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Limits on variations of the quark masses, QCD scale, and fine structure constant

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We perform calculations of the dependence of nuclear magnetic moments on quark masses and obtain limits on the variation of the fine structure constant α and (m_q/Λ_{QCD}) from recent measurements of hydrogen hyperfine (21 cm) and molecular rotational transitions in quasar absorption systems, atomic clock experiments with hyperfine transitions in H, Rb, Cs, Yb⁺, Hg⁺, and optical transition in Hg⁺. Experiments with Cd⁺, deuterium/hydrogen, molecular SF₆, and Zeeman transitions in ³He/Xe are also discussed.

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I. INTRODUCTION

Interest in the temporal and spatial variation of major constants of physics has been recently revived by astronomical data which seem to suggest a variation of the electromagnetic constant $\alpha = e^2/\hbar c$ at the 10^{-5} level for the time scale 10 billion years, see Ref. [1] (a discussion of other limits can be found in the review [2] and references therein). However, an independent experimental confirmation is needed.

The hypothetical unification of all interactions implies that the variation of the electromagnetic interaction constant α should be accompanied by the variation of masses and the strong interaction constant. Specific predictions need a model. For example, the grand unification model discussed in Ref. [3] predicts that the quantum chromodynamic (QCD) scale Λ_{QCD} (defined as the position of the Landau pole in the logarithm for the running strong coupling constant) is modified as follows: $\delta \Lambda_{QCD} / \Lambda_{QCD} \approx 34 \ \delta \alpha / \alpha$. The variation of quark and electron masses in this model is given by $\delta m/m$ $\sim 70 \ \delta \alpha / \alpha$. This gives an estimate for the variation of the dimensionless ratio

$$\frac{\delta(m/\Lambda_{QCD})}{(m/\Lambda_{QCD})} \sim 35 \frac{\delta \alpha}{\alpha}.$$
 (1)

This result is strongly model dependent (for example, the coefficient may be an order of magnitude smaller and even of opposite sign [4]). However, the large coefficients in these expressions are generic for grand unification models, in which modifications come from high-energy scales: they appear because the running strong-coupling constant and Higgs constants (related to mass) run faster than α . This means that if these models are correct the variation of masses and the strong interaction scale may be easier to detect than the variation of α .

One can only measure the variation of dimensionless quantities and therefore we want to extract from the measurements the variation of the dimensionless ratio m_q/Λ_{QCD} -where m_q is the quark mass (with the dependence on the renormalization point removed). A number of

limits on the variation of m_q/Λ_{QCD} have been obtained recently from consideration of big bang nucleosynthesis, quasar absorption spectra, and the Oklo natural nuclear reactor, which was active about 1.8 billion years ago [5–8] (see also Refs. [9–13]). Below we consider the limits on various combinations of the quark masses and the fine structure constant which follow from quasar absorption radio spectra and laboratory atomic clock comparisons. Laboratory limits with a time base of the order 1 yr are especially sensitive to oscillatory variations of fundamental constants. A number of relevant measurements have been performed already and even larger numbers have been started or are planned. The increase in precision is happening very fast.

It has been pointed out by Karshenboim [14] that measurements of ratios of hyperfine structure intervals in different atoms are sensitive to any variation of nuclear magnetic moments. First rough estimates of the dependence of nuclear magnetic moments on m_q/Λ_{QCD} and limits on the variation of this ratio with time were obtained in Ref. [5]. Using H, Cs, and Hg⁺ measurements [15,16], we obtained a limit on the variation of m_q/Λ_{QCD} of about 5×10^{-13} per year. Below we calculate the dependence of nuclear magnetic moments on m_q/Λ_{QCD} and obtain the limits from recent atomic clock experiments with hyperfine transitions in H, Rb, Cs, Yb⁺, Hg⁺, and the optical transition in Hg⁺. It is convenient to assume that the strong interaction scale Λ_{QCD} does not vary, so we will speak about the variation of masses (this means that we measure masses in units of Λ_{QCD}). We shall restore the explicit appearance of Λ_{QCD} in the final answers.

The hyperfine structure constant can be presented in the following form:

$$A = \text{const} \times \left(\frac{m_e e^4}{\hbar^2}\right) \left[\alpha^2 F_{rel}(Z\alpha)\right] \left(\mu \frac{m_e}{m_p}\right).$$
(2)

The factor in the first set of brackets is an atomic unit of energy. The second "electromagnetic" set of brackets determines the dependence on α . An approximate expression for

the relativistic correction factor (Casimir factor) for an *s*-wave electron is the following:

$$F_{rel} = \frac{3}{\gamma(4\gamma^2 - 1)},\tag{3}$$

where $\gamma = \sqrt{1 - (Z\alpha)^2}$ and Z is the nuclear charge. Variation of α leads to the following variation of F_{rel} [15]:

$$\frac{\delta F_{rel}}{F_{rel}} = K \frac{\delta \alpha}{\alpha},\tag{4}$$

$$K = \frac{(Z\alpha)^2 (12\gamma^2 - 1)}{\gamma^2 (4\gamma^2 - 1)}.$$
 (5)

More accurate numerical many-body calculations [17] of the dependence of the hyperfine structure on α have shown that the coefficient *K* is slightly larger than that given by this formula. For Cs (*Z*=55) *K*=0.83 (instead of 0.74), for Rb *K*=0.34 (instead of 0.29), and finally for Hg⁺ *K*=2.28 (instead of 2.18).

