for the International Space Station external facilities on the Columbus module, for the study of giant electrical discharges (lightning) in the high-altitude atmosphere above thunderstorms. The discharges are seen as optical-, X- and Gamma ray flashes in the stratosphere and mesosphere. The optical emissions are dubbed "red sprites", "blue jets", and "elves" or "Transient Luminous Events" (TLEs). The X- and Gamma ray emissions are dubbed "Terrestrial Gamma ray Flashes" (TGFs). The ASIM mission comprises therefore two main scientific instruments: The Miniature Multi-spectral Imaging Array (MMIA) composed of 6 cameras and 6 photometers and the Modular X- and Gamma ray Sensor (MXGS) The ASIM payload is shown in Fig. 1.



Fig. 1. The ASIM payload for the ISS Colombus external platform. The left tower supports the MMIA Nadir instrumentation. The right tower supports MXGS also pointing towards Nadir. The MMIA Limb optical instruments are placed below MXGS.

The ASIM mission will:

- Provide the most comprehensive survey of the occurrence of TLEs on a global scale and establish the relationship between TLEs and the thunderstorm source.
- Survey the occurrence of TGFs and establish the relationship between the TGFs, TLEs and other parameters characterizing atmospheric activity.
- Establish the relative importance of mechanisms proposed for the generation of sprites and jets.
- Quantify the effects of TLEs and TGFs on the upper atmosphere (stratosphere, mesosphere, lower thermosphere, and ionosphere).
- Quantify the interaction of TGF and lightninginduced electromagnetic radiation with the radiation belts.

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TABLE I MXGS Performance Specifications	
ENERGY RANGE	10-500 keV
ENERGY RESOLUTION	<10% @ 60 keV
QUANT. EFF.	>0.95@100 keV
ABSOLUTE TIME ACCURACY	<100 µs
BURST RATE CAPABILITY	1000 cts/ms
SENSITIVE AREA	1024 cm ²
TOTAL DETECTOR BACKGROUND	~2 cts/ms (10-500 keV)
Field of View	80°*80°
Angular Resolution	2°.

II. THE MODULAR X- AND GAMMA-RAY SENSOR (MXGS)

The MXGS performance specifications are given in Table I. The detector plane consists of a 1024cm² array of CZT detector crystals. It is protected against the background radiation by a passive graded shield surrounding the detector housing. A Hopper shaped collimator defines the 80°x80° field of view for MXGS and shields the detector plane against the Cosmic X-ray Background. The Detector Front End Electronics (DFEE) is mounted in the housing below the detectors. The electronics contains also the High Voltage Power Supply (HVPS) and Low Voltage Power Supply (LVPS) as well as the MXGS Data Processing Unit (DPU).



Fig. 2. MXGS Detector Module.

The MXGS instrument DFEE design consists of 4 Detector Assembly Units (DAU), and each DAU consists of 16 Detector Modules (DM) and one Detector Assembly Board (DAB). The DAB holds the read-out electronics and the Readout Control Unit (RCU). The RCUs interface to the Data Processing Unit (DPU). The purpose of the DAU is to read out the events and transfer the data to the DPU.

The DM consists of two separable units, a CZT sensor and an ASIC. The sensor comprises four $20x20x5mm^3$ CZT detectors tiled together on a PCB. Each detector is pixelated into 64 pixels (2.5 mm pixel pitch), making a 16 x 16-pixel array in total. The detector unit is stacked onto the ASIC unit via three connectors (see Fig. 2). The ASIC unit holds two XA 1.72 chips, each reading 128 pixels. The ASICs provide position and energy for each detected photon. The DM is operated through front end electronics and software developed at University of Bergen.

A. CZT Material Selection

At NSI we have conducted a campaign with the purpose of selecting the most suited CZT/CdTe detectors for MXGS. [1] The NSI test set-up comprises a 16 channel multi-parameter data acquisition system including 16 low-noise preamplifiers. A number of radioactive sources are available. A step motor facility can be loaded with these sources and fine scans with monochromatic beams can be performed for detailed analysis of the detector responses.

