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1. Fundamentals of mechanics

1.1 Lagrangian mechanics

[LL]

1.1.1 Generalized coordinates

- Particle: point mass
- Particle position vector \mathbf{r} . In Cartesian components $\mathbf{r} = (x, y, z)$.
- Particle velocity $\mathbf{v} = d\mathbf{r}/dt = \dot{\mathbf{r}}$. In Cartesian coordinates $v_x = dx/dt$, etc...
- Particle acceleration $\mathbf{a} = d^2\mathbf{r}/dt^2 = \dot{\mathbf{v}} = \ddot{\mathbf{r}}$. In Cartesian coordinates $a_x = d^2x/dt^2$, etc...
- N particles $\implies s = 3N$ degrees of freedom
- Generalized coordinates: any s quantities q_i that define the positions of the N -body system ($\mathbf{q} = q_1, \dots, q_s$)
- Generalized velocities: \dot{q}_i ($\dot{\mathbf{q}} = \dot{q}_1, \dots, \dot{q}_s$)
- We know from experience that, given \mathbf{q} and $\dot{\mathbf{q}}$ for all particles in the system at a given time, we are able to predict $\mathbf{q}(t)$ at any later time t . In other words, if all \mathbf{q} and $\dot{\mathbf{q}}$ are specified $\implies \ddot{\mathbf{q}}$ are known.
- Equations of motion are ODE for $\mathbf{q}(t)$ that relate $\ddot{\mathbf{q}}$ with \mathbf{q} and $\dot{\mathbf{q}}$. The solution $\mathbf{q}(t)$ is the path (orbit).

1.1.2 Principle of least action & Euler-Lagrange equations

- Given a mechanical system, we define the *Lagrangian* function $\mathcal{L}(\mathbf{q}, \dot{\mathbf{q}}, t)$. \mathcal{L} does not depend on higher derivatives, consistent with the idea that motion is determined if \mathbf{q} and $\dot{\mathbf{q}}$ are given.
- Given two instants t_1 and t_2 , we define the *action* $S = \int_{t_1}^{t_2} \mathcal{L}(\mathbf{q}, \dot{\mathbf{q}}, t) dt$
- The system occupies positions \mathbf{q}_1 and \mathbf{q}_2 at time t_1 and t_2 , respectively. Note that in this formalism instead of fixing position and velocity at the initial time t_1 , we fix positions at the initial and final times.

→ *Principle of least action* (or Hamilton's principle): from t_1 to t_2 the system moves in such a way that S is a minimum (extremum) over all paths, i.e. (for 1 degree of freedom)

$$\delta S = \delta \int_{t_1}^{t_2} \mathcal{L}(q, \dot{q}, t) dt = \int_{t_1}^{t_2} [\mathcal{L}(q + \delta q, \dot{q} + \delta \dot{q}, t) - \mathcal{L}(q, \dot{q}, t)] dt = \int_{t_1}^{t_2} \left(\frac{\partial \mathcal{L}}{\partial q} \delta q + \frac{\partial \mathcal{L}}{\partial \dot{q}} \delta \dot{q} \right) dt = 0,$$

Now, we have

$$\frac{\partial \mathcal{L}}{\partial q} \delta q + \frac{\partial \mathcal{L}}{\partial \dot{q}} \frac{d(\delta q)}{dt} = \frac{\partial \mathcal{L}}{\partial q} \delta q + \frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{q}} \delta q \right) - \frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{q}} \right) \delta q,$$

so the above equation can be rewritten as

$$\delta S = \left[\frac{\partial \mathcal{L}}{\partial \dot{q}} \delta q \right]_{t_1}^{t_2} + \int_{t_1}^{t_2} \left(\frac{\partial \mathcal{L}}{\partial q} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{q}} \right) \delta q dt = 0,$$

which is verified for all δq only when

$$\frac{\partial \mathcal{L}}{\partial q} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{q}} = 0,$$

because $\delta q(t_1) = \delta q(t_2) = 0$, as all possible paths are such that $q(t_1) = q_1$ and $q(t_2) = q_2$.

→ Generalizing to the case of s degrees of freedom we have the Euler-Lagrange (E-L) equations:

$$\frac{\partial \mathcal{L}}{\partial q_i} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{q}_i} = 0, \quad i = 1, \dots, s$$

→ Transformations like $\mathcal{L} \rightarrow A\mathcal{L}$, with A constant, or $\mathcal{L} \rightarrow \mathcal{L} + dF/dt$, where $F = F(\mathbf{q}, t)$ do not affect the particles' motion, because

$$\mathcal{L}' = A\mathcal{L} \implies \delta S' = \delta AS = A\delta S = 0 \iff \delta S = 0,$$

and

$$\mathcal{L}' = \mathcal{L} + dF/dt \implies S' = S + \int_{t_1}^{t_2} \frac{dF}{dt} dt = S + F(\mathbf{q}_2, t_2) - F(\mathbf{q}_1, t_1) = S + C,$$

where C is a constant (independent of $\mathbf{q}, \dot{\mathbf{q}}$).

1.1.3 Inertial frames

→ *Inertial reference frame*: such that space is homogeneous and isotropic and time is homogeneous. For instance, in any inertial reference frame a particle that is at rest at a given time will remain at rest at all later times.

→ Galileo's relativity principle: laws of motion are the same in all inertial reference frames (moving at constant velocity w.r.t. one another)

→ Free particle: particle subject to no force.

→ Lagrangian of a free particle cannot contain explicitly the position vector \mathbf{r} (space is homogeneous) or the time t (time is homogeneous) and cannot depend on the direction of \mathbf{v} (space is isotropic) $\implies \mathcal{L} = \mathcal{L}(v^2)$

→ More specifically, it can be shown (see LL) that for a free particle

$$\mathcal{L} = \frac{1}{2}mv^2,$$

where m is particle mass. $T = (1/2)mv^2$ is the particle kinetic energy.