The last set of brackets in Eq. (2) contains the dimensionless nuclear magnetic moment μ [i.e., the nuclear magnetic moment $M = \mu(e\hbar/2m_pc)$], electron mass m_e and proton mass m_p . We may also include a small correction arising from the finite nuclear size. However, its contribution is insignificant.

Recent experiments measured the time dependence of the ratios of the hyperfine structure intervals of ¹⁹⁹Hg⁺ and H [15], ¹³³Cs and ⁸⁷Rb [18], and the ratio of the optical frequency in Hg⁺ to the hyperfine frequency of ¹³³Cs [20]. In the ratio of two hyperfine structure constants for different atoms' time dependence may appear from the ratio of the factors F_{rel} (depending on α) as well as from the ratio of nuclear magnetic moments (depending on m_q/Λ_{QCD}). Magnetic moments in a single-particle approximation (one unpaired nucleon) are

$$\mu = [g_s + (2j-1)g_l]/2 \tag{6}$$

for j = l + 1/2,

$$\mu = \frac{j}{2(j+1)} \left[-g_s + (2j+3)g_l \right] \tag{7}$$

for j=l-1/2. Here the orbital g factors are $g_l=1$ for a valence proton and $g_l=0$ for a valence neutron. The present values of the spin g factors g_s are $g_p=5.586$ for protons and $g_n=-3.826$ for neutrons. They depend on m_q/Λ_{QCD} . The light quark masses are only about 1% of the nucleon mass $[m_q=(m_u+m_d)/2\approx 5 \text{ MeV}]$ and the nucleon magnetic moment remains finite in the chiral limit, $m_u=m_d=0$. Therefore one might think that the corrections to g_s arising from the finite quark masses would be very small. However, through the mechanism of spontaneous chiral symmetry breaking, which leads to contributions to hadron properties from Goldstone boson loops, one may expect some enhancement of the effect of quark masses [19]. The natural framework for dis-

cussing such corrections is chiral perturbation theory and we discuss these chiral corrections next.

II. CHIRAL PERTURBATION THEORY RESULTS FOR NUCLEON MAGNETIC MOMENTS AND MASSES

In recent years there has been tremendous progress in the calculation of hadron properties using lattice QCD. Moore's Law, in combination with sophisticated algorithms, means that one can now make extremely accurate calculations for light quark masses (m_q) larger than 50 MeV. However, in order to compare with experimental data, it is still necessary to extrapolate quite a long way as a function of quark mass. This extrapolation is rendered nontrivial by the spontaneous breaking of chiral symmetry in QCD, which leads to Goldstone boson loops and, as a direct consequence, nonanalytic behavior as a function of quark mass [21,22]. Fortunately the most important nonanalytic constraint on the extrapolation procedure.

In the past few years the behavior of hadron properties as a function of quark mass has been studied over a much wider range than one needs for the present purpose [22–28]. One can therefore apply the successful extrapolation formulas developed in the context of lattice QCD with considerable confidence.

The key qualitative feature learned from the study of lattice data is that Goldstone boson loops are strongly suppressed once the Compton wavelength of the boson is smaller than the source. Inspection of lattice data for a range of observables, from masses to charge radii and magnetic moments, reveals that the relevant mass scale for this transition is $m_q \sim 50$ MeV—i.e., $m_\pi \sim 400-500$ MeV [22,29]. The challenge of chiral extrapolation is therefore to incorporate the correct, model independent nonanalytic behavior dictated by chiral symmetry while ensuring excellent convergence properties of the chiral expansion in the large mass region, as well as maintaining the model independence of the results of the extrapolation. Considerable study of this problem has established that the use of a finite range regulator (FRR) fulfills all of these requirements [30–32]. Indeed, in the case of the mass of the nucleon, it has been shown that the extrapolation from $m_{\pi}^2 \sim 0.25 \text{ GeV}^2$ to the physical pion mass—a change of m_a by a factor of 10—can be carried out with a systematic error less than 1% [31]. In the following we apply this same method to calculate the change in the nucleon mass, corresponding to quark mass changes at the level of 0.1% or less, as required in the present context.

A. Variation of the nucleon mass with quark mass

The expansion for the mass of the nucleon given in Refs. [31,32] is

$$M_N = a_0 + a_2 m_\pi^2 + a_4 m_\pi^4 + a_6 m_\pi^6 + \sigma_{N\pi} + \sigma_{\Delta\pi} + \sigma_{\text{tad}},$$
(8)

where the chiral loops which given rise, respectively, to the leading and next-to-leading nonanalytic (LNA and NLNA) behavior are

$$\sigma_{N\pi} = -\frac{3}{32\pi f_{\pi}^2} g_A^2 I_M(m_{\pi}, \Delta_{NN}, \Lambda),$$
(9)

$$\sigma_{\Delta\pi} = -\frac{3}{32\pi f_{\pi}^2} \frac{32}{25} g_A^2 I_M(m_{\pi}, \Delta_{N\Delta}, \Lambda),$$
(10)

$$\sigma_{\rm tad} = -\frac{3}{16\pi^2 f_{\pi}^2} c_2 m_{\pi}^2 I_T(m_{\pi}, \Lambda), \tag{11}$$

