# Hard spectator interactions in $B \rightarrow \pi \pi$ at order $\alpha_{s}^{2}$ 

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## Zusammenfassung

In der vorliegenden Arbeit diskutiere ich die hard spectator interaction Amplitude von $B \rightarrow \pi \pi$ zur nächstführenden Ordnung in QCD (d.h. $\mathcal{O}\left(\alpha_{s}^{2}\right)$ ). Dieser spezielle Teil der Amplitude, dessen führende Ordnung bei $\mathcal{O}\left(\alpha_{s}\right)$ beginnt, ist im Rahmen der QCD Faktorisierung definiert. QCD Faktorisierung ermöglicht, in führender Ordnung in einer Entwicklung in $\Lambda_{\mathrm{QCD}} / m_{b}$ die kurz- und die langreichweitige Physik zu trennen, wobei die kurzreichweitige Physik in einer störungstheoretischen Entwicklung in $\alpha_{s}$ berechnet werden kann. Gegenüber anderen Teilen der Amplitude erfahren hard spectator interactions formal eine Verstärkung durch die zusätzlich zur $m_{b}$-Skala hinzutretende hartkollineare Skala $\sqrt{\Lambda_{\mathrm{QCD}} m_{b}}$, die zu einem größeren numerischen Wert von $\alpha_{s}$ führt.

Aus rechentechnischer Sicht liegen die hauptsächlichen Herausforderungen dieser Arbeit in der Tatsache begründet, dass die Feynmanintegrale, mit denen wir es zu tun haben, bis zu fünf äußere Beine haben und drei unabhängige Skalenverhältnisse enthalten. Diese Feynmanintegrale müssen in Potenzen in $\Lambda_{\mathrm{QCD}} / m_{b}$ entwickelt werden. Ich werde integration by parts Identitäten vorstellen, mit denen die Anzahl der Masterintegrale reduziert werden kann. Ebenso werde ich diskutieren, wie man mit Differenzialgleichungsmethoden die Entwicklung der Masterintegrale in $\Lambda_{\mathrm{QCD}} / m_{b}$ erhält. Im Anhang ist eine konkrete Implementierung der integration by parts Identitäten für ein Computeralgebrasystem vorhanden.

Schließlich diskutiere ich numerische Sachverhalte, wie die Abhängigkeit der Amplituden von der Renormierungsskala und die Größe der Verzweigungsverhältnisse. Es wird sich herausstellen das die nächstführende Ordnung der hard spectator interactions wichtig jedoch klein genug ist, so dass die Gültigkeit der Störungstheorie bestehen bleibt.

## Abstract

In the present thesis I discuss the hard spectator interaction amplitude in $B \rightarrow \pi \pi$ at NLO i.e. at $\mathcal{O}\left(\alpha_{s}^{2}\right)$. This special part of the amplitude, whose LO starts at $\mathcal{O}\left(\alpha_{s}\right)$, is defined in the framework of QCD factorization. QCD factorization allows to separate the short- and the long-distance physics in leading power in an expansion in $\Lambda_{\mathrm{QCD}} / m_{b}$, where the short-distance physics can be calculated in a perturbative expansion in $\alpha_{s}$. Compared to other parts of the amplitude hard spectator interactions are formally enhanced by the hard collinear scale $\sqrt{\Lambda_{\mathrm{QCD}} m_{b}}$, which occurs next to the $m_{b}$-scale and leads to an enhancement of $\alpha_{s}$.

From a technical point of view the main challenges of this calculation are due to the fact that we have to deal with Feynman integrals that come with up to five external legs and with three independent ratios of scales. These Feynman integrals have to be expanded in powers of $\Lambda_{\mathrm{QCD}} / m_{b}$. I will discuss integration by parts identities to reduce the number of master integrals and differential equations techniques to get their power expansions. A concrete implementation of integration by parts identities in a computer algebra system is given in the appendix.

Finally I discuss numerical issues like scale dependence of the amplitudes and branching ratios. It will turn out that the NLO contributions of the hard spectator interactions are important but small enough for perturbation theory to be valid.

## Chapter 1

## Introduction

The present situation of particle physics is the following. On the one hand we have got an extremely successful standard model that describes physics up to energy scales current accelerators are able to reach. On the other hand it has limitations and problems, e.g. the arbitrariness of the standard model parameters, the fact that the Higgs particle has not yet been found, the stabilisation of the Higgs mass under loop corrections (fine tuning problem) or the question why the electroweak scale is so much lower than the Planck scale (hierarchy problem). So most particle physicists expect new physics to show up at energy scales that are beyond the range of present accelerators but will be reached by future colliders. Within the next year LHC at CERN will start running and in the following years will collect data from proton proton collisions at a centre of mass energy of 14 TeV . This will allow us to obtain information about new physics by producing not yet observed particles directly. On the other hand physics beyond the standard model can be discovered by precision measurements of low energy quantities which are influenced by new physics particles because of quantum effects. The currently running experiments BaBar and Belle and after the start of LHC also LHCb are dedicated to examine decays of $B$-mesons, where new physics is expected to be seen in CP asymmetries.

However in order to find new physics by indirect search, some parameters of the standard model have to be determined more precisely. To this end LHCb will make an important contribution. Above all the Wolfenstein parameters $\bar{\rho}$ and $\bar{\eta}[1,2]$ that occur in the parametrisation of the Cabibbo-Kobayashi-Maskawa (CKM) matrix and determine its complex phase, which leads to CP asymmetry in the standard model, are up to now only very imprecisely known [3]:

$$
\begin{equation*}
\bar{\rho}=0.182_{-0.047}^{+0.045} \quad \bar{\eta}=0.332_{-0.036}^{+0.032} . \tag{1.1}
\end{equation*}
$$

These parameters, which influence weak interactions of quarks, can be determined with higher accuracy by $B$-meson decays. In order to reduce their large uncertainties on the experimental side better statistics is needed, which is expected to be improved in the next few years, and on the theoretical side we have to get hadronic physics under control. This is due to the fact that weak interactions of quarks, from which $\bar{\rho}$ and $\bar{\eta}$ are measured, are always spoiled by non-perturbative strong interactions, because quarks are bound in hadronic states like mesons. Hadronic physics, however,
is governed by the energy scale $\Lambda_{\mathrm{QCD}}$, where QCD cannot be handled perturbatively.
There are several advantages in observing $B$-meson decays. One of them is due to the production of $B$-mesons itself: There exists an extremely clean source to produce $B$-mesons: The resonance $\Upsilon(4 S)$, a bound state of a $b \bar{b}$ pair, has a mass that is only slightly larger than twice the mass of the $B$-meson and decays nearly completely into $B \bar{B}$ pairs. This resonance is used at the $B$-factories BaBar and Belle.

Another advantage is the possibility to obtain clean information about the complex phase of the CKM matrix by measurement of quantum mechanical oscillations in the $B-\bar{B}$ system. The lifetime of $B$-mesons, which is about 1.5 ps , is large enough to observe those oscillations in the detector [4, 5. By measuring the time dependent CP violation it is possible to obtain the CKM angles $\alpha, \beta$ and $\gamma$, which determine $\bar{\rho}$ and $\bar{\eta}$, with small hadronic uncertainties (see e.g. chapter 1 of [6]). The "golden channel" $B \rightarrow J / \psi K_{S}$, where the dependence on hadronic quantities is strongly suppressed by small CKM parameters, allows a quite precise determination of $\sin (2 \beta)=0.687 \pm 0.032$ [7]. In the same way the decay $B \rightarrow \pi \pi$ could be used for a precise determination of the angle $\alpha$. However other than in the "golden channel" in the case of $B \rightarrow \pi \pi$ hadronic physics plays a subdominant but non-negligible role.

At this point another convenient property of the $B$-meson comes into play. The mass of the $b$-quark introduces a hard scale, at which $\alpha_{s}$ is small enough to make perturbation theory possible. However the bound state of the $b$-quark in the $B$ meson is dominated by physics of the soft scale $\Lambda_{\mathrm{QCD}}$, where perturbation theory breaks down. While inclusive decays can be handled in the framework of operator product expansion, for exclusive decays, which the present thesis deals with, the framework of QCD factorization has been proposed [8, 9]. This framework makes it possible to disentangle the soft and hard physics at leading power in an expansion in $\Lambda_{\mathrm{QCD}} / m_{b}$. Decay amplitudes are then obtained in perturbative expansions, which come with hadronic parameters that have to be determined in experiment or by non perturbative methods like QCD sum rules or lattice QCD.

Whereas the $\alpha_{s}$ corrections for the transition matrix elements of $B \rightarrow \pi \pi$ have been calculated in [10], the present thesis deals with the $\mathcal{O}\left(\alpha_{s}^{2}\right)$ contribution of a specific part of the amplitude. This part, which consists of the hard spectator interaction Feynman diagrams, will be defined in the next section. There I will also argue, that it is reasonable to consider the hard spectator interactions separately. The calculation of the rest of the $\mathcal{O}\left(\alpha_{s}^{2}\right)$ corrections has been partly performed by Guido Bell in his PhD thesis [11, 12]. There the complete imaginary part and a preliminary result of the real part of the amplitude is given. My calculation of the hard spectator scattering amplitude is not the first one as it has been calculated recently by [13, 14]. It is however the first pure QCD calculation, whereas [13, 14] used the framework of soft-collinear effective theory (SCET) [15, 16, 17] an effective theory, where the expansion in $\Lambda_{\mathrm{QCD}} / m_{b}$ is performed at the level of the Lagrangian rather than of Feynman integrals. It is the main result of this thesis to confirm the results of [13, 14] and to show by explicit calculation that pure QCD and SCET lead to the same result in this special case.

From a technical point of view the calculation in this thesis consists of the evaluation of about 60 one-loop Feynman diagrams. The challenges of this task are due to the fact that these diagrams come with up to five external legs and three independent ratios of scales. In order to reduce the number of master integrals and to perform power expansions of the Feynman integrals, integration by parts methods and differential equation techniques will prove appropriate tools. Most parts of the calculation will be performed by a computer algebra system, whereas the algorithms and the necessary steps to obtain input results for the programs will be discussed in detail. As I did not obtain the completed $\mathcal{O}\left(\alpha_{s}^{2}\right)$ corrections, the phenomenological part of this thesis is restricted to the reproduction of the branching ratios numerically obtained in [13]. Other observables like CP asymmetries are not improved by my partial result alone.

This thesis is organised as follows: In chapter 2 I start with an introduction to QCD factorization and define in this framework the hard spectator scattering amplitude. After defining my notations I demonstrate the calculation of the LO of the hard spectator interactions and end the chapter by explaining the technical details of the integration by parts methods and differential equation techniques.

Chapter 3 is the most technical of all. There all of the Feynman diagrams that contribute are listed and their evaluation is discussed in detail. Furthermore NLO corrections to the wave functions and evanescent operators occurring at this order are dealt with.

After presenting the complete analytical results and the numerical analysis in chapter 4 I end up with the conclusions.

## Chapter 2

## Preliminaries

### 2.1 Hard spectator interactions and QCD factorization

Though the decay of the $B$-meson is caused by weak interactions, strong interactions play a dominant role. It is however not possible to handle the QCD effects completely perturbatively. This is due to the energy scales that are contained in the $B$-meson: Whereas $\alpha_{s}$ at the mass of the $b$-quark is a small parameter, the bound state of the quarks leads to an energy scale of $\mathcal{O}\left(\Lambda_{\mathrm{QCD}}\right)$ which spoils perturbation theory. The idea of QCD factorization [8, 9] is to separate these scales. At leading power in $\Lambda_{\mathrm{QCD}} / m_{b}$ we obtain the amplitude for $B \rightarrow \pi \pi$ in the following form:

$$
\begin{align*}
\langle\pi \pi| \mathcal{H}|B\rangle \sim & F^{B \rightarrow \pi} \int_{0}^{1} d x T^{\mathrm{I}}(x) f_{\pi} \phi_{\pi}(x)+ \\
& \int_{0}^{1} d x d y d \xi T^{\mathrm{II}}(x, y, \xi) f_{B} \phi_{B 1}(\xi) f_{\pi} \phi_{\pi}(x) f_{\pi} \phi_{\pi}(y) \tag{2.1}
\end{align*}
$$

Two different types of quantities enter this formula. On the one hand the hadronic physics is contained in the form factor $F^{B \rightarrow \pi}$ and the wave functions $\phi_{B 1}$ and $\phi_{\pi}$, which will be defined in the next section more precisely. These quantities contain the information about the bound states of the mesons. They have to be determined by non-perturbative methods like QCD sum rules or lattice calculations. Alternatively, because they are at least partly process independent, they might be extracted in the future from experiment. On the other hand the hard scattering kernels $T^{\mathrm{I}}$ and $T^{\mathrm{II}}$ contain the physics of the hard scale $\mathcal{O}\left(m_{b}\right)$ and the hard collinear scale $\mathcal{O}\left(\sqrt{m_{b} \Lambda_{\mathrm{QCD}}}\right)$ and can be calculated perturbatively.

Here I would like to make two remarks to (2.1):
First I want to note that (2.1) is only valid in leading power in the expansion in $\Lambda_{\mathrm{QCD}} / m_{b}$. Higher orders in this expansion lead to endpoint singularities i.e. the integrals over the variables $x, y$ and $\xi$ diverge at the endpoints. This leads to a mixture of the physics of the soft scale into $T^{\mathrm{I}}$ and $T^{\mathrm{II}}$ and spoils QCD factorization. However there are corrections that are formally of subleading power but numerically


Figure 2.1: Tree level, vertex correction and penguin contraction. These diagrams contribute to $T^{\mathrm{I}}$.


Figure 2.2: Hard spectator interactions at $\mathcal{O}\left(\alpha_{s}\right)$. This is the LO of $T^{\mathrm{II}}$
enhanced and cannot be handled within the framework of QCD factorization. They have to be estimated in the numerical analysis.

The second remark concerns the separation of the hard scattering kernel into $T^{\mathrm{I}}$ and $T^{\mathrm{II}}$. The Feynman diagrams that contribute to $B \rightarrow \pi \pi$ can be distributed into two different classes. The class of diagrams where there is no hard interaction of the spectator quark (fig. 2.1) contributes to $T^{\mathrm{I}}$. The hard spectator scattering diagrams, which are shown in LO in $\alpha_{s}$ in fig. 2.2, contribute to $T^{\mathrm{II}}$. Through the soft momentum $l$ of the constituent quark of the $B$-meson the hard collinear scale $\sqrt{\Lambda_{\mathrm{QCD}} m_{b}}$ comes into play. This leads to the fact that in contrast to $T^{\mathrm{I}}$, which is completely governed by the scale $m_{b}, T^{\mathrm{II}}$ has to be evaluated at the hardcollinear scale. This leads to an enhancement of $\alpha_{s}$ and makes the NLO corrections (i.e. $\mathcal{O}\left(\alpha_{s}^{2}\right)$ ) of the hard spectator interaction diagrams more important. These $\alpha_{s}^{2}$ corrections of $T^{\mathrm{II}}$ are the topic of the present thesis. Hard spectator scattering corrections to the penguin diagram (third diagram of fig. 2.1) are beyond the scope of this thesis. The cancellation of the dependence on the renormalisation scale of this class of diagrams is completely independent of the "tree amplitude" i.e. the diagrams of fig. 2.2 and higher order $\alpha_{s}$ corrections. For phenomenological applications, however, they should be taken into account.

### 2.2 Notation and basic formulas

### 2.2.1 Kinematics

For the process $B \rightarrow \pi \pi$ we will assign the momenta $p$ and $q$ to the pions which fulfil the condition

$$
\begin{equation*}
p^{2}, q^{2}=0 . \tag{2.2}
\end{equation*}
$$

This is the leading power approximation in $\Lambda_{\mathrm{QCD}} / m_{b}$ as we count the mass of the pion as $\mathcal{O}\left(\Lambda_{\mathrm{QCD}}\right)$. Let us define two Lorentz vectors $n_{+}, n_{-}$by:

$$
\begin{equation*}
n_{+}^{\mu} \equiv(1,0,0,1), \quad n_{-}^{\mu} \equiv(1,0,0,-1) . \tag{2.3}
\end{equation*}
$$

In the rest frame of the decaying meson $p$ can be defined to be in the direction of $n_{+}$and $q$ to be in the direction of $n_{-}$. Light cone coordinates for the Lorentz vector $z^{\mu}$ are defined by:

$$
\begin{equation*}
z^{+} \equiv \frac{z^{0}+z^{3}}{\sqrt{2}}, \quad z^{-} \equiv \frac{z^{0}-z^{3}}{\sqrt{2}}, \quad z_{\perp} \equiv\left(0, z^{1}, z^{2}, 0\right) \tag{2.4}
\end{equation*}
$$

So one can decompose $z^{\mu}$ into:

$$
\begin{equation*}
z^{\mu}=\frac{z \cdot p}{p \cdot q} q^{\mu}+\frac{z \cdot q}{p \cdot q} p^{\mu}+z_{\perp}^{\mu} \tag{2.5}
\end{equation*}
$$

such that

$$
\begin{equation*}
z_{\perp} \cdot p=z_{\perp} \cdot q=0 . \tag{2.6}
\end{equation*}
$$

We denote the mass of the $B$-meson with $m_{B}$ and the mass of the $b$-quark with $m_{b}$. The difference $m_{B}-m_{b}=\mathcal{O}\left(\Lambda_{\mathrm{QCD}}\right)$ such that we cannot distinguish those masses in leading power. However setting

$$
\begin{equation*}
m_{b}=m_{B} \tag{2.7}
\end{equation*}
$$

in Feynman integrals might lead to additional infrared divergences. So we have to perform the integral before we can make the substitution (2.7) unless we are sure that we do not produce infrared divergences. If we calculate Feynman integrals, it is convenient to set

$$
\begin{equation*}
m_{B}=1 \tag{2.8}
\end{equation*}
$$

such that $p \cdot q=\frac{1}{2}$. The dependence on $m_{B}$ can be reconstructed by giving the correct mass dimension to the physical quantities.

### 2.2.2 Colour factors

In our calculations we will use the following three colour factors, which arise from the $\mathrm{SU}(3)$ algebra:

$$
\begin{equation*}
C_{N}=\frac{1}{2}, \quad C_{F}=\frac{N_{c}^{2}-1}{2 N_{c}} \quad \text { and } \quad C_{G}=N_{c}, \tag{2.9}
\end{equation*}
$$

where $N_{c}=3$ is the number of colours.

### 2.2.3 Meson wave functions

The pion light cone distribution amplitude $\phi_{\pi}$ is defined by

$$
\begin{equation*}
\langle\pi(p)| \bar{q}(z)_{\alpha}[\ldots] q^{\prime}(0)_{\beta}|0\rangle_{z^{2}=0}=\frac{i f_{\pi}}{4}\left(\not p \gamma_{5}\right)_{\beta \alpha} \int_{0}^{1} d x e^{i x p \cdot z} \phi_{\pi}(x) . \tag{2.10}
\end{equation*}
$$

The ellipsis [...] stands for the Wilson line

$$
\begin{equation*}
[z, 0]=\mathrm{P} \exp \left(\int_{0}^{1} d t i g_{s} z \cdot A(z t)\right) \tag{2.11}
\end{equation*}
$$

which makes 2.10 gauge invariant. For the definition of the $B$-meson wave function $\phi_{B 1}$ we need the special kinematics of the process. Following [9] let us define

$$
\begin{equation*}
\Psi_{B}^{\alpha \beta}\left(z, p_{B}\right)=\langle 0| \bar{q}_{\beta}(z)[\ldots] b_{\alpha}(0)\left|B\left(p_{B}\right)\right\rangle=\int \frac{d^{4} l}{(2 \pi)^{4}} e^{-i l \cdot z} \phi_{B}^{\alpha \beta}\left(l, p_{B}\right) . \tag{2.12}
\end{equation*}
$$

In the calculation of matrix elements we get terms like:

$$
\begin{equation*}
\int \frac{d^{4} l}{(2 \pi)^{4}} \operatorname{tr}\left(A(l) \phi_{B}(l)\right)=\int \frac{d^{4} l}{(2 \pi)^{4}} \int d^{4} z e^{i l \cdot z} \operatorname{tr}\left(A(l) \Psi_{B}(z)\right) . \tag{2.13}
\end{equation*}
$$

We will only consider the case that the dependence of the amplitude $A$ on $l$ is like this:

$$
\begin{equation*}
A(l)=A(2 l \cdot p) \tag{2.14}
\end{equation*}
$$

In this case we can use the $B$-meson wave function on the light cone which is given by (9):

$$
\begin{align*}
& \left.\langle 0| \bar{q}_{\alpha}(z)[\ldots] b_{\beta}(0)\left|B\left(p_{B}\right)\right\rangle\right|_{z^{-}, z_{\perp}=0}  \tag{2.15}\\
& \quad=-\frac{i f_{B}}{4}\left[\left(\not p_{B}+m_{b}\right) \gamma_{5}\right]_{\beta \gamma} \int_{0}^{1} d \xi e^{-i \xi p_{B}^{-} z^{+}}\left[\Phi_{B 1}(\xi)+\not n_{+} \Phi_{B 2}(\xi)\right]_{\gamma \alpha}
\end{align*}
$$

where

$$
\begin{equation*}
\int_{0}^{1} d \xi \Phi_{B 1}(\xi)=1 \quad \text { and } \quad \int_{0}^{1} d \xi \Phi_{B 2}(\xi)=0 \tag{2.16}
\end{equation*}
$$

It is now straight forward to write down the momentum projector of the $B$-meson:

$$
\begin{align*}
& \int \frac{d^{4} l}{(2 \pi)^{4}} \operatorname{tr}(A(2 l \cdot p) \hat{\Psi}(l)) \\
& \quad=\frac{-i f_{B}}{4} \operatorname{tr}\left(\not \not_{B}+m_{B}\right) \gamma_{5} \int_{0}^{1} d \xi\left(\Phi_{B 1}(\xi)+\not n_{+} \Phi_{B 2}(\xi)\right) A\left(\xi m_{B}^{2}\right) \tag{2.17}
\end{align*}
$$

At this point we give the following definitions

$$
\begin{align*}
\frac{m_{B}}{\lambda_{B}} & \equiv \int_{0}^{1} \frac{d \xi}{\xi} \phi_{B 1}(\xi)  \tag{2.18}\\
\lambda_{n} & \equiv \frac{\lambda_{B}}{m_{B}} \int_{0}^{1} \frac{d \xi}{\xi} \ln ^{n} \xi \phi_{B 1}(\xi) \tag{2.19}
\end{align*}
$$

### 2.2.4 Effective weak Hamiltonian

The effective weak Hamiltonian which leads to $B \rightarrow \pi \pi$ is given by [18]:

$$
\begin{equation*}
\mathcal{H}_{\mathrm{eff}}=\frac{G_{F}}{\sqrt{2}} \sum_{p=u, c} \lambda_{p}^{\prime}\left[C_{1} \mathcal{O}_{1}+C_{2} \mathcal{O}_{2}+\sum_{i=3 \ldots 6} C_{i} \mathcal{O}_{i}+C_{8 g} \mathcal{O}_{8 g}\right]+\text { h.c. } \tag{2.20}
\end{equation*}
$$

where $\lambda_{p}^{\prime}=V_{p d}^{*} V_{p b}$ and

$$
\begin{align*}
\mathcal{O}_{1} & =(\bar{d} p)_{V-A}(\bar{p} b)_{V-A},  \tag{2.21}\\
\mathcal{O}_{2} & =\left(\bar{d}_{i} p_{j}\right)_{V-A}\left(\bar{p}_{j} b_{i}\right)_{V-A},  \tag{2.22}\\
\mathcal{O}_{3} & =(\bar{d} b)_{V-A} \sum_{q}(\bar{q} q)_{V-A},  \tag{2.23}\\
\mathcal{O}_{4} & =\left(\bar{d}_{i} b_{j}\right)_{V-A} \sum_{q}\left(\bar{q}_{j} q_{i}\right)_{V-A},  \tag{2.24}\\
\mathcal{O}_{5} & =(\bar{d} b)_{V-A} \sum_{q}(\bar{q} q)_{V+A},  \tag{2.25}\\
\mathcal{O}_{6} & =\left(\bar{d}_{i} b_{j}\right)_{V-A} \sum_{q}\left(\bar{q}_{j} q_{i}\right)_{V+A},  \tag{2.26}\\
\mathcal{O}_{8 g} & =\frac{g}{8 \pi^{2}} m_{b} \bar{d}_{i} \sigma^{\mu \nu}\left(1+\gamma_{5}\right) T_{i j}^{a} b_{j} G_{\mu \nu}^{a} . \tag{2.27}
\end{align*}
$$

Explicit expressions for the short-distance coefficients $C_{i}$ can be obtained from [18]. The decay amplitude of $B \rightarrow \pi \pi$ is given by

$$
\begin{equation*}
\mathcal{A}(B \rightarrow \pi \pi) \equiv\langle\pi \pi| \mathcal{H}_{\mathrm{eff}}|B\rangle \tag{2.28}
\end{equation*}
$$

For later convenience we define

$$
\begin{equation*}
\mathcal{A}(B \rightarrow \pi \pi) \equiv \mathcal{A}(B \rightarrow \pi \pi)^{\mathrm{I}}+\mathcal{A}(B \rightarrow \pi \pi)^{\mathrm{II}} \tag{2.29}
\end{equation*}
$$

where $\mathcal{A}^{\mathrm{I}}\left(\mathcal{A}^{\mathrm{II}}\right)$ belongs to the first (second) term of 2.1). Because $\mathcal{A}^{\mathrm{I}}$ and $\mathcal{A}^{\mathrm{II}}$ contain different hadronic quantities, the renormalisation scale dependence of both of them has to vanish separately. So we can set their scales to different values $\mu^{I}$ and $\mu^{\mathrm{II}}$. As in $\mathcal{A}^{\mathrm{I}}$ there occurs only the mass scale $m_{b}$ we can set $\mu^{\mathrm{I}}=m_{b}$. In $\mathcal{A}^{\text {II }}$ there occurs also the hard-collinear scale $\sqrt{\Lambda_{\mathrm{QCD}} m_{b}}$. As we will see this scale is an appropriate choice for $\mu^{\mathrm{II}}$.

In order to separate the QCD effects from the weak physics we write the matrix elements of the effective weak Hamiltonian in the following factorised form [10]:

$$
\begin{equation*}
\langle\pi \pi| \mathcal{H}_{\mathrm{eff}}|\bar{B}\rangle=\frac{G_{F}}{\sqrt{2}} \sum_{p=u, c} \lambda_{p}^{\prime}\langle\pi \pi| \mathcal{T}_{p}+\mathcal{T}_{p}^{\mathrm{ann}}|\bar{B}\rangle \tag{2.30}
\end{equation*}
$$



Figure 2.3: Annihilation topology. The gluon vertex that is marked by the cross can alternatively be attached to other crosses.
where

$$
\begin{align*}
\mathcal{T}_{p}= & a_{1} \delta_{p u}(\bar{u} b)_{V-A} \otimes(\bar{d} u)_{V-A} \\
& +a_{2} \delta_{p u}(\bar{d} b)_{V-A} \otimes(\bar{u} u)_{V-A} \\
& +a_{3} \sum_{q}(\bar{d} b)_{V-A} \otimes(\bar{q} q)_{V-A} \\
& +a_{4}^{p} \sum_{q}(\bar{q} b)_{V-A} \otimes(\bar{d} q)_{V-A} \\
& +a_{5} \sum_{q}(\bar{d} b)_{V-A} \otimes(\bar{q} q)_{V+A} \\
& +a_{6}^{p} \sum_{q}(-2)(\bar{q} b)_{S-P} \otimes(\bar{d} q)_{S+P} \tag{2.31}
\end{align*}
$$

Note that in contrast to [10] the electroweak corrections to the effective weak Hamiltonian are not included in the above equations as in the case of $B \rightarrow \pi \pi$ they can be safely neglected. $\mathcal{T}_{p}^{\text {ann }}$ stands for the contributions of the annihilation topologies, which are shown in fig. 2.3. These contributions do not occur in leading power and cannot be calculated in a model independent way in the framework of QCD factorization. For the exact definition of $\mathcal{T}_{p}^{\text {ann }}$ I refer to [10]. The matrix elements of the operators $j_{1} \otimes j_{2}$ are defined to be $\langle\pi \pi| j_{1} \otimes j_{2}|\bar{B}\rangle \equiv\langle\pi| j_{1}|\bar{B}\rangle\langle\pi| j_{2}|0\rangle$ or $\langle\pi| j_{2}|\bar{B}\rangle\langle\pi| j_{1}|0\rangle$ corresponding to the flavour structure of the $\pi$-mesons. The penguin contractions that are shown in the third diagram of fig. 2.1] and the contributions of the operators $\mathcal{O}_{3}-\mathcal{O}_{8 g}$ are by definition contained in the amplitudes $a_{3}-a_{6}^{p}$. As we take in the present thesis only the "tree amplitude" (fig. 2.2) into account, we only calculate $\alpha_{s}^{2}$ corrections to the amplitudes $a_{1}$ and $a_{2}$.

The decay amplitudes of $B \rightarrow \pi \pi$ can be written in terms of $a_{i}$ as follows [10]:

$$
\begin{align*}
-\mathcal{A}\left(\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}\right) & =\left[\lambda_{u}^{\prime} a_{1}+\lambda_{p}^{\prime}\left(a_{4}^{p}+r_{\chi}^{\pi} a_{6}^{p}\right)\right] A_{\pi \pi} \\
-\sqrt{2} \mathcal{A}\left(B^{-} \rightarrow \pi^{-} \pi^{0}\right) & =\lambda_{u}^{\prime}\left(a_{1}+a_{2}\right) A_{\pi \pi} \\
\mathcal{A}\left(\bar{B}^{0} \rightarrow \pi^{0} \pi^{0}\right) & =\left[-\lambda_{u}^{\prime} a_{2}+\lambda_{p}^{\prime}\left(a_{4}^{p}+r_{\chi}^{\pi} a_{6}^{p}\right)\right] A_{\pi \pi} \tag{2.32}
\end{align*}
$$

where

$$
A_{\pi \pi}=i \frac{G_{F}}{\sqrt{2}}\left(m_{B}^{2}-m_{\pi}^{2}\right) f_{+}^{B \pi} f_{\pi}
$$

and

$$
\begin{equation*}
r_{\chi}^{\pi}(\mu)=\frac{2 m_{\pi}^{2}}{\bar{m}_{b}(\mu)\left(\bar{m}_{u}(\mu)+\bar{m}_{d}(\mu)\right)} . \tag{2.33}
\end{equation*}
$$

For the LO and NLO results of the $a_{i}$ I refer to [10].
The annihilation contributions are parametrised in the following form [10]:

$$
\begin{align*}
-\mathcal{A}_{\mathrm{ann}}\left(\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}\right) & =\left[\lambda_{u}^{\prime} b_{1}+\left(\lambda_{u}^{\prime}+\lambda_{c}^{\prime}\right)\left(b_{3}+2 b_{4}\right)\right] B_{\pi \pi} \\
\mathcal{A}_{\mathrm{ann}}\left(B^{-} \rightarrow \pi^{-} \pi^{0}\right) & =0 \\
\mathcal{A}_{\mathrm{ann}}\left(\bar{B}^{0} \rightarrow \pi^{0} \pi^{0}\right) & =-\mathcal{A}_{\mathrm{ann}}\left(\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}\right) \tag{2.34}
\end{align*}
$$

where

$$
\begin{equation*}
B_{\pi \pi}=i \frac{G_{F}}{\sqrt{2}} f_{B} f_{\pi}^{2} \tag{2.35}
\end{equation*}
$$

The parameters $b_{i}$ can be further parametrised by the Wilson coefficients occurring in (2.20) and purely hadronic quantities:

$$
\begin{align*}
b_{1} & =\frac{C_{F}}{N_{c}^{2}} C_{1} A_{1}^{i} \\
b_{3} & =\frac{C_{F}}{N_{c}^{2}}\left[C_{3} A_{1}^{i}+C_{5}\left(A_{3}^{i}+A_{3}^{f}\right)+N_{c} C_{6} A_{3}^{f}\right] \\
b_{4} & =\frac{C_{F}}{N_{c}^{2}}\left[C_{4} A_{1}^{i}+C_{6} A_{2}^{i}\right] \tag{2.36}
\end{align*}
$$

where the quantities $A_{k}^{i(f)}$ are given by [10]:

$$
\begin{align*}
A_{1}^{i} & =\pi \alpha_{s}\left[18\left(X_{A}-4+\frac{\pi^{2}}{3}\right)+2 r_{\chi}^{\pi 2} X_{A}^{2}\right] \\
A_{2}^{i} & =A_{1}^{i} \\
A_{3}^{i} & =0 \\
A_{3}^{f} & =12 \pi \alpha_{s} r_{\chi}^{\pi}\left(2 X_{A}^{2}-X_{A}\right) \tag{2.37}
\end{align*}
$$

Here $r_{\chi}^{\pi}$ is defined as in (2.33) and $X_{A}$ parametrises an integral that is divergent because of endpoint singularities. In section 4.4.2 I will give an estimate of $X_{A}$ for numerical calculations.