→ In Cartesian coordinates $\mathbf{q} = \mathbf{r} = (x, y, z)$ and $\dot{\mathbf{q}} = \mathbf{v} = (v_x, v_y, v_z)$, so for a free particle the Lagrangian is $\mathcal{L} = \frac{1}{2}m(v_x^2 + v_y^2 + v_z^2)$. The E-L equations for a free particle are

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \mathbf{v}} = 0,$$

so for the x component

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial v_x} = m \frac{dv_x}{dt} = 0 \quad \implies \quad v_x = \text{const},$$

and similarly for y and z components. $\implies d\mathbf{v}/dt = 0$, which is the law of inertia (*Newton's first law of motion*).

1.1.4 Lagrangian of a free particle in different systems of coordinates

→ Let's write the Lagrangian of a free particle in different systems of coordinates. Note that, if dl is the infinitesimal displacement, $v^2 = (dl/dt)^2 = dl^2/dt^2$.

→ In Cartesian coordinates $dl^2 = dx^2 + dy^2 + dz^2$, so

$$\mathcal{L} = \frac{1}{2}m(\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$$

→ In cylindrical coordinates $dl^2 = dR^2 + R^2d\phi^2 + dz^2$, so

$$\mathcal{L} = \frac{1}{2}m(\dot{R}^2 + R^2\dot{\phi}^2 + \dot{z}^2)$$

→ In spherical coordinates $dl^2 = dr^2 + r^2d\theta^2 + r^2\sin^2\theta d\phi^2$, so

$$\mathcal{L} = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\theta}^2 + r^2\sin^2\theta\dot{\phi}^2)$$

→ The above can be derived also by taking the expression of the position vector \mathbf{r} , differentiating and squaring (see e.g. BT08 app. B). For instance, in cylindrical coordinates

$$\mathbf{r} = R\mathbf{e}_R + z\mathbf{e}_z$$

$$\frac{d\mathbf{r}}{dt} = \dot{R}\mathbf{e}_R + R\dot{\mathbf{e}}_R + \dot{z}\mathbf{e}_z,$$

because $\dot{\mathbf{e}}_z = 0$. Now,

$$\mathbf{e}_R = \cos\phi\mathbf{e}_x + \sin\phi\mathbf{e}_y$$

and

$$\mathbf{e}_\phi = -\sin\phi\mathbf{e}_x + \cos\phi\mathbf{e}_y$$

So

$$d\mathbf{e}_R = (-\sin\phi\mathbf{e}_x + \cos\phi\mathbf{e}_y)d\phi$$

$$d\mathbf{e}_R = \mathbf{e}_\phi d\phi$$

$$\dot{\mathbf{e}}_R = \dot{\phi}\mathbf{e}_\phi$$

so

$$\frac{d\mathbf{r}}{dt} = \dot{R}\mathbf{e}_R + R\dot{\phi}\mathbf{e}_\phi + \dot{z}\mathbf{e}_z,$$

$$v^2 = \left| \frac{d\mathbf{r}}{dt} \right|^2 = \dot{R}^2 + R^2\dot{\phi}^2 + \dot{z}^2.$$

→ Alternatively, we can derive the same expressions for the cylindrical and spherical coordinates, starting from the Cartesian coordinates. For instance, in spherical coordinates r , θ , ϕ we have

$$x = r \sin\theta \cos\phi, \quad y = r \sin\theta \sin\phi, \quad z = r \cos\theta,$$

so

$$\dot{x} = \dot{r} \sin\theta \cos\phi + r \cos\theta \cos\phi\dot{\theta} - r \sin\theta \sin\phi\dot{\phi},$$

$$\dot{y} = \dot{r} \sin\theta \sin\phi + r \cos\theta \sin\phi\dot{\theta} + r \sin\theta \cos\phi\dot{\phi},$$

$$\dot{z} = \dot{r} \cos\theta - r \sin\theta\dot{\theta},$$

then

$$v^2 = \dot{x}^2 + \dot{y}^2 + \dot{z}^2 = \dot{r}^2 + r^2\dot{\theta}^2 + r^2 \sin^2\theta\dot{\phi}^2.$$

→ The Lagrangian of a free particle in spherical coordinates is

$$\mathcal{L} = \frac{1}{2}m \left(\dot{r}^2 + r^2\dot{\theta}^2 + r^2 \sin^2\theta\dot{\phi}^2 \right).$$

1.1.5 Lagrangian of a system of particles

→ Additivity of the Lagrangian: take two dynamical systems A and B. If each of them were an isolated system, they would have, respectively, Lagrangians \mathcal{L}_A and \mathcal{L}_B . If they are two parts of the same system (but so distant that the interaction is negligible) the total Lagrangian must be $\mathcal{L} = \mathcal{L}_A + \mathcal{L}_B$.

→ So, for a *system of non-interacting particles*

$$\mathcal{L} = \sum_a \frac{1}{2}m_a v_a^2,$$

were the subscript a identifies the a -th the particle.

→ Closed system: system of particles that interact, but are not affected by external forces.

→ Lagrangian for a closed *system of interacting particles*:

$$\mathcal{L} = T - V$$

where T is the kinetic energy and V is potential energy.

→ The potential energy V depends only on the position of the particles: $V = V(\mathbf{q})$. This is a consequence of the assumption that the interaction is instantaneously propagated: a change in position of one of the particles instantaneously affects the force experienced by the other particles.

→ We have seen that in general $T = T(\mathbf{q}, \dot{\mathbf{q}})$: see, for instance, the expression of T in cylindrical or spherical coordinates. In other words, in generalized coordinates the kinetic energy can depend also on the \mathbf{q} , not only on $\dot{\mathbf{q}}$:

$$\mathcal{L} = \frac{1}{2} \sum_{i,k} A_{ik}(\mathbf{q}) \dot{q}_i \dot{q}_k - V(\mathbf{q}),$$

where $i = 1, \dots, s$ and $k = 1, \dots, s$, where s is the number of degrees of freedom ($s = 3N$ for a system of N particles).