and the relevant integrals are defined (in heavy baryon approximation) as

$$I_M(m_P, \Delta_{BB'}, \Lambda) = \frac{2}{\pi} \int_0^\infty dk \, \frac{k^4 u^2(k, \Lambda)}{\omega_k(\Delta_{BB'} + \omega_k)},\tag{12}$$

$$I_T(m_{\pi}, \Lambda) = \int_0^\infty dk \left(\frac{2k^2 u^2(k)}{\sqrt{k^2 + m_{\pi}^2}} \right) - t_0, \qquad (13)$$

with $\omega_k = \sqrt{k^2 + m_P^2}$ and $\Delta_{BB'}$ the relevant baryon mass difference (i.e., $M_{B'} - M_B$). We take the $\Delta - N$ mass splitting, $\Delta = M_{\Delta} - M_N$, to have its physical value (0.292 GeV), while $g_A = 1.26$. The regulator function $u(k,\Lambda)$ is taken to be a dipole with mass $\Lambda = 0.8$ GeV. In Eq. (13) t_0 , defined such that I_T vanishes at $m_{\pi} = 0$, is a local counter term introduced in FRR to ensure a linear relation for the renormalization of c_2 .

The model independence of the expansion given in Eq. (8) is ensured by fitting the unknown coefficients to the physical nucleon mass and lattice data from the CP-PACS Collaboration [33], yielding $a_0 = 1.22$, $a_2 = 1.76$, $a_4 = -0.829$, $a_6 = 0.260$ (with all parameters expressed in the appropriate powers of GeV). With these parameters fixed one can evaluate the rate of change of the mass of the nucleon with quark or pion mass at the physical pion mass:

$$m_q \frac{\partial}{\partial m_q} M_N = m_\pi^2 \frac{\partial}{\partial m_\pi^2} M_N = 0.035 \text{ GeV},$$
 (14)

a quantity commonly known as the pion-nucleon sigma commutator. Using Eq. (14) one finds the relationship (in terms of dimensionless quantities)

$$\frac{\delta M_N}{M_N} = \frac{m_\pi^2}{M_N} \frac{\partial M_N}{\partial m_\pi^2} \frac{\delta m_q}{m_q}$$
(15)

$$=0.037 \frac{\delta m_q}{m_q}.$$
 (16)

The extension of this procedure to the effect of a variation in the strange quark mass is similar, but one must include the variation arising from η -nucleon loops, as well as kaon loops with intermediate Σ or Λ baryons,

$$\sigma_{N\Sigma}^{K} + \sigma_{N\Lambda}^{K} + \sigma_{NN}^{\eta} \,. \tag{17}$$

These contributions can be expressed as

$$\sigma_{BB'}^{P} = -\frac{3}{32\pi f_{\pi}^{2}} G_{BB'}^{P} I_{M}(m_{P}, \Delta_{BB'}, \Lambda)$$
(18)

with $G_{BB'}^{P}$ the associated coupling squared. Once again we select the dipole regulator:

$$u(k,\Lambda) = \left(\frac{\Lambda^2}{\Lambda^2 + k^2}\right)^2.$$
 (19)

For the relevant diagrams, $N \rightarrow \Sigma K$, $N \rightarrow \Lambda K$, and $N \rightarrow N \eta$, we have

$$G_{N\Sigma}^{K} = \frac{1}{3}(D-F)^{2},$$

$$G_{N\Lambda}^{K} = \frac{1}{9}(3F+D)^{2},$$

$$G_{NN}^{\eta} = \frac{1}{9}(3F-D)^{2},$$
(20)

where we take F=0.50 and D=0.76. We use the Gell-Mann–Oakes–Renner relation in the SU(2) chiral limit to relate the variation of the kaon mass in the chiral SU(2) limit, $\tilde{m}_K = \sqrt{\mu_K^2 - \frac{1}{2}\mu_\pi^2} = 0.484$ GeV (with $\mu_{\pi\{K\}}$, the physical pion{kaon} mass), to the variation of the strange quark mass $(\delta \tilde{m}_K^2/\tilde{m}_K^2 = \delta m_s/m_s)$. Hence the variation of the nucleon mass with strange quark mass is given by

$$\frac{\delta M_N}{M_N} = \left\{ \frac{\tilde{m}_K^2}{M_N} \frac{\partial}{\partial \tilde{m}_K^2} (\sigma_{N\Sigma}^K + \sigma_{N\Lambda}^K + \sigma_{NN}^\eta) \right\} \frac{\delta m_s}{m_s}.$$
 (21)

Using the dipole regulator mass, $\Lambda\!=\!0.8$ GeV, Eq. (21) leads to the result

$$\frac{\delta M_N}{M_N} = 0.011 \frac{\delta m_s}{m_s}.$$
 (22)

B. Variation of proton and neutron magnetic moments with quark mass

The treatment of the mass dependence of the nucleon magnetic moments is very similar to that for the masses. Once again the loops which give rise to the LNA and NLNA behavior are evaluated with a FRR, while the smooth, analytic variation with quark mass is parametrized by fitting relevant lattice data with a finite number of adjustable constants.

For the lattice data we use the CSSM Lattice Collaboration results [34] of nucleon three-point functions. Results are obtained using established techniques in the extraction of form factor data [35]. Similar calculations have also been recently reported by the QCDSF Collaboration [28]. We use the two heaviest simulation results, $m_{\pi}^2 \sim 0.6-0.7 \text{ GeV}^2$ [34]. These simulations were performed with the FLIC fermion action [36] on a $20^3 \times 40$ lattice at a = 0.128 fm.