### 2.3 Hard spectator interactions at LO

The leading order of the hard spectator interactions which start at $\mathcal{O}\left(\alpha_{s}\right)$ is shown in fig. 2.2. The hard spectator scattering kernel $T^{\mathrm{II}}$, which does not depend on the wave functions, can be obtained by calculating the transition matrix element between free external quarks, to which we assign the momenta shown in fig. 2.2. The variables $x, \bar{x} \equiv 1-x, y, \bar{y} \equiv 1-y$ are the arguments of $T^{\mathrm{II}}$, which arise from the projection on the pion wave function 2.10 . In the sense of power counting we
count all components of $l$ of $\mathcal{O}\left(\Lambda_{\mathrm{QCD}}\right)$, while the components of $p$ and $q$ are $\mathcal{O}\left(m_{b}\right)$ or exactly zero. We define the following quantities

$$
\begin{align*}
\xi & \equiv \frac{l \cdot p}{p \cdot q} \\
\theta & \equiv \frac{l \cdot q}{p \cdot q} \tag{2.38}
\end{align*}
$$

We will see that in the end the dependence on $\theta$ vanishes in leading power such that we can use (2.17).

We consider the three cases $\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}, \bar{B}^{0} \rightarrow \pi^{0} \pi^{0}$ and $B^{-} \rightarrow \pi^{-} \pi^{0}$. In the case, that the external quarks come with the flavour content of $\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}$, the LO hard spectator amplitude for the effective operator $\mathcal{O}_{2}$ reads:

$$
\begin{align*}
& A_{\text {spect. }}^{(1)}\left(\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}\right) \equiv \\
& \quad\langle\bar{d}(\bar{x} p) u(x p) \bar{u}(\bar{y} q) d(y q)| \mathcal{O}_{2}|\bar{d}(l) b(p+q-l)\rangle_{\text {spect. }}= \\
& \quad 4 \pi \alpha_{s} C_{F} N_{c} \frac{1}{\bar{x} \xi m_{B}^{2}} \bar{d}(l) \gamma^{\mu} d(\bar{x} p) \bar{u}(x p) \gamma^{\nu}\left(1-\gamma_{5}\right) b(p+q-l) \\
& \quad \bar{d}(y q)\left(\frac{2 \not p g_{\mu \nu}}{\bar{y}}-\frac{p}{y \bar{y}} \gamma_{\mu} \gamma_{\nu}\right)\left(1-\gamma_{5}\right) u(\bar{y} q), \tag{2.39}
\end{align*}
$$

where the quark antiquark states in the input and output channels of the matrix element form colour singlets. The subscript "spect." means that only diagrams with a hard spectator interaction are taken into account. The amplitude of $\mathcal{O}_{1}$ vanishes to this order in $\alpha_{s}$. In the case of $\bar{B}^{0} \rightarrow \pi^{0} \pi^{0}$ we get the tree amplitude from the matrix element of $\mathcal{O}_{1}$. The case $B^{-} \rightarrow \pi^{-} \pi^{0}$ does not need to be considered separately, because from isospin symmetry follows [19, 10]:

$$
\begin{equation*}
\sqrt{2} \mathcal{A}\left(B^{-} \rightarrow \pi^{-} \pi^{0}\right)=\mathcal{A}\left(\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}\right)+\mathcal{A}\left(\bar{B}^{0} \rightarrow \pi^{0} \pi^{0}\right) \tag{2.40}
\end{equation*}
$$

On the other hand the full amplitude is the convolution of $T^{\mathrm{II}}$ with the wave functions, given by 2.1. To extract $T^{\mathrm{II}}$ from (2.39) we need the wave functions with the same external states we have used in (2.39), i.e. we have to calculate the matrix elements (2.10) and 2.12), where the pion or $B$-meson states are replaced by free external quark states. To the order $\mathcal{O}\left(\alpha_{s}^{0}\right)$ we get

$$
\begin{align*}
\phi_{\pi^{-} \alpha \beta}^{(0)}\left(y^{\prime}\right) & \equiv \int d(z \cdot q) e^{-i z \cdot q y^{\prime}}\langle\bar{u}(\bar{y} q) d(y q)| \bar{d}_{\beta}^{i}(z) u_{\alpha}^{i}(0)|0\rangle_{z^{-}, z_{\perp}=0} \\
& =2 \pi N_{c} \delta\left(y^{\prime}-y\right) \bar{d}_{\beta}(y q) u_{\alpha}(\bar{y} q) \\
\phi_{\pi^{+} \alpha \beta}^{(0)}\left(x^{\prime}\right) & =2 \pi N_{c} \delta\left(x^{\prime}-x\right) \bar{u}_{\beta}(x p) d_{\alpha}(\bar{x} p)  \tag{2.41}\\
\phi_{B \alpha \beta}^{(0)}\left(l^{\prime-}\right) & \equiv \int d z^{+} e^{i l^{\prime}-z^{+}}\langle 0| \bar{d}_{\beta}(z)^{i} b_{\alpha}(0)^{i}|\bar{d}(l) b(p+q-l)\rangle_{z^{-}, z_{\perp}=0} \\
& =2 \pi N_{c} \delta\left(l^{\prime-}-l^{-}\right) \bar{d}_{\beta}(l) b_{\alpha}(p+q-l)
\end{align*}
$$

By using

$$
\begin{equation*}
A_{\text {spect. }}^{(1)}=\int d x d y d l^{-} \phi_{\pi^{+} \alpha \alpha^{\prime}}^{(0)}(x) \phi_{\pi^{-} \beta \beta^{\prime}}^{(0)}(y) \phi_{B \gamma \gamma^{\prime}}^{(0)}\left(l^{-}\right) T^{\mathrm{II}(1)}\left(x, y, l^{-}\right)_{\alpha^{\prime} \alpha \beta^{\prime} \beta \gamma^{\prime} \gamma} \tag{2.42}
\end{equation*}
$$

we finally obtain:

$$
\begin{align*}
T^{\mathrm{II}(1)}\left(x, y, l^{-}\right)_{\alpha^{\prime} \alpha \beta^{\prime} \beta \gamma^{\prime} \gamma}= & 4 \pi \alpha_{s} \frac{C_{F}}{(2 \pi)^{3} N_{c}^{2}} \frac{1}{\xi \bar{x} m_{B}^{2}} \gamma_{\gamma^{\prime} \alpha}^{\mu}\left[\gamma^{\nu}\left(1-\gamma_{5}\right)\right]_{\alpha^{\prime} \gamma} \\
& {\left[\left(\frac{2 \not p g_{\mu \nu}}{\bar{y}}-\frac{\not p}{y \bar{y}} \gamma_{\mu} \gamma_{\nu}\right)\left(1-\gamma_{5}\right)\right]_{\beta^{\prime} \beta} } \tag{2.43}
\end{align*}
$$

It should be noted that only the first summand of the above equation contributes after performing the Dirac trace in four dimensions. The second summand is evanescent. This will be important, when we will calculate the NLO corrections of the wave functions (see section 3.3).

If we plug the hadronic wave functions defined by (2.10) and (2.17) into (2.42) i.e. we calculate the matrix element $(2.39)$ between meson states instead of free quark states, we get for the LO amplitude ${ }^{1}$ :

$$
\begin{equation*}
A_{\text {spect. }}^{(1)}=-\frac{i f_{\pi}^{2} f_{B} C_{F}}{4 N_{c}^{2}} 4 \pi \alpha_{s} \int_{0}^{1} d x d y d \xi \Phi_{B 1}(\xi) \phi_{\pi}(x) \phi_{\pi}(y) \frac{1}{\xi \bar{x} \bar{y}} . \tag{2.44}
\end{equation*}
$$

Following (2.1) and the conventions of [13] we write our amplitude in the form:

$$
\begin{equation*}
A_{\text {spect } . i}=-i m_{B}^{2} \int_{0}^{1} d x d y d \xi T_{i}^{\mathrm{II}}(x, y, \xi) f_{B} \Phi_{B 1}(\xi) f_{\pi} \phi_{\pi}(x) f_{\pi} \phi_{\pi}(y) \tag{2.45}
\end{equation*}
$$

where in the case of $\bar{B} \rightarrow \pi^{+} \pi^{-}$we define

$$
\begin{align*}
& A_{\text {spect.1 }}=\left\langle\mathcal{O}_{2}\right\rangle_{\text {spect. }} \\
& A_{\text {spect.2 }}=\left\langle\mathcal{O}_{1}\right\rangle_{\text {spect. }} \tag{2.46}
\end{align*}
$$

and in the case $\bar{B} \rightarrow \pi^{0} \pi^{0}$ we define

$$
\begin{align*}
A_{\text {spect. } 1} & =\left\langle\mathcal{O}_{1}\right\rangle_{\text {spect. }} \\
A_{\text {spect. } 2} & =\left\langle\mathcal{O}_{2}\right\rangle_{\text {spect. }} \tag{2.47}
\end{align*}
$$

Because we use the NDR-scheme which preserves Fierz transformations for $\mathcal{O}_{1}$ and $\mathcal{O}_{2}, T_{i}^{\text {II }}$ has the same form for both decay channels. From 2.44 and 2.45 we get:

$$
\begin{align*}
& T_{1}^{\mathrm{II}(1)}=4 \pi \alpha_{s} \frac{C_{F}}{4 N_{c}^{2}} \frac{1}{\xi \bar{x} \bar{y} m_{B}^{2}} \\
& T_{2}^{\mathrm{II}(1)}=0 \tag{2.48}
\end{align*}
$$

According to [10] the contribution of (2.44) to $a_{1}$ and $a_{2}$ (see (2.31) is given by:

$$
\begin{align*}
& a_{1, \mathrm{II}}=\frac{C_{2} C_{F} \pi \alpha_{s}}{N_{c}^{2}} H_{\pi \pi} \\
& a_{2, \mathrm{II}}=\frac{C_{1} C_{F} \pi \alpha_{s}}{N_{c}^{2}} H_{\pi \pi} \tag{2.49}
\end{align*}
$$

[^0]where
\[

$$
\begin{equation*}
H_{\pi \pi}=\frac{f_{B} f_{\pi}}{m_{B}^{2} f_{+}^{B \pi}} \int_{0}^{1} \frac{d \xi}{\xi} \Phi_{B 1}(\xi) \int_{0}^{1} \frac{d x}{\bar{x}} \phi_{\pi}(x) \int_{0}^{1} \frac{d y}{\bar{y}} \phi_{\pi}(y) \tag{2.50}
\end{equation*}
$$

\]

and the label II in 2.49) denotes the contribution to $\mathcal{A}^{\mathrm{II}}$ as defined in 2.29.

### 2.4 Calculation techniques for Feynman integrals

### 2.4.1 Integration by parts method

Integration by parts (IBP) identities were introduced in [20, 21]. An algorithm to reduce Feynman integrals by IBP-identities to master integrals is very well described in [22]. So I will only show the basic principles. Because the topic of my thesis is a one-loop calculation I will restrict to the one-loop case, the generalisation to multi loop is straight forward.

The most general form ${ }^{2}$ of a one-loop Feynman integral is

$$
\begin{align*}
& \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\left(k \cdot p_{j_{1}}\right)^{n_{1}} \ldots\left(k \cdot p_{j_{l}}\right)^{n_{l}}}{\left[\left(k+p_{i_{1}}\right)^{2}-M_{1}^{2}\right]^{m_{1}} \ldots\left[\left(k+p_{i_{t}}\right)^{2}-M_{t}^{2}\right]^{m_{t}}} \times \\
&  \tag{2.51}\\
& {\left[k \cdot p_{\tilde{i}_{1}}+\tilde{M}_{1}^{2}\right]^{\tilde{m}_{1}} \ldots\left[k \cdot p_{\tilde{i}_{u}}+\tilde{M}_{u}^{2}\right]^{\tilde{m}_{u}}} \\
&
\end{align*} \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}},
$$

where $n_{1} \ldots n_{l}, m_{1} \ldots m_{t}, \tilde{m}_{1} \ldots \tilde{m}_{u} \geq 0, j_{1} \ldots, j_{l}, i_{1} \ldots, i_{t}, \tilde{i}_{1}, \ldots, \tilde{i}_{u} \in\{1, \ldots, n\}$ and $p_{1}, \ldots, p_{n}$ are the momenta which appear in the internal propagator lines. Without loss of generality we can assume that there is no $k^{2}$ in the numerator as we can make the replacement

$$
\begin{equation*}
k^{2}=D_{1}+M_{1}^{2}-p_{i_{1}}^{2}-2 k \cdot p_{i_{1}} . \tag{2.52}
\end{equation*}
$$

Because of our special kinematics we have only three linearly independent momenta $p, q, l$, so all of the momenta $p_{1}, \ldots, p_{n}$ are linear combinations of $p, q, l$. This will simplify the reduction of the Feynman integrals. We will define

$$
\begin{equation*}
\mathbf{B} \equiv\left\{\tilde{p}_{1}, \ldots, \tilde{p}_{k}\right\} \tag{2.53}
\end{equation*}
$$

to be a basis of $\operatorname{span}\left\{p_{1}, \ldots, p_{n}\right\}$ where $k \leq 3$ and $\tilde{p}_{1}, \ldots, \tilde{p}_{k} \in\left\{p_{1}, \ldots, p_{n}\right\}$.
Following [23] we can reduce (2.51) by performing algebraic transformations on the integrands, which are defined in the following three rules:

Rule 1. Consider the case that there exist $\left\{c_{1}, \ldots, c_{t}\right\}$ such that

$$
\begin{equation*}
\sum_{j=1}^{t} c_{j} p_{i_{j}}=0 \quad \text { and } \quad \sum_{j=1}^{t} c_{j}=1 \tag{2.54}
\end{equation*}
$$

[^1]Now we can make the following simplification $(l \in\{1 \ldots, t\})$ :

$$
\begin{align*}
& \frac{k \cdot p_{i_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}}} \\
& \quad=\frac{1}{2} \frac{D_{l}-\sum_{j=1}^{t} c_{j}\left(D_{j}+M_{j}^{2}-p_{i_{j}}^{2}\right)+M_{l}^{2}-p_{i_{l}}^{2}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}}} \\
& \quad=\frac{1}{2} \sum_{j=1}^{t}\left(\delta_{j l}-c_{j}\right)\left[\frac{1}{D_{1}^{m_{1}} \ldots D_{j}^{m_{j}-1} \ldots D_{t}^{m_{t}}}+\frac{M_{j}^{2}-p_{i_{j}}^{2}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}}}\right] \tag{2.55}
\end{align*}
$$

and the scalar product $k \cdot p_{i_{l}}$ has disappeared from the numerator. If (2.54) cannot be fulfilled we use the identity

$$
\begin{equation*}
k \cdot p_{i_{l}}=\frac{1}{2}\left(D_{l}-D_{1}+M_{l}^{2}-M_{1}^{2}+p_{1}^{2}-p_{i_{l}}^{2}\right)+k \cdot p_{i_{1}} \tag{2.56}
\end{equation*}
$$

to reduce our set of integrals further. This identity does not reduce the total number of scalar products in the numerator but the number of different scalar products.
Rule 2. For scalar products of the form $k \cdot p_{\tilde{i}_{j}}$ we make the replacement

$$
\begin{equation*}
\frac{k \cdot p_{\tilde{i}_{j}}}{\tilde{D}_{j}^{\tilde{m}_{j}}}=\frac{1}{\tilde{D}_{j}^{\tilde{m}_{j}-1}}-\frac{\tilde{M}_{j}^{2}}{\tilde{D}_{j}^{\tilde{m}_{j}}} . \tag{2.57}
\end{equation*}
$$

Rule 3. Now consider the case that our integrand is of the form

$$
\begin{equation*}
\frac{k \cdot p_{k}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \tag{2.58}
\end{equation*}
$$

where $p_{k} \notin\left\{p_{i_{1}}, \ldots, p_{i_{t}}, p_{\tilde{i}_{1}}, \ldots, p_{\tilde{i}_{u}}\right\}$. In that case we use the following rule: Choose a set $\mathbf{b}_{1} \subset\left\{p_{\tilde{i}_{1}}, \ldots, p_{\tilde{i}_{u}}\right\}$ which is a basis of $\operatorname{span}\left\{p_{\tilde{i}_{1}}, \ldots, p_{\tilde{i}_{u}}\right\}$. Choose $\mathbf{b}_{2} \subset$ $\left\{p_{i_{1}}, \ldots, p_{i_{t}}\right\}$ such that $\mathbf{b}=\mathbf{b}_{1} \cup \mathbf{b}_{2}$ forms a basis of $\operatorname{span}\left\{p_{i_{1}}, \ldots, p_{i_{t}}, p_{\tilde{i}_{1}}, \ldots, p_{i_{u}}\right\}$. Complete $\mathbf{b}$ to a basis of $\operatorname{span}\left\{p_{1}, \ldots, p_{n}\right\}$ by adding elements of $\left\{p_{1}, \ldots, p_{n}\right\}$ to b. Then write $p_{k}$ as a linear combination of this new basis and apply (if possible) (2.55), 2.56) or (2.57) respectively.

For the following identities which are called integration by parts or IBP identities we will use the fact that in dimensional regularisation an integral over a total derivative with respect to the loop momentum vanishes. Using the definitions of (2.51) we get two further rules:

## Rule 4.

$$
\begin{align*}
0= & \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\partial}{\partial k^{\mu}} \frac{k^{\mu} s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \\
= & (d+s-2 r) \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{u}_{u}}} \\
& -\sum_{a=1}^{t} 2 m_{a} \int \frac{d^{d} k}{(2 \pi)^{d}}\left[\frac{\left(M_{a}^{2}-p_{i_{a}}^{2}\right) s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{a}^{m_{a}+1} \ldots D_{t}^{m_{t}}}-\frac{k \cdot p_{i_{a}} s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{a}^{m_{a}+1} \ldots D_{t}^{m_{t}}}\right] \times \\
& \frac{1}{\tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}}-\sum_{a=1}^{u} \tilde{m}_{a} \int \frac{d^{d} k}{(2 \pi)^{d}} \overline{\tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{a}^{\tilde{m}_{a}+1} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \tag{2.59}
\end{align*}
$$

where $s \equiv \sum_{i=1}^{l} n_{i}$ and $r \equiv \sum_{i=1}^{t} m_{i}$.
Another identity is:

## Rule 5.

$$
\begin{align*}
0= & \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\partial}{\partial k^{\mu}} \frac{p_{a}^{\mu} s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \\
= & \sum_{b=1}^{l} n_{b} p_{a} \cdot p_{j_{b}} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{b}^{n_{b}-1} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \\
& -\sum_{b=1}^{t} 2 m_{b} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}\left(k \cdot p_{a}+p_{i_{b}} \cdot p_{a}\right)}{D_{1}^{m_{1}} \ldots D_{b}^{m_{b}+1} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \\
& -\sum_{b=1}^{u} \tilde{m}_{b} p_{a} \cdot p_{\tilde{i}_{b}} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{b}^{\tilde{m}_{b}+1} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} \tag{2.60}
\end{align*}
$$

where $p_{a} \in \mathbf{B}$.
For the IBP identities 2.59 and 2.60 we have used the translation invariance of the dimensional regularised integral. We get another class of identities if we use the invariance under Lorentz transformations. From equation (2.9) of [24] we get

$$
\begin{align*}
0= & \left(p_{i_{1} \nu} \frac{\partial}{\partial p_{i_{1}}^{\mu}}-p_{i_{1} \mu} \frac{\partial}{\partial p_{i_{1}}^{\nu}}+\ldots+p_{i_{t} \nu} \frac{\partial}{\partial p_{i_{t}}^{\mu}}-p_{i_{t} \mu} \frac{\partial}{\partial p_{i_{t}}^{\nu}}+\right. \\
& p_{\tilde{i}_{1} \nu} \frac{\partial}{\partial p_{\tilde{i}_{1}}^{\mu}}-p_{\tilde{i}_{1} \mu} \frac{\partial}{\partial p_{\tilde{i}_{1}}^{\nu}}+\ldots+p_{\tilde{i}_{u} \nu} \frac{\partial}{\partial p_{\tilde{i}_{u}}^{\mu}}-p_{\tilde{i}_{u} \mu} \frac{\partial}{\partial p_{\tilde{i}_{u}}^{\nu}}+ \\
& \left.p_{j_{1} \nu} \frac{\partial}{\partial p_{j_{1}}^{\mu}}-p_{j_{1} \mu} \frac{\partial}{\partial p_{j_{1}}^{\nu}}+\ldots+p_{j_{l} \nu} \frac{\partial}{\partial p_{j_{l}}^{\mu}}-p_{j_{l} \mu} \frac{\partial}{\partial p_{j_{l}}^{\nu}}\right) \times \\
& \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}} . \tag{2.61}
\end{align*}
$$

We choose $p_{i}, p_{j} \in \mathbf{B}$. By multiplying of (2.61) with $p_{i}^{\mu} p_{j}^{\nu}$ we get

## Rule 6.

$$
\begin{align*}
0= & \int \frac{d^{d} k}{(2 \pi)^{d}}\left[\sum_{a=1}^{l} n_{a}\left(k \cdot p_{i} p_{j_{a}} \cdot p_{j}-k \cdot p_{j} p_{j_{a}} \cdot p_{i}\right) \frac{s_{1}^{n_{1}} \ldots s_{a}^{n_{a}-1} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}}-\right. \\
& \sum_{a=1}^{t} 2 m_{a}\left(k \cdot p_{i} p_{i_{a}} \cdot p_{j}-k \cdot p_{j} p_{i_{a}} \cdot p_{i}\right) \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{a}^{m_{a}+1} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}}- \\
& \left.\sum_{a=1}^{u} \tilde{m}_{a}\left(k \cdot p_{i} p_{\tilde{i}_{a}} \cdot p_{j}-k \cdot p_{j} p_{\tilde{i}_{a}} \cdot p_{i}\right) \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}} \tilde{D}_{1}^{\tilde{m}_{1}} \ldots \tilde{D}_{a}^{\tilde{m}_{a}+1} \ldots \tilde{D}_{u}^{\tilde{m}_{u}}}\right] \tag{2.62}
\end{align*}
$$

An implementation in Mathematica of the IBP identities can be found in appendix A.

### 2.4.2 Calculation of Feynman diagrams with differential equations

In this section I will discuss the extraction of subleading powers of Feynman integrals with the method of differential equations [25, 26, 24]. This method will prove to be easy to implement in a computer algebra system. The idea to obtain the analytic expansion of Feynman integrals by tracing them back to differential equations has first been proposed in [25]. This method, which is demonstrated in [25] by the oneloop two-point integral and in [26] by the two-loop sunrise diagram, uses differential equations with respect to the small or large parameter, in which the integral has to be expanded.

In contrast to [25, 26] I will discuss the case that setting the small parameter to zero gives rise to new divergences. In this case the initial condition is not given by the differential equation itself and also cannot be obtained by calculation of the simpler integral that is defined by setting the expansion parameter to zero. It is not possible to give a general proof, but it seems to be a rule, that one needs the leading power as a "boundary condition", which can be calculated by the method of regions [27, 28, 29, 30]. The subleading powers can be obtained from the differential equation. In the present section I will discuss which conditions the differential equation has to fulfil in order for this to work.

## Description of the method

We start with a (scalar) integral of the form

$$
\begin{equation*}
I\left(p_{1}, \ldots, p_{n}, m_{1}, \ldots, m_{n}\right)=\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{D_{1} \ldots D_{n}} \tag{2.63}
\end{equation*}
$$

where the propagators are of the form $D_{i}=\left(k+p_{i}\right)^{2}-m_{i}^{2}$. We assume that there is only one mass hierarchy, i.e. there are two masses $m \ll M$ such that all of the momenta and masses $p_{i}$ and $m_{i}$ are of $\mathcal{O}(m)$ or of $\mathcal{O}(M)$. We expand (2.63) in $\frac{m}{M}$ by replacing all small momenta and masses by $p_{i} \rightarrow \lambda p_{i}$ and expand in $\lambda$. After the expansion the bookkeeping parameter $\lambda$ can be set to 1 .

We obtain a differential equation for $I$ by differentiating the integrand in (2.63) with respect to $\lambda$. This gives rise to new Feynman integrals with propagators of the form $\frac{1}{D_{i}^{2}}$ and scalar products $k \cdot p_{i}$ in the numerator. Those Feynman integrals, however, can be reduced to the original integral and to simpler integrals (i.e. integrals that contain less propagators in the denominator) by using integration by parts identities.

Finally we obtain for $(2.63)$ a differential equation of the form

$$
\begin{equation*}
\frac{d}{d \lambda} I(\lambda)=h(\lambda) I(\lambda)+g(\lambda) \tag{2.64}
\end{equation*}
$$

where $h(\lambda)$ contains only rational functions of $\lambda$ and $g(\lambda)$ can be expressed by Feynman integrals with a reduced number of propagators. It is easy to see that $h$ and $g$ are unique if and only if $I$ and the integrals contained in $g$ are master integrals
with respect to IBP-identities, i.e. they cannot be reduced to simpler integrals by IBP-identities. If $I(\lambda)$ is divergent in $\epsilon=\frac{4-d}{2}, I, h$ and $g$ have to be expanded in $\epsilon$ :

$$
\begin{align*}
I & =\sum_{i} I_{i} \epsilon^{i} \\
h & =\sum_{i} h_{i} \epsilon^{i} \\
g & =\sum_{i} g_{i} \epsilon^{i} \tag{2.65}
\end{align*}
$$

Plugging (2.65) into 2.64 gives a system of differential equations for $I_{i}$, similar to (2.64). In the next paragraph we will consider an example for this case.

First let us assume that $h(\lambda)$ and $g(\lambda)$ have the following asymptotic behaviour in $\lambda$ :

$$
\begin{align*}
h(\lambda) & =h^{(0)}+\lambda h^{(1)}+\ldots \\
g(\lambda) & =\sum_{j} \lambda^{j} g^{(j)}(\ln \lambda) \tag{2.66}
\end{align*}
$$

i.e. $h$ starts at $\lambda^{0}$, and we allow that $g$ starts at a negative power of $\lambda$. We count $\ln \lambda$ as $\mathcal{O}\left(\lambda^{0}\right)$ so the $g^{(j)}$ may depend on $\ln \lambda$. This dependence, however, has to be such that

$$
\begin{equation*}
\lim _{\lambda \rightarrow 0} \lambda g^{(j)}(\ln \lambda)=0 \tag{2.67}
\end{equation*}
$$

The condition 2.67 is fulfilled, if the $g^{(j)}$ are of the form of a finite sum

$$
\begin{equation*}
\sum_{n=n_{0}}^{m} a_{n} \ln ^{n} \lambda \tag{2.68}
\end{equation*}
$$

The limit $m \rightarrow \infty$ however can spoil the expansion 2.66). E.g. $e^{-\ln \lambda}=\frac{1}{\lambda}$ so the condition 2.67 is not fulfilled, which is due to the fact that we must not change the order of the limits $\lambda \rightarrow 0$ and $m \rightarrow \infty$.

Further we assume that also $I(\lambda)$ starts at $\lambda^{0}$

$$
\begin{equation*}
I(\lambda)=I^{(0)}(\ln \lambda)+\lambda I^{(1)}(\ln \lambda)+\ldots \tag{2.69}
\end{equation*}
$$

and plug this into 2.64 such that we obtain an equation which gives $I^{(i)}$ recursively:

$$
\begin{equation*}
\lambda^{i} I^{(i)}=\int_{0}^{\lambda} d \lambda^{\prime} \lambda^{\prime i-1}\left(\sum_{j=0}^{i-1} h^{(j)} I^{(i-1-j)}\left(\ln \lambda^{\prime}\right)+g^{(i-1)}\left(\ln \lambda^{\prime}\right)\right) \tag{2.70}
\end{equation*}
$$

I want to stress that, because $h$ starts at $\mathcal{O}\left(\lambda^{0}\right), 2.70$ is a recurrence relation, i.e. $I^{(j)}$ does not mix into $I^{(i)}$ if $j \geq i$. As the integral is only well defined if $i \geq 1$, we need the leading power $I^{(0)}$ as "boundary condition" and 2.70 will give us all the higher powers in $\lambda$. It is easy to implement 2.70 in a computer algebra system, because we just need the integration of polynomials and finite powers of logarithms.

A modification is needed if $h$ starts at $\lambda^{-1}$ i.e.

$$
\begin{equation*}
h=-\frac{n}{\lambda} h^{(-1)}+\ldots \tag{2.71}
\end{equation*}
$$

By replacing $\bar{I} \equiv \lambda^{n} I$ we obtain the differential equation

$$
\begin{equation*}
\frac{d}{d \lambda} \bar{I}=\left(\frac{n}{\lambda}+h\right) \bar{I}+\lambda^{n} g \tag{2.72}
\end{equation*}
$$

which is similar to (2.64) and leads to

$$
\begin{equation*}
\lambda^{i+n} I^{(i)}=\int_{0}^{\lambda} d \lambda^{\prime} \lambda^{i+n-1}\left(\sum_{j=0}^{i+n-1} h^{(j)} I^{(i-1-j)}\left(\ln \lambda^{\prime}\right)+g^{(i-1)}\left(\ln \lambda^{\prime}\right)\right) \tag{2.73}
\end{equation*}
$$

which is valid for $i \geq 1-n$. So, if $I$ starts at $\mathcal{O}\left(\lambda^{-n}\right)$, the subleading powers result from the leading power.

## Examples

We start with a pedagogic example:

## Example 2.4.1.

$$
\begin{equation*}
I=\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}\left(k^{2}-\lambda\right)\left(k^{2}-1\right)} \tag{2.74}
\end{equation*}
$$

where $\lambda \ll 1$. The exact expression for this integral is given by:

$$
\begin{equation*}
I=\frac{i}{(4 \pi)^{2}} \frac{\ln \lambda}{1-\lambda}=\frac{i}{(4 \pi)^{2}} \ln \lambda\left(1+\lambda+\lambda^{2}+\ldots\right) . \tag{2.75}
\end{equation*}
$$

We see that $I$ diverges for $\lambda \rightarrow 0$. As described e.g. in 30] we can obtain the leading power by expanding the integrand in the regions $k \sim \sqrt{\lambda}$ and $k \sim 1$. This leads in the first region to

$$
\begin{equation*}
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{-1}{k^{2}\left(k^{2}-\lambda\right)}=-\frac{i}{(4 \pi)^{2-\epsilon}} \Gamma(1+\epsilon)\left(\frac{1}{\epsilon}+1-\ln \lambda\right) \tag{2.76}
\end{equation*}
$$

and in the second region to

$$
\begin{equation*}
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{4}\left(k^{2}-1\right)}=\frac{i}{(4 \pi)^{2-\epsilon}} \Gamma(1+\epsilon)\left(\frac{1}{\epsilon}+1\right) \tag{2.77}
\end{equation*}
$$

such that we finally obtain

$$
\begin{equation*}
I^{(0)}(\ln \lambda)=\frac{i}{(4 \pi)^{2}} \ln \lambda . \tag{2.78}
\end{equation*}
$$

This is the result we obtain from the leading power of (2.75). We write the derivative of $I$ with respect to $\lambda$ in the following form:

$$
\begin{equation*}
\frac{d}{d \lambda} I=\frac{1}{1-\lambda}\left[I-\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}\left(k^{2}-\lambda\right)^{2}}\right] \tag{2.79}
\end{equation*}
$$

We obtained the right hand side of 2.79 by decomposing $\frac{d}{d \lambda} I$ into partial fractions. Of course this decomposition is not unique which is due to the fact that $I$ itself is not a master integral but can be further simplified by partial fractioning. From (2.79) and (2.64) we get:

$$
\begin{align*}
h & =\frac{1}{1-\lambda}=1+\lambda+\lambda^{2}+\ldots \\
g & =\frac{i}{(4 \pi)^{2}} \frac{1}{\lambda(1-\lambda)}=\frac{i}{(4 \pi)^{2}}\left(\lambda^{-1}+1+\lambda+\ldots\right) \tag{2.80}
\end{align*}
$$

such that the coefficients in the expansion in $\lambda$ according to (2.66) do not depend on the power label $(k)$ :

$$
\begin{equation*}
h^{(k)}=1 \quad \text { and } \quad g^{(k)}=\frac{i}{(4 \pi)^{2}} . \tag{2.81}
\end{equation*}
$$

We obtain for the recurrence relation (2.70):

$$
\begin{equation*}
I^{(k)}=\frac{1}{\lambda^{k}} \int_{0}^{\lambda} d \lambda^{\prime} \lambda^{k-1}\left(\sum_{j=0}^{k-1} I^{(k-1-j)}\left(\ln \lambda^{\prime}\right)+\frac{i}{(4 \pi)^{2}}\right) . \tag{2.82}
\end{equation*}
$$

Using the initial value (2.78) it is easy to prove by induction

$$
\begin{equation*}
I^{(k)}(\ln \lambda)=\frac{i}{(4 \pi)^{2}} \ln \lambda \quad \forall k \geq 0 \tag{2.83}
\end{equation*}
$$

This result coincides with (2.75).
The first nontrivial example, we want to consider, is the following three-point integral:

## Example 2.4.2.

$$
\begin{equation*}
I=\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}\left(k+u n_{-}+l\right)^{2}\left(k+n_{+}+n_{-}\right)^{2}} . \tag{2.84}
\end{equation*}
$$

Here $n_{+}$and $n_{-}$are collinear Lorentz vectors, which fulfil $n_{+}^{2}=n_{-}^{2}=0$ and $n_{+} \cdot n_{-}=\frac{1}{2}, u$ is a real number between 0 and 1 and $l$ is a Lorentz vector with $l^{2}=0$ and $l^{\mu} \ll 1$. Furthermore we define

$$
\begin{equation*}
\xi=2 l \cdot n_{+} \quad \text { and } \quad \theta=2 l \cdot n_{-} . \tag{2.85}
\end{equation*}
$$

We expand $I$ in $l$, so we make the replacement $l \rightarrow \lambda l$ and differentiate $I$ with respect to $\lambda$. The integral is not divergent in $\epsilon$ such that we obtain a differential equation of the form (2.64) where the Taylor series of $h(\lambda)$ starts at $\lambda^{0}$ as in (2.66). In $g(\lambda)$ only two-point integrals occur, which are easy to calculate. I do not want to give the explicit expressions for $h$ and $g$ because they are complicated, their exact form is not needed to understand this example and they can be handled by a computer algebra system. Because the leading power of $I$ is of $\mathcal{O}\left(\lambda^{0}\right)$, 2.70) gives all of the subleading powers.