→ In Cartesian coordinates we have $\mathbf{q} = \mathbf{r}_a = (x_a, y_a, z_a)$ positions and $\dot{\mathbf{q}} = \mathbf{v}_a = (v_{x,a}, v_{y,a}, v_{z,a})$ (velocities), so the Lagrangian for a closed system of N particles is

$$\mathcal{L} = \frac{1}{2} \sum_{a=1, \dots, N} m_a v_a^2 - V(\mathbf{r}_1, \mathbf{r}_2, \dots, \mathbf{r}_N)$$

→ Applying Euler-Lagrange equations

$$\frac{\partial \mathcal{L}}{\partial \mathbf{r}_a} - \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \mathbf{v}_a} = 0,$$

we get the equations of motion:

$$m_a \dot{\mathbf{v}}_a = -\frac{\partial V}{\partial \mathbf{r}_a}, \quad \text{i.e.} \quad m_a \ddot{x}_a = -\frac{\partial V}{\partial x_a}, \quad \text{etc.}$$

i.e. $m_a \ddot{\mathbf{r}} = \mathbf{F}_a$ (*Newton's second law of motion*), where $\mathbf{F}_a = -\partial V / \partial \mathbf{r}_a$ is the force acting on the a -th particle.

→ For a particle moving in an external field

$$\mathcal{L} = \frac{1}{2} m v^2 - V(\mathbf{r}, t).$$

If the external field is uniform $\implies V = -\mathbf{F}(t) \cdot \mathbf{r}$ (\mathbf{F} dependent of time, but independent of position)

1.2 Conservation laws

→ *Constant of motion*: quantity that remains constant during the evolution of a mechanical system $C = C[\mathbf{q}(t), \dot{\mathbf{q}}(t), t] = \text{const}$ (i.e. $dC/dt = 0$).

→ *Integral of motion*: a constant of motion that depends only on \mathbf{q} and $\dot{\mathbf{q}}$ (in other words, it does not depend explicitly on time) $I = I[\mathbf{q}(t), \dot{\mathbf{q}}(t)] = \text{const}$ (i.e. $dI/dt = 0$). The value of the integral for a system equals the sum of the values for sub-systems that interact negligibly with one another.

→ Integrals of motion derive from fundamental properties (symmetries): isotropy/homogeneity of time and space. Among constants of motions, only integrals of motions are important in mechanics. Example of a constant of motion that is not an integral of motion: for a 1-D free particle $x(t) = x_0 + \dot{x}_0 t$ (where x_0 and \dot{x}_0 are the initial conditions and $\dot{x} = \dot{x}_0 = \text{const}$), so $x_0(x, t) = x(t) - \dot{x}_0 t$ is a constant of motion, but not an integral of motion (it depends explicitly on t).

→ There are seven integrals of motions: total energy E , momentum \mathbf{P} (3 components), angular momentum \mathbf{L} (3 components).

1.2.1 Energy

→ Homogeneity of time \implies Lagrangian of a closed system does not depend explicitly on time $\partial\mathcal{L}/\partial t = 0 \implies$

$$\frac{d\mathcal{L}}{dt} = \sum_i \left[\frac{\partial\mathcal{L}}{\partial q_i} \dot{q}_i + \frac{\partial\mathcal{L}}{\partial \dot{q}_i} \ddot{q}_i \right],$$

where $i = 1, \dots, s$, with s number of degrees of freedom. Using E-L equations:

$$\frac{d\mathcal{L}}{dt} = \sum_i \left[\frac{d}{dt} \left(\frac{\partial\mathcal{L}}{\partial \dot{q}_i} \right) \dot{q}_i + \frac{\partial\mathcal{L}}{\partial \dot{q}_i} \ddot{q}_i \right] = \sum_i \frac{d}{dt} \left(\frac{\partial\mathcal{L}}{\partial \dot{q}_i} \dot{q}_i \right)$$

$$\frac{d}{dt} \left(\sum_i \frac{\partial\mathcal{L}}{\partial \dot{q}_i} \dot{q}_i - \mathcal{L} \right) = 0$$

→ Energy

$$E \equiv \sum_i \frac{\partial\mathcal{L}}{\partial \dot{q}_i} \dot{q}_i - \mathcal{L}$$

$$\implies dE/dt = 0$$

→

$$\mathcal{L} = T - V \implies E = \sum_i \frac{\partial T}{\partial \dot{q}_i} \dot{q}_i - T + V$$

→ We have seen that T is a quadratic function of \dot{q}_i , so by Euler theorem on homogeneous functions $\dot{q}_i \partial T / \partial \dot{q}_i = 2T$, so

$$E = 2T - \mathcal{L} = T(\mathbf{q}, \dot{\mathbf{q}}) + V(\mathbf{q}),$$

i.e. total energy E is the sum of kinetic and potential energy.

→ Euler theorem on homogeneous functions: if $f(tx) = t^n f(x)$ then $xf'(x) = nf(x)$.

1.2.2 Momentum

Homogeneity of space \implies conservation of momentum. Lagrangian must be invariant if the system is shifted in space by ϵ .

Cartesian coordinates

\rightarrow Lagrangian must be invariant if the system is shifted in space by $\delta\mathbf{r}_a = \epsilon$:

$$\delta\mathcal{L} = \mathcal{L}(\mathbf{q} + d\mathbf{q}, \dot{\mathbf{q}}) - \mathcal{L}(\mathbf{q}, \dot{\mathbf{q}}) = \sum_a \frac{\partial\mathcal{L}}{\partial\mathbf{r}_a} \cdot \delta\mathbf{r}_a = \epsilon \cdot \sum \frac{\partial\mathcal{L}}{\partial\mathbf{r}_a}$$

\implies

$$\sum \frac{\partial\mathcal{L}}{\partial\mathbf{r}_a} = 0 \implies \frac{d}{dt} \sum \frac{\partial\mathcal{L}}{\partial\mathbf{v}_a} = 0$$

or

$$\frac{d\mathbf{P}}{dt} = 0,$$

where

$$\mathbf{P} \equiv \sum \frac{\partial\mathcal{L}}{\partial\mathbf{v}_a} = \sum m_a \mathbf{v}_a$$

is momentum

\rightarrow Momentum is additive $\mathbf{P} = \sum_a \mathbf{p}_a$, where $\mathbf{p}_a = \partial\mathcal{L}/\partial\mathbf{v}_a = m_a \mathbf{v}_a$ is the momentum of the individual particles.