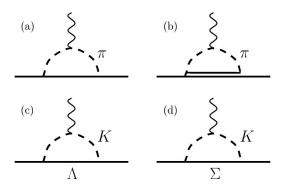


FIG. 1. Chiral corrections to the nucleon magnetic moments included in the present work.

In the magnetic moment case the formulas are a little more complicated, so we leave the details for the Appendix. Suffice it to say here that the relevant processes are shown in Fig. 1. Again we use a dipole form for the regulator with $\Lambda = 0.8$ GeV.

Having parametrized the neutron and proton magnetic moments as a function of m_{π} , the fractional change versus m_a or m_s is given by

$$\frac{\delta\mu}{\mu} = \left\{ \frac{m_{\pi}^2}{\mu} \frac{\partial\mu}{\partial m_{\pi}^2} \right\} \frac{\delta m_q}{m_q},\tag{23}$$

$$\frac{\delta\mu}{\mu} = \left\{ \frac{\tilde{m}_K^2}{\mu} \frac{\partial\mu}{\partial\tilde{m}_K^2} \right\} \frac{\delta m_s}{m_s}.$$
 (24)

The numerical results may then be summarized as

$$\frac{\delta\mu_p}{\mu_p} = -0.087 \frac{\delta m_q}{m_q},\tag{25}$$

$$\frac{\delta\mu_p}{\mu_p} = -0.013 \frac{\delta m_s}{m_s},\tag{26}$$

$$\frac{\delta\mu_n}{\mu_n} = -0.118 \frac{\delta m_q}{m_q},\tag{27}$$

$$\frac{\delta\mu_n}{\mu_n} = 0.0013 \frac{\delta m_s}{m_s},\tag{28}$$

$$\frac{\delta(\mu_p/\mu_n)}{(\mu_p/\mu_n)} = 0.031 \frac{\delta m_q}{m_q},\tag{29}$$

$$\frac{\delta(\mu_p/\mu_n)}{(\mu_p/\mu_n)} = -0.015 \frac{\delta m_s}{m_s}.$$
(30)

III. DEPENDENCE OF ATOMIC TRANSITION FREQUENCIES ON FUNDAMENTAL CONSTANTS

Using the results of the previous section we can now use Eqs. (6), (7) to study the variation of nuclear magnetic moments. For all even Z nuclei with valence neutron (¹⁹⁹Hg, ¹⁷¹Yb, ¹¹¹Cd, etc.) we obtain $\delta \mu/\mu = \delta g_n/g_n$. For ¹³³Cs we

have a valence proton with j = 7/2, l = 4, and

$$\frac{\delta\mu}{\mu} = 0.110 \frac{\delta m_q}{m_q} + 0.017 \frac{\delta m_s}{m_s}.$$
(31)

For ⁸⁷Rb we have valence proton with j = 3/2, l = 1, and

$$\frac{\delta\mu}{\mu} = -0.064 \frac{\delta m_q}{m_q} - 0.010 \frac{\delta m_s}{m_s}.$$
(32)

As an intermediate result it is convenient to present the dependence of the ratio of the hyperfine constant A to the atomic unit of energy $E = m_e e^4/\hbar^2$ (or the energy of the 1s-2s transition in hydrogen, which is equal to 3/8E) on a variation of the fundamental constants. We introduce a parameter V defined by the relation

$$\frac{\delta V}{V} \equiv \frac{\delta(A/E)}{A/E}.$$
(33)

We start from the hyperfine structure of 133 Cs which is used as a frequency standard. Using Eqs. (2), (31) we obtain

$$V(^{133}\text{Cs}) = \alpha^{2.83} \left(\frac{m_q}{\Lambda_{QCD}}\right)^{0.110} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{0.017} \frac{m_e}{m_p}.$$
 (34)

The factor m_e/m_p will cancel out in the ratio of hyperfine transition frequencies. However, it will survive in comparison between hyperfine and optical or molecular transitions (see below). According to Eqs. (16) and (22) the relative variation of the electron to proton mass ratio can be described by the parameter

$$X(m_e/m_p) = \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.037} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{-0.011} \frac{m_e}{\Lambda_{QCD}} \quad (35)$$

which can be substituted into Eq. (34) instead of m_e/m_p . This gives an expression which is convenient to use for comparison with optical and molecular vibrational or rotational transitions

$$V(^{133}\text{Cs}) = \alpha^{2.83} \left(\frac{m_q}{\Lambda_{QCD}}\right)^{0.073} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{0.006} \frac{m_e}{\Lambda_{QCD}}.$$
 (36)

The dependence on the strange quark mass is relatively weak. Therefore it may be convenient to assume that the relative variation of the strange quark mass is the same as the relative variation of the light quark masses (this assumption is motivated by the Higgs mechanism of mass generation) and to use an approximate expression $V(^{133}\text{Cs}) \approx \alpha^{2.83} (m_q / \Lambda_{QCD})^{0.13} (m_e / m_p)$.

