# MHD OBLIQUE STAGNATION-POINT FLOW OF A MICROPOLAR FLUID 

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#### Abstract

The steady two-dimensional oblique stagnation-point flow of an electrically conducting micropolar fluid in the presence of a uniform external electromagnetic field $\left(\mathbf{E}_{0}, \mathbf{H}_{0}\right)$ is analyzed and some physical situations are examined. In particular, if $\mathbf{E}_{0}$ vanishes, $\mathbf{H}_{0}$ lies in the plane of the flow, with a direction not parallel to the boundary, and the induced magnetic field is neglected. It is proved that the oblique stagnationpoint flow exists if, and only if, the external magnetic field is parallel to the dividing streamline. In all cases it is shown that the governing nonlinear partial differential equations admit similarity solutions and the resulting ordinary differential problems are solved numerically. Finally, the behaviour of the flow near the boundary is analyzed; this depends on the three dimensionless material parameters, and also on the Hartmann number if $\mathbf{H}_{0}$ is parallel to the dividing streamline. $76 \mathrm{~W} 05,76 \mathrm{D} 10$. Micropolar fluids, MHD flow, oblique stagnation-point flow.


## 1. Introduction

The micropolar fluids introduced by Eringen ([9]) physically represent fluids consisting of rigid randomly oriented particles suspended in a viscous medium which have an intrinsic rotational micromotion (for example biological fluids in thin vessels, polymeric suspensions, slurries, colloidal fluids). Extensive reviews of the theory and its applications can be found in [10] and [20].
The aim of this paper is to study how the steady two dimensional oblique stagnation-point flow of an electrically conducting micropolar fluid is influenced by a uniform external electromagnetic field $\left(\mathbf{E}_{0}, \mathbf{H}_{0}\right)$. The motions we find depend upon how the applied electromagnetic field is oriented relative to the flat boundary.
Oblique stagnation-point flow appears when a jet of fluid impinges obliquely on a rigid wall at an arbitrary angle of incidence. From a mathematical point of view, such a flow is obtained by combining orthogonal stagnation-point flow with a shear flow parallel to the wall. The steady two-dimensional oblique stagnation-point flow of a Newtonian fluid has been object of many investigations starting from the paper of Stuart in 1959 ([24]). We refer to [28] and [8] for a review.
The orthogonal plane and axially symmetric stagnation-point flow of a micropolar fluid have been treated by Guram and Smith ([17]), who reduced the equations to dimensionless form, including three dimensionless parameters and integrated them numerically. Previously Ahmadi ([1]) obtained self-similar solutions of the boundary layer equations for micropolar flow imposing restrictive conditions on the material parameters which make the equations to contain only one parameter. This restrictive approach has been followed in [18] and [19] in order to study the oblique stagnation-point flow in the absence of an external electromagnetic field.

[^0]In this paper we extend the results of [3] about Newtonian fluids to incompressible homogeneous micropolar fluids assuming potential flow far from the boundary and prescribing the strict adherence condition on the flat plane boundary.
First of all, we summarize the results of [3] concerning an inviscid fluid, and analyze three cases, which are significant from a physical point of view. In the first two cases, an external constant field, either electric or magnetic, is impressed parallel to the rigid wall. In both cases, we have found that an oblique stagnation-point flow exists, and we obtained the exact induced magnetic field. In the third case, we suppose that $\mathbf{E}_{0}$ vanishes and $\mathbf{H}_{0}$ lies in the plane of the flow, with a direction not parallel to the boundary. Under the hypothesis that the magnetic Reynolds number is small, we neglect the induced magnetic field, as it is customary in the literature. We have proved that the oblique stagnation-point flow exists if, and only if, $\mathbf{H}_{0}$ is parallel to the dividing streamline.

Then we consider the same problems for a micropolar fluid, assuming that at infinity, the velocity $\mathbf{v}$ approaches the flow of an inviscid fluid for which the stagnation-point is shifted from the origin ([8], [26], and [23]), and the microrotation $\mathbf{w}$ is given by $\mathbf{w}=\frac{1}{2} \nabla \times \mathbf{v}$, i.e. the micropolar fluid behaves like a classical fluid far from the wall. The coordinates of this stagnation-point contain two constants : $A$ and $B . A$ is determined as part of the solution of the orthogonal flow, and $B$ is free.
As far as the velocity and the microrotation are concerned, in the first two cases we find the same equations of the oblique stagnation-point flow in the absence of an electromagnetic field, while the induced magnetic field is obtained by direct integration. Hence, the external uniform electromagnetic field doesn't influence the flow, and modifies only the pressure $p$. Moreover $\nabla p$ has a constant component parallel to the wall proportional to $B-A$. This does not appear in the orthogonal stagnation-point flow. This component determines the displacement parallel to the boundary of the uniform shear flow. The flow is obtained for different values of $B$ and of the material parameters by numerical integration using a finitedifferences method.
We remark that the influence of the viscosity appears only in a layer lining the boundary whose thickness is larger than that in the orthogonal stagnation-point flow.
Finally, in the more general case in which $\mathbf{H}_{0}$ is parallel to the dividing streamline of the inviscid flow, we find that the flow has to satisfy an ordinary differential problem whose solution depend on $\mathbf{H}_{0}$ through the Hartmann number $M$. The numerical integration is provided using a finite-differences method. In this case, $A$ (and so the stagnation-point) depends on $M$ and decreases as $M$ is increased. Further, when the material parameters are fixed, the influence of the viscosity appears only in a layer near to the wall depending on $M$ whose thickness decreases as $M$ increases from zero. This is standard in magnetohydrodynamics.
Some numerical examples and pictures are given in order to illustrate the effects due to the magnetic field.

The paper is organized in this way:
In Section 2 we summarize the results in [3] for an inviscid fluid.
Section 3 is devoted to treat the same physical problems for a micropolar fluid. Theorems $1,2,3$ collect our results.
Further, we study the behaviour of the flow near the wall. We show that it depends on the three dimensionless material parameters, and also on the Hartmann number in the third case.
In Section 4, we numerically integrate the previous problems, and discuss some numerical results.

## 2. Inviscid Fluids

Consider the MHD steady plane stagnation-point flow of an inviscid, homogeneous, incompressible, electrically conducting fluid filling the region $\mathcal{S}$, given by

$$
\begin{equation*}
\mathcal{S}=\left\{\mathbf{x} \in \mathbb{R}^{3}:\left(x_{1}, x_{3}\right) \in \mathbb{R}^{2}, x_{2}>0\right\} \tag{1}
\end{equation*}
$$

The boundary of $\mathcal{S}$ is a rigid, fixed, non-electrically conducting wall.
The equations governing such a flow in the absence of external mechanical body forces are:

$$
\begin{align*}
& \rho \mathbf{v} \cdot \nabla \mathbf{v}=-\nabla p+\mu_{e}(\nabla \times \mathbf{H}) \times \mathbf{H} \\
& \nabla \cdot \mathbf{v}=0 \\
& \nabla \times \mathbf{H}=\sigma_{e}\left(\mathbf{E}+\mu_{e} \mathbf{v} \times \mathbf{H}\right), \\
& \nabla \times \mathbf{E}=\mathbf{0}, \quad \nabla \cdot \mathbf{E}=0, \quad \nabla \cdot \mathbf{H}=0, \quad \text { in } \mathcal{S} \tag{2}
\end{align*}
$$

where $\mathbf{v}$ is the velocity field, $p$ is the pressure, $\mathbf{E}$ and $\mathbf{H}$ are the electric and magnetic fields, respectively, $\rho$ is the mass density (constant $>0$ ), $\mu_{e}$ is the magnetic permeability, $\sigma_{e}$ is the electrical conductivity ( $\mu_{e}, \sigma_{e}=$ constants $>0$ ). We assume the region

$$
\mathcal{S}^{-}=\left\{\mathbf{x} \in \mathbb{R}^{3}:\left(x_{1}, x_{3}\right) \in \mathbb{R}^{2}, x_{2}<0\right\}
$$

to be a vacuum (free space), and $\mu_{e}$ equal to the magnetic permeability of free space.
To equations (2) we append the usual boundary condition for $\mathbf{v}$ :

$$
\begin{equation*}
v_{2}=0 \quad \text { at } \quad x_{2}=0 \tag{3}
\end{equation*}
$$

Further, suppose that the tangential components of $\mathbf{H}$ and $\mathbf{E}$ are continuous through the plane $x_{2}=0$.

We are interested in the oblique plane stagnation-point flow, so that

$$
\begin{equation*}
v_{1}=a x_{1}+b x_{2}, \quad v_{2}=-a x_{2}, \quad v_{3}=0, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+} \tag{4}
\end{equation*}
$$

with $a, b$ constants $(a>0)$.
As known, the streamlines of such a flow are hyperbolas whose asymptotes have the equations:

$$
x_{2}=0 \quad \text { and } \quad x_{2}=-\frac{2 a}{b} x_{1} .
$$

These two straight-lines are degenerate streamlines too.
We summarize our results ([3]) concerning the influence upon such a flow of a uniform external electromagnetic field $\left(\mathbf{E}_{0}, \mathbf{H}_{0}\right)$. To this end, we consider three cases which, from a physical point of view, are significant.