\rightarrow We also have

$$\sum \frac{\partial\mathcal{L}}{\partial\mathbf{r}_a} = 0 \implies \sum \frac{\partial V}{\partial\mathbf{r}_a} = \sum \mathbf{F}_a = 0,$$

where \mathbf{F}_a is the force acting on the a -th particle. When the bodies are two, is $\mathbf{F}_1 + \mathbf{F}_2 = 0$ or $\mathbf{F}_1 = -\mathbf{F}_2$ i.e. *Newton's third law of motion*.

Generalized coordinates

\rightarrow We define

$$p_i \equiv \frac{\partial\mathcal{L}}{\partial\dot{q}_i}$$

as *generalized momenta*. Note that $p_i = mv_i$ in Cartesian coordinates, but in general p_i depend on both q_i and \dot{q}_i .

\rightarrow The E-L equation can be written as

$$\frac{dp_i}{dt} = F_i,$$

where $F_i = \partial\mathcal{L}/\partial q_i$ is the generalized force.

\rightarrow Then

$$\frac{d}{dt} \sum_i p_i = \sum_i F_i = 0.$$

Centre of mass

→ For a system of particles there is a special inertial reference frame in which $\mathbf{P} = 0$: this is the reference frame in which the centre of mass is at rest.

$$\mathbf{P} = \sum m_a \mathbf{v}_a = \frac{d}{dt} \sum m_a \mathbf{r}_a = \frac{d}{dt} \mathbf{r}_{\text{cm}} \sum m_a = \sum m_a \frac{d\mathbf{r}_{\text{cm}}}{dt} = 0,$$

where

$$\mathbf{r}_{\text{cm}} \equiv \frac{\sum m_a \mathbf{r}_a}{\sum m_a}$$

is the position of the centre of mass.

→ In a general inertial frame the centre of mass moves with a velocity

$$\mathbf{v}_{\text{cm}} = \frac{d\mathbf{r}_{\text{cm}}}{dt} = \frac{\sum m_a \mathbf{v}_a}{\sum m_a} = \frac{\mathbf{P}}{\sum m_a} = \text{const}$$

→ If the total energy of the system in the centre-of-mass reference frame is E_{int} , in a general inertial frame the total energy is

$$E = \frac{1}{2} \sum_a m_a v_{\text{cm}}^2 + E_{\text{int}}$$

→ Note: the components of the centre of mass \mathbf{r}_{cm} are not constants of motion. The components of $\mathbf{r}_{\text{cm}}(0) = \mathbf{r}_{\text{cm}}(t) - t\mathbf{v}_{\text{cm}}$ are constants of motion, but they are not integrals of motion (they depend explicitly on time).

1.2.3 Angular momentum

→ Isotropy of space \implies conservation of angular momentum

→ The Lagrangian is invariant under rotation. Apply a rotation represented by a vector $\delta\phi$ (with magnitude $\delta\phi$, which is the angle of rotation, and direction along the rotation axis) $\implies \delta\mathbf{r} = \delta\phi \times \mathbf{r}$ and $\delta\mathbf{v} = \delta\phi \times \mathbf{v}$

$$\delta\mathcal{L} = \sum_a \left(\frac{\partial\mathcal{L}}{\partial\mathbf{r}_a} \cdot \delta\mathbf{r}_a + \frac{\partial\mathcal{L}}{\partial\mathbf{v}_a} \cdot \delta\mathbf{v}_a \right) = 0$$

$$\delta\mathcal{L} = \sum_a [\dot{\mathbf{p}}_a \cdot (\delta\phi \times \mathbf{r}_a) + \mathbf{p}_a \cdot (\delta\phi \times \mathbf{v}_a)] = 0.$$

Using the vector identity $\mathbf{A} \cdot (\mathbf{B} \times \mathbf{C}) = \mathbf{B} \cdot (\mathbf{C} \times \mathbf{A})$ we get

$$\delta\mathcal{L} = \delta\phi \cdot \sum (\mathbf{r}_a \times \dot{\mathbf{p}}_a + \mathbf{v}_a \times \mathbf{p}_a) = \delta\phi \cdot \frac{d}{dt} \sum \mathbf{r}_a \times \mathbf{p}_a = 0,$$

because

$$\frac{d}{dt} (\mathbf{r}_a \times \mathbf{p}_a) = \mathbf{r}_a \times \dot{\mathbf{p}}_a + \mathbf{v}_a \times \mathbf{p}_a$$

→ As this must be satisfied for all $\delta\phi$, we must have

$$\frac{d\mathbf{L}}{dt} = 0,$$

where

$$\mathbf{L} \equiv \sum \mathbf{r}_a \times \mathbf{p}_a$$

is the angular momentum, which (as well as linear momentum) is additive.