For hyperfine transition frequencies in other atoms we obtain

$$V(^{87}\text{Rb}) = \alpha^{2.34} \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.064} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{-0.010} \frac{m_e}{m_p}, \qquad (37)$$

$$V(^{1}\mathrm{H}) = \alpha^{2} \left(\frac{m_{q}}{\Lambda_{QCD}}\right)^{-0.087} \left(\frac{m_{s}}{\Lambda_{QCD}}\right)^{-0.013} \frac{m_{e}}{m_{p}},$$
(38)

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$$V(^{2}\mathrm{H}) = \alpha^{2} \left(\frac{m_{q}}{\Lambda_{QCD}}\right)^{-0.018} \left(\frac{m_{s}}{\Lambda_{QCD}}\right)^{-0.045} \frac{m_{e}}{m_{p}}, \quad (39)$$

$$V(^{199}\text{Hg}^{+}) = \alpha^{4.3} \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.118} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{0.0013} \frac{m_e}{m_p},$$
(40)

$$V(^{171}\mathrm{Yb}^{+}) = \alpha^{3.5} \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.118} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{0.0013} \frac{m_e}{m_p},$$
(41)

$$V(^{111}\text{Cd}^+) = \alpha^{2.6} \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.118} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{0.0013} \frac{m_e}{m_p}.$$
(42)

Note that the hyperfine frequencies of all even-Z atoms where the nuclear magnetic moment is determined by a valence neutron have the same dependence on quark masses.

IV. LIMITS ON VARIATION OF FUNDAMENTAL CONSTANTS

Now we can use these results to place limits on the possible variation of the fundamental constants from particular measurements. Let us start from the measurements of quasar absorption spectra. Comparison of the atomic H 21 cm (hyperfine) transition with molecular rotational transitions [9] gave limits for the variation of $Y_g \equiv \alpha^2 g_p$. In Refs. [5,37] it was suggested that one might use these limits to estimate the variation of m_q / Λ_{QCD} . According to Eqs. (25) and (26) the relative variation of $Y (\delta Y/Y = \delta Y_g / Y_g)$,

$$Y = \alpha^2 \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.087} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{-0.013}.$$
 (43)

Then the measurements in Ref. [9] lead to the following limits on the variation of *Y*: $\delta Y/Y = (-0.20 \pm 0.44) 10^{-5}$ for redshift z = 0.2467 and $\delta Y/Y = (-0.16 \pm 0.54) 10^{-5}$ for z = 0.6847.

The second limit corresponds to roughly t=6 billion years ago. There is also a limit on the variation of $X_m \equiv \alpha^2 g_p m_e/m_p$ obtained in Ref. [10]. This limit was interpreted as a limit on the variation of α or m_e/m_p . The relative variation of X_m can be replaced by the relative variation of

$$X = \alpha^2 \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.124} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{-0.024} \frac{m_e}{\Lambda_{QCD}}.$$
 (44)

The dependence on quark masses appears from both the proton *g* factor and the proton mass. The measurement in Ref. [10] leads to the following limit on the variation of *X*: $\delta X/X = (0.7 \pm 1.1)10^{-5}$ for z = 1.8.

Now let us discuss the limits obtained from the laboratory measurements of the time dependence of hyperfine structure intervals. The dependence of the ratio of frequencies $A(^{133}Cs)/A(^{87}Rb)$ can be presented in the following form:

$$X(\text{Cs/Rb}) = \frac{V(\text{Cs})}{V(\text{Rb})} = \alpha^{0.49} [m_q / \Lambda_{QCD}]^{0.174} [m_s / \Lambda_{QCD}]^{0.027}$$
(45)

and the result of the measurement in Ref. [18] may be presented as a limit on variation of the parameter *X*:

$$\frac{1}{X(\text{Cs/Rb})} \frac{dX(\text{Cs/Rb})}{dt} = (0.2 \pm 7) \times 10^{-16} / \text{yr.}$$
(46)

Note that if the relation (1) were correct, the variation of X(Cs/Rb) would be dominated by the variation of $[m_q/\Lambda_{QCD}]$. The relation (1) would give $X(Cs/Rb) \propto \alpha^8$.

For $\tilde{A}(^{133}Cs)/A(H)$ we have

$$X(Cs/H) = \frac{V(Cs)}{V(H)} = \alpha^{0.83} [m_q / \Lambda_{QCD}]^{0.196} [m_s / \Lambda_{QCD}]^{0.030}$$
(47)

and the result of the measurements in Ref. [16] may be presented as

$$\left|\frac{1}{X(\text{Cs/H})}\frac{dX(\text{Cs/H})}{dt}\right| < 5.5 \times 10^{-14}/\text{yr.}$$
(48)

For $A(^{199}\text{Hg})/A(\text{H})$ we have

$$X(\text{Hg/H}) = \frac{V(\text{Hg})}{V(\text{H})} \approx \alpha^{2.3} [m_q / \Lambda_{QCD}]^{-0.031} [m_s / \Lambda_{QCD}]^{0.015}.$$
(49)

The result of the measurement in Ref. [15] may be presented as

$$\left|\frac{1}{X(\text{Hg/H})}\frac{dX(\text{Hg/H})}{dt}\right| < 8 \times 10^{-14}/\text{yr.}$$
(50)

Note that because the dependence on masses and the strong interaction scale is very weak here, this experiment may be interpreted as a limit on the variation of α .