### 2.1. CASE I.

$$
\mathbf{E}_{0}=E_{0} \mathbf{e}_{3}, \quad \mathbf{H}_{0}=\mathbf{0}
$$

Let the induced electromagnetic field $\left(\mathbf{E}^{i}, \mathbf{H}^{i} \equiv \mathbf{H}\right)$ be in the form

$$
\begin{aligned}
& \mathbf{E}^{i}=E_{1}^{i} \mathbf{e}_{1}+E_{2}^{i} \mathbf{e}_{2}+E_{3}^{i} \mathbf{e}_{3} \\
& \mathbf{H}=h\left(x_{2}\right) \mathbf{e}_{1}
\end{aligned}
$$

where $\left(\mathbf{e}_{1}, \mathbf{e}_{2}, \mathbf{e}_{3}\right)$ is the canonical base in $\mathbb{R}^{3}$.
The boundary conditions require that

$$
\begin{align*}
& E_{1}^{i}=0, E_{3}^{i}=0 \text { at } x_{2}=0 \\
& h(0)=0 \tag{5}
\end{align*}
$$



Figure 1. Flow description in CASE I.

From (2) ${ }_{4},(2)_{3}$ follows that

$$
\begin{equation*}
h\left(x_{2}\right)=-\sigma_{e} E_{0} e^{-\frac{a x_{2}^{2}}{2 \eta_{e}}} \int_{0}^{x_{2}} e^{\frac{a t^{2}}{2 \eta_{e}}} d t, \quad x_{2} \in \mathbb{R}^{+} \tag{6}
\end{equation*}
$$

with $\eta_{e}=\frac{1}{\sigma_{e} \mu_{e}}=$ electrical resistivity.
As far as the pressure field is concerned, from (2) $)_{1}$ we get

$$
p=-\frac{1}{2} \rho a^{2}\left(x_{1}^{2}+x_{2}^{2}\right)-\frac{\mu_{e}}{2} h^{2}\left(x_{2}\right)+p_{0}, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+}
$$

where $h$ is given by (6) and $p_{0}$ is the pressure in the stagnation point.

### 2.2. CASE II.



Figure 2. Flow description in CASE II.

$$
\mathbf{E}_{0}=\mathbf{0}, \quad \mathbf{H}_{0}=H_{0} \mathbf{e}_{1} .
$$

Let the induced electromagnetic field $\left(\mathbf{E}^{i} \equiv \mathbf{E}, \mathbf{H}^{i}\right)$ be in the form

$$
\begin{aligned}
& \mathbf{E}=E_{1} \mathbf{e}_{1}+E_{2} \mathbf{e}_{2}+E_{3} \mathbf{e}_{3}, \\
& \mathbf{H}=\left[h\left(x_{2}\right)+H_{0}\right] \mathbf{e}_{1} .
\end{aligned}
$$

We append the boundary conditions (5).
In this case we proved in [3] that

$$
\begin{aligned}
& \mathbf{E}=\mathbf{0} \\
& \mathbf{H}=H_{0} e^{-\frac{a x_{2}^{2}}{2 \eta_{e}}} \mathbf{e}_{1}, \\
& p\left(x_{1}, x_{2}\right)=-\frac{1}{2} \rho a^{2}\left(x_{1}^{2}+x_{2}^{2}\right)-\frac{\mu_{e}}{2} H_{0}^{2} e^{-\frac{a x_{2}^{2}}{\eta_{e}}}+p_{0}, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+} .
\end{aligned}
$$

### 2.3. CASE III.



Figure 3. Flow description in CASE III.

$$
\mathbf{E}_{0}=\mathbf{0}, \quad \mathbf{H}_{0}=H_{0}\left(\cos \vartheta \mathbf{e}_{1}+\sin \vartheta \mathbf{e}_{2}\right)
$$

with $\vartheta$ fixed in $(0, \pi)$.
Under the hypothesis that the magnetic Reynolds number is small, we neglect the induced magnetic field, as it is customary in the literature. We proved that $\mathbf{E}=\mathbf{0}$, and that the MHD oblique stagnation-point flow is possible if and only if $\mathbf{H}_{0}$ is parallel to the dividing streamline, i.e.

$$
\begin{equation*}
\tan \vartheta=-\frac{2 a}{b} \tag{7}
\end{equation*}
$$

Moreover the pressure field has the form

$$
p=-\frac{1}{2} \rho a^{2}\left(x_{1}^{2}+x_{2}^{2}\right)-\frac{\sigma_{e} B_{0}^{2}}{4 a^{2}+b^{2}} \frac{a}{2}\left(2 a x_{1}+b x_{2}\right)^{2}+p_{0}, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+} .
$$

Remark 1. In order to study oblique stagnation-point flow for micropolar fluids, it is convenient to consider a more general flow. More precisely, we suppose the fluid obliquely impinging on the flat plane $x_{2}=A$ and

$$
\begin{equation*}
v_{1}=a x_{1}+b\left(x_{2}-B\right), \quad v_{2}=-a\left(x_{2}-A\right), \quad v_{3}=0, \quad x_{1} \in \mathbb{R}, \quad x_{2} \geq A \tag{8}
\end{equation*}
$$

with $A, B=$ positive constants.
In this way, the stagnation point is $\left(\frac{b}{a}(B-A), A\right)$.
In this case, the streamlines are the hyperbolas whose asymptotes are $x_{2}=-\frac{2 a}{b} x_{1}+2 B-A$ and $x_{2}=A$.
As it is easy to verify, in the absence of $(\mathbf{E}, \mathbf{H})$, the pressure field is given by:

$$
p=-\frac{1}{2} \rho a^{2}\left\{\left[x_{1}-\frac{b}{a}(B-A)\right]^{2}+\left(x_{2}-A\right)^{2}\right\}+p_{0}
$$

We underline that under these new assumptions, all previous results continue to hold by replacing $x_{1}, x_{2}$ with $x_{1}-\frac{b}{a}(B-A), x_{2}-A$ respectively.

## 3. Micropolar fluids

Consider the steady two-dimensional oblique stagnation-point flow of an electrically conducting homogeneous incompressible micropolar fluid towards a flat surface coinciding with the plane $x_{2}=0$, the flow being confined to the region $\mathcal{S}$. In the absence of external mechanical body forces and body couples, the MHD equations for such a fluid are ([20])

$$
\begin{align*}
& \mathbf{v} \cdot \nabla \mathbf{v}=-\frac{1}{\rho} \nabla p+\left(\nu+\nu_{r}\right) \triangle \mathbf{v}+2 \nu_{r}(\nabla \times \mathbf{w})+\frac{\mu_{e}}{\rho}(\nabla \times \mathbf{H}) \times \mathbf{H} \\
& \nabla \cdot \mathbf{v}=0, \\
& I \mathbf{v} \cdot \nabla \mathbf{w}=\lambda \triangle \mathbf{w}+\lambda_{0} \nabla(\nabla \cdot \mathbf{w})-4 \nu_{r} \mathbf{w}+2 \nu_{r}(\nabla \times \mathbf{v})=\mathbf{0} \tag{9}
\end{align*}
$$

together with $(2)_{3}-(2)_{6}$, where $\mathbf{w}$ is the microrotation field, $\nu$ is the kinematic newtonian viscosity coefficient, $\nu_{r}$ is the microrotation viscosity coefficient, $\lambda, \lambda_{0}$ (positive constants) are material parameters related to the coefficient of angular viscosity and $I$ is the microinertia coefficient.
We notice that in [9], [10], eqs. (9) are slightly different, as they are deduced as a special case of much more general model of microfluids. For the details, we refer to [20], p.23.
As far as the boundary conditions are concerned, of course, we modify condition (3) and prescribe the appropriate boundary condition for the microrotation $\mathbf{w}$, i.e.

$$
\begin{equation*}
\left.\mathbf{v}\right|_{x_{2}=0}=\mathbf{0},\left.\quad \mathbf{w}\right|_{x_{2}=0}=\mathbf{0} \text { (strict adherence condition). } \tag{10}
\end{equation*}
$$

Other boundary conditions are possible. We refer to Eringen ([9], p.17-18) for a complete discussion. In our studies we will always assume the strict adherence condition.
We search $\mathbf{v}, \mathbf{w}$ in the following form

$$
\begin{align*}
& v_{1}=a x_{1} f^{\prime}\left(x_{2}\right)+b g\left(x_{2}\right), \quad v_{2}=-a f\left(x_{2}\right), \quad v_{3}=0, \\
& w_{1}=0, \quad w_{2}=0, \quad w_{3}=x_{1} F\left(x_{2}\right)+G\left(x_{2}\right), \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+} \tag{11}
\end{align*}
$$

where $f, g, F, G$ are unknown functions.
The conditions (10) supply

$$
\begin{align*}
& f(0)=0, \quad f^{\prime}(0)=0, \quad g(0)=0 \\
& F(0)=0, \quad G(0)=0 \tag{12}
\end{align*}
$$

Moreover, as is customary when studying the oblique plane stagnation-point flow for viscous
fluids, we assume that at infinity, the flow approaches the flow of an inviscid fluid given by (8) ([8], [23], and [26]).