→ The angular momentum in a reference frame in which the system is at rest ($\mathbf{P} = 0$) is \mathbf{L}_{int} (intrinsic angular momentum)

→ In a general inertial frame the angular momentum is

$$\mathbf{L} = \mathbf{L}_{int} + \mathbf{r}_{cm} \times \left(\sum m_a \right) \mathbf{v}_{cm} = \mathbf{L}_{int} + \mathbf{r}_{cm} \times \sum m_a \mathbf{v}_a = \mathbf{L}_{int} + \mathbf{r}_{cm} \times \mathbf{P}$$

1.3 Integration of the equations of motion

1.3.1 Motion in one dimension

→ One dimension = one degree of freedom = one coordinate q

→ Lagrangian:

$$\mathcal{L} = \frac{1}{2}a(q)\dot{q}^2 - V(q).$$

If $q = x$ is a Cartesian coordinate

$$\mathcal{L} = \frac{1}{2}m\dot{x}^2 - V(x).$$

→ Energy is integral of motion:

$$E = \frac{1}{2}m\dot{x}^2 + V(x),$$

then (taking $\dot{x} \geq 0$)

$$\frac{dx}{dt} = \sqrt{\frac{2}{m}[E - V(x)]},$$

$$dt = \frac{\sqrt{m}}{\sqrt{2[E - V(x)]}} dx,$$

$$t = t_0 + \sqrt{\frac{m}{2}} \int_{x_0}^x \frac{1}{\sqrt{[E - V(x)]}} dx.$$

→ Motion in region of space such that $V(x) < E$. If this interval is bounded, motion is finite. From the above equations and the E-L equations (equations of motion) it is clear that motion is oscillatory (\dot{x} changes sign only at turning points x such that $V(x) = E$) \implies motion is periodic with period

$$T = \sqrt{2m} \int_{x_1}^{x_2} \frac{1}{\sqrt{[E - V(x)]}} dx,$$

where x_1 and x_2 are the turning points at which $E = V(x_1) = V(x_2)$. T is twice the time to go from x_1 to x_2 (see Fig. 6 LL. FIG CM1.1).

1.3.2 Motion in a central field

→ Motion in a central field: motion of a single particle in an external field such that its potential energy depends only on the distance r from a fixed point: $V = V(r) \implies$

$$\mathbf{F} = -\frac{\partial V}{\partial \mathbf{r}} = -\frac{dV}{dr} \frac{\mathbf{r}}{r}$$

→ For instance, in Cartesian coordinates:

$$F_x = -\frac{\partial V}{\partial x} = -\frac{dV}{dr} \frac{\partial r}{\partial x} = -\frac{dV}{dr} \frac{x}{r},$$

etc., because

$$\frac{\partial r}{\partial x} = \frac{\partial \sqrt{x^2 + y^2 + z^2}}{\partial x} = \frac{2x}{2\sqrt{x^2 + y^2 + z^2}} = \frac{x}{r}$$

→ Take center of the field as origin: angular momentum \mathbf{L} is conserved (even in the presence of the field), because the field does not have component orthogonal to position vector.

$$\frac{d\mathbf{L}}{dt} = \dot{\mathbf{r}} \times \mathbf{p} + \mathbf{r} \times \dot{\mathbf{p}} = 0.$$

→ $\mathbf{L} = \mathbf{r} \times \mathbf{p}$ is conserved and is orthogonal to \mathbf{r} , so \mathbf{r} stays always in the same plane \implies motion is planar.

→ Using polar coordinates (r, ϕ) in the plane of the motion, the Lagrangian reads (see kinetic energy in cylindrical coordinates: Section 1.1.4)

$$\mathcal{L} = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\phi}^2) - V(r)$$

→ *Motion in ϕ* . E-L equations for coordinate $\phi \implies$

$$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{d(mr^2\dot{\phi})}{dt} = 0,$$

where

$$L_z = L = mr^2\dot{\phi} = \text{const}$$

is the modulus of the angular momentum. ϕ is cyclic coordinate: it does not appear in \mathcal{L} . Associated generalized momentum is constant.

→ *Kepler's second law*: let's define an infinitesimal sector bounded by the path as

$$dA = \frac{1}{2}r^2 d\phi$$

(show Fig. 8 LL FIG CM1.2). $dA/dt = r^2(d\phi/dt)/2 = L/(2m) = \text{const}$ is the sectorial velocity \implies the particle's position vector sweeps equal areas in equal times (Kepler's second law).

→ *Motion in r.*

$$E = 2T - \mathcal{L} = \text{const},$$

$$E = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\phi}^2) + V(r) = \frac{1}{2}m\dot{r}^2 + \frac{L^2}{2mr^2} + V(r),$$

⇒

$$\frac{dr}{dt} = \sqrt{\frac{2}{m}[E - V_{\text{eff}}(r)]}$$

$$dt = \frac{dr}{\sqrt{\frac{2}{m}[E - V_{\text{eff}}(r)]}}$$

(time t as a function of r), where

$$V_{\text{eff}}(r) = V(r) + \frac{L^2}{2mr^2}.$$

→ The radial part of the motion behaves like a motion in one-dimension with effective potential energy $V_{\text{eff}}(r)$, defined above, where $L^2/2mr^2$ is called the centrifugal energy.

→ The radii r such that $E = V_{\text{eff}}(r)$ are the radial turning points, corresponding to $\dot{r} = 0$: if motion is finite, pericentre (r_{peri}) and apocentre (r_{apo}). If motion is infinite $r_{\text{apo}} = \infty$

→ Substituting $mr^2d\phi/L = dt$

$$d\phi = \frac{Ldr}{r^2\sqrt{2m[E - V_{\text{eff}}(r)]}}$$

(angle ϕ as a function of r , i.e. path or trajectory).

→ Consider variation of ϕ for finite motion in one radial period:

$$\Delta\phi = 2 \int_{r_{\text{peri}}}^{r_{\text{apo}}} \frac{Ldr}{r^2\sqrt{2m[E - V_{\text{eff}}(r)]}}$$

→ Closed orbit only if $\Delta\phi = 2\pi m/n$ with m, n integers. In general orbit is not closed (rosette). All orbits are closed only when $V \propto 1/r$ (Kepler's potential) or $V \propto r^2$ (harmonic potential).

1.4 Hamiltonian mechanics

[LL; VK]

1.4.1 Hamilton's equations

→ In Lagrangian mechanics generalized coordinates (q_i) and generalized velocities (\dot{q}_i), $i = 1, \dots, s$, where s is the number of degrees of freedom.

→ In Hamiltonian mechanics generalized coordinates (q_i) and generalized momenta (p_i), $i = 1, \dots, s$.