In Ref. [14] a limit was obtained on the variation of the ratio of hyperfine transition frequencies¹⁷¹Yb⁺/¹³³Cs (this limit is based on the measurements of Ref. [38]). Using Eqs. (34), (41) we can present the result as a limit on X(Yb/Cs) = $\alpha^{0.7} [m_q / \Lambda_{QCD}]^{-0.228} [m_s / \Lambda_{QCD}]^{-0.015}$:

$$\frac{1}{X(Yb/Cs)} \frac{dX(Yb/Cs)}{dt} \approx -1(2) \times 10^{-13}/yr.$$
 (51)

The optical clock transition energy E(Hg) ($\lambda = 282$ nm) in the Hg⁺ ion can be presented in the following form:

$$E(\mathrm{Hg}) = \mathrm{const} \times \left(\frac{m_e e^4}{\hbar^2}\right) F_{rel}(Z\alpha).$$
 (52)

Numerical calculation of the relative variation of E(Hg) has given [17]

$$\frac{\delta E(\text{Hg})}{E(\text{Hg})} = -3.2 \frac{\delta \alpha}{\alpha}.$$
(53)

This corresponds to $V(\text{Hg Opt}) = \alpha^{-3.2}$. Variation of the ratio of the Cs hyperfine splitting A(Cs) to this optical transition energy is described by X(Opt) = V(Cs)/V(Hg Opt):

$$X(\text{Opt}) = \alpha^6 \left(\frac{m_q}{\Lambda_{QCD}}\right)^{0.073} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{0.006} \left(\frac{m_e}{\Lambda_{QCD}}\right). \quad (54)$$

Here we used Eq. (36) for V(Cs). The work of Ref. [20] gives the limit on variation of this parameter:

$$\left|\frac{1}{X(\text{Opt})}\frac{dX(\text{Opt})}{dt}\right| < 7 \times 10^{-15}/\text{yr.}$$
(55)

Molecular vibrational transitions frequencies are proportional to $(m_e/m_p)^{1/2}$. Based on Eq. (35) we may describe the relative variation of vibrational frequencies by the parameter

$$V(\text{vib}) = \left(\frac{m_q}{\Lambda_{QCD}}\right)^{-0.018} \left(\frac{m_s}{\Lambda_{QCD}}\right)^{-0.005} \left(\frac{m_e}{\Lambda_{QCD}}\right)^{0.5}.$$
 (56)

Comparison of the Cs hyperfine standard with SF₆ molecular vibration frequencies was discussed in Ref. [39]. In this case $X(\text{Cs/Vib}) = \alpha^{2.8} [m_e / \Lambda_{QCD}]^{0.5} [m_q / \Lambda_{QCD}]^{0.091} (m_s / \Lambda_{QCD})^{0.011}$.

The measurements of hyperfine constant ratios in different isotopes of the same atom depends on the ratio of magnetic moments and is therefore sensitive to m_q/Λ_{QCD} . For example, it would be interesting to measure the rate of change for hydrogen/deuterium ratio where $X(H/D) = [m_q/\Lambda_{QCD}]^{-0.068} [m_s/\Lambda_{QCD}]^{0.032}$.

Walsworth has suggested that one might measure the ratio of the Zeeman transition frequencies in noble gases in order to explore the time dependence of the ratio of nuclear magnetic moments. Consider, for example ¹²⁹Xe/³He. For ³He the magnetic moment is very close to that of neutron. For other noble gases the nuclear magnetic moment is also given by the valence neutron, however, there are significant manybody corrections. For ¹²⁹Xe the valence neutron is in an $s_{1/2}$ state, which corresponds to the single-particle value of the nuclear magnetic moment, $\mu = \mu_n = -1.913$. The measured value is $\mu = -0.778$. The magnetic moment of the nucleus changes most efficiently through the spin-spin interaction, because the valence neutron transfers a part of its spin, $\langle s_z \rangle$, to the core protons and the proton magnetic moment is large and has the opposite sign. In this approximation $\mu = (1$ $(-b)\mu_n + b\mu_p$. This gives b = 0.24 and the ratio of magnetic moments $Y \equiv \mu ({}^{129}\text{Xe})/\mu ({}^{3}\text{He}) \approx 0.76 + 0.24g_p/g_n$. Using Eqs. (25)-(28) we obtain an estimate for the relative variation of $\mu(^{129}\text{Xe})/\mu(^{3}\text{He})$, which can be presented as variation of $X = [m_q / \Lambda_{QCD}]^{-0.027} [m_s / \Lambda_{QCD}]^{0.012}$. Here again $\delta Y/Y = \delta X/X$.

Note that the accuracy of the results presented in this paper depends strongly on the fundamental constant under study. The accuracy for the dependence on α is a few percent. The accuracy for m_q/Λ_{QCD} is about 30%—being limited mainly by the accuracy of the single-particle approximation for nuclear magnetic moments. (For comparison, the estimated systematic error associated with the calculation of the effect of the quark mass variation is less than 10%.) Finally, we stress that the relation (1) between the variation of α and m/Λ_{QCD} has been used solely for purposes of illustration.

TABLE I. Chiral coefficients for various diagrams contributing to proton and neutron magnetic moments. We use SU(6) symmetry to relate the meson couplings to the $\pi N\Delta$ vertex, C = -2D.