The tests were typically conducted on 10x10x5 mm CZT/CdTe crystals with a pixel pitch of 2.5 mm similar to that of the DMs. As a result of these investigations we chose Redlen CZT material as the baseline for MXGS. Fig. 3 shows a photo of investigated Redlen detectors. The pixel gap is only 100 micrometers. Our investigations [1] have shown that a minimum pixel gap reduce the influence of pixel charge sharing and pixel gap charge losses.



Fig. 3. Redlen CZT semples (10 mm x 10 mm x 5 mm).

It is important that the Redlen material consistently showed very good charge collection properties with an electron mobility lifetime product $(\mu\tau)_e > 10^{-2} \text{ cm}^2/\text{V}$. This material is therefore well suited for detectors with fine energy resolution.



Fig. 4. Redlen pixel detector response to 137Cs photons. The left plot gives the bi-parametric distribution for the pixel signal versus the ratio between the cathode and the pixel signal. The right figure displays the resulting 137Cs spectrum after application of depth of interaction correction.

An example of the achievable quality is shown in Fig. 4. The presented results were obtained with a 137 Cs source. The left plot gives the bi-parametric distribution for the pixel

signal versus the ratio between the common cathode and the pixel signal. The right figure displays the resulting ¹³⁷Cs spectrum after application of depth of interaction correction (DOI) as described in [2] and [3].

III. MEASUREMENTS ON THE FIRST DMS

Redlen has recently delivered the first set of CZT detector modules. Measurements at the University of Bergen on one of the DM connected to two XA 1.72 ASICs demonstrate that the detectors exhibit the expected spectral performance. This is illustrated in Fig 4 where the upper figure is a 2D image of 128 ⁵⁷ Co spectra from one of the ASICs. The middle plot shows a typical spectrum. The overall energy resolution measured for the 122 keV line is 6.2% (FWHM) at 20 °C and improved slightly to 5.5% (FWHM) at a detector temperature of 5°C.



Fig. 5. DM measurements at UB. The upper figure is a 2D image of 128 ⁵⁷ Co spectra from one of the ASICs. The middle plot shows a typical spectrum. Pixel-to-pixel count rate variations are shown in the lower plot.

However it was found that the detector modules showed large unexplained pixel-to-pixel count rate variations as it can be seen pixel count plot in the lower left part of Fig. 5. Here the pixel count distribution has a rms scatter of 20% with a max to min pixel count of more than a factor of 2. This behavior is problematic since it might indicate that the DMs suffer from pixel varying inefficiencies and that the DM therefore has an overall efficiency which is lower than the predicted one.

IV. X-Y SCANS

We decided to investigate the Redlen pixel detectors and the count rate variations in more detail performing X-Y scans with fine X-ray beams. The early detectors shown in. Fig. 3 showed a similar pixel to pixel variation. We decided to start with these detectors since they were easier to interface to the NSI data acquisition set-up than the DMs. The NSI facility comprises also a 2D step motor equipment which allows to scan (position accuracy < 5 micrometers) the detectors with collimated X-ray beams



Fig. 6 The upper part displays the 16 pixel spectra obtained when illuminating detector 2da uniformly with ¹⁰⁹Cd X-rays. The ¹⁰⁹Cd scan was performed with a 0.5 mm beam and with (x, y) steps of 0.25 mm. The results are shown in the lower part as a 2D count-image.

The investigations were performed using two radioactive sources: a ¹⁰⁹Cd source for low energy scans (22.1 keV and 25

108

keV) and a 57 Co source for high energy scans (122 keV and 136 keV).

All measurements were performed with the common planar cathode biased at -500 V and the pixel anodes at 0 V. The pixel detector was illuminated from the side of the common planar electrode.

The upper part of Fig 6 displays the 16 pixel spectra obtained when illuminating detector 2da uniformly with ¹⁰⁹Cd X-rays. The data analysis s/w was operated such that only so-called single pixel events are displayed. In other words, shared events for which two or more pixel were triggered by the same event are rejected. The spectra clearly show pixel-to-pixel count rate variations also given by the total counts in top of each spectrum. Remark the difference of almost 25% between the counts for pixel 10 and pixel 11. These pixels are both central pixels and are not influenced by edge effect.