→ The idea is to transform $\mathcal{L}(\mathbf{q}, \dot{\mathbf{q}}, t)$ into a function of $(\mathbf{q}, \mathbf{p}, t)$, where $p_i = \partial\mathcal{L}/\partial\dot{q}_i$ are the generalized momenta. This can be accomplished through a *Legendre transform*.

→ Example: Legendre transform for functions of two variables. Start from $f = f(x, y)$, $u \equiv \partial f / \partial x$ and $v \equiv \partial f / \partial y$. The total differential of f is

$$df = udx + vdy.$$

We want to replace y with v , so we use

$$d(vy) = vdy + ydv$$

so

$$df = udx + d(vy) - ydv,$$

$$d(vy - f) = -udx + ydv,$$

so $g(x, v) \equiv vy - f(x, v)$ with $\partial g / \partial x = -u$ and $\partial g / \partial v = y$.

→ We can do the same starting from \mathcal{L} :

$$d\mathcal{L} = \sum_i \frac{\partial \mathcal{L}}{\partial q_i} dq_i + \sum_i \frac{\partial \mathcal{L}}{\partial \dot{q}_i} d\dot{q}_i + \frac{\partial \mathcal{L}}{\partial t} dt = \sum_i \dot{p}_i dq_i + \sum_i p_i d\dot{q}_i + \frac{\partial \mathcal{L}}{\partial t} dt$$

$$d\mathcal{L} = \sum_i \dot{p}_i dq_i + \sum_i d(p_i \dot{q}_i) - \sum_i \dot{q}_i dp_i + \frac{\partial \mathcal{L}}{\partial t} dt$$

$$d\left(\sum_i p_i \dot{q}_i - \mathcal{L}\right) = \sum_i \dot{q}_i dp_i - \sum_i \dot{p}_i dq_i - \frac{\partial \mathcal{L}}{\partial t} dt$$

→ So the Legendre transform of \mathcal{L} is the Hamiltonian:

$$\mathcal{H}(p, q, t) \equiv \sum_i p_i \dot{q}_i - \mathcal{L}$$

→ The differential of \mathcal{H} is:

$$d\mathcal{H} = \sum_i \dot{q}_i dp_i - \sum_i \dot{p}_i dq_i - \frac{\partial \mathcal{L}}{\partial t} dt$$

→ It follows

$$\dot{q}_i = \frac{\partial \mathcal{H}}{\partial p_i}, \quad \dot{p}_i = -\frac{\partial \mathcal{H}}{\partial q_i},$$

which are called *Hamilton's equations* or canonical equations. We have replaced s 2nd-order equations with $2s$ first-order equations. We also have

$$\frac{\partial \mathcal{H}}{\partial t} = -\frac{\partial \mathcal{L}}{\partial t}$$

→ \mathbf{p} and \mathbf{q} are called canonical coordinates.

→ The time derivative of \mathcal{H} is

$$\frac{d\mathcal{H}}{dt} = \frac{\partial \mathcal{H}}{\partial t} + \sum_i \frac{\partial \mathcal{H}}{\partial q_i} \dot{q}_i + \sum_i \frac{\partial \mathcal{H}}{\partial p_i} \dot{p}_i = \frac{\partial \mathcal{H}}{\partial t}$$

→ Hamiltonian is constant if \mathcal{H} does not depend explicitly on time. This is the case for closed system, for which \mathcal{L} does not depend explicitly on time. This is a reformulation of energy conservation, because we recall that

$$E \equiv \sum_i \frac{\partial \mathcal{L}}{\partial \dot{q}_i} \dot{q}_i - \mathcal{L} = \sum_i p_i \dot{q}_i - \mathcal{L} = \mathcal{H}$$

1.4.2 Canonical transformations

[VK 4.7-4.8]

→ Given a set of canonical coordinates (\mathbf{p}, \mathbf{q}) , we might want to change to another set of coordinates (\mathbf{P}, \mathbf{Q}) to simplify our problem.

→ We can consider general transformations of the form $\mathbf{Q} = \mathbf{Q}(\mathbf{p}, \mathbf{q}, t)$ and $\mathbf{P} = \mathbf{P}(\mathbf{p}, \mathbf{q}, t)$: it is not guaranteed that Hamilton's equations are unchanged.

→ A transformation $\mathbf{Q} = \mathbf{Q}(\mathbf{p}, \mathbf{q}, t)$ and $\mathbf{P} = \mathbf{P}(\mathbf{p}, \mathbf{q}, t)$ is called *canonical* if in the new coordinates

$$\dot{Q}_i = \frac{\partial \mathcal{H}'}{\partial P_i}, \quad \dot{P}_i = -\frac{\partial \mathcal{H}'}{\partial Q_i},$$

with some Hamiltonian $\mathcal{H}' = \mathcal{H}'(\mathbf{P}, \mathbf{Q}, t)$.

→ In any canonical coordinate system the variation of the action is null

$$\begin{aligned} \delta S &= \delta \int_{t_1}^{t_2} \mathcal{L} dt = \delta \int_{t_1}^{t_2} \left(\sum_i p_i \dot{q}_i - \mathcal{H} \right) dt = 0 \\ \delta S' &= \delta \int_{t_1}^{t_2} \mathcal{L}' dt = \delta \int_{t_1}^{t_2} \left(\sum_i P_i \dot{Q}_i - \mathcal{H}' \right) dt = 0 \end{aligned}$$

This means that the difference between the two Lagrangians $\mathcal{L} - \mathcal{L}' = dF/dt$ must be a total time derivative, because

$$\delta \int_{t_1}^{t_2} \frac{dF}{dt} dt = \delta [F]_{t_1}^{t_2} = 0.$$

→ Let us impose the above condition, i.e.,

$$\begin{aligned} \mathcal{L} &= \mathcal{L}' + \frac{dF}{dt}, \\ \mathcal{L} dt &= \mathcal{L}' dt + dF, \\ \left(\sum_i p_i \frac{dq_i}{dt} - \mathcal{H} \right) dt &= \sum_i P_i dQ_i - \mathcal{H}' dt + dF \\ \sum_i p_i dq_i - \mathcal{H} dt &= \sum_i P_i dQ_i - \mathcal{H}' dt + dF, \end{aligned}$$

where $F = F(\mathbf{q}, \mathbf{Q}, t)$.