We obtain the leading power as follows: First we have to identify the regions, which contribute at leading power. If we decompose $k$ into

$$
\begin{equation*}
k^{\mu}=2 k \cdot n_{+} n_{-}^{\mu}+2 k \cdot n_{-} n_{+}^{\mu}+k_{\perp}^{\mu} \tag{2.86}
\end{equation*}
$$

we note that the only regions, which remain at leading power, are the hard region $k^{\mu} \sim 1$ and the hard-collinear region

$$
\begin{align*}
k \cdot n_{+} & \sim 1 \\
k \cdot n_{-} & \sim \lambda \\
k_{\perp}^{\mu} & \sim \sqrt{\lambda} . \tag{2.87}
\end{align*}
$$

The soft region $k^{\mu} \sim \lambda$ leads at leading power to a scaleless integral, which vanishes in dimensional regularisation. In the hard region we expand the integrand to

$$
\begin{equation*}
\frac{1}{k^{2}\left(k+u n_{-}\right)^{2}\left(k+n_{+}+n_{-}\right)^{2}} . \tag{2.88}
\end{equation*}
$$

By introducing a convenient Feynman parametrisation we obtain for the ( $4-2 \epsilon$ )dimensional integral over (2.88):

$$
\begin{equation*}
\frac{i}{(4 \pi)^{2-\epsilon}} \Gamma(1+\epsilon) \exp (i \pi \epsilon) \frac{1}{u}\left(\frac{\ln (1-u)}{\epsilon}-\frac{1}{2} \ln ^{2}(1-u)\right) \tag{2.89}
\end{equation*}
$$

In the hard-collinear region we expand the integrand to

$$
\begin{equation*}
\frac{1}{k^{2}\left(k+u n_{-}+\theta n_{+}\right)^{2}\left(2 k \cdot n_{+}+1\right)} \tag{2.90}
\end{equation*}
$$

The integral over (2.90) gives:

$$
\begin{align*}
\frac{i}{(4 \pi)^{2-\epsilon}} \Gamma(1+\epsilon) \exp (i \pi \epsilon) \frac{1}{u}( & \frac{-\ln (1-u)}{\epsilon}+2 \operatorname{Li}_{2}(u)+\frac{1}{2} \ln ^{2}(1-u) \\
& +\ln u \ln (1-u)+\ln (1-u) \ln \theta) \tag{2.91}
\end{align*}
$$

So adding (2.89) and (2.91) together we get the leading power of 2.84 ):

$$
\begin{equation*}
I^{(0)}=\frac{i}{(4 \pi)^{2}} \frac{1}{u}\left(2 \operatorname{Li}_{2}(u)+\ln u \ln (1-u)+\ln (1-u) \ln \theta\right) . \tag{2.92}
\end{equation*}
$$

By plugging (2.92) into (2.70) we obtain $I$ at $\mathcal{O}(\lambda)$ :

$$
\begin{align*}
& I^{(1)}= \\
& \frac{i}{(4 \pi)^{2}} \frac{1}{u}\left[\theta\left(-2+\ln u+\frac{\ln (1-u) \ln \theta}{u}+\frac{\ln (1-u) \ln u}{u}+\ln \xi+\frac{2 \operatorname{Li}_{2}(u)}{u}\right)-\right. \\
& \\
& \quad \xi\left(\frac{\ln u}{1-u}+2 \frac{\ln (1-u)}{u}+\frac{\ln (1-u) \ln \theta}{u}+\frac{\ln (1-u) \ln u}{u}+\right.  \tag{2.93}\\
& \left.\left.\quad \frac{\ln \xi}{1-u}+\frac{2 \operatorname{Li}_{2}(u)}{u}\right)\right] .
\end{align*}
$$

Now we want to consider the following four-point integral

## Example 2.4.3.

$$
\begin{equation*}
I=\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}\left(k+n_{-}\right)^{2}\left(k+l-n_{+}\right)^{2}\left(k+l-u n_{+}\right)^{2}} \tag{2.94}
\end{equation*}
$$

where we used the same variables, which were introduced in 2.84. This example is very special, because in this case our method will allow us to obtain not only the subleading but also the leading power in $l . I$ is divergent in $\epsilon$ such that we obtain after the expansion (2.65) a system of differential equations of the following form:

$$
\begin{align*}
\frac{d}{d \lambda} I_{-1} & =h_{0} I_{-1}+g_{-1} \\
\frac{d}{d \lambda} I_{0} & =h_{0} I_{0}+h_{1} I_{-1}+g_{0} \tag{2.95}
\end{align*}
$$

It turns out that in our example $h$ takes the simple form

$$
\begin{equation*}
h=-\frac{2+2 \epsilon}{\lambda} \tag{2.96}
\end{equation*}
$$

such that analogously to (2.72) we can transform (2.95) into

$$
\begin{align*}
\frac{d}{d \lambda}\left(\lambda^{2} I_{-1}\right) & =\lambda^{2} g_{-1} \\
\frac{d}{d \lambda}\left(\lambda^{2} I_{0}\right) & =-2 \lambda I_{-1}+\lambda^{2} g_{0} \tag{2.97}
\end{align*}
$$

This system of differential equations can easily be integrated to:

$$
\begin{align*}
I_{-1}^{(i)} & =\frac{1}{\lambda^{i+2}} \int_{0}^{\lambda} d \lambda^{\prime} \lambda^{i+1} g_{-1}^{(i-1)} \\
I_{0}^{(i)} & =\frac{1}{\lambda^{i+2}} \int_{0}^{\lambda} d \lambda^{\prime} \lambda^{i+1}\left(-2 I_{-1}^{(i)}+g_{0}^{(i-1)}\right) \tag{2.98}
\end{align*}
$$

where the superscript ( $i$ ) denotes the order in $\lambda$ as in (2.66) and (2.69). Both $I_{-1}$ and $I_{0}$ start at $\mathcal{O}\left(\lambda^{-1}\right)$. Because 2.98 is valid for $i \geq-1$, it gives us the leading power expression, which reads:

$$
\begin{equation*}
I^{(-1)}=\frac{i}{(4 \pi)^{2-\epsilon}} \Gamma(1+\epsilon) \frac{2}{u \xi}\left(\frac{1}{\epsilon}-1-\frac{\ln u}{1-u}-\ln \xi\right) \tag{2.99}
\end{equation*}
$$

where $\xi=2 l \cdot n_{+}$as in the example above. The exact expression for (2.94) can be obtained from (31). Thereby (2.99) can be tested.

## A simplification for the calculation of the leading power

In the last paragraph I want to return to Example 2.4.2. I will show how we can use differential equations to prove that the integral (2.84) depends in leading power only on the soft kinematical variable $\theta=2 l \cdot n_{-}$and not on $\xi=2 l \cdot n_{+}$. We
need derivatives of the integral with respect to $\xi$ and $\theta$, which we have to express through derivatives with respect to $l^{\mu}$. These derivatives can be applied directly to the integrand, whose dependence on $l^{\mu}$ is obvious. We start from the following identities:

$$
\begin{align*}
n_{+}^{\mu} \frac{\partial}{\partial l^{\mu}} I & =\frac{\partial}{\partial \theta} I+\xi \frac{\partial}{\partial l^{2}} I \\
n_{-}^{\mu} \frac{\partial}{\partial l^{\mu}} I & =\frac{\partial}{\partial \xi} I+\theta \frac{\partial}{\partial l^{2}} I  \tag{2.100}\\
l^{\mu} \frac{\partial}{\partial l^{\mu}} I & =\xi \frac{\partial}{\partial \xi} I+\theta \frac{\partial}{\partial \theta} I+2 l^{2} \frac{\partial}{\partial l^{2}} I
\end{align*}
$$

which lead to

$$
\begin{align*}
\xi \frac{\partial}{\partial \xi} I & =\frac{1}{2}\left(-\theta n_{+}^{\mu}+\xi n_{-}^{\mu}+l^{\mu}\right) \frac{\partial}{\partial l^{\mu}} I \\
\theta \frac{\partial}{\partial \theta} I & =\frac{1}{2}\left(\theta n_{+}^{\mu}-\xi n_{-}^{\mu}+l^{\mu}\right) \frac{\partial}{\partial l^{\mu}} I \tag{2.101}
\end{align*}
$$

where we have set $l^{2}=0$ in 2.101). Using (2.101) we can show that in leading power (2.84) depends only on $\theta$ and not on $\xi$. So we can simplify the calculation of the leading power by making the replacement $l^{\mu} \rightarrow \theta n_{+}^{\mu}$. The proof goes as follows: From 2.69 we see that the statement " $I^{(0)}$ does not depend on $\xi$ " is equivalent to

$$
\begin{equation*}
\xi \frac{\partial}{\partial \xi} I(\xi \lambda, \theta \lambda)=\mathcal{O}(\lambda) \tag{2.102}
\end{equation*}
$$

Using the first equation of (2.101) we get

$$
\begin{equation*}
\xi \frac{\partial}{\partial \xi} I(\xi \lambda, \theta \lambda)=\mathcal{O}(\lambda) I(\xi \lambda, \theta \lambda)+\mathcal{O}(\lambda) \tag{2.103}
\end{equation*}
$$

Because we know (e.g. from power counting) that $I(\xi \lambda, \theta \lambda)$ starts at $\lambda^{0}$, 2.102) is proven.

## Chapter 3

## Calculation of the NLO

### 3.1 Notation

### 3.1.1 Dirac structure

In this thesis I used the NDR scheme [32] such that $\gamma_{5}$-matrices are anticommuting. We get for the matrix elements of $\mathcal{O}_{1}$ and $\mathcal{O}_{2}$ Dirac structures of the following type:

$$
\begin{equation*}
\left\langle\mathcal{O}_{1,2}\right\rangle=\bar{q}_{1}(l) \Gamma_{1} q_{1}(\bar{x} p) \bar{q}_{2}(x p) \Gamma_{2} b(p+q-l) \bar{q}_{3}(y q) \Gamma_{3} q_{4}(\bar{y} q) \tag{3.1}
\end{equation*}
$$

where $q_{i}$ are $u$ - or $d$-quarks. To avoid to specify the flavour, which depends on the decay mode, I introduce for (3.1) the following short notation:

$$
\begin{equation*}
\Gamma_{1} \tilde{\otimes} \Gamma_{2} \otimes \Gamma_{3} . \tag{3.2}
\end{equation*}
$$

The equations of motion lead to:

$$
\begin{align*}
\not \Gamma_{1} \tilde{\otimes} \Gamma_{2} \otimes \Gamma_{3} & =0 \\
\Gamma_{1} \not p \tilde{\otimes} \Gamma_{2} \otimes \Gamma_{3} & =0 \\
\Gamma_{1} \tilde{\otimes} \not p \Gamma_{2} \otimes \Gamma_{3} & =0 \\
\Gamma_{1} \tilde{\otimes} \Gamma_{2}\left(\not p+\not q-\nmid+m_{b}\right) \otimes \Gamma_{3} & =0 \\
\Gamma_{1} \tilde{\otimes} \Gamma_{2} \otimes \not q \Gamma_{3} & =0 \\
\Gamma_{1} \tilde{\otimes} \Gamma_{2} \otimes \Gamma_{3} \not q & =0 . \tag{3.3}
\end{align*}
$$

### 3.1.2 Imaginary part of the propagators

Unless otherwise stated propagators in Feynman integrals always contain an term $+i \eta$ where $\eta>0$ and we take the limit $\eta \rightarrow 0$ after the integration. For example an integral of the form

$$
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+p)^{2}}
$$

is just an abbreviation for

$$
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{\left(k^{2}+i \eta\right)\left((k+p)^{2}+i \eta\right)} .
$$

If the propagator does not contain the integration momentum quadratically, the $i \eta$ will always be given explicitly.

### 3.2 Evaluation of the Feynman diagrams

The diagrams that contribute to $T_{1}^{\mathrm{II}}$ at NLO are listed in fig. 3.1-3.7. Fig. 3.1 shows the gluon self energy. This is a subclass of diagrams which factorizes separately. After adding the counter term for the gluon propagator the contribution to $T_{1}^{\mathrm{II}(2)}$ is:

$$
\begin{align*}
T_{\mathrm{gs}}^{\mathrm{II}(2)}= & \alpha_{s}^{2} \frac{C_{F}}{4 N_{c}^{2}} \frac{1}{\xi \bar{x} \bar{y}}\left[C_{N}\left(-\frac{20}{3} \ln \frac{\mu^{2}}{m_{b}^{2}}+\frac{16}{3} \ln \xi+\frac{16}{3} \ln \bar{x}-\frac{80}{9}\right)\right. \\
& \left.+C_{F} C_{G}\left(\frac{5}{3} \ln \frac{\mu^{2}}{m_{b}^{2}}-\frac{5}{3} \ln \xi-\frac{5}{3} \ln \bar{x}+\frac{31}{9}\right)\right] . \tag{3.4}
\end{align*}
$$

For (3.4) we have set the number of active quark flavours to $n_{f}=5$ and set the mass of the $u-, d-, s$ - and $c$-quark to zero.

For the rest of this section I will consider only those diagrams, for which the calculation of Feynman integrals is not straight forward. I will only show how those Feynman integrals can be calculated in leading power. For higher powers I refer to the methods shown in section 2.4.2. It is important to note, that though for the evaluation of the Feynman integrals arguments depending on power counting have been used, all of the integrals occuring in this section have been tested by methods that do not depend on power counting.

The first class of diagrams with non trivial Feynman integrals are the diagrams in fig. 3.3. The two diagrams from fig. 3.3(a) read

$$
\begin{align*}
& \mathrm{aII} 1=-i g_{s}^{4} N_{c} C_{F}\left(C_{F}-\frac{1}{2} C_{G}\right) \\
& \quad \times \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma_{\mu}(\not k-\not p) \gamma_{\tau} \tilde{\otimes} \gamma_{\nu}\left(1-\gamma_{5}\right) \otimes \gamma^{\tau}(\not k+\bar{x} \not p+y \not q-\not p) \gamma^{\nu}\left(1-\gamma_{5}\right)(\bar{y} \not q-\not k) \gamma^{\mu}}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}(k+\bar{x} p+y q-l)^{2}} \tag{3.5}
\end{align*}
$$

$$
\begin{align*}
& \text { aII } 2=-i g_{s}^{4} N_{c} C_{F}^{2} \\
& \quad \times \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma_{\mu}(\not k-\not k) \gamma_{\tau} \tilde{\otimes} \gamma_{\nu}\left(1-\gamma_{5}\right) \otimes \gamma^{\mu}(y \not q-\not k) \gamma^{\nu}\left(1-\gamma_{5}\right)(\not k+\bar{y} \not q+\bar{x} \not p-\not p) \gamma^{\tau}}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-y q)^{2}(k+\bar{x} p+\bar{y} q-l)^{2}} . \tag{3.6}
\end{align*}
$$

The denominators of (3.5) and (3.6) are identical up to the substitution $y \rightarrow \bar{y}$. So the Feynman integrals we have to calculate are the same. We can reduce our fivepoint integrals to four-point integrals by expanding the denominator of the integrand





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Figure 3.1: Gluon self energy




Figure 3.2: Diagrams aI


Figure 3.3: Diagrams aII


Figure 3.4: Diagrams aIII


Figure 3.5: Diagrams aIV


Figure 3.6: Diagrams aV


Figure 3.7: Nonabelian diagrams
into partial fractions

$$
\begin{gather*}
\frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}(k+\bar{x} p+y q-l)^{2}}= \\
\frac{1}{y(\bar{x}-\xi)}\left[\frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}}+\right. \\
\frac{y}{\bar{y}} \frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}}- \\
\frac{1}{k^{2}(k-l)^{2}(k-\bar{y} q)^{2}(k+\bar{x} p+y q-l)^{2}}- \\
\left.\frac{y}{\bar{y}} \frac{1}{(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}(k+\bar{x} p+y q-l)^{2}}\right] \tag{3.7}
\end{gather*}
$$

Only the first two summands of the right hand side of this equation give leading power contributions to aII1 and aII2:

$$
\begin{align*}
t_{I} & \equiv \frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}}  \tag{3.8}\\
t_{I I} & \equiv \frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}} \tag{3.9}
\end{align*}
$$

The third one

$$
\begin{equation*}
t_{I I I} \equiv \frac{1}{k^{2}(k-l)^{2}(k-\bar{y} q)^{2}(k+\bar{x} p+y q-l)^{2}} \tag{3.10}
\end{equation*}
$$

gives only a leading power contribution in the hard-collinear region

$$
\begin{align*}
k \cdot p & \sim 1 \\
k \cdot q & \sim \lambda \\
k_{\perp}^{\mu} & \sim \sqrt{\lambda} \tag{3.11}
\end{align*}
$$

and the soft region

$$
\begin{equation*}
k^{\mu} \sim \lambda, \tag{3.12}
\end{equation*}
$$

where we introduced the counting $l^{\mu} \sim \lambda$ and set $m_{B}=1$. In both regions the leading power of the numerators of aII1 and aII2 vanishes because of equations of motion. The fourth summand of (3.7)

$$
\begin{equation*}
t_{I V} \equiv \frac{1}{(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}(k+\bar{x} p+y q-l)^{2}} \tag{3.13}
\end{equation*}
$$

does not give a leading power contribution at all.
The topologies $t_{I}$ and $t_{I I}$ are exactly the denominators of the Feynman integrals
of the last four diagrams of fig. 3.3.

$$
\begin{align*}
& \mathrm{aII} 3=i g_{s}^{4} N_{c} C_{F}^{2} \frac{1}{\bar{x} y-\bar{x} \xi-y \theta} \int \frac{d^{d} k}{(2 \pi)^{d}} \gamma_{\mu}(\not k-\not \subset) \gamma_{\tau}  \tag{3.14}\\
& \tilde{\otimes} \frac{\gamma_{\nu}\left(1-\gamma_{5}\right) \otimes \gamma^{\tau}(\not k+y \not q+\bar{x} p p-\not p) \gamma^{\mu}(y \not q+\bar{x} \not p-\not l) \gamma^{\nu}\left(1-\gamma_{5}\right)}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}} \\
& \text { aII4 }=i g_{s}^{4} N_{c} C_{F}\left(C_{F}-\frac{1}{2} C_{G}\right) \frac{1}{\bar{x} \bar{y}-\bar{x} \xi-\bar{y} \theta} \int \frac{d^{d} k}{(2 \pi)^{d}} \gamma_{\mu}(\not k-\not k) \gamma_{\tau}  \tag{3.15}\\
& \tilde{\otimes} \frac{\gamma_{\nu}\left(1-\gamma_{5}\right) \otimes \gamma^{\nu}\left(1-\gamma_{5}\right)(\bar{x} \not p+\bar{y} \not q-\not l) \gamma^{\mu}(\not k+\bar{x} \not p+\bar{y} \not q-\not l) \gamma^{\tau}}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+\bar{y} q-l)^{2}} \\
& \text { aII5 }=i g_{s}^{4} N_{c} C_{F}\left(C_{F}-\frac{1}{2} C_{G}\right) \frac{1}{\bar{x} y-\bar{x} \xi-y \theta} \int \frac{d^{d} k}{(2 \pi)^{d}} \gamma_{\mu}(\not k-\nless) \gamma_{\tau}  \tag{3.16}\\
& \tilde{\otimes} \frac{\gamma_{\nu}\left(1-\gamma_{5}\right) \otimes \gamma^{\mu}(y q \underline{q}-\not k) \gamma^{\tau}(\bar{x} \not p+y \not q-\not p) \gamma^{\nu}\left(1-\gamma_{5}\right)}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-y q)^{2}} \\
& \text { aII6 }=i g_{s}^{4} N_{c} C_{F}^{2} \frac{1}{\bar{x} \bar{y}-\bar{x} \xi-\bar{y} \theta} \times  \tag{3.17}\\
& \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma_{\mu}(\not k-\not p) \gamma_{\tau} \tilde{\otimes} \gamma_{\nu}\left(1-\gamma_{5}\right) \otimes \gamma^{\nu}\left(1-\gamma_{5}\right)(\bar{x} \not p+\bar{y} \underline{q}-\not p) \gamma^{\tau}(\bar{y} \not q-\not k) \gamma^{\mu}}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(k-\bar{y} q)^{2}},
\end{align*}
$$

where the Feynman diagrams are given in the same order as they occur in fig. 3.3. So we have reduced all the Feynman integrals of fig. 3.3 to the topologies $t_{I}$ and $t_{I I}$. Regarding $t_{I}$ we need the scalar integral

$$
\begin{equation*}
D_{0 I}(l) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} t_{I} \tag{3.18}
\end{equation*}
$$

By following the procedure of [23] the tensor integrals $\int \frac{d^{d} k}{(2 \pi)^{d}} k^{\mu} t_{I}, \int \frac{d^{d} k}{(2 \pi)^{d}} k^{\mu} k^{\nu} t_{I}$ and $\int \frac{d^{d} k}{(2 \pi)^{d}} k^{\mu} k^{\nu} k^{\tau} t_{I}$ can be reduced to 3.18 and to the two-point master integrals that are listed in appendix B.1. The leading power of (3.18) gets only contributions from the region where $k$ is soft i.e. all components of $k$ are of $\mathcal{O}\left(\Lambda_{\mathrm{QCD}}\right)$. In this region we can expand the integrand of (3.18) to

$$
\begin{equation*}
\frac{1}{\left(k^{2}+i \eta\right)\left((k-l)^{2}+i \eta\right)(2 k \cdot p-\xi+i \eta)(-2 k \cdot q+i \eta) \bar{x} \bar{y}} . \tag{3.19}
\end{equation*}
$$

Now we can obtain the leading power of (3.18) by integrating (3.19) over all momenta $k$ because there is no other region, which gives a leading power contribution, besides where $k$ is soft. The integration of (3.19) can be easily performed by using the following Feynman parametrisation [33]:

$$
\begin{equation*}
\frac{1}{A_{0} \ldots A_{n}}=\int_{0}^{\infty} d^{n} \lambda \frac{n!}{\left(A_{0}+\sum_{i=1}^{n} \lambda_{i} A_{i}\right)^{n+1}} \tag{3.20}
\end{equation*}
$$

Finally we obtain for the leading power of (3.18)

$$
\begin{align*}
& D_{0 I} \doteq \frac{i}{(4 \pi)^{2}} \frac{\Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}}{\bar{x} \bar{y} \xi \theta}\left(\frac{2}{\epsilon^{2}}-\frac{2 \ln \xi+2 \ln \theta+2 i \pi}{\epsilon}\right. \\
&\left.-\pi^{2}+\ln ^{2} \xi+\ln ^{2} \theta+2 \ln \xi \ln \theta+2 \pi i(\ln \xi+\ln \theta)\right) \tag{3.21}
\end{align*}
$$

Regarding the topology $t_{I I}$ we need the tensor integrals $\int \frac{d^{d} k}{(2 \pi)^{d}} k^{\mu} t_{I I}, \int \frac{d^{d} k}{(2 \pi)^{d}} k^{\mu} k^{\nu} t_{I I}$ and $\int \frac{d^{d} k}{(2 \pi)^{d}} k^{\mu} k^{\nu} k^{\tau} t_{I I}$. This topology can be reduced to two-point master integrals and to the four-point master integral

$$
\begin{equation*}
D_{0 I I}(l) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} t_{I I} \tag{3.22}
\end{equation*}
$$

In order to calculate this integral we decompose $t_{I I}$ into

$$
\begin{align*}
t_{I I}= & \frac{1}{y}\left[\frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p-l)^{2}(2 k \cdot q+\bar{x}-\theta)}\right. \\
& \left.-\frac{1}{k^{2}(k-l)^{2}(k+\bar{x} p+y q-l)^{2}(2 k \cdot q+\bar{x}-\theta)}\right] \tag{3.23}
\end{align*}
$$

where only the integral over the first summand of (3.23) gives a leading power contribution. This integration is straight forward if one uses the Feynman parametrisation (3.20).

Finally we get the leading power of 3.22 :

$$
\begin{align*}
& D_{0 I I} \doteq \\
& \qquad-\frac{i}{(4 \pi)^{2}} \frac{\Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}}{\bar{x}^{2} y \xi}\left[\frac{2}{\epsilon^{2}}-\frac{2}{\epsilon}(\ln \bar{x}+\ln \xi)-\frac{\pi^{2}}{3}+\ln ^{2} \bar{x}+\ln ^{2} \xi+2 \ln \bar{x} \ln \xi\right] . \tag{3.24}
\end{align*}
$$

Actually it turns out that the sum of the diagrams in fig. 3.3 vanishes in leading power.

The diagrams of fig. 3.4 are straight forward to calculate. It is easy to see that in leading power $l$ does not occur within a loop integral. So there are only two linearly independent momenta in the Feynman integrals, which, using similar relations like (3.7) and IBP identities, can be reduced to the master integrals that are listed in appendix B. 1 .

The diagrams of fig. 3.5(a),(b) are easy to calculate, because they contain only
three-point integrals. The next two (fig. 3.5(c)) are given by

$$
\begin{align*}
& \operatorname{aIVm} 1=i g_{s}^{4} N_{c} C_{F}\left(C_{F}-\frac{1}{2} C_{G}\right) \int \frac{d^{d} k}{(2 \pi)^{d}} \gamma^{\mu}(\not k+\bar{x} \not p) \gamma^{\tau} \tilde{\otimes}  \tag{3.25}\\
& \quad \frac{\gamma^{\nu}\left(1-\gamma_{5}\right)\left(\not k+\not p+\not q-\not p+m_{b}\right) \gamma_{\tau} \otimes \gamma_{\nu}\left(1-\gamma_{5}\right)(\not k-\not k+\bar{x} \not p+\bar{y} \not q) \gamma_{\mu}}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+\bar{y} q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)} \\
& \operatorname{aIVm2}=-i g_{s}^{4} N_{c} C_{F}\left(C_{F}-\frac{1}{2} C_{G}\right) \int \frac{d^{d} k}{(2 \pi)^{d}} \gamma^{\mu}(\not k+\bar{x} \not p) \gamma^{\tau} \tilde{\otimes}  \tag{3.26}\\
& \quad \frac{\gamma^{\nu}\left(1-\gamma_{5}\right)\left(\not k+\not p+\not q-\not x+m_{b}\right) \gamma_{\tau} \otimes \gamma_{\mu}(\not k-\not k+\bar{x} \not p+y \not q) \gamma_{\nu}\left(1-\gamma_{5}\right)}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)} .
\end{align*}
$$

By expanding the denominator of (3.25) or (3.26) into partial fractions we get:

$$
\begin{gather*}
\frac{1}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)}= \\
\frac{1}{\bar{x}-\bar{x} \xi-\theta}\left[-\frac{1}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}}+\right. \\
\frac{1}{y} \frac{1}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)}- \\
\frac{\bar{y}}{y} \frac{1}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p+y q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)}+ \\
\frac{x}{\bar{x}} \frac{1}{k^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)}- \\
\left.\frac{x}{\bar{x}} \frac{1}{(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)}\right] \tag{3.27}
\end{gather*}
$$

where we get leading power contributions only from:

$$
\begin{align*}
t_{I} & =\frac{1}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}}  \tag{3.28}\\
t_{I I} & =\frac{1}{k^{2}(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)}  \tag{3.29}\\
t_{I I I} & =\frac{1}{(k+\bar{x} p)^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+y q-l)^{2}\left(k^{2}+2 k \cdot(p+q-l)\right)} \tag{3.30}
\end{align*}
$$

The leading power of (3.28) can be taken from (3.21). We obtain the leading power of (3.29) by making the replacement $l \rightarrow \xi q$. Alternatively we can use (B.18) in appendix B. 2 and take the leading power afterwards. For (3.30) we obtain the leading power by making the replacement $l \rightarrow \theta p$. In contrast to (3.29) this replacement is not so obvious and to calculate this integral exactly is very involved. However we can start with the value, we obtained by this prescription, and show afterwards by solving a partial differential equation that it is correct. We define

$$
\begin{equation*}
I(x, y, \lambda \xi, \lambda \theta) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} t_{I I I}(x, y, p, q, \lambda l) \tag{3.31}
\end{equation*}
$$

and derive two differential equations by deriving $I$ with respect to $x$ and to $y$. As a boundary condition for our differential equations we calculate $I$ at the point $x=0$ and $y=1$. This can be done by decomposing $t_{I I I}$ into partial fractions and using IBP identities. We use the fact that the limits $x \rightarrow 0$ and $y \rightarrow 1$ do not lead to extra divergences in $\epsilon$ and in $\lambda$. So we can solve our differential equations order by order in $\epsilon$ and $\lambda$. Defining

$$
\begin{equation*}
I \equiv \sum_{j, k} I_{j}^{(k)} \epsilon^{j} \lambda^{k} \tag{3.32}
\end{equation*}
$$

and using the fact that we only need the leading power in $\lambda$ we obtain differential equations of the form

$$
\begin{align*}
\frac{\partial}{\partial x} I_{j}^{(-1)} & =h_{x 0} I_{j}^{(-1)}+h_{x 1} I_{j-1}^{(-1)}+g_{x j}^{(-1)} \\
\frac{\partial}{\partial y} I_{j}^{(-1)} & =h_{y_{0}} I_{j}^{(-1)}+g_{y j}^{(-1)} . \tag{3.33}
\end{align*}
$$

The coefficients $h_{x 0}, h_{x 1}, g_{x_{j}}{ }^{(-1)}, h_{y_{0}}$ and $g_{y j}^{(-1)}$ are straight forward to calculate using IBP identities and the master integrals are given in appendix B.1. It turns out that the leading power integral we obtained by the prescription $l \rightarrow \theta p$ fulfils our boundary condition for $x=0$ and $y=1$ as well as the set of differential equations (3.33).

The sum of the following two diagrams (fig. 3.5(d)) is

$$
\begin{align*}
& -i g_{s}^{4} N_{c} C_{F}\left(C_{F}-\frac{1}{2} C_{G}\right) \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma^{\tau} \nmid k \cdot \gamma^{\mu} \tilde{\otimes} \gamma_{\tau}(\not k+x \not k-\not p) \gamma^{\nu}\left(1-\gamma_{5}\right)}{k^{2}(k-l)^{2}(k-\bar{x} p)^{2}(k+x p-l)^{2}} \otimes \\
& \left(\gamma_{\mu} \frac{y \not q+\bar{x} \not p-\not k}{(k-\bar{x} p-y q)^{2}} \gamma_{\nu}-\gamma_{\nu} \frac{\bar{y} \not q+\bar{x} \not p-\not k}{(k-\bar{x} p-\bar{y} q)^{2}} \gamma_{\mu}\right)\left(1-\gamma_{5}\right) . \tag{3.34}
\end{align*}
$$

As in the previous case the denominator can be decomposed into partial fractions and the remaining four-point master integrals can be calculated in leading power by the replacement $l \rightarrow \xi q$. Alternatively we can use the explicit formulas for four-point integrals given in [31] and derive the leading power (and higher powers if necessary) afterwards. This is what I have done in order to get an independent test of my master integrals.

The last two diagrams of this class (fig. 3.5(e)) are given by

$$
\begin{align*}
& i g_{s}^{4} N_{c} C_{F}^{2} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\gamma^{\mu} \not k \cdot \gamma^{\tau} \tilde{\otimes} \gamma_{\tau} \nmid k \gamma^{\nu}\left(1-\gamma_{5}\right)}{k^{2}(k+\bar{x} p)^{2}(k+p)^{2}(k+l)^{2}} \otimes \\
& \left(\gamma_{\mu} \frac{\not k+\not x+y \not q}{(k+l+y q)^{2}} \gamma_{\nu}-\gamma_{\nu} \frac{\not k+\not \subset+\bar{y} \not q}{(k+l+\bar{y} q)^{2}} \gamma_{\mu}\right)\left(1-\gamma_{5}\right) . \tag{3.35}
\end{align*}
$$

As in the case before we obtain the leading power of the four-point master integrals by the replacement $l \rightarrow \xi q$ or by using the formulas in 31. The denominator of (3.35) however cannot by decomposed into partial fractions. So we additionally need the five-point master integral given in appendix B.3.