Therefore, to (11) we must append also the following conditions

$$
\begin{align*}
& \lim _{x_{2} \rightarrow \infty} f^{\prime}\left(x_{2}\right)=1, \quad \lim _{x_{2} \rightarrow \infty} g^{\prime}\left(x_{2}\right)=1, \\
& \lim _{x_{2} \rightarrow \infty} F\left(x_{2}\right)=0, \quad \lim _{x_{2} \rightarrow \infty} G\left(x_{2}\right)=-\frac{b}{2} . \tag{13}
\end{align*}
$$

Conditions $(13)_{3,4}$ mean that at infinity, $\mathbf{w}=\frac{1}{2} \nabla \times \mathbf{v}$, i.e. the micropolar fluid behaves like a classical fluid.
In all the following cases, when we will refer to inviscid fluid, all results have to be modified by replacing $x_{1}, x_{2}$ with $x_{1}-\frac{b}{a}(B-A), x_{2}-A$ respectively.
In particular the asymptotic behaviour of $f$ and $g$ at infinity is related to the constants $A, B$, in the following way:

$$
\begin{equation*}
f \sim x_{2}-A, \quad g \sim x_{2}-B \quad \text { as } x_{2} \rightarrow \infty \tag{14}
\end{equation*}
$$

As we will see, $A$ is determined as part of the solution of the orthogonal flow, while $B$ is a free parameter.
In order to study the influence of a uniform external electromagnetic field, we consider the three cases analyzed in the previous section.
3.1. CASE I-M. By proceeding as for an inviscid fluid, from $(2)_{3},(2)_{4}$ and boundary conditions for electromagnetic field, we obtain $\mathbf{E}=E_{0} \mathbf{e}_{3}$ and the induced magnetic field $h\left(x_{2}\right)$ satisfies

$$
\begin{equation*}
h^{\prime}+\frac{a}{\eta_{e}} f h=-\eta_{e} E_{0}, \quad x_{2}>0, \quad h(0)=0 \tag{15}
\end{equation*}
$$

If we regard $f$ as a known function, we arrive at

$$
\begin{equation*}
h\left(x_{2}\right)=-\sigma_{e} E_{0} e^{-\frac{a}{\eta_{e}} \int_{0}^{x_{2}} f(t) d t} \int_{0}^{x_{2}} e^{\frac{a}{\eta_{e}} \int_{0}^{s} f(t) d t} d s, \quad x_{2} \in \mathbb{R}^{+} \tag{16}
\end{equation*}
$$

As is easy to verify, the induced magnetic fields given by (16) and (6) have the same asymptotic behaviour at infinity $\left(\sim-\frac{\eta_{e} E_{0} \sigma_{e}}{a\left(x_{2}-A\right)}\right)$.
In order to determine $p, f, g, F, G$ we substitute (11) in (9) $)_{1,3}$. After some calculations, we arrive at

$$
\begin{align*}
& p=p\left(x_{1}, x_{2}\right) \\
& a x_{1}\left[\left(\nu+\nu_{r}\right) f^{\prime \prime \prime}+a f f^{\prime \prime}-a f^{\prime 2}+\frac{2 \nu_{r}}{a} F^{\prime}\right]+b\left[\left(\nu+\nu_{r}\right) g^{\prime \prime}+a\left(f g^{\prime}-f^{\prime} g\right)+\frac{2 \nu_{r}}{b} G^{\prime}\right]=\frac{1}{\rho} \frac{\partial p}{\partial x_{1}}, \\
& \left(\nu+\nu_{r}\right) a f^{\prime \prime}+a^{2} f^{\prime} f+2 \nu_{r} F+\frac{\mu_{e}}{\rho} h^{\prime} h=-\frac{1}{\rho} \frac{\partial p}{\partial x_{2}} \\
& x_{1}\left[\alpha F^{\prime \prime}+I a\left(F^{\prime} f-F f^{\prime}\right)-2 \nu_{r}\left(2 F+a f^{\prime \prime}\right)\right]+\lambda G^{\prime \prime}+I\left(a G^{\prime} f-b F g\right)-2 \nu_{r}\left(2 G+b g^{\prime}\right)=0 . \tag{17}
\end{align*}
$$

Then, by integrating $(17)_{3}$, we find

$$
p\left(x_{1}, x_{2}\right)=-\frac{1}{2} \rho a^{2} f^{2}\left(x_{2}\right)-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right)-2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s-\frac{\mu_{e}}{2} h^{2}\left(x_{2}\right)+P\left(x_{1}\right)
$$

where the function $P\left(x_{1}\right)$ is determined supposing that, far from the wall, the pressure $p$ has the same behaviour as for an inviscid electroconducting fluid, whose velocity is given by (8).

Therefore, since the induced electromagnetic fields given by (16), (6) have the same asymptotic behaviour, under the assumption $F \in L^{1}([0,+\infty)$ ), by virtue of (13), (14), we get

$$
P\left(x_{1}\right)=-\rho \frac{a^{2}}{2}\left[x_{1}-\frac{b}{a}(B-A)\right]^{2}+p_{0}+\rho a\left(\nu+\nu_{r}\right) .
$$

Finally, the pressure field assumes the form

$$
\begin{align*}
p\left(x_{1}, x_{2}\right)= & -\rho \frac{a^{2}}{2}\left[x_{1}^{2}-2 \frac{b}{a}(B-A) x_{1}+f^{2}\left(x_{2}\right)\right]-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right) \\
& -2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s-\frac{\mu_{e}}{2} h^{2}\left(x_{2}\right)+p_{0}^{*} \tag{18}
\end{align*}
$$

with $p_{0}^{*}=p_{0}+\rho a\left(\nu+\nu_{r}\right)-\rho \frac{b^{2}}{2}(B-A)^{2}$.
In consideration of (18), we obtain the ordinary differential system

$$
\begin{align*}
& \frac{\nu+\nu_{r}}{a} f^{\prime \prime \prime}+f f^{\prime \prime}-f^{\prime 2}+1+\frac{2 \nu_{r}}{a^{2}} F^{\prime}=0 \\
& \frac{\nu+\nu_{r}}{a} g^{\prime \prime}+f g^{\prime}-g f^{\prime}+\frac{2 \nu_{r}}{a b} G^{\prime}=B-A \\
& \lambda F^{\prime \prime}+a I\left(f F^{\prime}-f^{\prime} F\right)-2 \nu_{r}\left(2 F+a f^{\prime \prime}\right)=0 \\
& \lambda G^{\prime \prime}+I\left(a f G^{\prime}-b g F\right)-2 \nu_{r}\left(2 G+b g^{\prime}\right)=0 \tag{19}
\end{align*}
$$

To these equations we append the boundary conditions (12), (13).
We remark that the system (19) governs the oblique stagnation-point flow of an inert, electromagnetic micropolar fluid, as is easy to verify. In literature, such a flow has been studied in [18], and [19] under restrictive assumptions upon the material parameters, and following a different approach.

Remark 2. If $\nu_{r}=0$, then $(19)_{1}$ and $(19)_{2}$ are the equations governing the oblique stagnation-point flow of a Newtonian fluid.
We observe that $(19)_{1}$ and $(19)_{3}$ have the same form as the equations found by Guram and Smith ([17]) for the orthogonal stagnation-point flow of a micropolar fluid.

Theorem 1. Let a homogeneous, incompressible, electrically conducting micropolar fluid occupy the region $\mathcal{S}$. The steady MHD oblique plane stagnation-point flow of such a fluid has the following form when an external uniform electric field $\mathbf{E}_{0}=E_{0} \mathbf{e}_{3}$ is impressed:

$$
\begin{aligned}
\mathbf{v}= & {\left[a x_{1} f^{\prime}\left(x_{2}\right)+b g\left(x_{2}\right)\right] \mathbf{e}_{1}-a f\left(x_{2}\right) \mathbf{e}_{2}, \quad \mathbf{H}=h\left(x_{2}\right) \mathbf{e}_{1}, \quad \mathbf{E}=E_{0} \mathbf{e}_{3}, } \\
\mathbf{w}= & {\left[x_{1} F\left(x_{2}\right)+G\left(x_{2}\right)\right] \mathbf{e}_{3}, } \\
p= & -\rho \frac{a^{2}}{2}\left[x_{1}^{2}-2 \frac{b}{a}(B-A) x_{1}+f^{2}\left(x_{2}\right)\right]-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right) \\
& -2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s-\frac{\mu_{e}}{2} h^{2}\left(x_{2}\right)+p_{0}^{*} \\
& x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+},
\end{aligned}
$$

where $(f, g, F, G)$ satisfies the problem (19), (12), and (13), provided $F \in L^{1}([0,+\infty))$, and $h\left(x_{2}\right)$ is given by (16).