α	eta^p_{\mulpha}	eta_{\mulpha}^n
(a)	$-(F+D)^2$	$(F+D)^2$
(b)	$-\frac{2}{9}C^2$	$\frac{2}{9}C^2$
(c)	$-\frac{1}{6}(D+3F)^2$	0
(d)	$-\frac{1}{2}(D-F)^2$	$-(D-F)^2$

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APPENDIX MAGNETIC MOMENTS

As explained in the text, we explicitly include the processes shown in Fig. 1, which give rise to the leading and next-to-leading nonanalytic behavior as a function of quark mass.

We describe the quark mass dependence of the magnetic moments as

$$\mu = \frac{\alpha_0}{1 + \alpha_2 m_\pi^2} + M^{\mathrm{L}},\tag{A1}$$

where M^{L} denotes the chiral loop corrections given by

$$M^{\mathrm{L}} = \chi_{\mu(a)} I_{\mu}(m_{\pi}, 0, \Lambda) + \chi_{\mu(b)} I_{\mu}(m_{\pi}, \Delta_{N\Delta}, \Lambda) + \chi_{\mu(c)} I_{\mu}(m_{K}, \Delta_{N\Lambda}, \Lambda) + \chi_{\mu(d)} I_{\mu}(m_{K}, \Delta_{N\Sigma}, \Lambda).$$
(A2)

The chiral coefficients of the loop integrals $\chi_{\mu\alpha}$ are given by

$$\chi_{\mu\alpha} = \beta_{\mu\alpha} \frac{M_N}{8\pi f_\pi^2} \tag{A3}$$

and are summarized in Table I [40-42]. Note that the required analytic terms in the chiral expansion to this order have been placed in a Padé approximant designed to reproduce the Dirac moment behavior of the nucleon at moderate quark mass.

The corresponding loop integral is given by

$$I_{\mu}(m,\Delta,\Lambda) = -\frac{4}{3\pi} \int_0^\infty dk \frac{(\Delta+2\omega_k)k^4 u^2(k,\Lambda)}{2\omega_k^3 (\Delta+\omega_k)^2},$$
(A4)

where the various terms have been defined in Sec. II. We note that in the limit where the mass splitting vanishes this integral is normalized such that the leading nonanalytic contribution is m.

With the coefficients of the loop integrals defined, we only require determination of the parameters α_0 and α_2 in Eq. (A1) to constrain the variation with quark mass. We note also that this form assumes no analytic dependence on the strange quark mass, beyond what is implicitly included in the loop diagrams (c,d). We determine $\alpha_{0,2}$ for both the proton and neutron by fitting the physical magnetic moment as well as the lattice QCD data. We fit only to the two heaviest simulation results of the CSSM Lattice Collaboration [34], $m_{\pi}^2 \sim 0.6-0.7 \text{ GeV}^2$. These simulations were performed with the FLIC fermion action [36] on a $20^3 \times 40$ lattice at a = 0.128 fm. We select the heaviest two data points, where the effects of quenching are anticipated to be small [43,44].

The best fits to the physical values and the lattice data give

$$\alpha_0^p = 2.17 \ \mu_N, \ \alpha_2^p = 0.817 \ \text{GeV}^{-2},$$
 (A5)

$$\alpha_0^n = -1.33 \ \mu_N, \ \alpha_2^n = 0.758 \ \text{GeV}^{-2}.$$
 (A6)

- J.K. Webb *et al.*, Phys. Rev. Lett. **82**, 884 (1999); **87**, 091301 (2001); M.T. Murphy, J.K. Webb, and V.V. Flambaum, Mon. Not. R. Astron. Soc. **345**, 609 (2003).
- [2] J-P. Uzan, Rev. Mod. Phys. 75, 403 (2003).
- [3] P. Langacker, G. Segre, and M.J. Strassler, Phys. Lett. B 528, 121 (2002); see also X. Calmet and H. Fritzsch, Eur. Phys. J. C 24, 639 (2002); W.J. Marciano, Phys. Rev. Lett. 52, 489 (1984).
- [4] T. Dent, hep-ph/0305026.
- [5] V.V. Flambaum and E.V. Shuryak, Phys. Rev. D 65, 103503 (2002).
- [6] K.A. Olive et al., Phys. Rev. D 66, 045022 (2002).
- [7] V.F. Dmitriev and V.V. Flambaum, Phys. Rev. D 67, 063513 (2003).
- [8] V.V. Flambaum and E.V. Shuryak, Phys. Rev. D 67, 083507 (2003).
- [9] M.T. Murphy *et al.*, Mon. Not. R. Astron. Soc. **327**, 1244 (2001).
- [10] L.L. Cowie and A. Songalia, Astrophys. J. 453, 596 (1995).
- [11] A.I. Shlyakhter, Nature (London) 264, 340 (1976); T. Damour and F.J. Dyson, Nucl. Phys. B480, 37 (1996); Y. Fujii *et al.*, *ibid.* B573, 377 (2000).
- [12] H. Oberhummer, R. Pichler, and A. Csoto, nucl-th/9810057.
- [13] S.R. Beane and M.J. Savage, Nucl. Phys. A713, 148 (2003).
- [14] S.G. Karshenboim, Can. J. Phys. 78, 639 (2000).
- [15] J.D. Prestage, R.L. Tjoelker, and L. Maleki, Phys. Rev. Lett. 74, 3511 (1995).
- [16] N. A. Demidov et al., in Proceedings of the 6th European Frequency and Time Forum, Noordwijk, the Netherlands, 1992 (European Space Agency, Noordwijk, 1992), pp. 409–414; L. A. Breakiron, in Proceedings of the 25th Annual Precise Time Interval Applications and Planning Meeting, NASA Conference Publication No. 3267 [U.S. Naval Observatory Time Service Department (TSS1), Washington, DC, 1993], pp. 401– 412.
- [17] V.A. Dzuba, V.V. Flambaum, and J.K. Webb, Phys. Rev. A 59,