Regarding the diagrams in fig. 3.6 there do not occur any subtleties we have not yet considered in the paragraph above. So we directly switch to the non-abelian diagrams in fig. 3.7. The first four (fig. 3.7(a),(b)) do not lead to any problems because they contain only three-point integrals. The diagrams in fig. 3.7(c) are given by

$$
\begin{align*}
& \overline{\sum^{2}}+\overline{\sum^{2}}=  \tag{3.36}\\
& -i g_{s}^{4} N_{c} C_{F} C_{G} \frac{1}{\bar{x} \xi} \int \frac{d^{d} k}{(2 \pi)^{d}}\left(g^{\mu \lambda}(l-k)^{\tau}+2 g^{\lambda \tau} k^{\mu}-g^{\tau \mu}(k+2 l-2 \bar{x} p)^{\lambda}\right) \times+ \\
& \frac{\gamma_{\mu} \tilde{\otimes} \gamma_{\tau}(\not p-\not p-\not k) \gamma^{\nu}\left(1-\gamma_{5}\right)}{k^{2}(k+l-p)^{2}(k+l-\bar{x} p)^{2}} \otimes\left[\frac{\gamma_{\lambda}(\not k+y \not q) \gamma_{\nu}}{(k+y q)^{2}}-\frac{\gamma_{\nu}(\not k+\bar{y} \not q) \gamma_{\lambda}}{(k+\bar{y} q)^{2}}\right]\left(1-\gamma_{5}\right) .
\end{align*}
$$

The four-point integral we have to solve is nearly the same as (2.94) of example 2.4.3. We can reduce this integral to a solution of a differential equation and get in this way every power in $\Lambda_{\mathrm{QCD}} / m_{b}$.

The last diagrams we will consider are those from fig. 3.7(d). In leading power they read:

$$
\begin{align*}
& \overline{\text { 慮 }}+\overline{\frac{1}{4}}=  \tag{3.37}\\
& i g_{s}^{4} N_{c} C_{F} C_{G} \frac{1}{\bar{x} \xi} \int \frac{d^{d} k}{(2 \pi)^{d}}\left(g^{\mu \lambda}(k-\bar{x} p)^{\tau}-2 g^{\lambda \tau} k^{\mu}+g^{\tau \mu}(k+2 \bar{x} p)^{\lambda}\right) \times \\
& \frac{\gamma_{\mu} \tilde{\otimes} \gamma^{\nu}\left(1-\gamma_{5}\right)(\not k+\not p+\not q+1) \gamma_{\lambda}}{k^{2}(k+\bar{x} p-l)^{2}\left((k+p+q)^{2}-1\right)} \otimes \\
& {\left[\frac{\gamma_{\nu}(\not k+\bar{x} \not p \not p+\bar{y} q q-\not l) \gamma_{\tau}}{(k+\bar{x} p+\bar{y} q-l)^{2}}-\frac{\gamma_{\tau}(\not k+\bar{x} \not p+y \not q-\not p) \gamma_{\nu}}{(k+\bar{x} p+y q-l)^{2}}\right]\left(1-\gamma_{5}\right)}
\end{align*}
$$

The scalar master integral

$$
\begin{equation*}
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+\bar{x} p-l)^{2}(k+\bar{x} p+\bar{y} q-l)^{2}\left((k+p+q)^{2}-1\right)} \tag{3.38}
\end{equation*}
$$

can be calculated in leading power by setting $\theta=0$ i.e. we make the replacement $l^{\mu} \rightarrow \xi q^{\mu}$. This can be seen as follows: Counting soft momenta as $\mathcal{O}(\lambda)$ and hard momenta as $\mathcal{O}(1)$ (remember $m_{B}=1$ ) the regions of space where (3.38) gives a leading power contribution are

$$
\begin{aligned}
& k^{\mu} \sim 1 \\
& k^{\mu} \sim \lambda \\
& k \cdot p \sim \lambda \quad k_{\perp}^{\mu} \sim \sqrt{\lambda} \quad k \cdot q \sim 1
\end{aligned}
$$

In these regions $l^{\mu}$ occurs only in the combination $l \cdot p$. So we can make the replacement $l^{\mu} \rightarrow \xi q^{\mu}$. Those people who do not believe these arguments are invited to use the exact expression (B.17) for the four-point integral with one massive propagator line, which is given in appendix B.2. After taking the leading power it can easily be seen that we get the same result as by just making the replacement $l^{\mu} \rightarrow \xi q^{\mu}$.

The diagrams which contribute to $T_{2}^{\mathrm{II}}$ are those of fig. 3.3 and fig. 3.4 . The other diagrams drop out because their colour trace is zero. As in the case of $T_{1}^{\mathrm{II}}$ the diagrams of fig. 3.3 cancel each other in leading power. The remaining diagrams are easy to calculate because their Feynman integrals are the same as in the case of $T_{1}^{\mathrm{II}}$.

### 3.3 Wave function contributions

### 3.3.1 General remarks

It has already been demonstrated in section 2.3 how in principle we can extract the scattering kernel $T^{\mathrm{II}}$ of 2.1 from the amplitude if we know the wave functions. $T^{\mathrm{II}}$ does not depend on the hadronic physics and on the form of the wave function $\phi_{\pi}$ and $\phi_{B}$ in particular, so we can get $T^{\text {II }}$ by calculating the matrix elements of the effective operators between free quark states carrying the momenta shown in fig. 2.2 on page 6. Because we calculate $T^{\mathrm{II}}$ in NLO we need unlike as in section 2.3 the wave functions up to NLO. Let us write the second term of (2.1) in the following formal way:

$$
\begin{equation*}
\mathcal{A}_{\text {spect. }}=\phi_{\pi} \otimes \phi_{\pi} \otimes \phi_{B} \otimes T^{\mathrm{II}} \tag{3.39}
\end{equation*}
$$

All of the objects arising in (3.39) have their perturbative series in $\alpha_{s}$, so 3.39) becomes

$$
\begin{align*}
\mathcal{A}_{\text {spect. }}^{(1)}= & \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}  \tag{3.40}\\
\mathcal{A}_{\text {spect. }}^{(2)}= & \phi_{\pi}^{(1)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}+\phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(1)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}+ \\
& \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(1)} \otimes T^{\mathrm{II}(1)}+\phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(2)}
\end{align*}
$$

where the superscript $(i)$ denotes the order in $\alpha_{s}{ }^{1}$. In order to get $T^{\mathrm{II}(2)}$ we have to calculate $\mathcal{A}_{\text {spect. }}^{(2)} \phi_{\pi}^{(1)}$ and $\phi_{B}^{(1)}$ for our final states. Then $T^{\mathrm{II}(2)}$ is given by

$$
\begin{align*}
& \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(2)}=  \tag{3.41}\\
& \quad \mathcal{A}_{\text {spect. }}^{(2)}-\phi_{\pi}^{(1)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}-\phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(1)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}- \\
& \quad \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(1)} \otimes T^{\mathrm{II}(1)}
\end{align*}
$$

At this point a subtlety occurs. Let us have a closer look to the factorization formula (2.1.) By calculating the first order in $\alpha_{s}$ of the partonic form factor $F^{B \rightarrow \pi,(1)}$, which

[^2]

Figure 3.8: Example for diagrams which obviously belong to the form factor
is defined by free quark states instead of hadronic external states, we see that it can be written in the form

$$
\begin{equation*}
F^{B \rightarrow \pi,(1)}=\phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T_{\text {formfact. }}^{(1)} . \tag{3.42}
\end{equation*}
$$

But $T_{\text {formfact. }}^{(1)}$ is no part of $T^{\mathrm{II}}$. So we have to modify 3.41 insofar as we have to subtract the right hand side of (3.42) from the right hand side of (3.41):

$$
\begin{align*}
& \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(2)}=  \tag{3.43}\\
& \quad \mathcal{A}_{\text {spect. }}^{(2)}-\phi_{\pi}^{(1)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}-\phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(1)} \otimes \phi_{B}^{(0)} \otimes T^{\mathrm{II}(1)}- \\
& \quad \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(1)} \otimes T^{\mathrm{II}(1)}-\phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T_{\text {formfact. }}^{(1)} \otimes T^{\mathrm{I}(1)}
\end{align*}
$$

In (3.43) we did not include the term $F^{B \rightarrow \pi,(2)} \otimes T^{I(0)}$, because it is obviously identical with the diagrams where the gluons do not interact with the emitted pion (e.g. those of fig. (3.8). Those diagrams where not considered in the last section. So we do not have to consider them here.

The wave functions for free external quark states are given at LO by (2.41). At NLO there exist three possible contractions: The two external quark states can be connected by a gluon propagator or one of the external quarks can be connected to the eikonal Wilson line of the wave function (fig. 3.9). The diagrams of fig. 3.9(a),(b) and (c) give $\mathcal{O}\left(\alpha_{s}\right)$ of the "pion wave function for free quarks", i.e. we have replaced the pion final state $\langle\pi(p)|$ in $(2.10)$ by the free quark state $\left\langle\bar{q}^{\prime}(\bar{x} p) q(x p)\right|$. The Fourier transformed wave function $\phi_{\pi}^{(1)}\left(x^{\prime}\right)$ is defined analogously to 2.41. For the diagrams in fig. 3.9 (a), (b) and (c) respectively we get:

$$
\begin{align*}
& \phi_{\pi \alpha \beta}^{(\mathrm{a}),(1)}\left(x^{\prime}\right)=8 \pi^{2} i \alpha_{s} C_{F} N_{c} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\delta\left(x^{\prime}-x-\frac{k^{+}}{p^{+}}\right)-\delta\left(x^{\prime}-x\right)}{k^{2} k^{+}} \bar{q}_{\beta}(x p)\left[\frac{1}{\not k-\bar{x} \not p} \gamma^{+} q^{\prime}(\bar{x} p)\right]_{\alpha} \\
& \phi_{\pi \alpha \beta}^{(\mathrm{b}),(1)}\left(x^{\prime}\right)=8 \pi^{2} i \alpha_{s} C_{F} N_{c} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\delta\left(x-x^{\prime}-\frac{k^{+}}{p^{+}}\right)-\delta\left(x-x^{\prime}\right)}{k^{2} k^{+}}\left[\bar{q}(x p) \gamma^{+} \frac{1}{\not k-x \not p}\right]_{\beta} q_{\alpha}^{\prime}(\bar{x} p) \\
& \phi_{\pi \alpha \beta}^{(\mathrm{c}),(1)}\left(x^{\prime}\right)=8 \pi^{2} i \alpha_{s} C_{F} N_{c} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\delta\left(x^{\prime}-x+\frac{k^{+}}{p^{+}}\right)}{k^{2}}\left[\bar{q}(x p) \gamma^{\tau} \frac{1}{x \not p-\not k}\right]_{\beta}\left[\frac{1}{\not k+\overline{x p p}} \gamma_{\tau} q^{\prime}(\bar{x} p)\right]_{\alpha} \tag{3.44}
\end{align*}
$$



Figure 3.9: NLO contributions to the meson wave functions. The dashed line stands for the eikonal Wilson line which makes the wave functions gauge invariant.

### 3.3.2 Evanescent operators

At NLO the convolution of the wave functions with the tree level kernel $T^{\mathrm{II},(1)}$ gives rise to new Dirac structures, which, however, can in four dimensions be reduced to the tree level Dirac structures. So we obtain the tree level Dirac structures plus further evanescent structures, which vanish for $d=4$ but give finite contributions if they are multiplied by UV-poles. We define our renormalisation scheme such that we subtract the UV-poles and these finite parts of the evanescent structures.

The tree level kernel (2.43) contains two Dirac structures where the second one is evanescent (after the the projection on the wave functions). We write $T^{\mathrm{II}(1)}$ in the following form (that must not be mixed up with the notation (3.2)):

$$
\begin{equation*}
T^{\mathrm{II}(1)}\left(x, y, l^{-}\right) \equiv \frac{1}{\bar{x} l^{-}} \gamma^{\mu} \tilde{\otimes} \gamma^{\nu}\left(1-\gamma_{5}\right) \otimes\left(\frac{2 \not p g_{\mu \nu}}{\bar{y}}-\frac{\not p \gamma_{\mu} \gamma_{\nu}}{y \bar{y}}\right)\left(1-\gamma_{5}\right) \tag{3.45}
\end{equation*}
$$

where the symbol $\tilde{\otimes}$ stands for the "wrong contraction" of the Dirac indices i.e. the Dirac indices are given by

$$
\begin{equation*}
\left[\Gamma^{1} \tilde{\otimes} \Gamma^{2} \otimes \Gamma^{3}\right]_{\alpha^{\prime} \alpha \beta^{\prime} \beta \gamma^{\prime} \gamma}=\Gamma_{\gamma^{\prime} \alpha}^{1} \Gamma_{\alpha^{\prime} \gamma}^{2} \Gamma_{\beta^{\prime} \beta}^{3} \tag{3.46}
\end{equation*}
$$

as in (2.43). The "right contraction" is defined by the symbol $\otimes$ i.e. writing the Dirac indices explicitly

$$
\begin{equation*}
\left[\Gamma^{1} \otimes \Gamma^{2} \otimes \Gamma^{3}\right]_{\alpha^{\prime} \alpha \beta^{\prime} \beta \gamma^{\prime} \gamma}=\Gamma_{\alpha^{\prime} \alpha}^{1} \Gamma_{\gamma^{\prime} \gamma}^{2} \Gamma_{\beta^{\prime} \beta}^{3} \tag{3.47}
\end{equation*}
$$

In $d=4$ the wrong and the right contraction are related by Fierz transformations. It is convenient and commonly used to define the renormalised wave functions in
terms of the right contraction, i.e. to define $\phi_{\pi}^{\text {ren. }}$ by renormalising the operator $\bar{q}(z) \gamma^{\mu} \gamma_{5} q^{\prime}(0)$ instead of $\bar{q}(z)_{\beta} q^{\prime}(0)_{\alpha}$. This is why we define our renormalisation scheme such that only the UV-finite part of the right contraction operators remains: Using the notation of (3.45) - (3.47) we define the following operators:

$$
\begin{align*}
\mathcal{O}_{0}\left(x, y, l^{-}\right) & \equiv-\frac{1}{2 l^{-} \bar{x}} \gamma^{\mu}\left(1-\gamma_{5}\right) \otimes \gamma_{\mu}\left(1-\gamma_{5}\right) \otimes \frac{2 \not p}{\bar{y}}\left(1-\gamma_{5}\right)  \tag{3.48}\\
\mathcal{O}_{1}\left(x, y, l^{-}\right) & \equiv \frac{1}{l^{-} \bar{x}} \gamma^{\mu} \tilde{\otimes} \gamma_{\mu}\left(1-\gamma_{5}\right) \otimes \frac{2 \not p}{\bar{y}}\left(1-\gamma_{5}\right)  \tag{3.49}\\
\mathcal{O}_{2}\left(x, y, l^{-}\right) & \equiv \frac{1}{l^{-} \bar{x}} \gamma^{\mu} \tilde{\otimes} \gamma^{\nu}\left(1-\gamma_{5}\right) \otimes \frac{-\not p \gamma_{\mu} \gamma_{\nu}}{y \bar{y}}\left(1-\gamma_{5}\right) \tag{3.50}
\end{align*}
$$

The matrix elements of these operators are defined analogously to (2.42):

$$
\begin{equation*}
\left\langle\mathcal{O}_{i}\right\rangle \equiv \int d x^{\prime} d y^{\prime} d l^{--} \phi_{\pi \alpha \alpha^{\prime}}\left(x^{\prime}\right) \phi_{\pi \beta \beta^{\prime}}\left(y^{\prime}\right) \phi_{B \gamma \gamma^{\prime}}\left(l^{\prime-}\right) \mathcal{O}_{i \alpha^{\prime} \alpha \beta^{\prime} \beta \gamma^{\prime} \gamma}\left(x^{\prime}, y^{\prime}, l^{--}\right) \tag{3.51}
\end{equation*}
$$

Note that $\left\langle\mathcal{O}_{1}+\mathcal{O}_{2}\right\rangle$ is just the convolution of the tree level kernel (3.45) with the wave functions. Furthermore by using Fierz identities it is easy to prove that we have in four dimensions

$$
\begin{align*}
\left\langle\mathcal{O}_{0}\right\rangle & =\left\langle\mathcal{O}_{1}\right\rangle  \tag{3.52}\\
\left\langle\mathcal{O}_{2}\right\rangle & =0 . \tag{3.53}
\end{align*}
$$

So we define the following evanescent operators:

$$
\begin{align*}
E_{1} \equiv & \mathcal{O}_{2}  \tag{3.54}\\
E_{2} \equiv & \mathcal{O}_{1}-\mathcal{O}_{0}  \tag{3.55}\\
E_{3} \equiv & \frac{1}{\bar{x} \bar{y} l^{-}}\left(\gamma^{\mu} \gamma^{\nu} \gamma^{\rho} \tilde{\otimes} \gamma_{\rho} \gamma_{\nu} \gamma_{\mu}\left(1-\gamma_{5}\right)+\frac{(2-d)^{2}}{2} \gamma^{\mu}\left(1-\gamma_{5}\right) \otimes \gamma_{\mu}\left(1-\gamma_{5}\right)\right) \\
& \otimes 2 \not p\left(1-\gamma_{5}\right)  \tag{3.56}\\
E_{4} \equiv & \frac{1}{\bar{x} y \bar{y} l^{-}} \gamma^{\mu} \gamma^{\lambda} \gamma^{\tau} \tilde{\otimes} \gamma_{\tau} \gamma_{\lambda} \gamma^{\nu}\left(1-\gamma_{5}\right) \otimes \not p \gamma_{\mu} \gamma_{\nu}\left(1-\gamma_{5}\right) \tag{3.57}
\end{align*}
$$

where we have defined $E_{3}$ and $E_{4}$ for later convenience. Using those operator definitions we define our renormalisation scheme such that we subtract the UV-pole of $\left\langle\mathcal{O}_{0}\right\rangle$ and the finite parts of $\left\langle E_{i}\right\rangle$ i.e. terms of the form $\frac{1}{\epsilon_{\mathrm{UV}}}\langle E\rangle$, where $\langle E\rangle$ is an arbitrary evanescent structure. It is important to note that we do not subtract IR-poles, because they depend not only on the operator but also on the external states the operator is sandwiched in between. They have to vanish in (3.43) such that the hard scattering kernel is finite. Finally we obtain the same result as if we had regularised the IR-divergences by small quark and gluon masses because the evanescent structures vanish in $d=4$.

In the next step we will calculate the convolution integral of $T^{\mathrm{II},(1)}$ with the NLO wave functions given by (3.44), i.e. we have to calculate the renormalised matrix elements of $\mathcal{O}_{1}+\mathcal{O}_{2}$ at NLO.

### 3.3.3 Wave function of the emitted pion

First we consider the renormalisation of the emitted pion wave function: Because the contribution of the wave functions $\phi_{\pi}^{(a)}$ and $\phi_{\pi}^{(b)}$ does not change the Dirac structure of the operators, we do not need to consider evanescent operators when we calculate the diagrams of fig. 3.9(a),(b). So for the emitted pion wave function these diagrams give after renormalisation:

$$
\begin{align*}
\left\langle\mathcal{O}_{1}^{\text {ren. }}+\mathcal{O}_{2}^{\text {ren. }}\right\rangle_{\text {emitted }}^{(1),(\mathrm{a}),(\mathrm{b})}=\frac{2 \alpha_{s}}{4 \pi} C_{F}[ & \left(-\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right) \frac{\ln \bar{y}+2 y}{y}\left\langle\mathcal{O}_{1}\right\rangle^{(0)} \\
& \left.-\left(\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \mu_{\mathrm{IR}}\right)(2+\ln y+\ln \bar{y})\left\langle\mathcal{O}_{2}\right\rangle^{(0)}\right] \tag{3.58}
\end{align*}
$$

where the LO matrix elements $\left\langle\mathcal{O}_{i}\right\rangle^{(0)}$ can be obtained from 2.39. Note that we kept the IR-pole times the evanescent matrix element $\left\langle\mathcal{O}_{2}\right\rangle^{(0)}$. This is needed for consistency because we also kept similar terms in the QCD-calculation of $\mathcal{A}_{\text {spect. }}$. Furthermore it allows us to show that all IR-divergences vanish.

The diagram in fig. 3.9 (c) mixes different Dirac structures. So we have to include evanescent operators in the renormalisation. In the case of the emitted pion wave function the operator $\mathcal{O}_{1}$ does not mix under renormalisation with the evanescent operator $E_{1}(3.54)$. We obtain for the renormalised matrix element:

$$
\begin{equation*}
\left\langle\mathcal{O}_{1}^{\mathrm{ren}}\right\rangle_{\mathrm{emitted}}^{(1),(\mathrm{c})}=-\frac{2 \alpha_{s}}{4 \pi} C_{F} \frac{\bar{y} \ln \bar{y}}{y}\left(-\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right)\left\langle\mathcal{O}_{1}\right\rangle^{(0)} \tag{3.59}
\end{equation*}
$$

The matrix element of $E_{1}$ however has an overlap with $\mathcal{O}_{1}$ :

$$
\begin{align*}
\left\langle E_{1}\right\rangle_{\text {emitted }}^{(1),(\mathrm{c})}=\frac{\alpha_{s}}{4 \pi} C_{F}\left(\frac{1}{\epsilon_{\mathrm{UV}}}-\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right) & {\left[(-2 y \ln y-2 \bar{y} \ln \bar{y})\left\langle E_{1}\right\rangle^{(0)}\right.} \\
& \left.-4 \epsilon\left(\ln y+\frac{\bar{y} \ln \bar{y}}{y}\right)\left\langle\mathcal{O}_{1}\right\rangle^{(0)}\right] \tag{3.60}
\end{align*}
$$

The renormalisation prescription tells us to subtract the UV-pole and the UV-finite part of $\left\langle E_{1}\right\rangle$. So we obtain after renormalisation:

$$
\begin{align*}
\left\langle E_{1}^{\text {ren. }}\right\rangle_{\text {emitted }}^{(1),(\mathrm{c})}=\frac{\alpha_{s}}{4 \pi} C_{F}[ & \left(-\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right)(-2 y \ln y-2 \bar{y} \ln \bar{y})\left\langle E_{1}\right\rangle^{(0)} \\
& \left.+4\left(\ln y+\frac{\bar{y} \ln \bar{y}}{y}\right)\left\langle\mathcal{O}_{1}\right\rangle^{(0)}\right] . \tag{3.61}
\end{align*}
$$

Note that the evanescent operator $E_{1}$ leads to a finite term $4\left(\ln y+\frac{\bar{y} \ln \bar{y}}{y}\right)\left\langle\mathcal{O}_{1}\right\rangle^{(0)}$, which we would have missed if we had just dropped the evanescent operators.

### 3.3.4 Wave function of the recoiled pion

In the next step we consider the NLO contribution of the recoiled pion wave function. As in the case of the emitted pion the diagrams fig. 3.9(a),(b) do not lead to a mixing between the operators. Therefore we get:

$$
\begin{align*}
&\left\langle\mathcal{O}_{1}^{\text {ren. }}+\mathcal{O}_{2}^{\text {ren. }}\right\rangle_{\text {reccoiled }}^{(1),(\mathrm{a}),(\mathrm{b})}=\frac{\alpha_{s}}{4 \pi} C_{F} \frac{2 \ln \bar{x}+4 x}{x} {[ } \\
&\left(-\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right)\left\langle\mathcal{O}_{1}\right\rangle^{(0)}  \tag{3.62}\\
&\left.-\left(\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \mu_{\mathrm{IR}}\right)\left\langle\mathcal{O}_{2}\right\rangle^{(0)}\right] .
\end{align*}
$$

Other than in the case of the emitted pion the operators $\mathcal{O}_{1}$ and $\mathcal{O}_{2}$ mix the spinors of the recoiled pion and the $B$-meson. Therefore we have to work in the operator basis of $\mathcal{O}_{0}$ and the evanescent operators and define our renormalisation scheme such that the finite parts of the matrix elements of the evanescent operators vanish. The diagram fig. 3.9.(c) contributes to the matrix element of the renormalised operator $\mathcal{O}_{0}^{\text {ren. }}$ :

$$
\begin{equation*}
\left\langle\mathcal{O}_{0}^{\text {ren. }}\right\rangle_{\text {recoiled }}^{(1),(\mathrm{c})}=2 \frac{\alpha_{s}}{4 \pi} C_{F}\left(\frac{1}{\epsilon_{\mathrm{IR}}}-2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right) \frac{\bar{x} \ln \bar{x}}{x}\left\langle\mathcal{O}_{0}\right\rangle^{(0)} . \tag{3.63}
\end{equation*}
$$

In the case of the evanescent operators we keep the IR-pole:

$$
\begin{align*}
\left\langle E_{1}^{\text {ren. }}\right\rangle_{\text {recoiled }}^{(1),(\mathrm{c})} & =\frac{1}{2} \frac{\alpha_{s}}{4 \pi} C_{F}\left(\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \mu_{\mathrm{IR}}\right) \frac{\bar{x} \ln \bar{x}}{x}\left\langle E_{4}\right\rangle^{(0)}  \tag{3.64}\\
\left\langle E_{2}^{\text {ren. }}\right\rangle_{\text {recoiled }}^{(1),(\mathrm{c})} & =\frac{1}{4} \frac{\alpha_{s}}{4 \pi} C_{F}\left(\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \mu_{\mathrm{IR}}\right) \frac{\bar{x} \ln \bar{x}}{x}\left\langle E_{3}\right\rangle^{(0)} \tag{3.65}
\end{align*}
$$

At the end of the day we obtain a contribution from diagram fig. 3.9(c):

$$
\begin{align*}
& \left\langle\mathcal{O}_{0}^{\text {ren. }}+E_{1}^{\text {ren. }}+E_{2}^{\text {ren. }}\right\rangle_{\text {recoiled }}^{(1),(\mathrm{c})}= \\
& \left.\frac{1}{\frac{1}{2} \frac{\alpha_{s}}{4 \pi} C_{F} \frac{\bar{x} \ln \bar{x}}{x}}\left[\begin{array}{l}
\left(\frac{1}{\epsilon_{\mathrm{IR}}}-2 \ln \frac{\mu_{\mathrm{UV}}}{\mu_{\mathrm{IR}}}\right)\left\langle\frac { 1 } { \overline { x } l ^ { - } } \gamma ^ { \mu } \gamma ^ { \lambda } \gamma ^ { \tau } \tilde { \otimes } \gamma _ { \tau } \gamma _ { \lambda } \gamma ^ { \nu } \otimes \left(\frac{2 \not p g_{\mu \nu}}{\bar{y}}-\not \not \not \gamma_{\mu} \gamma_{\nu}\right.\right. \\
y \bar{y}
\end{array}\right)\right\rangle^{(0)} \\
& \left.\quad+8\left\langle\mathcal{O}_{1}\right\rangle^{(0)}\right] . \tag{3.66}
\end{align*}
$$

The very complicated but also very explicit form, in which the above equation was given, is rather convenient, because on the QCD side the diagrams in fig. 3.5(e) on page 28 come with the same Dirac structure and cancel the IR-pole of 3.66 .

### 3.3.5 Wave function of the $B$-meson

The $\alpha_{s}$ corrections of the wave function of the $B$-meson are given by the second row of fig. 3.9. For the diagrams (d), (e) and (f) respectively they read:

$$
\begin{aligned}
& \phi_{B \alpha \beta}^{(\mathrm{d})(1)}\left(l^{\prime-}\right)= \\
& 8 \pi^{2} i \alpha_{s} N_{c} C_{F} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\delta\left(l^{\prime-}-l^{-}-k^{-}\right)-\delta\left(l^{--}-l^{-}\right)}{k^{2} k^{-}}\left[\bar{q}(l) \gamma^{-} \frac{1}{\not k+\not l}\right]_{\beta} b_{\alpha}(p+q-l) \\
& \phi_{B \alpha \beta}^{(\mathrm{e})(1)}\left(l^{\prime-}\right)= \\
& 8 \pi^{2} i \alpha_{s} N_{c} C_{F} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\delta\left(l^{\prime-}-l^{-}-k^{-}\right)-\delta\left(l^{\prime-}-l^{-}\right)}{k^{2} k^{-}} \times \\
& \bar{q}_{\beta}(l)\left[\frac{1}{\not k-\not p-\not q+\not q+m_{b}} \gamma^{-} b(p+q-l)\right]_{\alpha} \\
& \phi_{B \alpha \beta}^{(\mathrm{f}),(1)}\left(l^{\prime-}\right)= \\
& 8 \pi^{2} i \alpha_{s} N_{c} C_{F} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{\delta\left(l^{\prime-}-l^{-}-k^{-}\right)}{k^{2}} \times \\
& {\left[\bar{q}(l) \gamma^{\mu} \frac{1}{\not k+\nmid}\right]_{\beta}\left[\frac{1}{\not p+\not q-\not q-\not k-m_{b}} \gamma_{\mu} b(p+q-l)\right]_{\alpha}}
\end{aligned}
$$

In the case of the $B$-meson only the diagrams in fig. 3.9(d),(e) give rise to UVpoles. Those diagrams however do not lead to a mixing of $\mathcal{O}_{0}$ and the evanescent operators and we do not have to deal with evanescent operators.

First let's have a look at the convolution integral which belongs to the diagram in fig. 3.9(f) ${ }^{2}$ :

$$
\begin{align*}
& \left\langle\mathcal{O}_{1}+\mathcal{O}_{2}\right\rangle_{B}^{(1),(\mathrm{f})}=  \tag{3.68}\\
& \quad(4 \pi)^{2} i \alpha_{s}^{2} N_{c} C_{F}^{2} \frac{1}{\bar{x}} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{2(k+l) \cdot p k^{2}} \times \\
& \quad \gamma^{\tau} \frac{\not k+\not l}{(k+l)^{2}} \gamma^{\mu} \tilde{\otimes} \gamma^{\nu}\left(1-\gamma_{5}\right) \frac{\not p+\not q-\not l-\not k+m_{b}}{k^{2}-2 k \cdot(p+q-l)} \gamma_{\tau} \otimes\left(\frac{2 \not p}{\bar{y}} g_{\mu \nu}-\frac{\not p}{y \bar{y}} \gamma_{\mu} \gamma_{\nu}\right)\left(1-\gamma_{5}\right) .
\end{align*}
$$

In leading power $(3.68)$ is identical to the contribution of the two diagrams shown in fig. 3.10, which is given by:

$$
\begin{align*}
- & (4 \pi)^{2} i \alpha_{s}^{2} N_{c} C_{F}^{2} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+l-\bar{x} p)^{2}} \\
& \times \gamma^{\tau} \frac{\not k+\not l}{(k+l)^{2}} \gamma^{\mu} \tilde{\otimes} \gamma^{\nu}\left(1-\gamma_{5}\right) \frac{\not p+\not q-\not q-\not k+m_{b}}{k^{2}-2 k \cdot(p+q-l)} \gamma_{\tau} \\
& \otimes\left(\gamma_{\mu} \frac{y q d+\bar{x} \not p-\not k-\nless}{(y q+\bar{x} p-k-l)^{2}} \gamma_{\nu}-\gamma_{\nu} \frac{\bar{y} q q+\bar{x} \not p-\not ้-\nless}{(\bar{y} q+\bar{x} p-k-l)^{2}} \gamma_{\mu}\right)\left(1-\gamma_{5}\right) . \tag{3.69}
\end{align*}
$$

[^3]

Figure 3.10: Two diagrams which correspond in leading power to the contribution of the $B$-meson wavefunction (3.68).


Figure 3.11: $\alpha_{s}$ contributions to the form factor

In (3.69) the leading power comes from the region where $k$ is soft. In this region of space the integrand gets the form of the integrand in (3.68), so both contributions cancel. As we did not include the diagrams of fig. 3.10 in the last section we can skip the contribution of (3.68) here.