Now we write the system (19), together with the conditions (12) (13), in dimensionless form. To this end we put

$$
\begin{aligned}
& \eta=\sqrt{\frac{a}{\nu+\nu_{r}}} x_{2}, \quad \phi(\eta)=\sqrt{\frac{a}{\nu+\nu_{r}}} f\left(\sqrt{\frac{\nu+\nu_{r}}{a}} \eta\right) \\
& \gamma(\eta)=\sqrt{\frac{a}{\nu+\nu_{r}}} g\left(\sqrt{\frac{\nu+\nu_{r}}{a}} \eta\right), \quad \Phi(\eta)=\frac{2 \nu_{r}}{a^{2}} \sqrt{\frac{a}{\nu+\nu_{r}}} F\left(\sqrt{\frac{\nu+\nu_{r}}{a}} \eta\right), \\
& \Gamma(\eta)=\frac{2 \nu_{r}}{b\left(\nu+\nu_{r}\right)} G\left(\sqrt{\frac{\nu+\nu_{r}}{a}} \eta\right), \quad \Psi(\eta)=\frac{1}{\eta_{e} E_{0}} \sqrt{\frac{a}{\nu+\nu_{r}}} h\left(\sqrt{\frac{\nu+\nu_{r}}{a}} \eta\right) .
\end{aligned}
$$

So system (19) and equation (15) can be written as

$$
\begin{align*}
& \phi^{\prime \prime \prime}+\phi \phi^{\prime \prime}-\phi^{2}+1+\Phi^{\prime}=0 \\
& \gamma^{\prime \prime}+\phi \gamma^{\prime}-\phi^{\prime} \gamma+\Gamma^{\prime}=\beta-\alpha \\
& \Phi^{\prime \prime}+c_{3}\left(\phi \Phi^{\prime}-\phi^{\prime} \Phi\right)-c_{2} \Phi-c_{1} \phi^{\prime \prime}=0 \\
& \Gamma^{\prime \prime}+c_{3}\left(\phi \Gamma^{\prime}-\Phi \gamma\right)-c_{2} \Gamma-c_{1} \gamma^{\prime}=0 \\
& \Psi^{\prime}+R_{m} \phi \Psi=-1 \tag{20}
\end{align*}
$$

where

$$
\begin{aligned}
& c_{1}=\frac{4 \nu_{r}^{2}}{\lambda a}, \quad c_{2}=\frac{4 \nu_{r}\left(\nu+\nu_{r}\right)}{\lambda a}, \quad c_{3}=\frac{I}{\lambda}\left(\nu+\nu_{r}\right) \\
& \alpha=\sqrt{\frac{a}{\nu+\nu_{r}}} A, \quad \beta=\sqrt{\frac{a}{\nu+\nu_{r}}} B, \quad R_{m}=\frac{\nu+\nu_{r}}{\eta_{e}}=\text { magnetic Reynolds number. }
\end{aligned}
$$

The boundary conditions in dimensionless form become:

$$
\begin{align*}
& \phi(0)=0, \quad \phi^{\prime}(0)=0, \quad \gamma(0)=0, \\
& \Phi(0)=0, \quad \Gamma(0)=0, \\
& \Psi(0)=0, \\
& \lim _{\eta \rightarrow+\infty} \phi^{\prime}(\eta)=1, \quad \lim _{\eta \rightarrow+\infty} \gamma^{\prime}(\eta)=1 \\
& \lim _{\eta \rightarrow+\infty} \Phi(\eta)=0, \quad \lim _{\eta \rightarrow+\infty} \Gamma(\eta)=-\frac{c_{1}}{c_{2}} . \tag{21}
\end{align*}
$$

The last equation in (20), if we regard $\phi$ as a known function, can be formally integrated to give

$$
\Psi(\eta)=-e^{-R_{m} \int_{0}^{\eta} \phi(s) d s} \int_{0}^{\eta} e^{R_{m} \int_{0}^{t} \phi(s) d s} d t, \quad \eta \in \mathbb{R}^{+}
$$

The remaining equations have to be integrated numerically.

Remark 3. Along the wall $x_{2}=0$, there are three important coordinates: the origin $x_{1}=0$ towards which the dividing streamline at infinity is pointed, the point $x_{1}=x_{p}$ of maximum pressure, and the point $x_{1}=x_{s}$ of zero tangential stress (zero skin friction) where the dividing streamline of equation

$$
\begin{equation*}
\xi \phi(\eta)+\frac{b}{a} \int_{0}^{\eta} \gamma(s) d s=0, \quad \xi=\sqrt{\frac{\nu+\nu_{r}}{a}} x_{1} \tag{22}
\end{equation*}
$$

meets the boundary.
In consideration of (18), we see that

$$
\begin{equation*}
x_{p}=b \sqrt{\frac{\nu+\nu_{r}}{a^{3}}}(\beta-\alpha) \tag{23}
\end{equation*}
$$

and so $x_{p}$ does not depend on $h$.
The wall shear stress is given by

$$
\tau=\left.\rho\left(\nu+\nu_{r}\right) \frac{\partial v_{1}}{\partial x_{2}}\right|_{x_{2}=0}
$$

the position $x_{s}$ is obtained by putting $\tau=0$. Hence

$$
\begin{equation*}
x_{s}=-b \sqrt{\frac{\nu+\nu_{r}}{a^{3}}} \frac{\gamma^{\prime}(0)}{\phi^{\prime \prime}(0)} . \tag{24}
\end{equation*}
$$

We note that the ratio $\frac{x_{p}}{x_{s}}=(\alpha-\beta) \frac{\phi^{\prime \prime}(0)}{\gamma^{\prime}(0)}$ is the same for all angles of incidence.
Finally, we recall that studying the small- $\eta$ behaviour of $\frac{\int_{0}^{\eta} \gamma(s) d s}{\phi(\eta)}$, the slope of the dividing streamline at the wall is given by:

$$
m_{s}=-\frac{3 a\left[\phi^{\prime \prime}(0)\right]^{2}}{b\left\{\left[\beta-\alpha-\Gamma^{\prime}(0)\right] \phi^{\prime \prime}(0)+\left[1+\Phi^{\prime}(0)\right] \gamma^{\prime}(0)\right\}}
$$

and does not depend on the kinematic viscosities. Thus, the ratio of this slope to that of the dividing streamline at infinity $\left(m_{i}=-\frac{2 a}{b}\right)$ is the same for all oblique stagnation-point flows and is given by

$$
\begin{equation*}
\frac{m_{s}}{m_{i}}=\frac{3}{2} \frac{\left[\phi^{\prime \prime}(0)\right]^{2}}{\left[\beta-\alpha-\Gamma^{\prime}(0)\right] \phi^{\prime \prime}(0)+\left[1+\Phi^{\prime}(0)\right] \gamma^{\prime}(0)} \tag{25}
\end{equation*}
$$

This ratio is independent of $a$ and $b$, depending only upon the constant pressure gradient parallel to the boundary through $B-A$, as with Newtonian fluids ([7]).
3.2. CASE II-M. By proceeding as one would with an inviscid fluid, from $(2)_{3},(2)_{4}$ and boundary conditions for electromagnetic field, we get

$$
\begin{equation*}
h^{\prime}+\frac{a}{\eta_{e}} f h=-\frac{a}{\eta_{e}} f H_{0}, \quad x_{2}>0, \quad h(0)=0 . \tag{26}
\end{equation*}
$$

The integration of (26) leads to

$$
\begin{equation*}
h\left(x_{2}\right)=H_{0}\left(e^{-\frac{a}{\eta_{e}} \int_{0}^{x_{2}} f(t) d t}-1\right), \quad x_{2} \in \mathbb{R}^{+}, \tag{27}
\end{equation*}
$$

so that

$$
\mathbf{H}=H_{0} e^{-\frac{a}{\eta_{e}} \int_{0}^{x_{2}} f(s) d s} \mathbf{e}_{1}
$$

The pressure field, as is easy to verify, becomes

$$
\begin{align*}
p\left(x_{1}, x_{2}\right)= & -\rho \frac{a^{2}}{2}\left[x_{1}^{2}-2 \frac{b}{a}(B-A) x_{1}+f^{2}\left(x_{2}\right)\right]-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right) \\
& -2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s-\frac{\mu_{e}}{2}\left[h\left(x_{2}\right)+H_{0}\right]^{2}+p_{0}^{*} \tag{28}
\end{align*}
$$

where $(f, g, F, G)$ satisfies system (19), together with boundary conditions (12) and (13). Therefore, in this case as well, the uniform external electromagnetic field does not influence the flow.

Thus, we obtain the following:
Theorem 2. Let a homogeneous, incompressible, electrically conducting micropolar fluid occupy the region $\mathcal{S}$. The steady MHD oblique plane stagnation-point flow of such a fluid has the following form when a uniform external magnetic field $\mathbf{H}_{0}=H_{0} \mathbf{e}_{1}$ is impressed:

$$
\begin{aligned}
\mathbf{v}= & {\left[a x_{1} f^{\prime}\left(x_{2}\right)+b g\left(x_{2}\right)\right] \mathbf{e}_{1}-a f\left(x_{2}\right) \mathbf{e}_{2}, \quad \mathbf{H}=\left[H_{0}+h\left(x_{2}\right)\right] \mathbf{e}_{1}, \quad \mathbf{E}=E_{0} \mathbf{e}_{3}, } \\
\mathbf{w}= & {\left[x_{1} F\left(x_{2}\right)+G\left(x_{2}\right)\right] \mathbf{e}_{3}, } \\
p= & -\rho \frac{a^{2}}{2}\left[x_{1}^{2}-2 \frac{b}{a}(B-A) x_{1}+f^{2}\left(x_{2}\right)\right]-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right)-2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s \\
& -\frac{\mu_{e}}{2}\left[h\left(x_{2}\right)+H_{0}\right]^{2}+p_{0}^{*}, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+},
\end{aligned}
$$

where $(f, g, F, G)$ satisfies the problem (19), (12), and (13), provided $F \in L^{1}([0,+\infty))$, and $h\left(x_{2}\right)$ is given by (27).