The ¹⁰⁹Cd scan was performed with a 0.5 mm beam and with (x, y) steps of 0.25 mm. The results are shown in the lower part of Fig. 6 as a 2D count-image. Again only single pixel events are displayed which has the consequence that the count rate drops when the beam passes over the boarders between pixels where charge sharing is present. The pixel boundaries are therefore clearly visible in the map as given by the dark boarder lines with fewer counts. The dark feature seen for pixel 13 is an exception since it is caused by the Ag filled conducting glue which is used for the contact to the common planar electrode. However, the important thing to remark is that sizes of the pixels are varying and that the pixel size is strongly correlated with the pixel counts in the upper plots, e.g. compare pixel 10 and pixel 11. It should also be remarked that the count rate variations within the pixels are very small and that the count rate inside the pixel is constant without pixel-to-pixel variations. We can conclude that the pixel size variations seen in the lower part of Fig. 6 fully explains the pixel-to-pixel count rate variations for uniform illumination given in the upper part of Fig. 6 The present measurements do also not show in-efficiencies and charge collection problems as those reported by [4] for a number of pixelated CZT detectors.

The X-ray photons of the ¹⁰⁹Cd source are stopped within 0.1 mm below the common planar electrode surface. In order to investigate whether the effective area variations are surface correlated we repeated the measurements with a ⁵⁷Co photon beam (122 and 136 keV) which has a mean free path of ~1.7 mm inside the detector CZT volume.

The results are displayed in Fig. 7. The upper plots show the 16 pixels spectra obtained by uniformly illumination of the detector with photons from the 57 Co source. The total number of detected counts are given in the top of each pixel plot. The pixel to pixel count variation is very similar to the one found when illuminating the detector with 109 Cd X-rays. The count difference between pixel 10 and pixel 11 is again ~25%.

The lower part image of Fig. 7 displays the results of the 57 Co scans. In this case we could operate with a somewhat finer beam of 0.3 mm diameter; but again with steps (x,y) of 0.25 mm. The image displays the number of counts of single

pixel events as explained for Fig. 6. The pixel boarders are clearly visible and the effective pixel sizes are almost identical to those measured with the ¹⁰⁹Cd source



Fig. 7 The upper part displays the 16 pixel spectra obtained when illuminating detector 2da uniformly with photons form the 57 Co source. The scan was performed with a 0.3 mm beam and with (x,y) steps of 0.25 mm. The results are shown in the lower part as a 2D count-image.

The preliminary conclusion is therefore that the effective pixel shape(size) does not depend on the photon interaction depth. This is surprising since we believe that the shape variations are caused by non-uniform electric fields inside the detector volume. Such fields can be formed by crystal defects and cause image distortions as demonstrated in [5], [6]. From the present results, it cannot be excluded that field distortions close to the pixel electrodes are responsible for the image distortions. Issues associated with attachment/assembly (delamination, smearing of low temperature solder, misalignment) could also possibly contribute to the scanned pixel size variation as well. In order to clarify this we plan to perform a full 3D analysis employing the DOI method [2], [3] to determine also the depth coordinate.

V. CONCLUSION

We have conducted pixel to pixel uniformity measurements on Redlen CZT detectors. Result for full illumination and scan consistently shown large count rate variations between pixels. We believe that the pixel-to-pixel non-uniformity is related to the non-uniform internal electric field profile. It is at present unclear what is disturbing the internal electric field. Other groups [5], [6] have found similar effects for CZT pixel detectors. The explanation in [5], [6] relates to extended defects and residual strains existing inside the crystal.

We plan to perform a full 3D analysis employing the DOI method on the CZT detector to reveal the electron drift path close to the pixels function of beam position.

The discussed pixel-to-pixel non-uniformities of the Redlen CZT pixel detectors are not a serious concern for the ASIM MXGS instrument. MSGS imaging will be performed with a coded mask with 10 mm x 10 mm holes such that the variations will be smoothed over 16 detector pixels.

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