The remaining contributions are the diagrams in fig. 3.9(d) and (e). Together they read:

$$
\begin{align*}
& \left\langle\mathcal{O}_{1}+\mathcal{O}_{2}\right\rangle_{B}^{(1),(\mathrm{d}),(\mathrm{e})}=  \tag{3.70}\\
& \quad \alpha_{s}^{2} N_{c} C_{F}^{2} \frac{1}{\xi \bar{x}} \gamma^{\mu} \tilde{\otimes} \gamma^{\nu}\left(1-\gamma_{5}\right) \otimes\left(\frac{2 \not p}{\bar{y}} g_{\mu \nu}-\frac{\not p}{y \bar{y}} \gamma_{\mu} \gamma_{\nu}\right) \times \\
& \quad\left(\left(\frac{1}{\epsilon_{\mathrm{UV}}}+2 \ln \frac{\mu_{\mathrm{UV}}}{m_{b}}\right)(4+2 \ln \xi)-2\left(\frac{1}{\epsilon_{\mathrm{IR}}}+2 \ln \frac{\mu_{\mathrm{IR}}}{m_{b}}\right)+4-\frac{2 \pi^{2}}{3}-2 \ln ^{2} \xi\right) .
\end{align*}
$$

### 3.3.6 Form factor contribution

Finally we have to calculate the contribution of (3.42). It is given by

$$
\begin{align*}
\mathcal{A}_{\text {formfact. }} \equiv & \phi_{\pi}^{(0)} \otimes \phi_{\pi}^{(0)} \otimes \phi_{B}^{(0)} \otimes T_{\text {formfact. }}^{(1)} \otimes T^{\mathrm{I}(1)} \equiv  \tag{3.71}\\
& \frac{C_{F} \alpha_{s}}{4 \pi} f^{(1), \nu} \bar{q}_{e}(y q) \gamma_{\nu}\left(1-\gamma_{5}\right) q_{e}^{\prime}(\bar{y} q) T^{(1)}(y) .
\end{align*}
$$

The form factor $f^{(1), \nu}$ is the $\alpha_{s}$ correction of the matrix element

$$
\begin{equation*}
\left\langle\bar{q}^{\prime}(\bar{x} p) q(x p)\right| \bar{q} \gamma^{\nu}\left(1-\gamma_{5}\right) b\left|b(p+q-l) \bar{q}^{\prime}(l)\right\rangle \tag{3.72}
\end{equation*}
$$

where $\bar{q}_{e}^{\prime}$ and $q_{e}$ are the spinors with the flavour quantum numbers of the emitted pion and $T^{(1)}(y)$ is given by [10]:

$$
\begin{align*}
T^{(1)}(y)= & -6\left(\frac{1}{\epsilon}+\ln \frac{\mu^{2}}{m_{b}^{2}}\right)-18+3\left(\frac{1-2 y}{\bar{y}} \ln y-i \pi\right)+  \tag{3.73}\\
& {\left[2 \operatorname{Li}_{2}(y)-\ln ^{2} y+\frac{2 \ln y}{\bar{y}}-(3+2 i \pi) \ln y-(y \leftrightarrow \bar{y})\right] . }
\end{align*}
$$

We get $f^{(1), \nu}$ by evaluating the diagrams in fig. 3.11. Using the notation of (3.2) we finally obtain:

$$
\begin{align*}
\mathcal{A}_{\text {formfact. }}= & \alpha_{s}^{2} N_{c} C_{F}^{2} \frac{1}{\bar{x} \xi} \gamma^{\mu} \tilde{\otimes}\left(\gamma_{\mu} \frac{\not}{\xi} \gamma^{\nu}\left(1-\gamma_{5}\right)-\gamma^{\nu}\left(1-\gamma_{5}\right) \frac{x \not p+\not p+1}{\bar{x}} \gamma_{\mu}\right) \otimes \\
& \gamma_{\nu}\left(1-\gamma_{5}\right) T^{(1)}(y) . \tag{3.74}
\end{align*}
$$

## Chapter 4

## NLO results

### 4.1 Analytical results for $T_{1}^{\mathrm{II}}$ and $T_{2}^{\mathrm{II}}$

After the analysis of the last chapter we finally obtain the $\mathcal{O}\left(\alpha_{s}^{2}\right)$ results for the hard spectator scattering kernels $T_{1,2}^{\mathrm{II}}$ which are defined by 2.45 ). Those expressions appear in convolution integrals with wave functions, where $x, y$ and $\xi$ are the integration variables as defined in (2.45). The ultraviolet divergences are renormalised in the $\overline{\mathrm{MS}}$-scheme. The infrared divergences drop out after subtracting the wave function contributions from the amplitude. The infrared finiteness together with the finiteness of the convolution integrals ensures that the framework of QCDfactorization works at this order in $\alpha_{s}$.

The explicit $\mathcal{O}\left(\alpha_{s}^{2}\right)$ contributions for $T_{1,2}^{\mathrm{II}}$ read (see next page):

$$
\begin{align*}
& \operatorname{Re}_{1}^{\mathrm{II}(2)}=-\frac{\alpha_{s}^{2} C_{F}}{4 N_{c}^{2} m_{B}^{2} \xi} \times  \tag{4.1}\\
& {\left[C_{N}\left(-\frac{16 \ln \xi}{3 \bar{x} \bar{y}}-\frac{16 \ln \bar{x}}{3 \bar{x} \bar{y}}+\frac{40 \ln \frac{\mu}{m_{b}}}{3 \bar{x} \bar{y}}+\frac{80}{9 \bar{x} \bar{y}}\right)\right.} \\
& +C_{F}\left(\left(\frac{4 \ln \xi}{\bar{x} \bar{y}}+\frac{4 \ln \bar{x}}{\bar{x} \bar{y}}+\frac{4 \ln \bar{y}}{\bar{x} \bar{y}}+\frac{30}{\bar{x} \bar{y}}\right) \ln \frac{\mu}{m_{b}}\right. \\
& -\frac{\ln ^{2} \xi}{\bar{x} \bar{y}}+\ln \xi\left(-\frac{2 \ln x}{\bar{x}^{2} \bar{y}}-\frac{2 \ln \bar{x}}{\bar{x} \bar{y}}-\frac{5}{\bar{x} \bar{y}}\right) \\
& +\left(-\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{4 \bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}-\frac{2 x}{(y-x) \bar{x}}-\frac{2}{y \bar{x}^{2}}+\frac{2(5 x-2)}{\bar{y} \bar{x}^{2}}\right) \mathrm{Li}_{2} x \\
& +\left(-\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{4 \bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}+\frac{2 x}{(y-x) \bar{x}}-\frac{4}{\bar{x}}+\frac{2}{y \bar{x}^{2}}+\frac{4}{\bar{y} \bar{x}^{2}}\right) \mathrm{Li}_{2} y \\
& +\left(\frac{2(x-2)}{\bar{x}^{2} \bar{y}}+\frac{2}{\bar{x}}\right) \operatorname{Li}_{2}(x y) \\
& +\left(-\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{4 \bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}\right) \operatorname{Li}_{2}\left(-\frac{x y}{\bar{x}}\right) \\
& +\left(\frac{2 x}{(y-x) \bar{x}}+\frac{2}{\bar{x} \bar{y}}\right) \mathrm{Li}_{2}\left(-\frac{y \bar{x}}{\bar{y}}\right)+\left(-\frac{2}{\bar{x}}+\frac{2}{\bar{x}^{2} \bar{y}}+\frac{2}{\bar{x}^{2} y}\right) \mathrm{Li}_{2}(x \bar{y}) \\
& +\left(-\frac{2 x}{(y-x) \bar{x}}-\frac{2}{\bar{x} \bar{y}}\right) \mathrm{Li}_{2}\left(-\frac{x \bar{y}}{\bar{x}}\right) \\
& +\left(\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}+\frac{4 \bar{x}}{(y-\bar{x})^{2}}+\frac{2}{y-\bar{x}}\right) \operatorname{Li}_{2}\left(-\frac{\bar{x} \bar{y}}{y}\right) \\
& +\left(-\frac{2}{\bar{y} \bar{x}}-\frac{2}{\bar{x}}\right) \ln x \ln y+\frac{2(3 x-2) \ln x \ln \bar{x}}{\bar{x}^{2} \bar{y}}+\left(\frac{2}{\bar{x}}+\frac{2}{\bar{y} \bar{x}^{2}}\right) \ln x \ln \bar{y} \\
& +\left(-\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{4 \bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}-\frac{2}{\bar{x}}+\frac{2}{y \bar{x}^{2}}+\frac{2 x}{\bar{y} \bar{x}^{2}}\right) \ln y \ln \bar{y} \\
& -\frac{2 \ln \bar{x} \ln \bar{y}}{\bar{x} \bar{y}}+\frac{\ln ^{2} x}{\bar{x} \bar{y}}+\frac{\ln ^{2} y}{\bar{x} \bar{y}}-\frac{\ln ^{2} \bar{x}}{\bar{x} \bar{y}}-\frac{2 \ln ^{2} \bar{y}}{\bar{x} \bar{y}} \\
& +\left(-\frac{4-3 x}{\bar{x}^{2} \bar{y}}-\frac{3}{\bar{x}}\right) \ln x+\left(\frac{2(3 x-2)}{\bar{x}^{2} \bar{y}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{3}{y-\bar{x}}+\frac{3}{\bar{x}}\right) \ln y \\
& +\left(\frac{-9 x-1}{x \bar{x} \bar{y}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{3}{y-\bar{x}}-\frac{1-3 x}{x^{2} y \bar{x}}\right) \ln \bar{x} \\
& +\left(-\frac{1}{\bar{x} \bar{y}}-\frac{4}{\bar{x}^{2} y}\right) \ln \bar{y}+\left(\frac{4}{x \bar{x} \bar{y}}+\frac{4}{x \bar{x}^{2} y}\right) \ln (1-x y) \\
& +\left(-\frac{3 x-1}{x^{2} y \bar{x}}+\frac{-3 x^{2}-2 x-1}{x^{2}(y-\bar{x})}-\frac{3}{\bar{x}}-\frac{2}{\bar{x}^{2} \bar{y}}+\frac{2\left(x^{2}-1\right)}{x(y-\bar{x})^{2}}\right) \ln (1-x \bar{y}) \\
& \left.+\frac{\pi^{2} \bar{x}^{2}}{3(y-\bar{x})^{3}}+\frac{2 \pi^{2} \bar{x}}{3(y-\bar{x})^{2}}+\frac{\pi^{2}}{3(y-\bar{x})}+\frac{\pi^{2}}{3 \bar{x}}-\frac{2\left(2 \pi^{2} x+63 x-63\right)}{3 \bar{y} \bar{x}^{2}}\right)
\end{align*}
$$

$$
\begin{aligned}
-\frac{1}{2} C_{G}( & \frac{80 \ln \frac{\mu}{m_{b}}}{3 \bar{x} \bar{y}}+\left(-\frac{2 \ln x}{\bar{x}^{2} \bar{y}}-\frac{22}{3 \bar{x} \bar{y}}\right) \ln \xi \\
& +\left(-\frac{2 x}{(y-x) \bar{x}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{4}{y-\bar{x}}+\frac{2(5 x-2)}{\bar{x}^{2} \bar{y}}-\frac{2}{y \bar{x}^{2}}\right) \mathrm{Li}_{2} x \\
& +\left(-\frac{2(x-3)}{\bar{x}^{2} \bar{y}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{4}{y-\bar{x}}+\frac{2 x}{(y-x) \bar{x}}-\frac{4}{\bar{x}}+\frac{2}{y \bar{x}^{2}}\right) \mathrm{Li}_{2} y \\
& +\left(\frac{2(x-2)}{\bar{x}^{2} \bar{y}}+\frac{2}{\bar{x}}\right) \mathrm{Li}_{2}(x y)+\left(\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{4}{y-\bar{x}}+\frac{2}{\bar{x}}\right) \mathrm{Li}_{2}\left(-\frac{x y}{\bar{x}}\right) \\
& +\left(\frac{2 x}{(y-x) \bar{x}}+\frac{2}{\bar{x}}\right) \mathrm{Li}_{2}\left(-\frac{y \bar{x}}{\bar{y}}\right)+\left(-\frac{2}{\bar{x}}+\frac{2}{\bar{x}^{2} \bar{y}}+\frac{2}{\bar{x}^{2} y}\right) \mathrm{Li}_{2}(x \bar{y}) \\
& +\left(-\frac{2 x}{(y-x) \bar{x}}-\frac{2}{\bar{x}}\right) \operatorname{Li}_{2}\left(-\frac{x \bar{y}}{\bar{x}}\right) \\
& +\left(-\frac{2 \bar{x}}{(y-\bar{x})^{2}}-\frac{4}{y-\bar{x}}-\frac{2}{\bar{x}}\right) \operatorname{Li}_{2}\left(-\frac{\bar{x} \bar{y}}{y}\right) \\
& -\frac{2 \ln x \ln y}{\bar{x} \bar{y}}+\frac{2(3 x-2) \ln x \ln \bar{x}}{\bar{x}^{2} \bar{y}}+\left(\frac{2}{\bar{x} \bar{y}}-\frac{2}{\bar{x}}\right) \ln y \ln \bar{x} \\
& +\frac{2 \ln x \ln \bar{y}}{\bar{x}^{2} \bar{y}}+\left(\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{4}{y-\bar{x}}-\frac{2}{\bar{x}}+\frac{2}{y \bar{x}^{2}}+\frac{2}{\bar{y} \bar{x}^{2}}\right) \ln y \ln \bar{y} \\
& +\frac{2 \ln \bar{x} \ln \bar{y}}{\bar{x}}+\frac{\ln x}{\bar{x} \bar{y}}+\frac{\ln { }^{2} y}{\bar{x}}-\frac{\ln 2 \bar{x}}{\bar{x} \bar{y}}-\frac{\ln { }^{2} \bar{y}}{\bar{x}} \\
& -\frac{(4-3 x) \ln x}{\bar{x}^{2} \bar{y}}+\left(-\frac{3-5 x}{\bar{x}^{2} \bar{y}}-\frac{2}{y-\bar{x}}\right) \ln y \\
& +\left(\frac{2}{x \bar{x} \bar{y}}+\frac{2}{x \bar{x}^{2} y}\right) \ln (1-x y)+\left(\frac{2}{x y \bar{x}}-\frac{31}{3 \bar{x} \bar{y}}-\frac{2}{y-\bar{x}}\right) \ln \bar{x} \\
& -\frac{2 \ln \bar{y}}{y \bar{x}^{2}}+\left(\frac{2(x+1)}{x(y-\bar{x})}-\frac{2}{x y \bar{x}}-\frac{2}{\bar{x}^{2} \bar{y}}\right) \ln (1-x \bar{y}) \\
& \left.\left.-\frac{2\left(3 \pi^{2} x+166 x+3 \pi^{2}-166\right)}{9 \bar{x}^{2} \bar{y}}-\frac{\pi^{2} \bar{x}}{3(y-\bar{x})^{2}}-\frac{2 \pi^{2}}{3(y-\bar{x})}+\frac{\pi^{2}}{3 \bar{x}}\right)\right]
\end{aligned}
$$

$$
\begin{align*}
\operatorname{Im} T_{1}^{\mathrm{II}(2)}= & -\frac{2 \pi \alpha_{s}^{2} C_{F}}{4 N_{c}^{2} m_{B}^{2} \xi} \times  \tag{4.2}\\
& {\left[C _ { F } \left(-\frac{x \ln x}{\bar{x}^{2} \bar{y}}+\left(\frac{\bar{x}^{2}}{(y-\bar{x})^{3}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{1}{y-\bar{x}}+\frac{1}{\bar{y} \bar{x}}\right) \ln y\right.\right.} \\
& +\left(-\frac{\bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{2 \bar{x}}{(y-\bar{x})^{2}}-\frac{1}{y-\bar{x}}-\frac{x}{(y-x) \bar{x}}-\frac{1}{\bar{y} \bar{x}}\right) \ln \bar{x} \\
& \left.\quad+\left(\frac{x}{(y-x) \bar{x}}+\frac{1}{\bar{x} \bar{y}}\right) \ln \bar{y}-\frac{\bar{x}}{(y-\bar{x})^{2}}-\frac{3}{2(y-\bar{x})}+\frac{2}{\bar{y} \bar{x}}\right)
\end{align*}
$$

$$
\begin{aligned}
-\frac{1}{2} C_{G}( & -\frac{x \ln x}{\bar{x}^{2} \bar{y}}+\left(-\frac{\bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}\right) \ln y \\
& +\left(-\frac{x}{(y-x) \bar{x}}+\frac{\bar{x}}{(y-\bar{x})^{2}}+\frac{2}{y-\bar{x}}-\frac{1}{\bar{x} \bar{y}}\right) \ln \bar{x} \\
& \left.\left.+\frac{x \ln \bar{y}}{(y-x) \bar{x}}+\frac{3}{2 \bar{x} \bar{y}}+\frac{1}{y-\bar{x}}\right)\right]
\end{aligned}
$$

$$
\begin{align*}
\operatorname{Re} T_{2}^{\mathrm{II}}= & -\frac{\alpha_{s}^{2} C_{F} C_{N}}{4 N_{c}^{2} m_{B}^{2} \xi} \times  \tag{4.3}\\
& {\left[\frac{12 \ln \frac{\mu}{m_{b}}}{\bar{x} \bar{y}}+\left(-\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{4 \bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}-\frac{2 x}{(y-x) \bar{x}}-\frac{2}{\bar{y} \bar{x}}\right) \mathrm{Li}_{2} x\right.} \\
& +\frac{2 x \operatorname{Li}_{2} y}{(y-x) \bar{x}}+\left(-\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{4 \bar{x}}{(y-\bar{x})^{2}}-\frac{2}{y-\bar{x}}\right) \operatorname{Li}_{2}\left(-\frac{x y}{\bar{x}}\right) \\
& +\left(\frac{2 x}{(y-x) \bar{x}}+\frac{2}{\bar{x} \bar{y}}\right) \operatorname{Li}_{2}\left(-\frac{y \bar{x}}{\bar{y}}\right) \\
& +\left(\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}+\frac{4 \bar{x}}{(y-\bar{x})^{2}}+\frac{2}{y-\bar{x}}+\frac{2}{\bar{y} \bar{x}}\right) \operatorname{Li}_{2} \bar{y} \\
& +\left(-\frac{2 x}{(y-x) \bar{x}}-\frac{2}{\bar{x} \bar{y}}\right) \operatorname{Li}_{2}\left(-\frac{x \bar{y}}{\bar{x}}\right) \\
& +\left(\frac{2 \bar{x}^{2}}{(y-\bar{x})^{3}}+\frac{4 \bar{x}}{(y-\bar{x})^{2}}+\frac{2}{y-\bar{x}}\right) \operatorname{Li}_{2}\left(-\frac{\bar{x} \bar{y}}{y}\right) \\
& -\frac{2 \ln x \ln y}{\bar{x} \bar{y}}+\frac{2 \ln x \ln \bar{y}}{\bar{x} \bar{y}}+\frac{\ln y}{\bar{x} \bar{y}}-\frac{\ln 2 \bar{y}}{\bar{x} \bar{y}}-\frac{(2-3 x) \ln x}{\bar{x}^{2} \bar{y}} \\
& +\left(\frac{x-2}{\bar{x}^{2} \bar{y}^{2}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{3}{y-\bar{x}}+\frac{x}{\bar{x}^{2} \bar{y}}\right) \ln y \\
& +\left(\frac{2}{x \bar{x} \bar{y}}+\frac{2}{x \bar{x}^{2} y}\right) \ln (1-x y)+\left(\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{3}{y-\bar{x}}\right) \ln \bar{x} \\
& +\left(\frac{-3 x^{2}-2 x-1}{x^{2}(y-\bar{x})}-\frac{1}{x^{2} \bar{x}^{2} \bar{y}}+\frac{2\left(x^{2}-1\right)}{x(y-\bar{x})^{2}}+\frac{1}{x \bar{x}^{2} \bar{y}^{2}}\right) \ln (1-x \bar{y}) \\
& \left.+\left(-\frac{3}{\bar{x} \bar{y}}-\frac{2}{\bar{x}^{2} y}\right) \ln \bar{y}+\frac{16}{\bar{x} \bar{y}}\right]
\end{align*}
$$

$$
\begin{align*}
\operatorname{Im} T_{2}^{\mathrm{II}}= & -\frac{2 \pi \alpha_{s}^{2} C_{F} C_{N}}{4 N_{c}^{2} m_{B}^{2} \xi} \times  \tag{4.4}\\
& {\left[\left(\frac{\bar{x}^{2}}{(y-\bar{x})^{3}}+\frac{2 \bar{x}}{(y-\bar{x})^{2}}+\frac{1}{y-\bar{x}}+\frac{1}{\bar{y} \bar{x}}\right) \ln y\right.} \\
& +\left(-\frac{\bar{x}^{2}}{(y-\bar{x})^{3}}-\frac{2 \bar{x}}{(y-\bar{x})^{2}}-\frac{1}{y-\bar{x}}-\frac{x}{(y-x) \bar{x}}-\frac{1}{\bar{y} \bar{x}}\right) \ln \bar{x} \\
& \left.+\frac{x \ln \bar{y}}{(y-x) \bar{x}}-\frac{\bar{x}}{(y-\bar{x})^{2}}-\frac{3}{2(y-\bar{x})}+\frac{3}{2 \bar{y} \bar{x}}\right]
\end{align*}
$$

The $\alpha_{s}^{2}$ corrections of the hard spectator interactions have already been calculated in [13, 14]. However both of these calculations have been performed in the framework of SCET, while my result is a pure QCD calculation. In order to compare (4.1)-(4.4) to [13, 14] we have to take into account the definition of $\lambda_{B}$. The SCET calculation naturally uses the $\lambda_{B}$, defined by the HQET field for the $b$-meson, while I define $\lambda_{B}$ by QCD-fields. Those two definitions differ at $\mathcal{O}\left(\alpha_{s}\right)$, which has been discussed in appendix C. The difference in the logarithmic moments of the $B$-meson wave function does not play a role, because these moments occur first at NLO. Using (C.14) I figured out with the help of a computer algebra system, that (4.1)-(4.4) reproduce the results of [13, 14].

The main difference between the present QCD calculation and the framework of SCET is the way how to make the expansion in $\Lambda_{\mathrm{QCD}} / m_{b}$. While in the QCD calculation this expansion takes place at the level of the amplitude and of Feynman integrals, in SCET the Lagrangian is expanded in powers of $\Lambda_{\mathrm{QCD}} / m_{b}$. This leads to the fact that the structure of the calculation of [13, 14 is completely different from the present calculation such that comparing intermediate results like single Feynman diagrams is not possible. So it is allowed to state that the analytical coincidence of the present result with the results published before gives more independent test of [13, 14] than than a SCET calculation could provide.

### 4.2 Scale dependence

It is instructive to have a closer look to the scale dependence of (4.1)- (4.4). As stated in section 2.2.4 the scale dependence of $\mathcal{A}^{\mathrm{II}}$ vanishes. In our case we can prove this up to $\mathcal{O}\left(\alpha_{s}^{2}\right)$ i.e.

$$
\begin{equation*}
\frac{d}{d \ln \mu} \mathcal{A}^{\mathrm{II}}=\mathcal{O}\left(\alpha_{s}^{3}\right) . \tag{4.5}
\end{equation*}
$$

We need the scale dependence of the following quantities, which I took from [18]:

$$
\begin{align*}
\frac{d}{d \ln \mu} C_{1} & =\frac{\alpha_{s}}{4 \pi}\left(12 C_{N} C_{2}+\left(12 C_{F}-6 C_{G}\right) C_{1}\right)  \tag{4.6}\\
\frac{d}{d \ln \mu} C_{2} & =\frac{\alpha_{s}}{4 \pi}\left(12 C_{N} C_{1}+\left(12 C_{F}-6 C_{G}\right) C_{2}\right)  \tag{4.7}\\
\frac{d}{d \ln \mu} \alpha_{s} & =-\frac{\alpha_{s}^{2}}{4 \pi} 2 \beta_{0} \tag{4.8}
\end{align*}
$$

where $\beta_{0}=\frac{11 C_{G}-4 n_{f} C_{N}}{3}$ and $n_{f}=5$ the number of active flavours. Regarding the wave functions we need the scale dependence of their convolution integrals with the LO kernels:

$$
\begin{align*}
\frac{d}{d \ln \mu} \int_{0}^{1} \frac{d x}{\bar{x}} \phi_{\pi}(x, \mu) & =\frac{\alpha_{s}}{\pi} C_{F} \int_{0}^{1} d x \frac{3+2 \ln \bar{x}}{2 \bar{x}} \phi_{\pi}(x, \mu)  \tag{4.9}\\
\frac{d}{d \ln \mu} \int_{0}^{1} \frac{d \xi}{\xi} \phi_{B 1}(\xi, \mu) & =\frac{\alpha_{s}}{4 \pi} C_{F} \int_{0}^{1} \frac{d \xi}{\xi}(4 \ln \xi+6) \phi_{B 1}(\xi, \mu) \tag{4.10}
\end{align*}
$$

(4.9) can be obtained using the renormalisation group equation (RGE) for lightcone distribution amplitudes, which can be found in [9, 34]. We get 4.10) from [35], where the RGE for the $B$-meson light-cone wave function, defined in the framework of HQET, is given. We get the RGE for the pure QCD defined wave function by matching the nonlocal heavy to light current with QCD. The matching coefficient is given in Appendix C.

In the case of $\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}$we obtain for the $\mu$-dependent part of $\mathcal{A}^{\mathrm{II}}$ :

$$
\begin{align*}
& \mathcal{A}^{\mathrm{II}}\left(\bar{B}^{0} \rightarrow \pi^{+} \pi^{-}\right)=-i \frac{G_{F}}{\sqrt{2}} \lambda_{u}^{\prime} f_{\pi}^{2} f_{B} \int_{0}^{1} d x d y d \xi \phi_{\pi}(x) \phi_{\pi}(y) \phi_{B 1}(\xi) \frac{1}{\bar{x} \bar{y} \xi} \times \\
& {\left[\frac{\pi \alpha_{s} C_{F}}{N_{c}^{2}} C_{2}-\frac{\alpha_{s}^{2} C_{F}}{4 N_{c}^{2}}\left(\operatorname { l n } \mu \left(12 C_{N} C_{1}\right.\right.\right.} \\
& \left.\quad+C_{2}\left(\frac{40}{3} C_{N}+C_{F}(30+4 \ln \bar{x}+4 \ln \bar{y}+4 \ln \xi)-\frac{40}{3} G_{G}\right)\right) \\
& \quad+(\ldots))] \tag{4.11}
\end{align*}
$$

where the ellipsis (. . .) stands for $\mu$-independent terms. Using (4.6)-(4.10) it is easily seen that (4.5) is fulfilled. In the case of $\bar{B}^{0} \rightarrow \pi^{0} \pi^{0}$ one just has to interchange $C_{1}$ and $C_{2}$.

### 4.3 Convolution integrals and factorizability

By looking at the hard scattering kernels of (4.1)-(4.4) it is not obvious that there remain no singularities in the convolution integrals over wave functions (2.45). It is however possible to perform the integration analytically, which proves the factorizabilty.

Regarding the $B$-meson wave function we will obtain the result in terms of the quantities $\lambda_{B}$ and $\lambda_{n}$, which are defined in 2.18) and 2.19. The $\pi$-meson wave function is given in terms of Gegenbauer polynomials:

$$
\begin{equation*}
\phi_{\pi}(x)=6 x \bar{x}\left[1+\sum_{n=1}^{\infty} a_{n}^{\pi} C_{n}^{(3 / 2)}(2 x-1)\right] . \tag{4.12}
\end{equation*}
$$

Because of the symmetry properties of the pion we set $a_{2 n-1}^{\pi}=0$, furthermore we neglect $a_{n}^{\pi}$ for $n>2$. So we need only the second Gegenbauer polynomial which is
given by:

$$
\begin{equation*}
C_{2}^{(3 / 2)}(x)=\frac{15}{2} x^{2}-\frac{3}{2} . \tag{4.13}
\end{equation*}
$$

Using (2.45) we get for the NLO of $A_{\text {spect }}$ :

$$
\begin{align*}
& A_{\text {spect. } 1}^{(2)}=\alpha_{s}^{2} \frac{i f_{\pi}^{2} f_{B}}{4 N_{c}^{2}} C_{F} \frac{m_{B}}{\lambda_{B}} \times \\
& \begin{aligned}
& {\left[C_{N}\left(120 \ln \frac{\mu}{m_{b}}-48 \lambda_{1}+152\right)\right.} \\
&+ C_{F}\left(\left(162+36 \lambda_{1}\right) \ln \frac{\mu}{m_{b}}-9 \lambda_{2}+\left(-54+6 \pi^{2}\right) \lambda_{1}+\frac{1566}{5}-\frac{1008}{5} \zeta(3)+27 \pi^{2}\right. \\
&\left.+i\left(-9 \pi+\frac{18}{5} \pi^{3}\right)\right) \\
&-\frac{1}{2} C_{G}\left(240 \ln \frac{\mu}{m_{b}}+\left(-102+6 \pi^{2}\right) \lambda_{1}+\frac{2101}{5}-\frac{1008}{5} \zeta(3)+18 \pi^{2}\right. \\
&\left.\quad+i\left(9 \pi+\frac{18}{5} \pi^{3}\right)\right) \\
&+a_{2}^{\pi}\left\{C_{N}\left(240 \ln \frac{\mu}{m_{b}}-96 \lambda_{1}+404\right)\right. \\
&+C_{F}\left(\left(174+72 \lambda_{1}\right) \ln \frac{\mu}{m_{b}}-18 \lambda_{2}+\left(-\frac{741}{2}+42 \pi^{2}\right) \lambda_{1}-\frac{14809}{35}\right. \\
&\left.\quad-\frac{45072}{35} \zeta(3)+204 \pi^{2}+i\left(-338 \pi+\frac{1362}{35} \pi^{3}\right)\right) \\
& \quad-\frac{1}{2} C_{G}\left(480 \ln \frac{\mu}{m_{b}}+\left(-504+42 \pi^{2}\right) \lambda_{1}+\frac{22299}{35}-\frac{43992}{35} \zeta(3)+161 \pi^{2}\right. \\
&\left.\left.\left.\quad+i\left(-292 \pi+\frac{1482}{35} \pi^{3}\right)\right)\right\}\right]
\end{aligned}
\end{align*}
$$

and

$$
\begin{align*}
& A_{\text {spect. 2 }}^{(2)}=\alpha_{s}^{2} \frac{i f_{\pi}^{2} f_{B}}{4 N_{c}^{2}} C_{F} C_{N} \frac{m_{B}}{\lambda_{B}} \times \\
& \quad\left[108 \ln \frac{\mu}{m_{b}}+\frac{1467}{10}+\frac{252}{5} \zeta(3)-6 \pi^{2}+i\left(54 \pi-\frac{12}{5} \pi^{3}\right)\right. \\
& \left.\quad+a_{2}^{\pi}\left(216 \ln \frac{\mu}{m_{b}}+\frac{40281}{140}+\frac{29268}{35} \zeta(3)-112 \pi^{2}+i\left(118 \pi-\frac{108}{35} \pi^{3}\right)\right)\right] . \tag{4.15}
\end{align*}
$$

The finiteness of the above equations proves factorization of the hard spectator interactions at NLO.

| CKM-parameters |  |  |  |  | $\gamma$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $V_{u d}$ [3] | $V_{\text {cd }}$ | $V_{\text {cb }}$ [3] |  | \| 3 |  |  |
| 0.974 | -0.23 | 0.041 | 0.09 | 0.025 |  | $70 \pm 20) \mathrm{deg}$ |
| Parameters of the $B$-meson |  |  |  |  |  |  |
| $m_{B}$ | $f_{B}$ [36] | $\frac{f_{B}}{f^{B \lambda^{\prime}}{ }^{\text {a }} \text { [ }}$ [7] | $\lambda_{1}$ [13] | $\lambda_{2}$ 13] | $\tau_{B^{ \pm}}$ | $\tau_{B^{0}}$ |
| 5.28 GeV | $(210 \pm 19) \mathrm{MeV}$ | $1.56 \pm 0.17$ | $-3.2 \pm 1$ | $11 \pm 4$ | 1.67 ps | 1.54 ps |
| Parameters of the $\pi$-meson |  |  |  |  |  |  |
| $f_{+}^{B \pi}$ [38, | 39, 40 | $f_{\pi}$ | $m_{\pi}$ | $a_{1}^{\pi}$ |  | $a_{2}^{\pi}$ [41, 42] |
| $0.28 \pm$ | 0.0513 | MeV | 130 MeV | 0 |  | $0.3 \pm 0.15$ |
| Quark and W-boson masses |  |  |  |  |  |  |
| $m_{b}\left(m_{b}\right)$ |  | $m_{b}$ ) | $m_{t}($ | [10] |  | $M_{W}$ |
| 4.2 GeV | (1.3土 | 2) GeV |  | GeV |  | 80.4 GeV |
| Coupling constants |  |  |  |  |  |  |
| $\Lambda \frac{(5)}{\mathrm{MS}}$ |  |  | $G_{F}$ |  |  |  |
| 225 MeV |  |  | $1.16639 \times 10^{-5} \mathrm{GeV}^{-2}$ |  |  |  |

Table 4.1: Input parameters, which were used in the numerical analysis. All parameters given without explicit citation can be found in [43]. Unless otherwise stated scale dependent quantities are given at $\mu=1 \mathrm{GeV}$.