In dimensionless form, $h\left(x_{2}\right)$ becomes

$$
\begin{equation*}
\Psi(\eta)=e^{-R_{m} \int_{0}^{\eta} \phi(t) d t}-1, \quad \eta \in \mathbb{R}^{+} \tag{29}
\end{equation*}
$$

where

$$
\Psi(\eta)=\frac{1}{H_{0}} h\left(\sqrt{\frac{\nu+\nu_{r}}{a}} \eta\right)
$$

Of course, remark 3 continues to hold in this case.
3.3. CASE III-M. Taking into account the results obtained for an inviscid fluid, we assume

$$
\mathbf{H}_{0}=\frac{H_{0}}{\sqrt{4 a^{2}+b^{2}}}\left(-b \mathbf{e}_{1}+2 a \mathbf{e}_{2}\right), \quad \mathbf{E}_{0}=\mathbf{0}
$$

As in CASE III, for inviscid fluid, we deduce

$$
\mathbf{E}=\mathbf{0} \Rightarrow \nabla \times \mathbf{H}=\sigma_{e} \mu_{e}(\mathbf{v} \times \mathbf{H})
$$

Further, we neglect the induced magnetic field, replacing (9) $)_{1}$ with

$$
\begin{equation*}
\mathbf{v} \cdot \nabla \mathbf{v}=-\frac{1}{\rho} \nabla p+\left(\nu+\nu_{r}\right) \triangle \mathbf{v}+2 \nu_{r}(\nabla \times \mathbf{w})+\frac{\mu_{e}}{\rho}\left(\mathbf{v} \times \mathbf{H}_{\mathbf{0}}\right) \times \mathbf{H}_{\mathbf{0}} \tag{30}
\end{equation*}
$$

This approximation is motivated by physical arguments for MHD flow at small magnetic Reynolds numbers.

We substitute (11) into (30) to determine $p, f, g, F, G$. This yields

$$
\begin{align*}
& p=p\left(x_{1}, x_{2}\right), \\
& a x_{1}\left[\left(\nu+\nu_{r}\right) f^{\prime \prime \prime}+a f f^{\prime \prime}-a f^{\prime 2}+\frac{2 \nu_{r}}{a} F^{\prime}-4 a^{2} \frac{\sigma_{e}}{\rho} \frac{B_{0}^{2}}{4 a^{2}+b^{2}} f^{\prime}\right]+ \\
& +b\left[\left(\nu+\nu_{r}\right) g^{\prime \prime}+a\left(f g^{\prime}-f^{\prime} g\right)+\frac{2 \nu_{r}}{b} G^{\prime}-2 a^{2} \frac{\sigma_{e}}{\rho} \frac{B_{0}^{2}}{4 a^{2}+b^{2}}(2 g-f)\right]=\frac{1}{\rho} \frac{\partial p}{\partial x_{1}} \\
& \left(\nu+\nu_{r}\right) a f^{\prime \prime}+a^{2} f^{\prime} f+2 \nu_{r} F+\frac{\sigma_{e}}{\rho} \frac{B_{0}^{2}}{4 a^{2}+b^{2}}\left[2 a^{2} b x_{1} f^{\prime}+a b^{2}(2 g-f)\right]=-\frac{1}{\rho} \frac{\partial p}{\partial x_{2}} \tag{31}
\end{align*}
$$

The integration of $(31)_{3}$ gives

$$
\begin{aligned}
p\left(x_{1}, x_{2}\right)= & -\frac{1}{2} \rho a^{2} f^{2}\left(x_{2}\right)-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right)-2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s \\
& -\sigma_{e} \frac{B_{0}^{2}}{4 a^{2}+b^{2}}\left\{2 a^{2} b x_{1} f\left(x_{2}\right)+a b^{2} \int_{0}^{x_{2}}[2 g(s)-f(s)] d s\right\}+P\left(x_{1}\right)
\end{aligned}
$$

where $P\left(x_{1}\right)$ has to be found as in CASES I-M, II-M.
After some calculations, we obtain

$$
P\left(x_{1}\right)=-\rho \frac{a^{2}}{2}\left(1+\frac{4 a}{\rho} \frac{\sigma_{e} B_{0}^{2}}{4 a^{2}+b^{2}}\right)\left[x_{1}-\frac{b}{a}(B-A)\right]^{2}+p_{0}^{\prime}
$$

with $p_{0}^{\prime}$ constant.
So the pressure field is:

$$
\begin{align*}
p\left(x_{1}, x_{2}\right)= & -\rho \frac{a^{2}}{2}\left[x_{1}^{2}-2 \frac{b}{a}(B-A) x_{1}+f^{2}\left(x_{2}\right)\right]-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right) \\
& -2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s-\frac{\sigma_{e} B_{0}^{2}}{4 a^{2}+b^{2}}\left\{2 a^{2} b x_{1} f\left(x_{2}\right)+\int_{0}^{x_{2}}[2 g(s)-f(s)] d s\right. \\
& \left.+2 a^{3}\left[x_{1}^{2}-\frac{2 b}{a}(B-A) x_{1}\right]\right\}+p_{0}^{*}, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+} . \tag{32}
\end{align*}
$$

The constant $p_{0}^{*}$ represents the pressure at the stagnation point.
Then, (31) ${ }_{2}$ supplies

$$
\begin{align*}
& \frac{\nu+\nu_{r}}{a} f^{\prime \prime \prime}+f f^{\prime \prime}-f^{\prime 2}+1+\frac{2 \nu_{r}}{a^{2}} F^{\prime}+M^{2}\left(1-f^{\prime}\right)=0 \\
& \frac{\nu+\nu_{r}}{a} g^{\prime \prime}+f g^{\prime}-g f^{\prime}+\frac{2 \nu_{r}}{a b} G^{\prime}+M^{2}(f-g)=\left(1+M^{2}\right)(B-A) \tag{33}
\end{align*}
$$

where $M^{2}=4 a \frac{\sigma_{e} B_{0}^{2}}{\rho\left(4 a^{2}+b^{2}\right)}$ is the Hartmann number.
Of course $(f, g, F, G)$ also satisfies the equations $(19)_{3,4}$. We append boundary conditions (12) and (13) to the system in (33) and (19) $)_{3,4}$.

We remark that, unlike the previous cases, the external electromagnetic field modifies the flow; if $M=0$, then the system (33) and (19) $)_{3,4}$ reduces to the system (19).

Theorem 3. Let a homogeneous, incompressible, electrically conducting micropolar fluid occupy the region $\mathcal{S}$. If we impress the external magnetic field

$$
\mathbf{H}_{0}=\frac{H_{0}}{\sqrt{4 a^{2}+b^{2}}}\left(-b \mathbf{e}_{1}+2 a \mathbf{e}_{2}\right)
$$

and if we neglect the induced magnetic field, then the steady MHD oblique plane stagnationpoint flow of such a fluid has the form

$$
\begin{aligned}
\mathbf{v}= & {\left[a x_{1} f^{\prime}\left(x_{2}\right)+b g\left(x_{2}\right)\right] \mathbf{e}_{1}-a f\left(x_{2}\right) \mathbf{e}_{2}, } \\
\mathbf{w}= & {\left[x_{1} F\left(x_{2}\right)+G\left(x_{2}\right)\right] \mathbf{e}_{3}, \quad \mathbf{E}=\mathbf{0}, } \\
p= & -\rho \frac{a^{2}}{2}\left[x_{1}^{2}-2 \frac{b}{a}(B-A) x_{1}+f^{2}\left(x_{2}\right)\right]-\rho a\left(\nu+\nu_{r}\right) f^{\prime}\left(x_{2}\right) \\
& -2 \nu_{r} \rho \int_{0}^{x_{2}} F(s) d s-\frac{\sigma_{e} B_{0}^{2}}{4 a^{2}+b^{2}}\left\{2 a^{2} b x_{1} f\left(x_{2}\right)+\int_{0}^{x_{2}}[2 g(s)-f(s)] d s\right. \\
& \left.+2 a^{3}\left[x_{1}^{2}-\frac{2 b}{a}(B-A) x_{1}\right]\right\}+p_{0}^{*}, \quad x_{1} \in \mathbb{R}, \quad x_{2} \in \mathbb{R}^{+},
\end{aligned}
$$

where $(f, g, F, G)$ satisfies problem (33), (19) $)_{3,4}$, (12), and (13), provided $F \in L^{1}([0,+\infty)$ ).
In dimensionless form, we arrive at the following ordinary differential problem:

$$
\begin{align*}
& \phi^{\prime \prime \prime}+\phi \phi^{\prime \prime}-\phi^{\prime 2}+1+\Phi^{\prime}+M^{2}\left(1-\phi^{\prime}\right)=0 \\
& \gamma^{\prime \prime}+\phi \gamma^{\prime}-\phi^{\prime} \gamma+\Gamma^{\prime}+M^{2}(\phi-\gamma)=\left(1+M^{2}\right)(\beta-\alpha), \\
& \Phi^{\prime \prime}+c_{3}\left(\phi \Phi^{\prime}-\phi^{\prime} \Phi\right)-c_{2} \Phi-c_{1} \phi^{\prime \prime}=0, \\
& \Gamma^{\prime \prime}+c_{3}\left(\phi \Gamma^{\prime}-\Phi \gamma\right)-c_{2} \Gamma-c_{1} \gamma^{\prime}=0 . \tag{34}
\end{align*}
$$