→ F is called the generating function of the transformation.

→ Taking F in the form $F = F(\mathbf{q}, \mathbf{Q}, t)$, the above equation can be written as

$$\begin{aligned} \sum_i p_i dq_i - \mathcal{H} dt &= \sum_i P_i dQ_i - \mathcal{H}' dt + \frac{\partial F}{\partial t} dt + \sum_i \frac{\partial F}{\partial q_i} dq_i + \sum_i \frac{\partial F}{\partial Q_i} dQ_i, \\ \frac{\partial F}{\partial t} dt + \sum_i \frac{\partial F}{\partial q_i} dq_i + \sum_i \frac{\partial F}{\partial Q_i} dQ_i &= \sum_i p_i dq_i + (\mathcal{H}' - \mathcal{H}) dt - \sum_i P_i dQ_i \end{aligned}$$

Clearly the above is verified when

$$p_i = \frac{\partial F}{\partial q_i}$$

$$P_i = -\frac{\partial F}{\partial Q_i}$$

$$H' = H + \frac{\partial F}{\partial t}.$$

The above relations can be combined (and when necessary inverted) to give $\mathbf{Q} = \mathbf{Q}(\mathbf{q}, \mathbf{p}, t)$ and $\mathbf{P} = \mathbf{P}(\mathbf{q}, \mathbf{p}, t)$, i.e. the canonical transformation in terms of the generating function F .

→ Sometimes is it convenient to have a generating function that is not in the form $F = F(\mathbf{q}, \mathbf{Q}, t)$, but depends on other combinations of new and old canonical coordinates: other possible choices are $(\mathbf{q}, \mathbf{P}, t)$, $(\mathbf{p}, \mathbf{Q}, t)$, $(\mathbf{p}, \mathbf{P}, t)$.

→ We distinguish four classes of generating functions F , differing by the variables on which F depends:

$$F = F_1(\mathbf{q}, \mathbf{Q}, t), \quad F = F_2(\mathbf{q}, \mathbf{P}, t), \quad F = F_3(\mathbf{p}, \mathbf{Q}, t), \quad F = F_4(\mathbf{p}, \mathbf{P}, t).$$

→ We derive here the canonical transformation for a generating function F_2 , depending on $(\mathbf{q}, \mathbf{P}, t)$. In order to do so we use the Legendre transform. Start from

$$dF_1 = \sum_i p_i dq_i - \sum_i P_i dQ_i + (\mathcal{H}' - \mathcal{H})dt$$

where $F_1(\mathbf{q}, \mathbf{Q}, t)$ is the generating function considered above.

$$dF_1 = \sum_i p_i dq_i - \sum_i d(P_i Q_i) + \sum_i Q_i dP_i + (\mathcal{H}' - \mathcal{H})dt$$

$$d(F_1 + \sum_i P_i Q_i) = \sum_i p_i dq_i + \sum_i Q_i dP_i + (\mathcal{H}' - \mathcal{H})dt$$

so the generating function is now

$$F_2 = F_2(\mathbf{q}, \mathbf{P}, t) \equiv F_1(\mathbf{q}, \mathbf{Q}, t) + \sum_i P_i Q_i$$

and the change of variables is as follows

$$p_i = \frac{\partial F_2}{\partial q_i}$$

$$Q_i = \frac{\partial F_2}{\partial P_i}$$

$$\mathcal{H}' = \mathcal{H} + \frac{\partial F_2}{\partial t},$$

where $F_2 = F_2(\mathbf{q}, \mathbf{P}, t)$

→ Similarly (exploiting Legendre transform) we can obtain transformation equations for $F = F_3(\mathbf{p}, \mathbf{Q}, t) = F_1 - \sum_i q_i p_i$ and $F = F_4(\mathbf{p}, \mathbf{P}, t) = F_1 + \sum_i Q_i P_i - \sum_i q_i p_i$. In summary the canonical transformations are

$$\begin{aligned} F = F_1(\mathbf{q}, \mathbf{Q}, t), & \quad p_i = \frac{\partial F_1}{\partial q_i} & \quad P_i = -\frac{\partial F_1}{\partial Q_i} \\ F = F_2(\mathbf{q}, \mathbf{P}, t), & \quad p_i = \frac{\partial F_2}{\partial q_i} & \quad Q_i = \frac{\partial F_2}{\partial P_i} \\ F = F_3(\mathbf{p}, \mathbf{Q}, t), & \quad q_i = -\frac{\partial F_3}{\partial p_i} & \quad P_i = -\frac{\partial F_3}{\partial Q_i} \\ F = F_4(\mathbf{p}, \mathbf{P}, t), & \quad q_i = -\frac{\partial F_4}{\partial p_i} & \quad Q_i = \frac{\partial F_4}{\partial P_i}. \end{aligned}$$

In addition we have $\mathcal{H}' = \mathcal{H} + \partial F_i / \partial t$, for $i = 1, \dots, 4$.