Including the contributions of $A_{\text {spect. } 1}^{(2)}$ and $A_{\text {spect. } 2}^{(2)}$ the quantities $a_{1, \text { II }}$ and $a_{2, \text { II }}$ defined in (2.31) and (2.49) are

$$
\begin{align*}
& a_{1, \text { II }}=\frac{i}{f_{\pi} f_{+}^{B \pi} m_{B}^{2}}\left(C_{2} A_{\text {spect. 1 }}^{(1)}+C_{2} A_{\text {spect. 1 }}^{(2)}+C_{1} A_{\text {spect. 2 }}^{(2)}\right) \\
& a_{2, \text { II }}=\frac{i}{f_{\pi} f_{+}^{B \pi} m_{B}^{2}}\left(C_{1} A_{\text {spect. 1 }}^{(1)}+C_{1} A_{\text {spect. 1 }}^{(2)}+C_{2} A_{\text {spect. 2 }}^{(2)}\right), \tag{4.16}
\end{align*}
$$

where

$$
\begin{equation*}
A_{\text {spect. 1 }}^{(1)}=\frac{-i C_{F} \pi \alpha_{s}}{N_{c}^{2}} \frac{f_{B} f_{\pi}^{2} m_{B}}{\lambda_{B}} 9\left(1+a_{2}^{\pi}\right)^{2} \tag{4.17}
\end{equation*}
$$

### 4.4 Numerical analysis

### 4.4.1 Input parameters

For my numerical analysis I use the parameters given in table 4.1. The decay constant $f_{B}$ and the ratio $\frac{f_{B}}{f_{+}^{B \pi} \lambda_{B}}$ have been obtained by QCD sum rules in [36] and [37] respectively. The logarithmic moments $\lambda_{1}$ and $\lambda_{2}$ where calculated in [13] using model light-cone wave functions for the $B$-meson [44, 45, 46, 47]. For the form factor $f_{+}^{B \pi}$ I use the value from [40], which has been obtained by QCD sum rules.

This value is consistent with quenched and recent unquenched lattice calculations [38, 39]. The first Gegenbauer moment of the pion wave function is zero due to G-parity while the second moment has been obtained by lattice simulations [41, 42].

### 4.4.2 Power suppressed contributions

In our numerical analysis we include two contributions which are suppressed in leading power but numerically enhanced: The twist- 3 contributions and the annihilation topologies.

The twist- 3 contributions come from the twist- 3 contributions of the pion wave functions and are suppressed by the factor

$$
r_{\chi}^{\pi}(\mu)=\frac{2 m_{\pi}^{2}}{\bar{m}_{b}(\mu)\left(\bar{m}_{u}(\mu)+\bar{m}_{d}(\mu)\right)}
$$

(see (2.33)), which is formally of $\mathcal{O}\left(\Lambda_{\mathrm{QCD}} / m_{b}\right)$ but numerically about 0.9 . These contributions modify the hard spectator scattering function $H_{\pi \pi}$ (2.50) to [10]

$$
\begin{equation*}
H_{\pi \pi}=\frac{f_{B} f_{\pi}}{m_{B}^{2} f_{+}^{B \pi}} \int_{0}^{1} \frac{d \xi}{\xi} \Phi_{B 1}(\xi) \int_{0}^{1} \frac{d x}{\bar{x}} \phi_{\pi}(x)\left(\int_{0}^{1} \frac{d y}{\bar{y}} \phi_{\pi}(y)+\frac{\bar{x}}{x} r_{\chi}^{\pi} X_{H}\right) \tag{4.18}
\end{equation*}
$$

where $X_{H}$ is parametrised by:

$$
\begin{equation*}
X_{H}=\left(1+\rho_{H} e^{i \phi_{H}}\right) \ln \frac{m_{b}}{\Lambda_{h}} \tag{4.19}
\end{equation*}
$$

with

$$
\begin{equation*}
0 \leq \rho_{H} \leq 1 \tag{4.20}
\end{equation*}
$$

For numerical calculations we set

$$
\begin{equation*}
\ln \frac{m_{b}}{\Lambda_{h}} \approx 2.4 \tag{4.21}
\end{equation*}
$$

The annihilation contributions are parametrised in (2.34)-(2.37). Analogously to 4.19) $X_{A}$ is parametrised by

$$
\begin{equation*}
X_{A}=\left(1+\rho_{A} e^{i \phi_{A}}\right) \ln \frac{m_{b}}{\Lambda_{h}} \tag{4.22}
\end{equation*}
$$

with

$$
\begin{equation*}
0 \leq \rho_{A} \leq 1 \tag{4.23}
\end{equation*}
$$

### 4.4.3 Amplitudes $a_{1}$ and $a_{2}$

The amplitudes $a_{1}$ and $a_{2}$ are defined in (2.31). Their hard scattering parts $a_{1, \text { II }}$ and $a_{2, \text { II }}$, i.e. the parts of $a_{1}$ and $a_{2}$, which contribute to $\mathcal{A}^{\text {II }}$ (see 2.29), are plotted in fig. 4.1 as functions of the renormalisation scale $\mu$. The strong dependence on $\mu$ of the real part of LO is reduced at NLO. Taking the twist-3 contributions into account does not increase the $\mu$-dependence too much. The imaginary part, which occurs


Figure 4.1: Contribution of the hard spectator corrections to $a_{1}$ and $a_{2}$ as a function of the renormalisation scale $\mu$. The upper two figures show the real part, where the LO is given by the dashed line, while the sum of LO and NLO is shown by the thick solid line. The twist-3 corrections are included in the graph given by the thin solid line. The third figure shows the imaginary part, which occurs first at $\mathcal{O}\left(\alpha_{s}^{2}\right)$. So no distinction between LO and NLO is made.
first at NLO, is strongly dependent on the renormalisation scale. An appropriate choice for the scale of the hard scattering amplitude is the hard collinear scale

$$
\begin{equation*}
\mu_{\mathrm{hc}}=1.5 \mathrm{GeV} \tag{4.24}
\end{equation*}
$$

In the following numerical calculations we will evaluate $a_{1, \text { II }}$ and $a_{2, \text { II }}$ at $\mu_{\mathrm{hc}}$. The vertex corrections $\mathcal{A}^{\mathrm{I}}$ will be evaluated at

$$
\begin{equation*}
\mu_{b}=4.8 \mathrm{GeV} \tag{4.25}
\end{equation*}
$$

Using the parameters of table 4.1 we obtain

$$
\begin{align*}
a_{1}= & 1.015+[0.039+0.018 i]_{V}+[-0.012]_{\mathrm{tw} 3}+[-0.029]_{\mathrm{LO}} \\
& +[-0.010-0.031 i]_{\mathrm{NLO}} \\
a_{2}= & 0.184+[-0.171-0.080 i]_{V}+[0.038]_{\mathrm{tw} 3}+[0.096]_{\mathrm{LO}} \\
& +[0.021+0.045 i]_{\mathrm{NLO}} . \tag{4.26}
\end{align*}
$$

These equations are given in a form similar to (61) and (62) in [13]. The first number gives the tree contribution, the vertex corrections are indicated by the label $V$, the twist-3 contributions, which come from the last part of (4.18), are labelled by tw3. The hard scattering part is separated into LO and NLO. The hadronic input parameters I used are slightly different from [13] and in contrast to [13] I evaluated all quantities, which belong to the hard scattering amplitude, at the hard collinear scale $\mu_{\mathrm{hc}}$. This is why the values I get for $a_{1}$ and $a_{2}$ are different from [13].

The hard scattering amplitudes $a_{1, \text { II }}$ and $a_{2, \text { II }}$ together with their numerical errors read:

$$
\begin{align*}
a_{1, \mathrm{II}}= & -0.051 \pm 0.011(\text { param. })_{-0.005}^{+0.026}(\text { scale }) \pm 0.012(\mathrm{tw} 3) \\
& +\left[-0.031 \pm 0.008(\text { param. })_{-0.031}^{+0.024}(\text { scale }) \pm 0.012(\text { tw } 3)\right] i \\
a_{2, \mathrm{II}}= & 0.15 \pm 0.03(\text { param. })_{-0.04}^{+0.01}(\text { scale }) \pm 0.04(\mathrm{tw} 3) \\
& +\left[0.045 \pm 0.012(\text { param. })_{-0.033}^{+0.040}(\text { scale }) \pm 0.038(\mathrm{tw} 3)\right] i \tag{4.27}
\end{align*}
$$

The first error comes from the error of the input parameters in table 4.1. The scale uncertainty is obtained by varying $\mu_{\text {hc }}$ between 1 GeV and 6 GeV . The error labelled by tw3 gives the error of the twist-3 contribution, which is obtained by varying $\rho_{H}$ between 0 and 1 and $\phi_{H}$ between 0 and $2 \pi$. Within the scale uncertainty (4.27) is compatible with [13].

It is important to remark that the result I obtained in QCD comes with formally large logarithms $\ln \Lambda_{\mathrm{QCD}} / m_{b}$. In contrast to the SCET calculation of [13, 14 it is not possible to resum these logarithms by a pure QCD calculation. Without resummation, however, these logarithms might spoil perturbation theory. Regarding this fact it is the more important that the error arising from the scale uncertainty in (4.27) is small enough for perturbation theory to be valid.

### 4.4.4 Branching ratios

The dependence of the CP-averaged branching ratios on the hard collinear scale is shown in fig. 4.2. It is obvious that the NLO corrections reduce this dependence significantly.


Figure 4.2: CP-averaged branching ratios as functions of the hard collinear scale $\mu_{\mathrm{hc}}$ in units of $10^{-6}$. In the graph with the dashed line only the leading order of the hard spectator scattering is contained, while in the solid line hard spectator scattering is taken into account up to NLO.

From the parameter set in table 4.1 we obtain the following CP-averaged branching ratios

$$
\begin{align*}
10^{6} \mathrm{BR}\left(B^{+} \rightarrow \pi^{+} \pi^{0}\right) & =6.05_{-1.98}^{+2.36}(\text { had. })_{-2.33}^{+2.90}(\mathrm{CKM})_{-0.31}^{+0.18}(\text { scale }) \pm 0.27 \text { (sublead.) } \\
10^{6} \mathrm{BR}\left(B^{0} \rightarrow \pi^{+} \pi^{-}\right) & =9.41_{-2.99}^{+3.56}(\text { had. })_{-3.46}^{+4.00}(\mathrm{CKM})_{-3.93}^{+1.07}(\text { scale })_{-0.70}^{+1.13}(\text { sublead. }) \\
10^{6} \mathrm{BR}\left(B^{0} \rightarrow \pi^{0} \pi^{0}\right) & =0.39_{-0.12}^{+0.14}(\text { had. })_{-0.17}^{+0.20}(\mathrm{CKM})_{-0.06}^{+0.17}(\text { scale })_{-0.08}^{+0.20}(\text { sublead. }) \tag{4.28}
\end{align*}
$$

The origin of the errors are the uncertainties of the hadronic parameters and the CKM parameters, the scale dependence and the subleading power contributions, i.e. twist-3 and annihilation contributions. The error arising from the scale dependence was estimated by varying $\mu_{b}$ between 2 GeV and 8 GeV and $\mu_{\mathrm{hc}}$ between 1 GeV and 6 GeV . If we compare (4.28) to the experimental values [7]:

$$
\begin{align*}
10^{6} \mathrm{BR}\left(B^{+} \rightarrow \pi^{+} \pi^{0}\right) & =5.5 \pm 0.6 \\
10^{6} \mathrm{BR}\left(B^{0} \rightarrow \pi^{+} \pi^{-}\right) & =5.0 \pm 0.4 \\
10^{6} \mathrm{BR}\left(B^{0} \rightarrow \pi^{0} \pi^{0}\right) & =1.45 \pm 0.29 \tag{4.29}
\end{align*}
$$

we note that $\operatorname{BR}\left(B^{+} \rightarrow \pi^{+} \pi^{0}\right)$ is in good agreement with the data. This quantity is almost independent of $\gamma$. The other branching ratios, which come with large errors, depend strongly on $\gamma$. This dependence is shown in fig. 4.3. The light-grey band gives the uncertainty that is defined in the same way as the errors in (4.28), where different errors are added in quadrature. The solid inner line gives the central value. The experimental values are represented by the horizontal band, whereas the vertical band gives the value of $\gamma$. It is obvious that the errors of the branching fractions are too large for a reasonable determination of $\gamma$.

For $B^{+} \rightarrow \pi^{+} \pi^{0}$ and $B^{0} \rightarrow \pi^{+} \pi^{-}$QCD-factorization is expected to work well, because at tree level Wilson coefficients occur in the so called colour allowed combination $C_{1}+C_{2} / N_{c} \sim 1$, while $B^{0} \rightarrow \pi^{0} \pi^{0}$ comes at tree level with $C_{2}+C_{1} / N_{c} \sim 0.2$ such that subleading power corrections are expected to be more important. On the other hand there are big uncertainties in the parameters occurring in the combinations $\left|V_{u b}\right| f_{+}^{B \pi}, \frac{f_{B}}{f_{+}^{B \lambda_{B}}}$ and $a_{2}^{\pi}$. In [48] and [13] these parameters were fitted by the experimental values 4.29) of $\mathrm{BR}\left(B^{+} \rightarrow \pi^{+} \pi^{0}\right)$ and $\mathrm{BR}\left(B^{0} \rightarrow \pi^{+} \pi^{-}\right)$. Setting

$$
\begin{equation*}
a_{2}^{\pi}(1 \mathrm{GeV})=0.39 \tag{4.30}
\end{equation*}
$$

leads to

$$
\begin{align*}
\left|V_{u b}\right| f_{+}^{B \pi} & \rightarrow 0.80\left(\left|V_{u b}\right| f_{+}^{B \pi}\right)_{\text {default }} \\
\frac{f_{B}}{f_{+}^{B \pi} \lambda_{B}} & \rightarrow 2.89\left(\frac{f_{B}}{f_{+}^{B \pi} \lambda_{B}}\right)_{\text {default }} . \tag{4.31}
\end{align*}
$$

This leads to the following branching ratios:

$$
\begin{align*}
10^{6} \mathrm{BR}\left(B^{+} \rightarrow \pi^{+} \pi^{0}\right) & =5.5 \pm 0.2(\text { param. })_{-0.3}^{+0.5}(\text { scale }) \pm 0.6 \text { (sublead.) } \\
10^{6} \mathrm{BR}\left(B^{0} \rightarrow \pi^{+} \pi^{-}\right) & \left.=5.0_{-0.9}^{+0.8} \text { (param. }\right)_{-0.2}^{+0.9}(\text { scale })_{-0.6}^{+0.9} \text { (sublead.) } \\
10^{6} \mathrm{BR}\left(B^{0} \rightarrow \pi^{0} \pi^{0}\right) & \left.=0.77 \pm 0.3(\text { param. })_{-0.3}^{+0.2} \text { (scale) }\right)_{-0.2}^{+0.3} \text { (sublead.). } \tag{4.32}
\end{align*}
$$



Figure 4.3: CP-averaged branching ratios as functions of the CKM-angle $\gamma$ in units of $10^{-6}$. The light-grey band gives the uncertainty from the errors of table 4.1 and from the twist- 3 and the annihilation contributions. The solid inner line gives the central value. The horizontal dark-grey band gives the experimental value according to [7] and the vertical grey band gives the value of $\gamma$ from table 4.1 within the error ranges.

The uncertainties of the quantities that occurred in (4.30) and 4.31) have not been considered in the estimation of the errors in (4.32). The $B^{0} \rightarrow \pi^{0} \pi^{0}$ branching ratio obtained in (4.32) is compatible with the value obtained in [13]. Though it is too low, due to the theoretical and experimental errors it is compatible with (4.29).

There are two different sources of errors. On the one hand for errors that are due to uncertainties of input parameters and the renormalisation scale there is at least in principle no lower limit. On the other hand errors arising from subleading power corrections, i.e. twist-3 and annihilation contributions, cannot be reduced in the framework of QCD-factorization. Fig. 4.4 shows the branching fractions of $B^{0} \rightarrow \pi^{+} \pi^{-}$and $B^{0} \rightarrow \pi^{0} \pi^{0}$ as functions of $\gamma$. The errors arising from subleading power contributions are represented by the dashed lines inside of the light-grey error band. While in the case of $B^{0} \rightarrow \pi^{+} \pi^{-}$this remaining error might be small enough for non-trivial phenomenological statements about $\gamma$, in the case of $B^{0} \rightarrow \pi^{0} \pi^{0}$ there remains an error of about $30 \%$.


Figure 4.4: CP-averaged branching ratios as functions of the CKM-angle $\gamma$ in units of $10^{-6}$ with the input parameters (4.30) and (4.31). The dashed lines inside of the light-grey band give the error coming from subleading power contributions, while the dashed lines at the border of the grey bands are included to lead the eye. The meaning of the other curves and bands is the same as in fig. 4.3 besides the fact that the parameters occurring in (4.30) and 4.31) were not included in the error estimation.

## Chapter 5

## Conclusions

In the last decades $B$ physics has proven a promising field to determine parameters of the flavour sector with high precision. It is expected that in the next few years the angles $\alpha$ and $\gamma$, which are directly connected to the complex phase of the CKM matrix, will be measured with an accuracy at the percent level. Furthermore the discovery of physics beyond the standard model will be possible.

On the theoretical side QCD factorization has turned out to be an appropriate tool to calculate $B$ decay modes from first principles, because it allows for systematic disentanglement of the perturbative physics and the non-perturbative physics. The present calculation showed that the hard spectator scattering amplitude factorizes up to and including $\mathcal{O}\left(\alpha_{s}^{2}\right)$, i.e. all infrared divergences cancel and there are no remaining endpoint singularities. The former point is obvious after the explicit calculation of $T^{\mathrm{II}}$ and the latter point was shown by evaluating the convolution integral (2.1) analytically. The explicit expressions for the hard spectator scattering kernel (4.1)-(4.4) confirmed the result of [13, 14]. So they are also a confirmation that the leading power of the amplitudes can be obtained by performing the power expansion at the level of Feynman integrals rather than at the level of the QCD Lagrangian using an effective theory like SCET, which was done in [13, 14].

The main challenges in the evaluation of Feynman integrals, which were made possible with the help of tools like integration by parts identities and differential equation techniques, were due to the fact that the Feynman integrals came with up to five external legs and three independent rations of scales. Many steps in the calculations of section 3.2 might look like cookery. However I dare say calculating Feynman integrals is cookery.

One motivation to calculate the $\mathcal{O}\left(\alpha_{s}^{2}\right)$ corrections of the hard spectator interactions separately is the fact, that the LO of this class of diagrams starts at $\mathcal{O}\left(\alpha_{s}\right)$ such that in order to fix the scale we need the NLO correction. The numerical results of section 4.4.3 show that the NLO reduces the scale dependence significantly. This is even more important with respect to large logarithms that arise because of the fact that next to the $m_{b}$-scale also the hard-collinear scale $\sqrt{\Lambda_{\mathrm{QCD}} m_{b}}$ enters the hard spectator scattering amplitude. In contrast to the effective theory ansatz the QCD calculation of this work does not allow the resummation of these logarithms. This is why it is a crucial point, that the NLO is numerically important but small enough
for perturbation theory to be valid.
Next to the scale dependence a main source of uncertainty is due to the fact that we do not know hadronic quantities well enough. This might be improved in the next few years by lattice calculations and even determination of the hadronic input parameters in experiment. Also a better control of power corrections would allow to obtain much more precise predictions from QCD factorization.

Finally it is important to note that the present calculation is not the complete order $\alpha_{s}^{2}$ result as the contributions of penguin contractions and the effective penguin operators where not considered in this thesis. Actually they play a dominant role in the branching ratios of $B \rightarrow K \pi$ and CP asymmetries of $B \rightarrow \pi \pi$ and should be taken into account in phenomenological applications. While writing down this thesis the order $\alpha_{s}^{2}$ of these contributions has been recently published in [49]. Also the $\mathcal{O}\left(\alpha_{s}^{2}\right)$ corrections of $T^{\mathrm{I}}$ were not part of this thesis. These contributions have been calculated in [11, 12 .

So the calculation of the present thesis is a small but very important tessera in the mosaic of theoretical $B$-physics.

## Appendix A

## CAS implementation of IBP identities

## A. 1 User manual

This section will give an introduction how to use my Mathematica packages lorentz.m and ibp.m. These packages use the rules of section 2.4.1] and the algorithm of [22]. You can download these files from

```
http://www.theorie.physik.uni-muenchen.de/~ pilipp
```

I assume that these files are located on your hard disk in the directory path. After you have started your Mathematica notebook with the two lines

```
<<path/lorentz.m;
<<path/ibp.m;
```

you have to set some variables. Because my program distinguishes between Lorentz vectors and scalars we have to define which variables are of the type vector. This is done with the function

```
AddMomenta[p1,..., pn]
```

which defines the variables $\mathrm{p} 1, \ldots, \mathrm{pn}$ to be of the type vector. The function

```
RemMomenta[p1,...,pn]
```

removes the attribute vector from $\mathrm{p} 1, \ldots, \mathrm{pn}$ and

## ShowMomenta[]

gives list of all vector variables. Per default the variables p, q and 1 are defined to be vector variables.

The syntax of defining scalar products is the same as in Tracer [50]. The OnShellcommand

```
OnShell[on,{p1,0},{p2,p3,m},...]
```

defines the scalar products $\mathrm{p} 1 \cdot \mathrm{p} 1=0$ and $\mathrm{p} 2 \cdot \mathrm{p} 3=\mathrm{m}$. By default there are the following definitions:

OnShell[on, \{p, 0\}, \{q, 0\}, \{1, 0\}, \{p, $q, 1 / 2\},\{p, 1, x i / 2\},\{q, 1$, theta/2\}]
To undo the onshell definition use the flag off instead of on.
An integral of the form (2.51) contains the set of momenta $\left\{p_{1}, \ldots, p_{n}\right\}$, which are in general linear combinations of basis momenta e.g. $\left\{0, p^{\mu}, p^{\mu}+y q^{\mu}\right\}$ where the basis momenta are $\left\{p^{\mu}, q^{\mu}\right\}$. To tell Mathematica which variables are the basis momenta we have to define the variable MomBasis. In our example we set:

$$
\text { MomBasis }=\{p, q\}
$$

After this definition the function

$$
\text { ExternalMomenta }\left[p_{1}, \ldots, p_{n}\right] ;
$$

has to be called to tell Mathematica that $p_{1}, \ldots, p_{n}$ are the momenta which appear in the Feynman integrals. In the above example:

$$
\text { ExternalMomenta }[0, \mathrm{p}, \mathrm{p}+\mathrm{y} * \mathrm{q}] ;
$$

Feynman integrals are represented by the function FInt. This function will be simplified applying rule 1, 2 and rule 3 of section 2.4.1. After the call of the function ExternalMomenta $\left[p_{1}, \ldots, p_{n}\right]$ the momentum $p_{i}$ is represented by the position $i$ at which it appears in the argument list. So the integral (2.51) is represented by

$$
\begin{aligned}
& \operatorname{FInt}\left[\left\{\left\{i_{1}, M_{1}^{2}, m_{1}\right\}, \ldots,\left\{i_{t}, M_{t}^{2}, m_{t}\right\}\right\},\left\{\left\{\tilde{i}_{1}, \tilde{M}_{1}^{2}, \tilde{m}_{1}\right\}, \ldots,\left\{\tilde{i}_{u}, \tilde{M}_{u}^{2}, \tilde{m}_{u}\right\}\right\},\right. \\
& \left.\quad\left\{\left\{j_{1}, s_{1}\right\}, \ldots,\left\{j_{l}, s_{l}\right\}\right\}\right]
\end{aligned}
$$

Because most integrals do not have propagators of the form $k \cdot p+M^{2}$, the second argument of FInt can be dropped such that

$$
\operatorname{FInt}\left[\left\{\left\{i_{1}, M_{1}^{2}, m_{1}\right\}, \ldots,\left\{i_{t}, M_{t}^{2}, m_{t}\right\}\right\},\left\{\left\{j_{1}, s_{1}\right\}, \ldots,\left\{j_{l}, s_{l}\right\}\right\}\right]
$$

represents the integral

$$
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{s_{1}^{n_{1}} \ldots s_{l}^{n_{l}}}{D_{1}^{m_{1}} \ldots D_{t}^{m_{t}}}
$$

For example: We want to represent the integral

$$
\begin{equation*}
\int \frac{d^{d} k}{(2 \pi)^{d}} \frac{k \cdot(p+y q)}{k^{2}(k+p)^{2}(k+p+y q)^{2}} . \tag{A.1}
\end{equation*}
$$

After the above call of ExternalMomenta $[0, \mathrm{p}, \mathrm{p}+\mathrm{y} * \mathrm{q}]$ the momenta $0, \mathrm{p}, \mathrm{p}+\mathrm{y} * \mathrm{q}$ are represented by the numbers 1,2 and 3 respectively. So A.1 is represented by

$$
\operatorname{FInt}[\{\{1,0,1\},\{2,0,1\},\{3,0,1\}\},\{\{3,1\}\}]
$$

which is transformed into

$$
-1 / 2 * y * \operatorname{FInt}[\{\{1,0,1\},\{2,0,1\},\{3,0,1\}\},\{ \}]
$$

because rule 1 and rule 3 are applied and scaleless integrals vanish in dimensional regularisation.

The identities from rule 4 to rule 6 are created by the function IBP. This function takes three or five arguments and is called by

```
IBP[Denom1, Denom2, la1, la2, lb]
```

or

$$
\operatorname{IBP}[\text { Denom, la, lb] }
$$

which is a shortcut for

$$
\operatorname{IBP}[\operatorname{Denom},\{ \}, 1 \mathrm{a},\{ \}, \mathrm{lb}] .
$$

The first two arguments of IBP Denom1 and Denom2 are lists that take the form $\left\{\left\{i_{1}, M_{1}^{2}\right\}, \ldots,\left\{i_{t}, M_{t}^{2}\right\}\right\}$ and $\left\{\left\{\tilde{i}, \tilde{M}_{1}^{2}\right\}, \ldots,\left\{\tilde{i}_{u}, \tilde{M}_{u}^{2}\right\}\right\}$ respectively and describe the topology i.e. they tell Mathematica to create identities of integrals which contain the propagators $D_{1}, \ldots, D_{t}$ and $\tilde{D}_{1}, \ldots, \tilde{D}_{u}$ respectively. The powers $m_{i}$ and $\tilde{m}_{i}$ to which the propagators $D_{i}$ and $\tilde{D}_{i}$ have to appear, are given by the next two arguments la1 and la2. These are lists whose elements are of the form $\{n, m\}$ : For all integrals with $n$ different propagators of the form $D$ and $\tilde{D}$ respectively identities of the form (2.59), 2.60 and 2.62 are created for all integrals, where $\sum_{k=1}^{n}\left(m_{k}-1\right) \leq m$ and $\sum_{k=1}^{n}\left(\tilde{m}_{k}-1\right) \leq m$ respectively. The third argument $l b$ has the same form as la1 and la2 and tells the program how many scalar products of the form $k \cdot p$ should be in the numerator. Here $n$ stands for the number of different propagators of the form $D$ and $\tilde{D}$ and the integrands have to fulfil the condition $\sum_{k=1}^{l} n_{j_{k}} \leq m$. For all of the lists la1, la2 and lb the default value for $m$ is 0 e.g. $\{\{2,1\},\{3,0\}\}$ and $\{\{2,1\}\}$ lead to the same result.

In our example a convenient call of IBP would be

$$
\text { subslist }=\operatorname{IBP}[\{\{1,0\},\{2,0\},\{3,0\}\},\{\{2,1\},\{3,0\}\},\{\{2,1\},\{3,0\}\}] ;
$$

which is equivalent

$$
\text { subslist }=\operatorname{IBP}[\{\{1,0\},\{2,0\},\{3,0\}\},\{\{2,1\}\},\{\{2,1\}\}] ;
$$

because the default value of the powers of propagators is 0 . The command

$$
\text { FInt }[\{\{1,0,1\},\{2,0,1\},\{3,0,1\}\},\{\{3,1\}\}] / / . \text { subslist }
$$

reduces our integral to

$$
-(1-2 * e) * \operatorname{FInt}[\{\{1,0,1\},\{3,0,1\}\},\{ \}] / 2 / e
$$

where $\mathrm{e}=(4-d) / 2$.

## A. 2 Implementation

In this section I will describe in detail how to implement the rules of section 2.4.1 in Mathematica . I will follow the algorithm described in [22], the reader is expected to be familiar with this algorithm.

First of all we have to tell our computer algebra system how to handle Lorentz vectors. We will write all the definitions into the file lorentz.m. It proves to be useful to distinguish between Lorentz vectors and scalar variables. In a CAS which does not know about type declarations of variables this is done by putting all the vector variables in a list we call MomList. So the first part of the file lorentz.m looks like this:

```
BeginPackage[ "LORENTZ`" ];
(*Pattern variables*)
Unprotect[a,b, c,mom,mom1,mom2,ip,rest];
Clear[a,b,c,mom,mom1,mom2,ip,rest];
Protect[a,b,c,mom,mom1,mom2,ip,rest];
Unprotect[d,e];
Clear[d,e];
d = 4 - 2*e;(*Dimension*)
Protect[d,e];
Unprotect[MomList]; MomList := {};
(*List of variables which are defined to be momenta;
all other variables are handled as scalars*)
Protect [MomList];
Unprotect [AddMomenta];
Clear [AddMomenta];
AddMomenta[qlist___] :=
    (Unprotect[MomList];
            MomList = Union[MomList, {qlist}];
            Protect[MomList];)
Protect[AddMomenta];
Unprotect [RemMomenta];
Clear [RemMomenta];
RemMomenta[qlist___] :=
    (Unprotect[MomList];
        MomList = Complement[MomList, {qlist}];
```

```
    Protect[MomList];)
Protect[RemMomenta];
Unprotect[ClearMomenta];
Clear[ClearMomenta];
ClearMomenta[] :=
    (Unprotect[MomList]; MomList = {}; Protect[MomList];)
Protect[ClearMomenta];
Unprotect [ShowMomenta];
Clear [ShowMomenta];
ShowMomenta[] := Return[MomList];
Protect[ShowMomenta];
```

After we have defined which variables are used as pattern variables and we have set the dimension variable $\mathrm{d}=4-2 *$ e, we introduce the protected list variable MomList which is manipulated by the functions AddMomenta, RemMomenta, ClearMomenta and ShowMomenta.

In the next step we define the function IsVector which tells us if a variable is of the type vector (i.e. it is contained in MomList) or scalar:

```
Unprotect[IsVector];
Clear[IsVector];
IsVector[a_ + b_] := IsVector[a] || IsVector[b];
IsVector[a_ b_] := IsVector[a] || IsVector[b];
IsVector[a_ /; MemberQ[MomList, a]] := True; (*that's the point*)
IsVector[a_] := False;
Protect[IsVector];
```

Following the conventions of "Tracer" [50] we define a scalar product between Lorentz vectors which gets the attribute Orderless. The scalar product is defined to be linear, which makes the distinction between scalar and vector variables necessary.

```
Unprotect [SP];
Clear[SP];
SetAttributes[SP, Orderless]; (*scalarproduct of two Lorentzvectors*);
SP[(a_ /; ! IsVector[a])*b_, c_] := a*SP[b, c];
SP[a_ + b_, c_] := SP[a, c] + SP[b, c];
SP[0, _] := 0;
Protect[SP];
```

The function OnShell is defined as in 50 i.e. we define the scalar products $p \cdot q=\frac{1}{2}$ and $p^{2}=0$ by the command OnShell [on, $\{\mathrm{p}, 0\},\{\mathrm{p}, \mathrm{q}, 1 / 2\}$ ].