To system (34), we append boundary conditions (21).
Problem (34) and (21) will be solved numerically in Section 4 for some values $c_{1}, c_{2}, c_{3}, M$.
Remark 4. The points $x_{1}=x_{p}$ of maximum pressure and $x_{1}=x_{s}$ of zero tangential stress on $x_{2}=0$ are formally the same as in CASE I-M and II-M. However, these points depend on $M$.
Finally the slope of the dividing streamline at the wall is given by:

$$
m_{s}=-\frac{3 a\left[\phi^{\prime \prime}(0)\right]^{2}}{b\left\{\left[(\beta-\alpha)\left(1+M^{2}\right)-\Gamma^{\prime}(0)\right] \phi^{\prime \prime}(0)+\left[1+M^{2}+\Phi^{\prime}(0)\right] \gamma^{\prime}(0)\right\}}
$$

## 4. Numerical results and discussion

In this section we discuss the numerical solutions of the problems studied in CASES I, II, III-M.
4.1. CASE I-M. We have solved problem (20), (21) numerically by using a difference finite algorithm.
The values of the parameters $c_{1}, c_{2}, c_{3}$ were choosen according to Guram and Smith ([17]) and are given in Table 1, where we also assign some values to $\beta$ (i.e. $\beta-\alpha=-\alpha, 0, \alpha$ ). The consequent values of $\alpha, \phi^{\prime \prime}(0), \gamma^{\prime}(0), \Phi^{\prime}(0), \Gamma^{\prime}(0), \frac{x_{p}}{x_{s}}, \frac{m_{s}}{m_{i}}, \bar{\eta}_{\phi}, \bar{\eta}_{\gamma}$ are reported in this table. We denote by $\bar{\eta}_{\phi}$ the value of $\eta$ at which $\phi^{\prime}=0.99$ ( so when $\eta>\bar{\eta}_{\phi}$, then $\phi \sim \eta-\alpha$ ), while $\bar{\eta}_{\gamma}$ is the values of $\eta$ at which $\gamma^{\prime}=0.99$ (if $\beta-\alpha \geq 0$ ). or $\gamma^{\prime}=1.01$ (if $\beta-\alpha<0$ ). So when $\eta>\bar{\eta}_{\gamma}$ then $\gamma \sim \eta-\beta$.

Table 1

| $c_{1}$ | $c_{2}$ | $c_{3}$ | $\alpha$ | $\beta-\alpha$ | $\phi^{\prime \prime}(0)$ | $\gamma^{\prime}(0)$ | $\Phi^{\prime}(0)$ | $\Gamma^{\prime}(0)$ | $\frac{x_{p}}{x_{s}}$ | $\frac{m_{s}}{m_{i}}$ | $\bar{\eta}_{\phi}$ | $\bar{\eta}_{\gamma}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.1 | 1.5 | 0.1 | 0.6446 | -0.6446 | 1.2218 | 1.3647 | -0.0532 | -0.0892 | 0.5771 | 3.6492 | 2.5556 | 3.0667 |
|  |  |  |  | 0 | $"$ | 0.5771 | $"$ | -0.0550 | 0 | 3.6495 | $"$ | 3.1944 |
|  |  |  |  | 0.6446 | $"$ | -0.2105 | $"$ | -0.0207 | 3.7415 | 3.6498 | $"$ | 3.2370 |
|  |  | 0.5 | 0.6448 | -0.6448 | 1.2231 | 1.3651 | -0.0510 | -0.0889 | 0.5777 | 3.6455 | 2.5556 | 3.0667 |
|  |  |  |  | 0 | $"$ | 0.5765 | $"$ | -0.0560 | 0 | 3.6455 | $"$ | 3.1944 |
|  | 3.0 | 0.1 | 0.6453 | -0.6448 | $"$ | -0.2121 | $"$ | -0.0231 | 3.7177 | 3.6455 | $"$ | 3.2370 |
|  |  |  |  | 0 | $" .2250$ | 1.3817 | -0.0444 | -0.0658 | 0.5721 | 3.6871 | 2.5556 | 3.0667 |
|  |  |  |  | 0.6453 | $"$ | 0.5912 | $"$ | -0.0372 | 0 | 3.6872 | $"$ | 3.1944 |
|  |  | 0.5 | 0.6454 | -0.6454 | 1.2256 | 1.3822 | -0.0434 | -0.0085 | 3.9656 | 3.6872 | $"$ | 3.2370 |
|  |  |  |  | 0 | $"$ | 0.5912 | $"$ | -0.0372 | 0.5723 | 3.6871 | 0.5556 | 3.0667 |
|  |  |  |  | 0.6454 | $"$ | -0.1998 | $"$ | -0.0092 | 3.9577 | 3.6871 | $"$ | 3.1944 |
| 0.5 | 0.1 | 0.6311 | -0.6311 | 1.1780 | 1.1970 | -0.2659 | -0.4280 | 0.6211 | 3.2546 | 2.2148 | 2.2370 |  |
|  |  |  |  | 0 | $"$ | 0.4534 | $"$ | -0.2602 | 0 | 3.2560 | $"$ | 2.7259 |
|  |  |  |  | 0.6311 | $"$ | -0.2903 | $"$ | -0.0923 | 2.5611 | 3.2574 | $"$ | 2.7259 |
|  |  | 0.5 | 0.6321 | -0.6321 | 1.1848 | 1.1987 | -0.2553 | -0.4275 | 0.6248 | 3.2381 | 2.2148 | 2.5556 |
|  |  |  | 0 | $"$ | 0.4498 | $"$ | -0.2661 | 0 | 3.2383 | $"$ | 2.7259 |  |
|  |  |  |  | 0.6321 | $"$ | -0.2991 | $"$ | -0.1047 | 2.5034 | 3.2385 | $"$ | 2.8963 |
|  | 3.0 | 0.1 | 0.6351 | -0.6351 | 1.1943 | 1.2825 | -0.2220 | -0.3200 | 0.5914 | 3.4426 | 2.3852 | 2.9389 |
|  |  |  |  | 0 | $"$ | 0.5240 | $"$ | -0.1790 | 0 | 3.4429 | $"$ | 3.0667 |
|  |  |  |  | 0.6351 | $"$ | -0.2345 | $"$ | -0.0380 | 3.2343 | 3.4433 | $"$ | 3.0667 |
|  |  | 0.5 | 0.6356 | -0.6356 | 1.1972 | 1.2846 | -0.2173 | -0.3174 | 0.5923 | 3.4426 | 2.3852 | 2.9389 |
|  |  |  |  | 0 | $"$ | 0.5237 | $"$ | -0.1793 | 0 | 3.4426 | $"$ | 3.1596 |
|  |  |  |  | 0.6356 | $"$ | -0.2372 | $"$ | -0.0412 | 3.2074 | 3.4427 | $"$ | 3.0667 |

We see that $\bar{\eta}_{\gamma}$ is always greater than $\bar{\eta}_{\phi}$, as in the Newtonian case ([3]). Hence the influence of the viscosity on the velocity appears only in a layer of thickness $\bar{\eta}_{\gamma}$ lining the boundary. We remark that the thickness of the layer affected by the viscosity is proportional to $\sqrt{\frac{\nu+\nu_{r}}{a}}$ and it is larger than that in the orthogonal stagnation-point flow.
From Table 1 it appears that if we fix two parameters among $c_{1}, c_{2}, c_{3}$, then the values of $\alpha, \phi^{\prime \prime}(0), \gamma^{\prime}(0), \Phi^{\prime}(0), \Gamma^{\prime}(0)$ have the following behaviour :
they increase as $c_{2}$ increases,
they lower as $c_{1}$ or $c_{3}$ increases.
Moreover, the influence of $c_{1}$ appears more considerable also on the other quantities quoted in the table.
We have displayed some representative graphs to elucidate the trends of the functions describing the velocities.
In particular, Figures 4,5 , and 6 show $\phi, \phi^{\prime}, \phi^{\prime \prime}, \Phi, \Phi^{\prime}, \gamma, \gamma^{\prime}, \Gamma, \Gamma^{\prime}$ for $c_{1}=0.5, c_{2}=$ $3.0, c_{3}=0.5$. The other choices of these parameters modify the trends of these functions very slightly.
Of course, the behaviour of $\phi, \Phi$ doesn't depend on $\beta-\alpha$, unlike $\gamma, \Gamma$.
If we compare the velocity profile with the solution for classical viscous flow ([3]), we note that the trend is very similar, as was found in [17] for orthogonal stagnation-point flow.
Figures 7, 8, and 9 elucidate the dependence of the functions $\phi^{\prime}, \gamma, \Phi, \Gamma$ on the parameters
$c_{1}, c_{2}, c_{3}$. We can see that the functions which appear most influenced by $c_{1}, c_{2}, c_{3}$ are $\Phi$, and $\Gamma$ - in other words the microrotation. More precisely the profile of $\Phi$ rises as $c_{3}$ or $c_{2}$ increases and $c_{1}$ decreases, while the profile of $\Gamma$ rises as $c_{2}$ increases and $c_{1}$ or $c_{3}$ decreases. Moreover $c_{1}$ is the parameter that most influences the microrotation. The other two functions, $\phi^{\prime}$, and $\gamma$, do not show considerable variations as $c_{1}, c_{2}, c_{3}$ assume different values.