Upon renormalization of the loop diagrams, the resultant magnetic moments in the SU(2) chiral limit are given by

$$\mu_0^p = 3.48 \ \mu_N$$
, and $\mu_0^n = -2.58 \ \mu_N$. (A7)

We now take derivatives of Eq. (A1) at the physical pion mass to determine the variation with quark mass. In particular, we have

$$\frac{\delta\mu}{\mu} = \left\{ \frac{m_{\pi}^2}{\mu} \frac{d\mu}{d\,m_{\pi}^2} \right\} \frac{\delta m_q}{m_q},\tag{A8}$$

$$\frac{\delta\mu}{\mu} = \left\{ \frac{\tilde{m}_K^2}{\mu} \frac{d\mu}{d\tilde{m}_K^2} \right\} \frac{\delta m_s}{m_s}.$$
 (A9)

This yields the results shown in the text.

230 (1999); V. A. Dzuba (private communication).

- [18] H. Marion et al., Phys. Rev. Lett. 90, 150801 (2003).
- [19] D.B. Leinweber, A.W. Thomas, and R.D. Young, Phys. Rev. Lett. 86, 5011 (2001).
- [20] S. Bize et al., Phys. Rev. Lett. 90, 150802 (2003).
- [21] C. Bernard, S. Hashimoto, D.B. Leinweber, P. Lepage, E. Pallante, S.R. Sharpe, and H. Wittig, Nucl. Phys. B, Proc. Suppl. 119, 170 (2003).
- [22] A.W. Thomas, Nucl. Phys. B, Proc. Suppl. 119, 50 (2003).
- [23] D.B. Leinweber, A.W. Thomas, K. Tsushima, and S.V. Wright, Phys. Rev. D 61, 074502 (2000).
- [24] D.B. Leinweber, D.H. Lu, and A.W. Thomas, Phys. Rev. D 60, 034014 (1999).
- [25] T.R. Hemmert, M. Procura, and W. Weise, Nucl. Phys. A721, 938 (2003).
- [26] T.R. Hemmert, M. Procura, and W. Weise, hep-lat/0301005.
- [27] V. Bernard, T.R. Hemmert, and U.G. Meissner, Nucl. Phys. A732, 149 (2004).
- [28] QCDSF Collaboration, M. Gockeler et al., hep-lat/0303019.
- [29] W. Detmold, D.B. Leinweber, W. Melnitchouk, A.W. Thomas, and S.V. Wright, Pramana, J. Phys. 57, 251 (2001).
- [30] J.F. Donoghue, B.R. Holstein, and B. Borasoy, Phys. Rev. D 59, 036002 (1999).
- [31] D.B. Leinweber, A.W. Thomas, and R.D. Young, Phys. Rev. Lett. (to be published).
- [32] R.D. Young, D.B. Leinweber, and A.W. Thomas, Prog. Part. Nucl. Phys. 50, 399 (2003).
- [33] CP-PACS Collaboration, A. Ali Khan *et al.*, Phys. Rev. D 65, 054505 (2002); 67, 059901(E) (2003).
- [34] J.M. Zanotti, D.B. Leinweber, A.G. Williams, and J.B. Zhang, Nucl. Phys. B (Proc. Suppl.) 129, 287 (2004).
- [35] D.B. Leinweber, R.M. Woloshyn, and T. Draper, Phys. Rev. D 43, 1659 (1991).
- [36] CSSM Lattice Collaboration, J.M. Zanotti *et al.*, Phys. Rev. D 65, 074507 (2002).

- [37] V.F. Dmitriev and V.V. Flambaum, Phys. Rev. D 67, 063513 (2003).
- [38] Chr. Tamm, D. Schnier, and A. Bauch, Appl. Phys. B: Lasers Opt. 60, 19 (1995); P.T.H. Fisk, M.J. Sellar, M.A. Lawn, and C. Coles, IEEE Trans. Ultrason. Ferroelectr. Freq. Control 44, 344 (1997); P.T.H. Fisk, Rep. Prog. Phys. 60, 761 (1997).
- [39] C. Chardonnet, Talk at ICOLS 03, Palm Cove, July (2003).
- [40] S. Theberge, G.A. Miller, and A.W. Thomas, Can. J. Phys. 60,

59 (1982); A.W. Thomas, Adv. Nucl. Phys. 13, 1 (1984).

- [41] E. Jenkins, M.E. Luke, A.V. Manohar, and M.J. Savage, Phys. Lett. B 302, 482 (1993); 388, 866(E) (1996).
- [42] D.B. Leinweber, Phys. Rev. D 69, 014005 (2004).
- [43] R.D. Young, D.B. Leinweber, A.W. Thomas, and S.V. Wright, Phys. Rev. D 66, 094507 (2002).
- [44] R.D. Young, D.B. Leinweber, and A.W. Thomas, Nucl. Phys. B (Proc. Suppl.) 128, 227 (2004).