Off [General::spell1];
Unprotect[OnShell];

```
Clear[OnShell];
OnShell[flag_, list___] :=
    (*Defined as in Tracer*)
    Module[{l, i},
        l = {list};
        Unprotect [SP];
        Switch[flag,
        on, For[i = 1, i <= Length[l], i++,
            Switch[Length[l[[i]]],
                2, SP[l[[i]][[1]], l[[i]][[1]]] = l[[i]][[2]],
                3, SP[1[[i]][[1]], l[[i]][[2]]] = l[[i]][[3]];
                ]
            ],
        off, For[i = 1, i <= Length[l], i++,
            Switch[Length[1[[i]]],
                2, SP[l[[i]][[1]], l[[i]][[1]]] =.,
                3, SP[1[[i]][[1]], l[[i]][[2]]] =.;
                ]
            ]
        ];
        Protect[SP];
        ];
Protect[OnShell];
```

The last function we introduce is Project. This function, applied to a linear combination of vector variables and a vector variable $p$, gives the coefficient of $p$ in that linear combination. E.g. assume $p$ and $q$ are vector variables and $x$ and $y$ are scalars then Project $[\mathrm{x} * \mathrm{p}+\mathrm{y} * \mathrm{q}, \mathrm{p}]$ is simplified to x . This is the definition of Project:

Unprotect [Project];
Clear[Project];
Project [a_+b_, mom_]/;
MemberQ[ShowMomenta[],mom]:=Project [a,mom] +Project [b,mom];
Project[a_*mom1_, mom2_]/;
MemberQ[ShowMomenta[] , mom2]\&\&IsVector [mom1] \&\&!IsVector [a]:=
a*Project[mom1,mom2];
Project[0,mom_]:=0;
Project[mom_, mom_]/;MemberQ[ShowMomenta[] ,mom] :=1;
Project[mom1_, mom2_]/;
MemberQ[ShowMomenta [] , mom2] \&\&MemberQ[ShowMomenta [] , mom1] : $=0$; Protect [Project];

The end of the file lorentz.m is special for the calculation in this thesis i.e. we introduce the Lorentz vectors $p, q$ and $l$ whose scalar products fulfil our kinematical conditions:

```
(*------------------other definitions------------------------*)
```

AddMomenta[p, q, l];
OnShell[on, \{p,p,0\}, \{q,q,0\}, \{p,q,1/2\}, \{l,0\}, \{p,1,xi/2\}, \{q,l,theta/2\}];

```
(*-------------------end other definitions-------------------*)
```


## EndPackage []

We will write the definitions which handle the reduction of Feynman integrals into the file ibp.m. The first part of this file covers the global variables which will be described in more detail when they will be used. The variable MomBasis which has to be defined by the user contains all the basis momenta i.e. all external momenta which appear in the integrals have to consist of a linear combination of the components of MomBasis.

```
(*Patterns*)
Unprotect [Denom,Denom1,Denom2,Num, arg,
    expr1, expr2,ip,i1p,i2p,i3p,i4p,intp,n1p,n2p,Mp,M1p,M2p,
    a,b,c,dp,a1,b1,c1,d1,e1, a2,b2,c2,d2,M1,M2,m1,m2,mp,p1,p2,L,sp];
Clear[Denom,Denom1,Denom2,Num, arg,
    expr1,expr2,ip,i1p,i2p,i3p,i4p,intp,n1p,n2p,Mp,M1p,M2p,
    a,b,c,dp,a1,b1,c1,d1,e1, a2,b2, c2, d2,M1,M2,m1,m2,mp,p1,p2,L,sp];
Protect[Denom,Denom1,Denom2,Num, arg,
    expr1, expr2,ip,i1p,i2p,i3p,i4p,n1p,n2p,intp,Mp,M1p,M2p,
    a,b,c,dp,a1,b1,c1,d1,e1,a2,b2,c2,d2,M1,M2,mp,m1,m2,p1,p2,L,sp];
(*all global variables*)
MomBasis = {p,q,l};
Unprotect[eqlist];
eqlist = {};(*IBP1 writes into eqlist*)
Protect[eqlist];
Unprotect [Mom] ;
Mom = {};
Protect[Mom];
(*global list of external momenta.
    This list is used by all other functions.
```

```
    This list is set by the function ExternalMomenta (see below),
    do not edit! *)
Unprotect [LinIndepMom];
LinIndepMom = {};
Protect[LinIndepMom];
(*Gives the position of the linearly independent momenta in Mom.
    This list is set by the function ExternalMomenta (see below),
    do not edit!*)
```

If we define a set of external momenta, which appear in our integrals, not all of those will be linearly independent. To find the linearly independent ones i.e. a basis of our external momenta we define the function FindBasis. This function applied on a list of momenta gives a list of the position of the linearly independent ones. The algorithm is as follows: Consider a set of vectors $\left\{v_{1}, \ldots, v_{n}\right\}$ from which we want to choose a minimal subset of linearly independent vectors which form a basis of $\operatorname{span}\left\{v_{1}, \ldots, v_{n}\right\}$. We look for a most general solution of the equation system

$$
\begin{equation*}
v_{i}^{j} c^{i}=0 \tag{A.2}
\end{equation*}
$$

If one of the $c^{i}$ is not necessarily 0 i.e. $v_{i}$ can be expressed by a linear combination of the other vectors we remove $v_{i}$ from the set $\left\{v_{1}, \ldots, v_{n}\right\}$ and repeat the procedure until A.2 gets the unique solution $c^{i}=0$ for all $i$. Usually the basis of $\operatorname{span}\left\{v_{1}, \ldots, v_{n}\right\}$ is not unique. In this case we have an ambiguity, which of the vectors we can remove. The function below removes this vector which comes first in the list $\left\{v_{1}, \ldots, v_{n}\right\}$ i.e. the vectors which are more behind have a higher precendence to be a member of the basis which is returned. This is the implementation in Mathematica:

```
Unprotect[FindBasis];
Clear[FindBasis];
FindBasis[{ip_}] := {}/;(Mom[[ip]] == 0);
FindBasis[p_List] :=
    (*gives the position of the linearly independend momenta in p
        reads the variable MomBasis. If this is not unique, a vector
        p[[i1]] has a higher precendence than p[[i2]] iff i1>i2. The
        vector p[[i]] is an integer s.t. Mom[[ p[[i]] ]] is the
        corresponding vector*)
    Catch[
        Module[{pc, Ilist, V, C, c, i, j},
            pc = Array[Function[{i}, {Mom[[ p[[i]] ]], i}], Length[p]];
            (*maps every momentum onto a position in the list p*)
                Ilist = Range[Length[p]];
```

```
        (*result list*)
    While[True,
        V = Array[
            Function[{i, j}, Project[pc[[i, 1]], MomBasis[[j]]] ],
            {Length[pc], Length[MomBasis]}];
    C = Array[c, {Length[pc]}];
    Off[Solve::svars];
    C = C /. (Solve[C.V == 0, C][[1]]);
    On[Solve::svars];
    For[i = 1, i <= Length[C], i++,
        (*Because we start with lowest value i=1 and increase i,
        we make sure that the vectors in the begining of p are
        dropped out first, i.e. p[[i1]] has a higher precendence
        than p[[i2]] iff i1 > i2.*)
        If[! MatchQ[C[[i]], 0],
            Ilist = Drop[Ilist, Position[Ilist, pc[[i, 2]]][[1]]];
            pc = Drop[pc, {i}];
            Break[];
            ];
        If[i >= Length[C], Throw[Ilist]];
        ];
        ];
    ]
];
Protect[FindBasis];
```

In the next part we define some global functions. The function MakeComb takes the integer arguments $n$ and $k$ and gives a list of all subsets of $\{1, \ldots, n\}$ with length k. The functions GetBasis, RepMom and ruleI are well commented such that we do not explain them. These functions are defined such that they remember the value they have already calculated. So there are the commands defGetBasis, defRepMom and defruleI which redefine those functions and which are evaluated in the function ExternalMomenta. The argument of ExternalMomenta is a list $\left\{p_{1}, \ldots, p_{n}\right\}$ of the external momenta which appear in the integrals. By calling the function ExternalMomenta $\left[p_{1}, \ldots, p_{n}\right]$ the variable Mom is set to $\left\{p 1, \ldots, p_{n}\right\}$ and to the variable LinIndepMom a list of the position of the linearly independent momenta in $\left\{p 1, \ldots, p_{n}\right\}$ is assigned.

Unprotect [MakeComb] ;
Clear [MakeComb] ;

```
MakeComb[n_, k_] :=
    (*generates all subsets of {1, ..., n} with k elements *)
    Catch[
    Module[{L0, L1, i, j, l},
        LO = Array[{#} &, n];(*LO = {{1}, ..., {n}}*)
        L1 = {};
        For[l = 1, l <= k - 1, l++,
                For[j = 1, j <= Length[LO], j++,
                For[i = Last[LO[[j]]] + 1, i <= n, i++,
                    L1 = Append[L1, Append[LO[[j]], i]];
                        ];
                ];
            LO = L1;
                L1 = {};
                ];
            Throw[LO];
            ];
    ];
Protect[MakeComb];
Unprotect[partfrac,defpartfrac];
(*This function gives a list {True,{c1,...,cn}} s.t.
c1+...+cn=0,
c1*Mom[[mom[[1]]]]+\ldots...cn*Mom[[mom[[n]]]]=0,
c1*(mom[1]-mass[1])+\ldots. . cn*(mom[n]-mass[n])=1
If this not possible it gives back {False,{}}
*)
Clear[partfrac,defpartfrac];
defpartfrac:=(
partfrac[mom_,mass_]:=
Catch[
Unprotect[partfrac];
Module[{res},
res=(
partfrac[mom,mass]=
Catch[
Module[{coeff,coefflist,solvelist,eql,i1,j1,i},
    Off[Solve::svars];
        i = Length[mom];
        coefflist = Array[coeff,i];
        eql =
            Array [Function[{j1},
```

```
            Sum[coeff[i1]*Project[Mom[[ mom[[i1]] ]], MomBasis[[j1]]],
                {i1, 1, i}] == 0], Length[MomBasis]];
eql = Append[eql, Sum[coeff[i1], {i1, 1, i}] == 0];
eql = Append[eql, Sum[coeff[i1]*(
    SP[Mom[[mom[[i1]]]],Mom[[mom[[i1]]]]]-mass[[i1]]
    ),{i1,1,i}] == 1];
solvelist = Solve[eql, Array[coeff, i]];
If [MatchQ[solvelist,{}],Throw[{False,{}}]];
solvelist=solvelist[[1]];
coefflist = coefflist /. solvelist;
coefflist = coefflist /. coeff[_] -> 0;
    On[Solve::svars];
    Throw[{True,coefflist}];
](*endMod*);
](*endCatch*)
);
Protect[partfrac];
Throw[res];
]; (*endMod*)
];(*endCatch*)
);
Protect[partfrac,defpartfrac];
Unprotect[defGetBasis,GetBasis];
Clear[defGetBasis,GetBasis];
defGetBasis :=
( GetBasis[Denom1_List,Denom2_List]:=
    (*Gets Denom1 and Denom2 of the form
        {{p1,M1^2,m1},...,{pt,Mt^2,mt}} and
        {{q1,Mq1^2,mq1},....,{qr,Mqr^2,mqr}} resp. and finds the
        linear independent momenta {q_k1,...,q_kl} of {q1,...,qr}
        and {p_i1,...,p_il} of {p1,....,pt}, which are completed
        to a basis s.t. as much as possible vectors of
        {q_k1,....,q_kl} are in the basis list*)
        Catch[
        Unprotect[GetBasis];
        Module[{res, p1,p2,p, bl1,bl2, i},
            p1=Array[ Denom1[[#,1]]&, Length[Denom1] ];
            p2=Array[ Denom2[[#,1]]&, Length[Denom2] ];
            p = Join[p1,p2];
            For[i=1,i<=Length[Mom],i++,
                If[!MemberQ[p,i],p=Prepend[p,i]];
            ];
```

```
            bl1 = FindBasis[p];
            bl2 = Array[ p[[ bl1[[#]] ]]&, Length[bl1] ];
            res = (GetBasis[Denom1,Denom2] = bl2);
            Protect[GetBasis];
            Throw[res];
            ] (*endMod*);
    ](*endCatch*);
);
Protect[defGetBasis,GetBasis];
Unprotect[RepMom,defRepMom];
Clear[RepMom, defRepMom];
defRepMom := (
    RepMom[mom_,basis_List] :=
    (*RepMom takes as arguments a momentum p and a basis {p1,..., pk}.
    It gives a list {c1,...,ck} s.t. p = c1*p1+....+ck*pk *)
        Catch[
            Module[{coeff,coefflist,solvelist,eql, i,i1,i2},
                Unprotect[RepMom];
            Off[Solve::svars];
            eql = Array[
                Sum [
                    coeff[i1]*Project[Mom[[ basis[[i1]] ]], MomBasis[[#]] ],
                    {i1,1,Length[basis]} ] ==
                    Project[Mom[[mom]],MomBasis[[#]] ]&,
                    Length[MomBasis]
                    ];
        eql = eql // .
                            {(expr1_ == expr2_) /; FreeQ[{expr1, expr2}, coeff[_]] :>
                    MatchQ[expr1, expr2]};
        (*An expression of the form "x==x" in a list of equations
                        is not automatically
            transformed into "True": This has to be done by hand*)
        coefflist = Array[ coeff, Length[basis] ];
        solvelist = Solve[eql, coefflist];
        coefflist = coefflist /. solvelist;
        coefflist = coefflist /. coeff[_] -> 0;
        On[Solve::svars];
    RepMom[p,basis] = coefflist;
    Throw[ coefflist[[1]] ];
    Protect[RepMom];
```

```
        Throw[res];
    ](*endMod*);
    ](*endCatch*);
);
Unprotect[defruleI,ruleI];
Clear[defruleI,ruleI];
defruleI := (
    ruleI[mom_List]:=
    (* ruleI gives for the topology {i1,...,ik} a list of
coefficients (c1,...,ck) such that c1+...+ck = 1 and
c1*Mom[[i1]]+...+ck*Mom[[ik]] = 0. The output is of the form
    {True,{c1,...,ck}} or {False} if this is not possible *)
    Catch[
            Module[{res, coefflist,coeff,eql, i,j,i1,j1},
                Off[Solve::svars];
                coefflist = Array[coeff, Length[mom]];
                eql =
                Array[Function[{j1},
                    Sum[coeff[i1]*Project[Mom[[ mom[[i1]] ]], MomBasis[[j1]]],
                            {i1, 1, Length[mom]}] == 0], Length[MomBasis]];
        eql = Append[eql, Sum[coeff[i1], {i1, 1, Length[mom]}] == 1];
        coefflist = coefflist /. Solve[eql, Array[coeff, Length[mom]]];
        coefflist = coefflist /. coeff[_] -> 0;
        (*Set all the coefficients which remain after solving the
                    system of equations to zero, if there is no solution of
                    this system of equations it follows
                    coefflist = {coeff[1], coeff[2], coeff[3]} -> {0, 0, 0}.
                    As the linear equation system is inhomogeneous, coefflist
                    can only be substituted to {0, 0, 0} if there is no
                    solution:
                    The momenta are linearly independend or the condition
                    sum coeff[i] = 1 cannot be satisfied*)
        On[Solve::svars];
        If[(coefflist //. {0, r___} -> {r}) == {},
            (*first case : k^2 not reducible*)
            res={False};,
            (*else - part : k^2 reducible*)
            coefflist = coefflist[[1]];
            res={True,coefflist};
        ];(*endif*)
        Unprotect[ruleI];
```

```
            ruleI[mom]=res;
            (*remember previously calculated values*)
            Protect[ruleI];
            Throw[res];
            ] (*endMod*);
            ] (*endCatch*);
);
Protect[defruleI,ruleI];
Clear[ExternalMomenta];
ExternalMomenta[L
                            __]
                ] :=
    (*sets the variables Mom and LinIndepMom*)
    Module[{i,i1,j,j1,j2,k,eql,coeff,coefflist,
                L1,L2,L3,L4, cmom,base, solvelist},
    Unprotect[Mom, LinIndepMom];
    Clear [Mom,LinIndepMom];
    Mom = {L};
    (*The variable Mom has to be set. Otherwise you are not allowed
    to use the function FindBasis *)
    LinIndepMom = FindBasis[Range[Length[Mom]]];
    Protect[Mom,LinIndepMom];
    Unprotect[partfrac];
    Clear[partfrac];
    defpartfrac;
    (*define the function partfrac*)
    Protect[partfrac];
    Unprotect[GetBasis];
    Clear[GetBasis];
    defGetBasis;
    Protect[GetBasis];
    Unprotect[RepMom];
    Clear [RepMom];
    defRepMom;
    Protect[RepMom];
    Unprotect[ruleI];
    Clear[ruleI];
    defruleI;
Protect[ruleI];
```

];
Protect [ExternalMomenta];
Feynman integrals are represented by the function FInt. This function accesses the function scaleless which gives true if the topology is scaleless. In this case FInt gives 0. FInt uses the functions GetBasis, RepMom and ruleI to get simplified according to rule 1.3 .

```
(**********I Passarino Veltman reduction **************)
Unprotect[Scaleless];
Clear[Scaleless];
Scaleless[Denom_List]:=
Catch[
    Module[{pi,pj,i,j},
        For[i=1,i<=Length[Denom],i++,
            If[Denom[[i,2]]!=0,Throw[False]];
            pi=Mom[[ Denom[[i,1]] ]] - Mom[[ Denom[[1,1]] ]];
            For[j=2,j<=Length[Denom],j++,
            pj=Mom[[ Denom[[j]][[1]] ]] - Mom[[ Denom[[1]][[1]] ]];
            If[!MatchQ[Simplify[SP[pi,pj]],0],Throw[False];];
        ];
        ];
        Throw[True];
    ];
];
```

Scaleless[Denom1_List, Denom2_List]:=
Catch [
Module[\{p1, pi, pj, Mi, i, j\},
If [Length [Denom1] >0,
p1 = Mom[[ Denom1[[1,1]] ]];
For $[i=1, i<=$ Length [Denom1] , $i++$,
If [Denom1[[i, 2] ]!=0,Throw[False]];
pi=Mom[[ Denom1[[i,1]] ]] - p1;
For $[j=2, j<=$ Length [Denom1] , $j++$,
$\mathrm{pj}=\operatorname{Mom}[[\operatorname{Denom1[[j,1]]}]]-\mathrm{p} 1$;
If [!MatchQ[Simplify [SP[pi, pj]],0], Throw[False];];
];
For $[\mathrm{j}=1, \mathrm{j}<=$ Length [Denom2], $\mathrm{j}++$,
pj=Mom[[ Denom2[[j,1]] ]];
If [!MatchQ[Simplify [SP[pi, pj]],0], Throw[False];];
];
];
, (*else part *) $\mathrm{p} 1=0 ; \mathrm{]}$ (*endIf*);

```
    For[i=1,i<=Length[Denom2],i++,
    pi=Mom[[ Denom2[[i,1]] ]];
    If[!MatchQ[tres=Simplify[Denom2[[i,2]]-SP[pi,p1]],0],Throw[False]];
    For[j=1,j<=Length[Denom2],j++,
        pj=Mom[[ Denom2[[j,1]] ]];
        If[!MatchQ[Simplify[SP[pi,pj]],0],Throw[False];];
    ](*endFor[j]*);
](*endFor[i]*);
    Throw[True];
    ](*endMod*);
](*endCatch*);
Protect[Scaleless];
Unprotect [FInt];
Clear[FInt];
FInt [Denom_List,Num_List]:=0/;Scaleless [Denom];
FInt[Denom1_List,Denom2_List,Num_List]:=0/;Scaleless[Denom1,Denom2];
(*scaleless integrals vanish*)
FInt[Denom_List,{},Num_List] := FInt[Denom,Num];
FInt[Denom_List, {a___, {ip_Integer, 0}, b___}] := FInt[Denom, {a, b}];
FInt[Denom1_List, Denom2_List, {a___, {ip_Integer, 0}, b___}] :=
FInt[Denom1, Denom2, {a, b}];
FInt[{a___, {ip_Integer, Mp_, 0}, b___}, Num_List] := FInt[{a, b}, Num];
FInt[Denom_List, {a__-, {ip_Integer, Mp_, 0}, b___}, Num_List] :=
FInt[Denom, {a, b}, Num];
FInt[{a___, {ip_Integer, Mp_, 0}, b___}, Denom_List, Num_List] :=
FInt[{a, b}, Denom, Num];
FInt[{a1___,{i1p_Integer,Mp_,i2p_Integer},
    {i1p_Integer,Mp_,i3p_Integer},b1___},
    Num_List]:=
FInt[{a1,{i1p,Mp,i2p+i3p},b1},Num];
FInt[{a1___,{i1p_Integer,Mp_,i2p_Integer},
        {i1p_Integer,Mp_,i3p_Integer},b1___},
        Denom_List, Num_List]:=
```

FInt[\{a1,\{i1p,Mp,i2p+i3p\},b1\}, Denom, Num];
FInt [Denom_List, \{a1___, \{i1p_Integer, Mp_, i2p_Integer\},
\{i1p_Integer, Mp_,i3p_Integer\}, b1___\},
Num_List]:=
FInt [Denom, $\{\mathrm{a} 1,\{\mathrm{i} 1 \mathrm{p}, \mathrm{Mp}, \mathrm{i} 2 \mathrm{p}+\mathrm{i} 3 \mathrm{p}\}, \mathrm{b} 1\}, \mathrm{Num}] ;$
FInt [Denom_List,\{a1__, $\left\{i 1 p_{1}\right.$ Integer,i2p_Integer\}, \{i1p_Integer,i3p_Integer\}, b1___\}]:=
FInt [Denom, \{a1, \{i1p, i2p+i3p\}, b1\}];

FInt [Denom1_List, Denom2_List, \{a1___,\{i1p_Integer,i2p_Integer\}, \{i1p_Integer,i3p_Integer\},b1___\}]:=
FInt [Denom1, Denom2, \{a1,\{i1p,i2p+i3p\},b1\}];
FInt[\{a1_-_, \{i1p_Integer, M1p_, i2p_Integer\},
\{i3p_Integer, M2p_, i4p_Integer\}, b1___\},
Num_List]/; i1p > i3p :=
FInt[\{a1,\{i3p,M2p,i4p\},\{i1p,M1p,i2p\},b1\},Num];
FInt [Denom_List, \{a1___, \{i1p_Integer,M1p_,i2p_Integer\},
\{i3p_Integer, M2p_, i4p_Integer\}, b1___\},
Num_List]/; i1p > i3p :=
FInt [Denom, \{a1,\{i3p,M2p,i4p\},\{i1p,M1p,i2p\},b1\},Num];
FInt[\{a1_-_, \{i1p_Integer, M1p_, i2p_Integer\},
\{i3p_Integer, M2p_,i4p_Integer\}, b1___\},
Denom_List, Num_List]/; i1p > i3p :=
FInt[\{a1,\{i3p,M2p,i4p\},\{i1p,M1p,i2p\},b1\}, Denom, Num];
FInt [Denom_List,
\{a__-, \{i1p_Integer, n1p_Integer\},
\{i2p_Integer, n2p_Integer\}, b__\}] /; i1p > i2p :=
FInt [Denom, \{a, \{i2p, n2p\}, \{i1p, n1p\}, b\}];
FInt[Denom1_List, Denom2_List,
\{a $\qquad$ , \{i1p_Integer, n1p_Integer\},
\{i2p_Integer, n2p_Integer\}, b_-_\}] /; i1p > i2p :=
FInt [Denom1, Denom2, \{a, \{i2p, n2p\}, \{i1p, n1p\}, b\}];

```
FInt[{{p1_, M1_, m1_}, Denom___}, {a___,{0, n_Integer},b___}] :=
        (*The entry {0, n} in the numerator denotes (k^2)^n in the
        integrand where k is the integration variable*)
        FInt[{Denom}, {a, {0, n - 1}, b}] +
        (M1 - SP[ Mom[[p1]], Mom[[p1]] ])*
            FInt[{{p1, M1, m1}, Denom}, {a, {0, n - 1}, b}] -
        2*FInt[{{p1, M1, m1}, Denom}, {{p1, 1}, a, {0, n - 1}, b}];
FInt[{{p1_, M1_, m1_}, Denom1___}, Denom2_,
        {a___,{0, n_Integer},b___}] :=
        (*The entry {0, n} in the numerator denotes (k^2)^n in the
        integrand where k is the integration variable*)
    FInt[{Denom1}, Denom2, {a, {0, n - 1}, b}] +
    (M1 - SP[ Mom[[p1]], Mom[[p1]] ])*
        FInt[{{p1, M1, m1}, Denom1}, Denom2, {a, {0, n - 1}, b}] -
        2*FInt[{{p1, M1, m1}, Denom1}, Denom2, {{p1, 1}, a, {0, n - 1}, b}];
FInt[Denom_List, Num_List]:=
(*Try to expand the denomintor into partial fractions*)
Module[{pf,lth},
Sum[pf[[2,i]]*
FInt[ReplacePart[Denom, Denom[[i, 3]]-1,{i,3}],Num],{i,1,lth}]/;
(lth=Length[Denom];
pf=partfrac[Array[Denom[[#,1]]&,lth],Array[Denom[[#,2]]&,lth]];
pf[[1]])
];
(*reduction of the HQET-Propagators*)
FInt[Denom_List,{a___,{ip_Integer,Mp_,mp_Integer},b___},
    {a1___,{ip_Integer,sp_Integer},b1___}]:=
FInt[Denom,{a,{ip,Mp,mp-1},b},{a1,{ip,sp-1},b1}]-
Mp*FInt[Denom, {a,{ip,Mp,mp},b},{a1,{ip,sp-1},b1}];
    (*reduction corresponding ruleI*)
FInt[{a___, {pl_Integer, Ml_, ml_Integer}, b___},
    {c___, {pl_Integer, nl_Integer}, dp___}] :=
Module[{res,unchanged,coefflist, coeff, denom, num, p, M, i, j, l, rI},
    (*Rule I*)
```

```
    res/;
Catch[
    res=
    Catch[
        l = Length[{a}] + 1; (*Position of pl, Ml, ml;
                i.e. denom[[l, 1]] = pl etc. with denom see below*)
        denom = {a, {pl, Ml, ml}, b};
        num = {c, {pl, nl}, dp};
        p = Array[denom[[#, 1]] &, Length[denom]];
        M = Array[denom[[#, 2]] &, Length[denom]];
        rI=ruleI[p];
        If[ !rI[[1]] ,
            (*first case : k^2 not reducible*)
            Throw[unchanged];,
            (*else - part : k^2 reducible*)
            Throw[1/2*Sum[(KroneckerDelta[l, j1 ] - rI[[2,j1]])*
                    (FInt[
                        ReplacePart[denom, denom[[j1, 3]] - 1, {j1, 3}],
                            {c, {pl, nl - 1}, dp}]
                            + (M[[j1]] - SP[ Mom[[ p[[j1]] ]],
                                    Mom[[ p[[j1]] ]] ])*
                                    FInt[denom, {c, {pl, nl - 1}, dp}]),
                    {j1, 1, Length[p]}]];
            ];(*endif*)
    ](*endCatch*);
    If [MatchQ[res,unchanged],Throw [False]];
    Throw[True];
    ](*endCatch*)
] (*endMod*);
FInt[{a___, {pl_Integer, Ml_, ml_Integer}, b___},Denom_List,
    {c___, {pl_Integer, nl_Integer}, dp___}] :=
Module[{res,unchanged,coefflist, coeff, denom, num,
            p, M, i, j, l, rI},
(*Rule I*)
res/;
Catch[
    res=
    Catch[
        l = Length[{a}] + 1; (*Position of pl, Ml, ml;
        i.e. denom[[l, 1]] = pl etc. with denom see below*)
```

```
denom = {a, {pl, Ml, ml}, b};
num = {c, {pl, nl}, dp};
p = Array[denom[[#, 1]] &, Length[denom]];
M = Array[denom[[#, 2]] &, Length[denom]];
```

rI=ruleI[p];
If [ ! $\mathrm{II}[$ [1]],
(*first case : k^2 not reducible*)
Throw[unchanged];
(*else - part : k^2 reducible*)
Throw[1/2*Sum[(KroneckerDelta[l, j1 ] - rI[[2,j1]])*
(FInt [
ReplacePart[denom, denom[[j1, 3]] - 1, \{j1, 3\}],
Denom, $\{c,\{p l, n l-1\}, d p\}]$
+ (M[[j1]] - SP[ Mom[[ p[[j1]] ]],
Mom[[ p[[j1]] ] ] $]$ *
FInt[denom, Denom, $\{c,\{p l, n l-1\}, d p\}])$,
\{j1, 1, Length[p]\}]];
]; (*endif*)
] (*endCatch*) ;
If [MatchQ[res, unchanged], Throw [False]];
Throw [True];
] (*endCatch*)
] (*endMod*) ;
FInt [Denom1_List, Denom2_List,Num_List]:=
(*decompose the momenta of Num into a unique set of momenta
given by Denom2, Denom1 and some futher momenta which complete
the momenta of Denom2 and Denom1 to a basis*)
Module[\{res,unchanged, base,pn,rep, i,j \},
res/;
Catch [
If [MatchQ[Num, \{\}], Throw[False]];
res=Catch[
base $=$ GetBasis[Denom1,Denom2];
$\mathrm{pn}=$ Array[Num[[\#,1]]\&,Length[Num]];
For $[i=1, i<=$ Length[Num], i++,
If [MemberQ[ base, pn[[i]] ], Continue[] ];
(*The momentum in the numerator already appears in
the corresponding basis*)

```
            rep = RepMom[pn[[i]],base];
            Throw [Sum[
                    rep[[j]]*
            FInt[ Denom1, Denom2,
                                    Prepend[
                                    ReplacePart[Num, Num[[i, 2]] - 1, {i, 2}],
                                    {base[[j]],1}]],{j, 1, Length[base]}] ];
            ](*endFor[i]*);
            Throw[unchanged];
        ](*endCatch*);
        If [MatchQ[res,unchanged],Throw[False]];
        Throw[True];
](*endCatch*)
](*endMod*);
FInt[Denom_List,Num_List]:=
(*decompose the momenta of Num into a unique set of momenta
    given by Denom and some futher momenta which complete the
    momenta of Denom to a basis*)
Module[{res,unchanged, base,pn,rep, i,j },
    res/;
Catch[
    If [MatchQ[Num, {}],Throw[False]];
    res=Catch[
        base = GetBasis[Denom,{}];
        pn = Array[Num[[#,1]]&,Length[Num]];
            For[i = 1, i <= Length[Num], i++,
                    If [MemberQ[ base, pn[[i]] ], Continue[] ];
                    (*The momentum in the numerator already appears in
                    the corresponding basis*)
                    rep = RepMom[pn[[i]],base];
                    Throw [Sum[
                    rep[[j]]*
                    FInt[ Denom,
                                    Prepend[
                                    ReplacePart[Num, Num[[i, 2]] - 1, {i, 2}],
                                    {base[[j]],1}]],{j, 1, Length[base]}] ];
            ](*endFor[i]*);
            Throw[unchanged];
        ](*endCatch*);
```

If [MatchQ[res, unchanged], Throw [False]];
Throw[True];
] (*endCatch*)
] (*endMod*);
The implementation of (2.56) is given by these two rules:
(*these two rules should be applied last*)
FInt[\{a1__-, $\left\{\mathrm{p} 1_{-}\right.$Integer, M1_, m1_Integer\}, b1___,
\{p2_Integer, M2_, m2_Integer\}, c1___\},
\{d1__-, \{p2_Integer, n2_Integer\}, e1___\}] := Catch [

Module[\{coefflist, denom, i, j\}, denom $=\{a 1,\{p 1, \mathrm{M} 1, \mathrm{~m} 1\}, \mathrm{b} 1,\{\mathrm{p} 2, \mathrm{M} 2, \mathrm{~m} 2\}, \mathrm{c} 1\} ;$

Throw[1/2*(
FInt [\{a1, \{p1, M1, m1\}, b1, \{p2, M2, m2 - 1\}, c1\}, \{d1, \{p2, n2 - 1\}, e1\}] -
FInt [\{a1, \{p1, M1, m1-1\}, b1, \{p2, M2, m2\},c1\},
$\{d 1,\{p 2, n 2-1\}, e 1\}]$

+ (M2 - M1 + SP[ Mom[[p1 ]] , Mom[[ p1 ]] ] SP[ Mom[[p2]] , Mom[[ p2 ]] ])
*FInt[denom, \{d1, \{p2, n2 - 1\}, e1\}]
$+2 *$ FInt[denom, \{\{p1, 1\}, d1, \{p2, n2 - 1\}, e1\}])];
] (*endMod*);
] (*endCatch*)/;
MemberQ[GetBasis[\{a1,\{p1,M1,m1\}, b1, \{p2, M2,m2\}, c1\},\{\}], p1];
FInt[\{a1__, \{p1_Integer, M1_, m1_Integer\}, b1__,
\{p2_Integer, M2_, m2_Integer\}, c1___\}, Denom_,
\{d1___, \{p2_Integer, n2_Integer\}, e1___\}] :=
Catch [
Module[\{coefflist, denom, i, j\},

```
denom = {a1,{p1, M1, m1}, b1, {p2, M2, m2}, c1};
```

Throw[1/2*(

```
FInt[{a1,{p1, M1, m1}, b1, {p2, M2, m2 - 1}, c1},
                                    Denom, {d1, {p2, n2 - 1}, e1}] -
    FInt[{a1,{p1, M1, m1 - 1}, b1, {p2, M2, m2},c1},
                            Denom, {d1, {p2, n2 - 1}, e1}]
    + (M2 - M1 + SP[ Mom[[p1 ]] , Mom[[ p1 ]] ] -
        SP[ Mom[[p2]] , Mom[[ p2 ]] ])
        *FInt[denom, Denom, {d1, {p2, n2 - 1}, e1}]
        + 2*FInt[denom, Denom, {{p1, 1}, d1, {p2, n2 - 1},
                e1}]
    )
```

```
        ];
    ](*endMod*);
](*endCatch*)/;
MemberQ[GetBasis[{a1,{p1,M1,m1},b1,{p2,M2,m2},c1},Denom],p1];
```