Figure 4. CASE I-M: plots showing the behaviour of $\phi, \phi^{\prime}, \phi^{\prime \prime}$ and $\Phi, \Phi^{\prime}$ respectively for $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$.


Figure 5. CASE I-M: Figures $5_{1}$ and $5_{2}$ show $\gamma$ and $\gamma^{\prime}$ for $c_{1}=0.5, c_{2}=$ $3.0, c_{3}=0.5$ and with, from above, $\beta-\alpha=-\alpha, 0, \alpha$, respectively.

Observing $\phi^{\prime \prime}(0), \gamma^{\prime}(0)$ in Table 1 we notice that $x_{s}$ (given by (24)) has the sign of $b$ if $\beta-\alpha>0$ and the sign of $-b$ if $\beta-\alpha \leq 0$. Moreover if $b$ is positive (negative) $x_{s}$ increases (decreases) as $\beta-\alpha$ increases. As far as $\left|x_{s}\right|$ is concerned, if $\beta-\alpha$ increases from a negative value to zero, $\left|x_{s}\right|$ decreases and so $x_{s}$ approaches the origin, otherwise, as $\beta-\alpha$ increases from zero to a positive value, $\left|x_{s}\right|$ increases and so $x_{s}$ departs from the origin. The same results were also found for Newtonian fluids in [3].


Figure 6. CASE I-M: Figures $6_{1}$ and $6_{2}$ show $\Gamma$ and $\Gamma^{\prime}$ for $c_{1}=0.5, c_{2}=$ $3.0, c_{3}=0.5$ and with, from above, $\beta-\alpha=-\alpha, 0, \alpha$, respectively.

Moreover from Table 1 we see that $x_{p}$ and $x_{s}$ lie on the same side of the origin, and $\frac{m_{s}}{m_{i}}$ is constant once $c_{1}, c_{2}, c_{3}$ are fixed.
Figure 10 shows the streamlines and the points

$$
\begin{equation*}
\xi_{p}=\sqrt{\frac{\nu}{a}} x_{p}, \quad \xi_{s}=\sqrt{\frac{\nu}{a}} x_{s} \tag{35}
\end{equation*}
$$

for $\frac{b}{a}=1$ and $\beta-\alpha=-\alpha, 0, \alpha$, respectively.
Finally, figure $11_{1}$ shows the behaviour of the induced magnetic field $\Psi$ with $R_{m}=1$, that is similar to the behaviour of $\Psi$ in CASE I (inviscid fluid), as we can see. Figure $11_{2}$ shows the behaviour of $\Psi$ with $R_{m}=10^{-6}$; for $\eta \in[0,4.6]$ the graph is approximately linear, because in this interval, for very small values of $R_{m}$, the equation $(27)_{3}$ reduces to $\Psi^{\prime} \sim-1$.
4.2. CASE II-M. In this case the ordinary differential problem governing $\phi, \gamma, \Phi, \Gamma$ is the same as in (20); we have only to compute $\Psi(\eta)$, given by (29).
Figure 12 shows that $\Psi$ has a similar behaviour as in CASE II, as we can see.
4.3. CASE III-M. We have solved the problem (34), (21) with the boundary conditions (21) by using a finite-differences method.

Table 2 shows the numerical results of the quantities listed in Table 1 for the same $c_{1}, c_{2}, c_{3}, \beta$, and choosing $M=1,2,5,10$.
If we fix $M$, we see that the considerations of CASE I-M continue to hold.
As far as the dependence on $M$ is concerned, we can see that $\alpha$ and $\Phi^{\prime}(0)$ decrease and $\phi^{\prime \prime}(0)$ increases as $M$ is increased from 0 , as we would expect physically.
As far as the dependence of $\gamma^{\prime}(0)$ and $\Gamma^{\prime}(0)$ on $M$ are concerned, from 2 we can see that their values increase as $M$ increases if $\beta-\alpha<0$, otherwise they decrease.

In Figure $13_{1}$, we have plotted the profiles $\phi, \phi^{\prime}, \phi^{\prime \prime}$ for $M=2$ and $c_{1}=0.5, c_{2}=$ $3.0, c_{3}=0.5$, while Figure 16 shows the behaviour of $\phi^{\prime}$ for different $M$ and the same values of $c_{1}, c_{2}, c_{3}$.
Figures $14_{1}, 14_{2}$ show the profiles of $\gamma(\eta), \gamma^{\prime}(\eta)$, for $M=2, c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$


Figure 7. CASE I-M: plots showing the behaviour of $\phi^{\prime}, \gamma, \Phi$ and $\Gamma$ for $c_{1}=0.5, c_{2}=3.0$ fixed, and for different values of $c_{3}$.
and for some values of $\beta-\alpha$, i.e. $\beta-\alpha=-\alpha, 0, \alpha$. In Figures 17 and 18, we provide the behaviour of $\gamma^{\prime}$ for different $M$ when $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and $\beta-\alpha$ is fixed.
In Figure $13_{2}$, we can see the profiles $\Phi, \Phi^{\prime}$ for $M=2$ and $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$.
Figures $15_{1}, 15_{2}$ show the graphics of $\Gamma(\eta), \Gamma^{\prime}(\eta)$, for $M=2, c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and for some values of $\beta-\alpha$, i.e. $\beta-\alpha=-\alpha, 0, \alpha$.
We have only plotted the profiles of $\phi, \phi^{\prime}, \phi^{\prime \prime}, \gamma, \gamma^{\prime}, \Phi, \Phi^{\prime}, \Gamma, \Gamma^{\prime}$ for $M=2$ and $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$, because they have an analogous behaviour for $M \neq 2$ and different $c_{1}, c_{2}, c_{3}$.

In Table 2, we also list the values of $\bar{\eta}_{\phi}, \bar{\eta}_{\gamma}$ beyond which $\phi \sim \eta-\alpha$ and $\gamma \sim \eta-\beta$ respectively. We note that $\bar{\eta}_{\gamma}$ is greater than the corresponding value of $\bar{\eta}_{\phi}$; so the influence of the viscosity appears only in the region $\eta<\bar{\eta}_{\gamma}$, i.e. $x_{2}<\sqrt{\frac{\nu+\nu_{r}}{a}} \bar{\eta}_{\gamma}$. Further we underline that the thickness of this layer depends on $M$ and it decreases as $M$ increases (as easily seen in Figures 16, 17, and 18). This effect is normal in magnetohydrodynamics. Finally, we notice that the points $x_{p}, x_{s}$, given by (23) and by (24), lie on the same side of


Figure 8. CASE I-M: plots showing the behaviour of $\phi^{\prime}, \gamma, \Phi$ and $\Gamma$ for $c_{1}=0.5, c_{3}=0.5$ fixed, and for different values of $c_{2}$.
the origin. Their location depends on $M, c_{1}, c_{2}, c_{3}$ and $\beta-\alpha$, as seen in Table 2. The Figure 19 shows the streamlines and the points $\xi_{p}, \xi_{s}$ for $\frac{b}{a}=1, c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$, $\beta-\alpha=-\alpha, 0, \alpha$, and $M=1,5$.

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Figure 9. CASE I-M: plots showing the behaviour of $\phi^{\prime}, \gamma, \Phi$ and $\Gamma$ for $c_{2}=3.0, c_{3}=0.5$ fixed, and for different values of $c_{1}$.
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Figure 10. CASE I-M: plots showing the streamlines and the points $\xi_{p}, \xi_{s}$ for $\frac{b}{a}=1, c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and $\beta-\alpha=-\alpha, 0, \alpha$, respectively.
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Figure 11. CASE I-M: plots showing $\Psi$ with $R_{m}=1,10^{-6}$.


Figure 12. CASE II-M: plot showing $\Psi$ for $R_{m}=1$

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Figure 13. CASE III-M: plots showing $\phi, \phi^{\prime}, \phi^{\prime \prime}$ and $\Phi, \Phi^{\prime}$ (respectively) for $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and $M=2$.



Figure 14. CASE III-M: plots showing $\gamma, \gamma^{\prime}$ with $M=2, c_{1}=0.5, c_{2}=$ $3.0, c_{3}=0.5$ and, from above, $\beta-\alpha=-\alpha, 0, \alpha$ respectively.