→ *Example of canonical transformations: extended point transformations.* This is of the kind $F = F_2(\mathbf{q}, \mathbf{P})$:

$$\mathbf{Q} = \mathbf{G}(\mathbf{q}), \quad F(\mathbf{q}, \mathbf{P}) = \sum_k P_k G_k(\mathbf{q}),$$

where $\mathbf{G} = (G_1, \dots, G_s)$, and the G_i are given functions. Then

$$\begin{aligned} p_i &= \frac{\partial F}{\partial q_i} = \sum_k P_k \frac{\partial G_k}{\partial q_i}(\mathbf{q}), \\ Q_i &= \frac{\partial F}{\partial P_i} = G_i(\mathbf{q}). \end{aligned}$$

For instance, for a system with 1 degree of freedom, we have

$$Q = G(q), \quad F(q, P) = PG(q),$$

so

$$\begin{aligned} p = \frac{\partial F}{\partial q} = P \frac{\partial G}{\partial q}(q) &\implies P = p \left(\frac{\partial G}{\partial q} \right)^{-1}, \\ Q = \frac{\partial F}{\partial P} &= G(q). \end{aligned}$$

→ *Special classes of Canonical Coordinates.* Among canonical coordinates there are two special classes that are particularly important:

- Sets of canonical coordinates in which both \mathbf{q} and \mathbf{p} are integrals of motion (they remain constant during the evolution of the system): these are the coordinates obtained by solving the *Hamilton-Jacobi equation*.
- Sets of canonical coordinates in which the \mathbf{q} are not constant, but they are cyclic coordinates (they can be interpreted as angles, so they are called *angles*), while the \mathbf{p} are integrals of motion (they are constants; they are called *actions*). These are called *angle-action coordinates*.

1.4.3 Hamilton-Jacobi equation

[VK 4.9]

→ The action $S = \int \mathcal{L}dt$ can be seen as a generating function. The corresponding canonical transformations are very useful because they are such that $\mathcal{H}' = 0$. We have

$$dS = \mathcal{L}dt = \sum_i p_i \frac{dq_i}{dt} dt - \mathcal{H}dt = \sum_i p_i dq_i - \mathcal{H}dt.$$

→ It follows that

$$p_i = \frac{\partial S}{\partial q_i},$$

$$\mathcal{H} = -\frac{\partial S}{\partial t}.$$

→ So $S = S(\mathbf{q}, t)$, but S can be seen also as a generating function $S = S(\mathbf{q}, \mathbf{P}, t)$, with P_i constants (i.e. $dP_i = 0$ for all i). So, we have

$$Q_i = \frac{\partial S}{\partial P_i}$$

→ The new Hamiltonian is null:

$$\mathcal{H}' = \mathcal{H} + \frac{\partial S}{\partial t} = 0,$$

consistent with the fact that the new canonical coordinates are constant:

$$\dot{P}_i = -\frac{\partial \mathcal{H}'}{\partial Q_i} = 0 \quad \Longrightarrow \quad P_i = \alpha_i = \text{const}$$

$$\dot{Q}_i = \frac{\partial \mathcal{H}'}{\partial P_i} = 0 \quad \Longrightarrow \quad Q_i = \beta_i = \text{const}.$$

→ So the action can be seen as a generating function in the form $S = S(\mathbf{q}, \boldsymbol{\alpha}, t)$ where $\boldsymbol{\alpha} = (\alpha_1, \dots, \alpha_s) = \mathbf{P}$ are constants.

→ Exploiting the fact that $p_i = \partial S / \partial q_i$, the above equation $\mathcal{H}' = 0$ can be written as

$$\mathcal{H}\left(q_i, \frac{\partial S}{\partial q_i}, t\right) + \frac{\partial S}{\partial t} = 0.$$

This is known as the *Hamilton-Jacobi equation*.

→ If the solution S to the H-J equation is obtained, the solution of the equations of motions can be written explicitly as follows. The variables (\mathbf{p}, \mathbf{q}) are related to $(\boldsymbol{\alpha}, \boldsymbol{\beta})$ by

$$p_i = \frac{\partial S}{\partial q_i}(\mathbf{q}, \boldsymbol{\alpha}, t), \quad \beta_i = \frac{\partial S}{\partial \alpha_i}(\mathbf{q}, \boldsymbol{\alpha}, t),$$

which can be combined and inverted to give $q_i = q_i(\boldsymbol{\alpha}, \boldsymbol{\beta}, t)$ and $p_i = p_i(\boldsymbol{\alpha}, \boldsymbol{\beta}, t)$.

→ If \mathcal{H} does not depend explicitly on time then $\mathcal{H} = E = \text{const}$ and the H-J equation can be written

$$\mathcal{H}\left(q_i, \frac{\partial S}{\partial q_i}\right) = E,$$

and we also have

$$\frac{\partial S}{\partial t} = -E.$$

→ *Example: free particle.* Take Cartesian coordinates x, y, z as generalized coordinates q_i and p_x, p_y, p_z as generalized momenta p_i . The Hamiltonian of a free particle of mass m is

$$\mathcal{H} = \frac{1}{2m} (p_x^2 + p_y^2 + p_z^2).$$

The H-J equation is

$$\frac{1}{2m} \left[\left(\frac{\partial S}{\partial x} \right)^2 + \left(\frac{\partial S}{\partial y} \right)^2 + \left(\frac{\partial S}{\partial z} \right)^2 \right] + \frac{\partial S}{\partial t} = 0.$$

Separation of variables $S(x, y, z, t) = X(x) + Y(y) + Z(z) + T(t)$ then

$$\frac{1}{2m} \left[\left(\frac{\partial X}{\partial x} \right)^2 + \left(\frac{\partial Y}{\partial y} \right)^2 + \left(\frac{\partial Z}{\partial z} \right)^2 \right] + \frac{\partial T}{\partial t} = 0,$$

so $X = \alpha_x x$, $Y = \alpha_y y$, $Z = \alpha_z z$ and $T = -(\alpha_x^2 + \alpha_y^2 + \alpha_z^2)t/2m$, therefore $p_x = \alpha_x$, $p_y = \alpha_y$, $p_z = \alpha_z$, $\beta_x = x - \alpha_x t/m$, $\beta_y = y - \alpha_y t/m$, $\beta_z = z - \alpha_z t/m$, which is the solution (the values of the constants depend on the initial conditions at $t = 0$).

Bibliography

→ Binney J., Tremaine S. 2008 , “Galactic dynamics”, Princeton University Press, Princeton (BT08)

→ Landau L.D., Lifshitz E.M., 1982, “Mechanics”, Butterworth-Heinemann (LL)

→ Valtonen M., Karttunen H., 2006 “The three body problem”, Cambridge University Press, Cambridge (VK)