Table 2


| $c_{1}$ | $c_{2}$ | $c_{3}$ | $M$ | $\alpha$ | $\beta-\alpha$ | $\phi^{\prime \prime}(0)$ | $\gamma^{\prime}(0)$ | $\Phi^{\prime}(0)$ | $\Gamma^{\prime}(0)$ | $\frac{x_{p}}{x_{s}}$ | $\frac{m_{s}}{m_{i}}$ | $\bar{\eta}_{\phi}$ | $\bar{\eta}_{\gamma}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.5 | 1.5 | 0.1 | 1 | 0.5290 | -0.5290 | 1.5335 | 1.2465 | -0.2913 | -0.4247 | 0.6508 | 3.0440 | 2.0444 | 2.3000 |
|  |  |  |  |  | 0 | " | 0.4353 | " | -0.2706 | 0 | 3.0442 | " | 2.4278 |
|  |  |  |  |  | 0.5290 | " | -0.3760 | " | -0.1165 | 2.1577 | 3.0443 | " | 2.4278 |
|  |  |  | 2 | 0.3874 | -0.3874 | 2.3005 | 1.3193 | -0.3315 | -0.4208 | 0.6755 | 2.9716 | 1.5333 | 1.6611 |
|  |  |  |  |  | 0 | " | 0.4281 | " | -0.2924 | 0 | 2.9716 | " | 1.9167 |
|  |  |  |  |  | 0.3874 | " | -0.4630 | " | -0.1639 | 1.9246 | 2.9716 | " | 1.9167 |
|  |  |  | 5 | 0.1896 | -0.1896 | 5.1166 | 1.4170 | -0.4022 | -0.4159 | 0.6847 | 2.9810 | 0.8519 | 0.9370 |
|  |  |  |  |  | 0 | " | 0.4467 |  | -0.3396 | 0.0000 | 2.9810 | " | 1.0222 |
|  |  |  |  |  | 0.1896 | " | -0.5236 | " | -0.2634 | 1.8533 | 2.9810 | " | 1.1074 |
|  |  |  | 10 | 0.0986 | -0.0986 | 10.0552 | 1.4581 | -0.4440 | -0.4145 | 0.6798 | 2.9931 | 0.4543 | 0.5111 |
|  |  |  |  |  | 0 | " | 0.4668 | " | -0.3707 | 0 | 2.9931 | " | 0.5537 |
|  |  |  |  |  | 0.0986 | " | -0.5245 | " | -0.3270 | 1.8901 | 2.9930 | " | 0.5963 |
|  |  | 0.5 | 1 | 0.5299 | -0.5299 | 1.5392 | 1.2469 | -0.2802 | -0.4267 | 0.6541 | 3.0373 | 2.0444 | 2.1722 |
|  |  |  |  |  | 0 | " | 0.4313 |  | -0.2783 | 0.0000 | 3.0373 | " | 2.4278 |
|  |  |  |  |  | 0.5299 | " | -0.3843 | " | -0.1298 | 2.1222 | 3.0373 | " | 2.4278 |
|  |  |  | 2 | 0.3878 | -0.3878 | 2.3048 | 1.3181 | -0.3199 | -0.4272 | 0.6781 | 2.9686 | 1.6611 | 1.6611 |
|  |  |  |  |  | 0 | " | 0.4242 | " | -0.3031 | 0 | 2.9686 | " | 1.7889 |
|  |  |  |  |  | 0.3878 | " | -0.4696 | " | -0.1790 | 1.9033 | 2.9686 | " | 1.9167 |
|  |  |  | 5 | 0.1897 | -0.1897 | 5.1184 | 1.4146 | -0.3921 | -0.4307 | 0.6864 | 2.9802 | 0.8519 | 0.9370 |
|  |  |  |  |  | 0 | " | 0.4437 | " | -0.3563 | 0 | 2.9802 | " | 1.0222 |
|  |  |  |  |  | 0.1897 | " | -0.5272 | " | -0.2819 | 1.8416 | 2.9802 | " | 1.1074 |
|  |  |  | 10 | 0.0986 | -0.0986 | 10.0559 | 1.4562 | -0.4368 | -0.4347 | 0.6808 | 2.9929 | 0.4685 | 0.5111 |
|  |  |  |  |  | 0 | " | 0.4648 | " | -0.3916 | 0 | 2.9929 | " | 0.5537 |
|  |  |  |  |  | 0.0986 | " | -0.5266 | " | -0.3486 | 1.8826 | 2.9929 | " | 0.5963 |
|  | 3.0 | 0.1 | 1 | 0.5317 | -0.5317 | 1.5475 | 1.3160 | -0.2487 | -0.3140 | 0.6253 | 3.1370 | 2.0444 | 2.5556 |
|  |  |  |  |  | 0 | " | 0.4931 | " | -0.1818 | 0 | 3.1371 | " | 2.5556 |
|  |  |  |  |  | 0.5317 | " | -0.3297 | " | -0.0496 | 2.4958 | 3.1371 | " | 2.7259 |
|  |  |  | 2 | 0.3886 | -0.3886 | 2.3108 | 1.3689 | -0.2922 | -0.3066 | 0.6560 | 3.0081 | 1.6611 | 1.9167 |
|  |  |  |  |  | 0 | " | 0.4708 | " | -0.1930 | 0 | 3.0081 | " | 2.0444 |
|  |  |  |  |  | 0.3886 | " | -0.4272 | " | -0.0795 | 2.1021 | 3.0081 | " | 2.0444 |
|  |  |  | 5 | 0.1898 | -0.1898 | 5.1209 | 1.4407 | -0.3738 | -0.2965 | 0.6746 | 2.9867 | 0.8519 | 0.9370 |
|  |  |  |  |  | 0 | " | 0.4689 | " | -0.2255 | 0 | 2.9867 | " | 1.0222 |
|  |  |  |  |  | 0.1898 | " | -0.5030 | " | -0.1546 | 1.9322 | 2.9867 | " | 1.1074 |
|  |  |  | 10 | 0.0986 | -0.0986 | 10.0568 | 1.4703 | -0.4257 | -0.2926 | 0.6744 | 2.9941 | 0.4685 | 0.5111 |
|  |  |  |  |  | 0 | " | 0.4787 | " | -0.2506 | 0 | 2.9941 | " | 0.5537 |
|  |  |  |  |  | 0.0986 | " | -0.5129 | " | -0.2086 | 1.9335 | 2.9941 | " | 0.5963 |
|  |  | 0.5 | 1 | 0.5321 | -0.5321 | 1.5501 | 1.3175 | -0.2434 | -0.3124 | 0.6261 | 3.1373 | 2.0444 | 2.5556 |
|  |  |  |  |  | 0 | " | 0.4926 | " | -0.1829 | 0 | 3.1373 | " | 2.7259 |
|  |  |  |  |  | 0.5321 | " | -0.3323 | " | -0.0534 | 2.4827 | 3.1373 | " | 2.7259 |
|  |  |  | 2 | 0.3889 | -0.3889 | 2.3129 | 1.3696 | -0.2862 | -0.3066 | 0.6567 | 3.0078 | 1.6611 | 1.9167 |
|  |  |  |  |  | 0 | " | 0.4701 | " | -0.1953 | 0 | 3.0078 | " | 2.0444 |
|  |  |  |  |  | 0.3889 | " | -0.4293 | " | -0.0840 | 2.0951 | 3.0078 | " | 2.0444 |
|  |  |  | 5 | 0.1898 | -0.1898 | 5.1219 | 1.4403 | -0.3678 | -0.3000 | 0.6750 | 2.9864 | 0.8519 | 0.9370 |
|  |  |  |  |  | 0 | " | 0.4681 | " | -0.2302 | 0 | 2.9864 | " | 1.0222 |
|  |  |  |  |  | 0.1898 | " | -0.5041 | " | -0.1604 | 1.9284 | 2.9864 | " | 1.1074 |
|  |  |  | 10 | 0.0986 | -0.0986 | 10.0572 | 1.4698 | -0.4209 | -0.2986 | 0.6747 | 2.9940 | 0.4685 | 0.5111 |
|  |  |  |  |  | 0 | " | 0.4781 |  | -0.2571 | 0 | 2.9940 | " | 0.5537 |
|  |  |  |  |  | 0.0986 | " | -0.5136 | " | -0.2156 | 1.9310 | 2.9940 | " | 0.5963 |



Figure 15. CASE III-M: plots showing $\Gamma, \Gamma^{\prime}$ with $M=2, c_{1}=0.5, c_{2}=$ $3.0, c_{3}=0.5$ and, from above, $\beta-\alpha=-\alpha, 0, \alpha$ respectively.


Figure 16. CASE III-M: plots showing $\phi^{\prime}$ with $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and for different $M$.


Figure 17. CASE III-M: plots showing $\gamma^{\prime}$ with $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and for different $M$. In the first picture $\beta-\alpha=-\alpha$, in the second $\beta-\alpha=0$.


Figure 18. CASE III-M: plot showing $\gamma^{\prime}$ with $c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$, $\beta-\alpha=\alpha$ and for different $M$.


Figure 19. CASE III-M: Figures (19) 1,3,5 show the streamlines and the points $\xi_{p}, \xi_{s}$ for $\frac{b}{a}=1, c_{1}=0.5, c_{2}=3.0, c_{3}=0.5$ and $\beta-\alpha=-\alpha, 0, \alpha$, respectively and $M=1$. Figures (19) 2,4,6 for $M=5$.


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