The next step is to implement rule $4 \cdot 6$. We define the function IBP1 which takes as a starting point an integral as in (2.51) and is given the two arguments $\{\{\mathrm{i} 1, \mathrm{M} 1 \wedge 2, \mathrm{~m} 1\}, \ldots,\{\mathrm{it}, \mathrm{Mt} \wedge 2, \mathrm{mt}\}\}$ and $\{\{\mathrm{j} 1, \mathrm{~s} 1\}, \ldots,\{\mathrm{jl}, \mathrm{sl}\}\}$ like FInt. For the integral defined by these arguments the identities (2.59, 2.60) and (2.62) are generated and written into the variable eqlist.

```
Unprotect[IBP1];
Clear[IBP1];
IBP1[Denom_List, Num_List] := IBP1[Denom, {}, Num];
IBP1[Denom1_List, Denom2_List, Num_List] :=
    Module[{expr, s, t, l, p1, p2, M1, M2, m1, m2, r, n,
        i1, i2, i, j, idl},
    Unprotect[eqlist];
    p1 = Array[Denom1[[#, 1]] &, Length[Denom1]];
M1 = Array[Denom1[[#, 2]] &, Length[Denom1]];
m1 = Array[Denom1[[#, 3]] &, Length[Denom1]];
p2 = Array[Denom2[[#, 1]] &, Length[Denom2]];
M2 = Array[Denom2[[#, 2]] &, Length[Denom2]];
m2 = Array[Denom2[[#, 3]] &, Length[Denom2]];
n = Array[Num[[#, 2]] &, Length[Num]];
s = Sum[Num[[i, 2]], {i, 1, Length[Num]}];
r = Sum[Denom1[[i, 3]], {i, 1, Length[Denom1]}];
t = Length[Denom1];
l = Length[Num];
```

```
(* Identity I *)
expr = (Collect[(d + s - 2*r)*FInt[Denom1, Denom2, Num] -
    Sum[2 *m1[[i]]*(
    (M1[[i]] - SP[Mom[[p1[[i]]]], Mom[[p1[[i]]]]])*
                                    FInt[ReplacePart[Denom1,
                                    Denom1[[i, 3]] + 1, {i, 3}],
                            Denom2, Num]-
            FInt[ReplacePart[Denom1,
                        Denom1[[i, 3]] + 1, {i, 3}],
```

```
                    Denom2, Append[Num, {p1[[i]], 1}]]),
    {i, 1, t}]-
        Sum[m2[[i]]*
            FInt [Denom1,ReplacePart [Denom2,
                                    Denom2[[i,3]]+1,{i,3}],
                    Append[Num,{p2[[i]],1}]],
            {i,1,Length[Denom2]}],
        FInt[___]]
    );
eqlist = {expr};
(* Identity II *)
For[i1 = 1, i1 <= Length[LinIndepMom], i1++,
    i = LinIndepMom[[i1]];
    expr =
        (Collect[
            Sum[
                n[[j]]*SP[ Mom[[i]], Mom[[ Num[[j, 1]] ]] ]*
                FInt[Denom1, Denom2,
                    ReplacePart[Num, Num[[j, 2]] - 1, {j, 2}]],
            {j, 1, l}]-
        Sum[
                2*m1[[j]]*
                (SP[ Mom[[i]], Mom[[ p1[[j]] ]] ]*
                    FInt[ReplacePart[Denom1, Denom1[[j, 3]] + 1, {j, 3}],
                    Denom2, Num] +
                    FInt[ReplacePart[Denom1, Denom1[[j, 3]] + 1, {j, 3}],
                        Denom2, Append[Num, {i, 1}]]),
            {j, 1, t}]-
        Sum[m2[[j]]*SP[ Mom[[i]], Mom[[p2[[j]] ]] ]*
                    FInt[Denom1,
                                    ReplacePart[Denom2, Denom2[[j, 3]] + 1, {j, 3}],
                                    Num],
            {j, 1, Length[Denom2]}],
        FInt[___]]);
    eqlist = Append[eqlist, expr];
    ];(*endFor[i1]*)
(* Identity III *);
idl = MakeComb[Length[LinIndepMom], 2];
For[i = 1, i <= Length[idl], i++,
    i1 = LinIndepMom[[ idl[[i, 1]] ]];
    i2 = LinIndepMom[[ idl[[i, 2]] ]];
    expr = (Collect[
        Sum[2*n[[j]]*(SP[Mom[[Num[[j, 1]]]], Mom[[i2]]]*
```

```
    FInt[Denom1, Denom2,
    Append[
        ReplacePart[Num,
            Num[[j, 2]] - 1, {j, 2}],
            {i1, 1}]] -
SP[Mom[[Num[[j, 1]]]], Mom[[i1]]]*
FInt [Denom1, Denom2,
    Append[
                            ReplacePart[Num,
                            Num[[j, 2]] - 1, {j, 2}],
                            {i2, 1}]]),
            {j, 1, Length[Num]}]-
        Sum[4*m1[[j]]*(
                            SP[Mom[[Denom1[[j, 1]]]], Mom[[i2]]]*
                    FInt[ReplacePart[Denom1,
                            Denom1[[j, 3]] + 1, {j, 3}],
                            Denom2, Append[Num, {i1, 1}]] -
                            SP[Mom[[Denom1[[j, 1]]]], Mom[[i1]]]*
                    FInt[ReplacePart[Denom1,
                            Denom1[[j, 3]] + 1, {j, 3}],
                            Denom2, Append[Num, {i2, 1}]]),
            {j, 1, Length[Denom1]}] -
        Sum[2*m2[[j]]*(
        SP[Mom[[Denom2[[j, 1]]]], Mom[[i2]]]*
        FInt [Denom1,
            ReplacePart[Denom2,
                    Denom2[[j, 3]] + 1, {j, 3}],
            Append[Num, {i1, 1}]] -
                SP[Mom[[Denom2[[j, 1]]]], Mom[[i1]]]*
                    FInt [Denom1,
                            ReplacePart[Denom2,
                            Denom2[[j, 3]] + 1, {j, 3}],
                    Append[Num, {i2, 1}]]),
            {j, 1, Length[Denom2]}],
        FInt[___]]
        );
    eqlist = Append[eqlist, expr];
    ](*endFor[i]*);
Protect[eqlist];
] (*endModule*);
```

Protect[IBP1];
The function IBP follows the algorithm of 22 to generate a list of the IBP identities. After generation a new set of identities by calling IBP1 it replaces more complex integrals by less complex ones. The complexity of integrals is defined in [22].

```
Unprotect [Hf1];
Clear[Hf1];
Hf1[1, Mp_] := {{Mp}};
Hf1[1_, O] := {Array[0 &, l]};
Hf1[1_, Mp_] :=
    (*generate all l - tuples {n1, ..., nl} s.t. n1 + ... + nl = Mp *)
```

        Catch [
        Module[\{L1, L2, i, j\},
            L2 = \{\};
            For [i = 0, i <= Mp, i++,
                L1 = Hf1[1 - 1, Mp - i];
                For \([j=1, j<=\) Length[L1], j++,
                    L2 = Append[L2, Join[\{i\}, L1[[j]]]];
                    ];
                ];
            Throw [L2];
            ] (*endMod*);
        ](*endCatch*);
    Protect [Hf1] ;
Unprotect [BT];
Clear [BT];
BT[L1_List, L2_List] :=
(*give an ordering to lists of integers*)
Catch [
Module[\{i\},
For $[\mathrm{i}=1, \mathrm{i}<=\operatorname{Min}[L e n g t h[L 1]$, Length[L2]], i++,
If [L1[[i]] > L2[[i]], Throw[True]];
If [L1[[i]] < L2[[i]], Throw[False]];
];
If [Length[L1] > Length[L2], Throw[True]];
Throw[False];
];
];
Protect [BT];

Unprotect [Verbose];
Clear [Verbose];
Unprotect[verboseflag];
verboseflag=False;
Verbose[flag_]:=
(
Unprotect[verboseflag];
If [MatchQ[flag,on], verboseflag=True; ];
If [MatchQ[flag,off], verboseflag=False;];
Protect[verboseflag]
);
Protect[verboseflag];

Protect [Verbose];
Unprotect [IBP];
Clear [IBP];
IBP[Denom_List, la_List, lb_List] := IBP [Denom, \{\}, la, \{\}, lb];

IBP [Denom1_List, Denom2_List, la1_List, la2_List, lb_List] := Catch [

Module[\{dummy, i1, i2, j, k, l, m1, m2, n1, n2,
Mp, Mp2, Md1, Md2, si, subslist,
cl1, cl2, hl, delist1, delist2,
momlist, num, denom1, denom2, maxf, subsrule, a1, a2, b\},

For $[i=0$, $i<=$ Length[Denom1], i++, a1[i] = 0;
]; (*Default value for a1*)
For[i = 0, i <= Length[Denom1]+Length[Denom2], i++, b[i] = 0;
]; (*Default value for b*)
For $[\mathrm{i}=0$, $\mathrm{i}<=$ Length[Denom2], $\mathrm{i}++$, a2[i] = -1;
];

```
a2[0] = 0;
(*Default value for a2*)
For[i = 1, i <= Length[la1], i++,
    a1[la1[[i, 1]]] = la1[[i, 2]];
    ];
For[i = 1, i <= Length[la2], i++,
    a2[la2[[i, 1]]] = la2[[i, 2]];
    ];
For[i = 1, i <= Length[lb], i++,
    b[lb[[i, 1]]] = lb[[i, 2]];
    ];
subslist = {}; (*FInt[ ...] -> ...*);
For[n1 = 0, n1 <= Length[Denom1], n1++,
    cl1 = MakeComb[Length[Denom1], n1];
    For[n2 = 0, n2 <= Length[Denom2], n2++,
    cl2 = MakeComb[Length[Denom2], n2];
    For[i1 = 1, i1 <= Length[cl1], i1++,
        For[i2 = 1, i2 <= Length[cl2], i2++,
            For [Mp = 0, Mp <= b[n1+n2], Mp++,
            momlist = Hf1[Length[Mom], Mp];
                For[l = 1, l <= Length[momlist], l++,
                    num = Array[{#, momlist[[l,#]]} &, Length[momlist[[l]] ] ];
                    num = num//.{r1___,{i1_Integer,0},r2___} -> {r1,r2};
                    For[Md1 = 0, Md1 <= a1[n1], Md1++,
                        delist1 = Hf1[n1, Md1];
                        For[Md2 = 0, Md2 <= a2[n2], Md2++,
                        delist2 = Hf1[n2, Md2];
                        For[m1 = 1, m1 <= Length[delist1], m1++,
                    denom1 =
                                    Array[Join[Denom1[[ cl1[[i1,#]] ]],
                                    {1+delist1[[m1, #]]}] &, n1];
                For[m2 = 1, m2 <= Length[delist2], m2++,
                    denom2 =
                                    Array[Join[Denom2[[ cl2[[i2,#]] ]],
```

\{1+delist2[[m2, \#]]\}] \&, n2];
(*Step 8*)
If [(MatchQ[denom2, \{\}]\&\&
MatchQ[FInt[denom1, num] / .FInt->FIntin, FIntin[denom1, num]])||
MatchQ[FInt[denom1, denom2, num]/.FInt->FIntin, FIntin[denom1, denom2, num] ],
If [verboseflag, Print[denom1, denom2, num]];
IBP1[denom1, denom2, num],
Continue[] ];
(*Create IBP-identities from topologies, which cannot be reduced by passarino veltman. *)

Unprotect [eqlist] ;
(*Step 9(a)*)

```
For[j = 1, j <= Length[eqlist], j++,
    eqlist[[j]] = eqlist[[j]]//.subslist;
    eqlist[[j]] = Collect[eqlist[[j]],
                                    HoldPattern[FInt[___]],
                                    Expand[ Together[#] ]&];
    (*substitude all the known identities into the new
    IBP identities *)
        hl = {};
        eqlist[[j]] //. {FInt[expr___] :> (
        hl = Append[hl, FInt[expr]];
        dummy)};
        (*extract all the Feynmanintegrals from eqlist[[j]]
            and write them into hl *)
        If [MatchQ[hl, {}], Continue []];
            (*Step 9(b) *)
            hl = hl /. {FInt[De1_, De2_, Nu_] :>
        FIntin[De1, De2, Nu,
            Join[{Length[De1]+Length[De2]},
                            {Sum[ De1[[k,3]], {k,1,Length[De1]}]+
                            Sum[ De2[[k,3]], {k,1,Length[De2]}]},
                {Sum[ De2[[k,3]], {k,1,Length[De2]}]},
                {Sum[ Nu[[k, 2]], {k,1,Length[Nu]}]},
                Array[De1[[#,1]]&,Length[De1]],
                Array[De1[[#,3]]&, Length[De1]],
                Array[De2[[#,1]]&,Length[De2]],
                Array[De2[[#,3]]&, Length[De2]],
```

Array [Nu[[\#,2]]\&,Length[Nu]]] ]\};
hl = hl /. \{FInt[De_, Nu_] :> FIntin[De, Nu, Join[\{Length[De]\},
\{Sum $[\operatorname{De}[\mathrm{k}, 3]],\{k, 1$, Length[De] $\}]\}$, \{0\},
\{Sum [ $\operatorname{Nu}[\mathrm{[k}, 2]],\{k, 1$, Length[Nu] $\}]\}$, Array [De[[\#, 1] ]\&, Length[De]], Array[De[[\#, 3]]\&, Length[De]],
Array [Nu[[\#,2]]\&,Length[Nu]]] ]\};
(*Give to the Feynman integrals a specific weight which measures the complexity of the integral*)
maxf $=1$; (*hl[[maxf]] will be the Feynmanintegral with the highest complexity*)
For [ k=2, k<=Length[hl], k++, If [BT [hl[[k]]/.FIntin[___, arg_]->arg, hl[[maxf]]/.FIntin[___, arg_]->arg], $\operatorname{maxf}=\mathrm{k}](* e n d I f *)$;
] (*endFor [k]*) ;
hl[[maxf]] = hl[[maxf]]/.FIntin[De__-,_]:> FInt [De];
(*Step 9(c) *)
If [Coefficient[eqlist[[j]],hl[[maxf]]]==0, Throw[\{eqlist[[j]],hl[[maxf]]\}]](*endif*);

```
subsrule = {hl[[maxf]] ->
```

                Collect[
                        (-eqlist[[j]]/.hl[[maxf]]->0)/
    Coefficient[eqlist[[j]],hl[[maxf]]],
HoldPattern[FInt [___]], Expand[Together] ]\};
subslist = Join[subsrule,subslist];
] (*endFor [j]*);
] (*endFor [m2] *);
](*endFor [m1]*);
] (*endFor [Md2]*);
] (*endFor [Md1]*);
] (*endFor [l] *);
](*endFor [Mp]*);

```
            ](*endFor[i2]*);
            ](*endFor[i1]*);
        ](*endFor[n2]*);
        ](*endFor[n1]*);
        Throw[subslist];
        ](*endMod*);
    ](*endCatch*);
Protect[IBP];
```


## Appendix B

## Master integrals

## B. 1 Integrals with up to three external lines

In this section I give explicit expression for the one-, two- and three-point master integrals, which occur in my calculations. They are calculated in $d=4-2 \epsilon$ dimensions.

There remains only one nonzero one-point integral:

$$
\begin{equation*}
A_{0} \equiv \mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}-m^{2}}=\frac{i}{(4 \pi)^{2}} \Gamma(1+\epsilon)\left(\frac{4 \pi \mu^{2}}{m^{2}}\right)^{\epsilon} m^{2}\left(\frac{1}{\epsilon}+1+\epsilon+\mathcal{O}\left(\epsilon^{2}\right)\right) \tag{B.1}
\end{equation*}
$$

The two-point integrals are:

$$
\begin{align*}
B_{s 1}(x, y) \equiv & \mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+x p+y q)^{2}} \\
= & \frac{i}{(4 \pi)^{2}} \Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}\left[\frac{1}{\epsilon}+2-\ln x-\ln y+i \pi+\right. \\
& \epsilon\left(4-\frac{2 \pi^{2}}{3}-2 \ln x-2 \ln y+\frac{1}{2} \ln ^{2} x+\frac{1}{2} \ln ^{2} y+\ln x \ln y\right. \\
& \left.+i \pi(2-\ln x-\ln y))+\mathcal{O}\left(\epsilon^{2}\right)\right] \tag{B.2}
\end{align*}
$$

$$
\begin{align*}
B_{s 1}(x, y, \xi, \theta) \equiv & \mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+x p+y q-l)^{2}} \\
= & \frac{i}{(4 \pi)^{2}} \Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}\left[\frac{1}{\epsilon}+2-\ln (x y-x \xi-y \theta)+i \pi+\right. \\
& \epsilon\left(4-\frac{2 \pi^{2}}{3}+\frac{1}{2} \ln ^{2}(x y-x \xi-y \theta)-2 \ln (x y-x \xi-y \theta)\right. \\
& \left.+i \pi(2-\ln (x y-x \xi-y \theta)))+\mathcal{O}\left(\epsilon^{2}\right)\right] \tag{B.3}
\end{align*}
$$

$$
\begin{align*}
& B_{s 2}(x, y) \equiv \mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+x p-y q)^{2}} \\
& =\frac{i}{(4 \pi)^{2}} \Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}\left[\frac{1}{\epsilon}+2-\ln x-\ln y+\right. \\
& \epsilon\left(4-\frac{\pi^{2}}{6}-2 \ln x-2 \ln y+\frac{1}{2} \ln ^{2} x+\frac{1}{2} \ln ^{2} y+\ln x \ln y\right) \\
& \left.+\mathcal{O}\left(\epsilon^{2}\right)\right]  \tag{B.4}\\
& B_{m}(x, y) \equiv \mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{\left(k^{2}+2 k \cdot(p+q)\right)(k+x p+y q)^{2}} \\
& =\frac{i}{(4 \pi)^{2}} \Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}\left[\frac{1}{\epsilon}+2+\frac{(x+y-x y) \ln (x+y-x y)}{\bar{x} \bar{y}}+\right. \\
& \epsilon\left(4+2 \frac{(x+y-x y) \ln (x+y-x y)}{\bar{x} \bar{y}}-\frac{(x+y-x y) \ln ^{2}(x+y-x y)}{\bar{x} \bar{y}}\right. \\
& \left.\left.-\frac{(x+y-x y) \operatorname{Li}_{2}(\bar{x} \bar{y})}{\bar{x} \bar{y}}\right)+\mathcal{O}\left(\epsilon^{2}\right)\right]  \tag{B.5}\\
& B_{m}(x, y, \xi, \theta) \equiv \mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{\left(k^{2}+2 k \cdot(p+q-l)\right)(k+x p+y q-l)^{2}} \\
& =\frac{i}{(4 \pi)^{2}} \Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}\left[\frac{1}{\epsilon}+2-\frac{(1-\xi-\theta) \ln (1-\xi-\theta)}{\bar{x} \bar{y}}\right. \\
& +\frac{(x+y-x y-\xi-\theta) \ln (x+y-x y-\xi-\theta)}{\bar{x} \bar{y}} \\
& +\epsilon\left(4-2 \frac{1-\xi-\theta}{\bar{x} \bar{y}} \ln (1-\xi-\theta)+\frac{1-\xi-\theta}{2 \bar{x} \bar{y}} \ln ^{2}(1-\xi-\theta)\right. \\
& +2 \frac{x+y-x y-\xi-\theta}{\bar{x} \bar{y}} \ln (x+y-x y-\xi-\theta) \\
& -\frac{x+y-x y-\xi-\theta}{2 \bar{x} \bar{y}} \ln ^{2}(x+y-x y-\xi-\theta) \\
& \left.+\frac{x+y-x y-\xi-\theta}{\bar{x} \bar{y}} \mathrm{Li}_{2} \frac{\bar{x} \bar{y}}{-x-y+x y-\xi-\theta}\right) \\
& \left.+\mathcal{O}\left(\epsilon^{2}\right)\right] \tag{B.6}
\end{align*}
$$

Note that $(\overline{B .2})$ and $(\overline{B .5})$ are the leading power of $(\overline{B .3})$ and ( $\overline{\text { B.6 }})$ resp.
The three-point integrals are:

$$
\begin{aligned}
C_{1}(x, y) & \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+x p+y q)^{2}\left(k^{2}+2 k \cdot(p+q)\right)} \\
& =-\frac{i}{(4 \pi)^{2}} \frac{1}{x-y}\left[\operatorname{Li}_{2} \frac{y(x-1)}{x}-\operatorname{Li}_{2} \frac{x(y-1)}{y}+\mathrm{Li}_{2}(\bar{x})-\operatorname{Li}_{2}(\bar{y})+\right.
\end{aligned}
$$

$$
\begin{align*}
& i \pi(\ln x-\ln y)+\mathcal{O}(\epsilon)] \\
& C_{2}(x, \xi) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+x p-\xi q)^{2}\left(k^{2}+2 k \cdot(p+q)\right)} \\
& =-\frac{i}{(4 \pi)^{2}} \frac{1}{x}\left[-\operatorname{Li}_{2}(x)+\frac{2 \pi^{2}}{3}+\frac{1}{2} \ln ^{2} x-\ln x \ln \bar{x}+\frac{1}{2} \ln ^{2} \xi-\ln \xi \ln x+\right. \\
& \mathcal{O}(\xi)+\mathcal{O}(\epsilon)] \\
& C_{3}(x, y, \theta) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+\theta p+y q)^{2}(k+x p+q)^{2}} \\
& =\frac{i}{(4 \pi)^{2}} \frac{1}{x y}\left(\ln \theta \ln \bar{y}+\ln y \ln \bar{y}-\ln x \ln \bar{y}+2 \mathrm{Li}_{2} y\right) \\
& +\mathcal{O}(\theta)+\mathcal{O}(\epsilon)  \tag{B.9}\\
& C_{4}(x, y, \xi) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{(k-x p)^{2}(k-p+\xi q)^{2}(k+y q)^{2}} \\
& =\frac{i}{(4 \pi)^{2}} \frac{1}{\bar{x} y}\left(\ln x \ln \bar{x}+\ln \xi \ln x-\ln x \ln y+2 \operatorname{Li}_{2}(\bar{x})+i \pi \ln x\right) \\
& +\mathcal{O}(\xi)+\mathcal{O}(\epsilon)  \tag{B.10}\\
& C_{5}(x, y, \theta) \equiv \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{(k+x p)^{2}(k+(x-\theta) p+y q)^{2}\left(k^{2}+2 k \cdot((x-\theta) p+q)\right)} \\
& =\frac{i}{(4 \pi)^{2}} \frac{1}{\bar{x} y}\left[-\frac{\pi^{2}}{6}+\ln ^{2} x+\ln y \ln \bar{y}-\frac{\ln ^{2} y}{2}-\ln x \ln (x+y-x y)\right. \\
& +\ln y \ln (x+y-x y)-\frac{\ln ^{2}(x+y-x y)}{2}-\ln x \ln \theta \\
& \left.+\ln \theta \ln (x+y-x y)-\mathrm{Li}_{2} \frac{x(y-1)}{y}+\mathrm{Li}_{2}(y)+\operatorname{Li}_{2} \frac{y(x-1)}{x}\right] \\
& +\mathcal{O}(\theta)+\mathcal{O}(\epsilon) \tag{B.11}
\end{align*}
$$

## B. 2 Massive four-point integral

We consider the following massive four-point integral in $d=4-2 \epsilon$ dimensions (fig. B.1):

$$
\begin{equation*}
I_{4}\left(p_{1}, p_{2}, p_{3}, p_{4}\right)=\mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{D_{1} D_{2} D_{3} D_{4}} \tag{B.12}
\end{equation*}
$$



Figure B.1: Basic one-loop four-point intergral. The massive line, which carries the mass $m$, is indicated by the thick line.
where

$$
\begin{align*}
& D_{1}=k^{2}+i \eta \\
& D_{2}=\left(k+p_{1}\right)^{2}+i \eta \\
& D_{3}=\left(k+p_{1}+p_{2}\right)^{2}+i \eta \\
& D_{4}=\left(k+p_{1}+p_{2}+p_{3}+p_{4}\right)^{2}-m^{2}+i \eta \tag{B.13}
\end{align*}
$$

Following [31] we introduce the external masses

$$
\begin{equation*}
p_{i}^{2}=m_{i}^{2} \quad(i=1,2,3,4) \tag{B.14}
\end{equation*}
$$

and the Mandelstam variables

$$
\begin{equation*}
s=\left(p_{1}+p_{2}\right)^{2}, \quad t=\left(p_{2}+p_{3}\right)^{2} . \tag{B.15}
\end{equation*}
$$

Furthermore we consider only the case, where

$$
\begin{equation*}
m_{2}^{2}=0 \quad \text { and } \quad m_{4}^{2}=m^{2} \tag{B.16}
\end{equation*}
$$

The integral (B.12) can be evaluated using the method of 31. This paper gives explicit expressions for massless one-loop box integrals. It is however possible to extend the single steps of this paper to our case.

So finally we obtain:

$$
\begin{align*}
& I_{4}\left(p_{1}, p_{2}, p_{3}, p_{4}\right) \equiv I_{4}\left(s, t, m_{1}^{2}, m_{3}^{2}, m^{2}\right)=\frac{i}{(4 \pi)^{2}} \frac{\Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}}{m^{2}\left(s-m_{1}^{2}\right)-s t+m_{1}^{2} m_{3}^{2}} \\
& \times\left[\begin{array}{l}
\frac{1}{\epsilon}\left(\ln (-s-i \eta)+\ln \left(m^{2}-t-i \eta\right)-\ln \left(m^{2}-m_{3}^{2}-i \eta\right)-\ln \left(-m_{1}^{2}-i \eta\right)\right) \\
\quad+\ln ^{2}\left(m^{2}-m_{3}^{2}-i \eta\right)+\ln ^{2}\left(-m_{1}^{2}-i \eta\right)-\ln ^{2}(-s-i \eta)-\ln ^{2}\left(m^{2}-t-i \eta\right) \\
\quad+\ln \left(m^{2}-i \eta\right)\left(\ln (-s-i \eta)+\ln \left(m^{2}-t-i \eta\right)-\ln \left(m^{2}-m_{3}^{2}-i \eta\right)-\ln \left(-m_{1}^{2}-i \eta\right)\right) \\
\quad+2 \operatorname{Li}_{2}\left(1-\frac{m^{2}-t-i \eta}{-m_{1}^{2}-i \eta}\right)-2 \operatorname{Li}_{2}\left(1-\frac{m^{2}-m_{3}^{2}-i \eta}{-s-i \eta}\right) \\
\quad+2 \operatorname{Li}_{2}\left(1-\left(m_{3}^{2}-m^{2}+i \eta\right) f^{m}\right)+2 \operatorname{Li}_{2}\left(1-\left(m_{1}^{2}+i \eta\right) f^{m}\right) \\
\left.\quad-2 \operatorname{Li}_{2}\left(1-\left(t-m^{2}+i \eta\right) f^{m}\right)-2 \operatorname{Li}_{2}\left(1-(s+i \eta) f^{m}\right)\right]
\end{array}\right.
\end{align*}
$$

where $f^{m}=\frac{s+t-m_{1}^{2}-m_{3}^{2}}{m^{2}\left(m_{1}^{2}-s\right)+s t-m_{1}^{2} m_{3}^{2}}$.
The case $m_{1}^{2}=0$ gives rise to further divergences and has to be considered separately:

$$
\begin{align*}
& I_{4}\left(s, t, m_{1}^{2}=0, m_{3}^{2}, m^{2}\right)=\frac{i}{(4 \pi)^{2}} \frac{\Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}}{s\left(m^{2}-t\right)} \\
& \times\left[-\frac{3}{2 \epsilon^{2}}+\frac{1}{\epsilon}\left(2 \ln \left(m^{2}-t-i \eta\right)-\frac{1}{2} \ln \left(m^{2}-i \eta\right)+\ln (-s-i \eta)-\ln \left(m^{2}-m_{3}^{2}-i \eta\right)\right)\right. \\
& \quad+\frac{2 \pi^{2}}{3}+\frac{1}{4} \ln ^{2}\left(m^{2}-i \eta\right)-\ln ^{2}\left(m^{2}-t-i \eta\right)+\ln ^{2}\left(m^{2}-m_{3}^{2}-i \eta\right)-\ln ^{2}(-s-i \eta) \\
& \quad+\ln \left(m^{2}-i \eta\right)\left(\ln (-s-i \eta)-\ln \left(m^{2}-m_{3}^{2}-i \eta\right)\right) \\
& \\
& \quad-2 \operatorname{Li}_{2}\left(1-\frac{m^{2}-m_{3}^{2}-i \eta}{-s-i \eta}\right)+2 \operatorname{Li}_{2}\left(1-\left(m_{3}^{2}-m^{2}+i \eta\right) f^{m}\right)  \tag{B.18}\\
& \\
& \left.\quad-2 \operatorname{Li}_{2}\left(1-\left(t-m^{2}+i \eta\right) f^{m}\right)-2 \operatorname{Li}_{2}\left(1-(s+i \eta) f^{m}\right)\right]
\end{align*}
$$

where $f^{m}=\frac{s+t-m_{3}^{2}}{s\left(t-m^{2}\right)}$.

## B. 3 Massless five-point integral

We consider the following massless five-point integral:

$$
\begin{equation*}
E_{0}=\mu^{2 \epsilon} \int \frac{d^{d} k}{(2 \pi)^{d}} \frac{1}{k^{2}(k+\bar{x} p)^{2}(k+p)^{2}(k+l)^{2}(k+l-y q)^{2}} . \tag{B.19}
\end{equation*}
$$

Despite the fact that there occur only three linearly independent momenta in $E_{0}$, (B.19) cannot be decomposed into partial fractions. However using a standard Feynman parametrisation we end up with integrals which can be calculated by computer algebra systems. Now the exact result is rather involved so I give just the leading power. Higher powers can be obtained by using the methods of section 2.4.2.

$$
\begin{align*}
E_{0}=\frac{i}{(4 \pi)^{2}} \frac{\Gamma(1+\epsilon)\left(4 \pi \mu^{2}\right)^{\epsilon}}{\bar{x} y \theta \xi^{2}}[ & -\frac{2}{\epsilon^{2}}+\frac{-4+2 \ln \xi+2 \ln \theta+2 i \pi}{\epsilon}-\ln ^{2} \xi-\ln ^{2} \theta \\
& -2 \ln \xi \ln \theta+4 \ln \xi+4 \ln \theta+\pi^{2} \\
& +2 i \pi(2-2 \ln \xi-2 \ln \theta)] \tag{B.20}
\end{align*}
$$

## Appendix C

## Matching of $\lambda_{B}$

In this thesis we consider a pure QCD calculation. This calculation is given in terms of $f_{B} / \lambda_{B}$. In order to compare our calculation to HQET or SCET results we have to match the expression $f_{B} / \lambda_{B}$ onto HQET. We use the definition of the $B$-meson wave function of [35], which is defined in the case of HQET by the heavy quark field of $h_{v}$. So we define

$$
\begin{equation*}
i f_{B} m_{B} \phi_{+}^{\mathrm{QCD}}(\omega) \equiv \frac{1}{2 \pi} \int_{0}^{\infty} d t e^{i \omega t}\langle 0| \bar{q}(z)[\ldots] \not \varkappa_{+} \gamma_{5} b(0)|\bar{B}\rangle_{z^{-}, z_{\perp}=0} \tag{C.1}
\end{equation*}
$$

in the case of QCD and

$$
\begin{equation*}
i \hat{f}_{B} m_{B} \phi_{+}^{\mathrm{HQET}}(\omega) \equiv \frac{1}{2 \pi} \int_{0}^{\infty} d t e^{i \omega t}\langle 0| \bar{q}(z)[\ldots] \not h_{+} \gamma_{5} h_{v}(0)|\bar{B}\rangle_{z^{-}, z_{\perp}=0} \tag{C.2}
\end{equation*}
$$

in the case of HQET. Here the integration goes over $t=v \cdot z$ where $v$ is the fourvelocity of the $B$-meson. We define the $B$-meson decay constant by

$$
\begin{equation*}
i f_{B} m_{B}=\langle 0| \bar{q}(0)[\ldots] \not \gamma_{5} b(0)|\bar{B}\rangle \tag{C.3}
\end{equation*}
$$

and analogously the HQET decay constant, which depends on the renormalisation scale $\mu$, by

$$
\begin{equation*}
i \hat{f}_{B}(\mu) m_{B}=\langle 0| \bar{q}(0)[\ldots] \not \ldots \gamma_{5} h_{v}(0)|\bar{B}\rangle . \tag{C.4}
\end{equation*}
$$

The matching coefficient is defined by

$$
\begin{equation*}
f_{B} \int_{0}^{\infty} d \omega \frac{\phi_{+}^{\mathrm{QCD}}(\omega)}{\omega} \equiv C_{\lambda_{B}} \hat{f}_{B} \int_{0}^{\infty} d \omega \frac{\phi_{+}^{\mathrm{HQET}}(\omega)}{\omega} \tag{C.5}
\end{equation*}
$$

such that we get writing the $\mu$-dependence explicitly

$$
\begin{equation*}
\frac{f_{B}}{\lambda_{B}^{\mathrm{QCD}}(\mu)}=C_{\lambda_{B}}(\mu) \frac{\hat{f}_{B}(\mu)}{\lambda_{B}^{\mathrm{HQET}}(\mu)} . \tag{C.6}
\end{equation*}
$$

We get $C_{\lambda_{B}}$ to $\mathcal{O}\left(\alpha_{s}\right)$ by calculating the convolution integrals occurring in C.5 in both QCD and HQET up to $\mathcal{O}\left(\alpha_{s}\right)$. The corresponding diagrams are shown in fig. C.1. As in section 3.3 we can use the wave functions (C.1), (C.2) defined by free
$\underbrace{}_{b}$
(a)

(b)

(c)

Figure C.1: NLO contributions to $\lambda_{B}$. The double line stands for the $b$-quark field.
quark states instead of hadronic states. We assign to the $b$-quark the momentum $v\left(m_{b}-\tilde{\omega}\right)(-v \tilde{\omega}$ resp.) in the case of pure QCD (HQET resp.) and $v \tilde{\omega}$ to the soft constituent quark, where $v$ is the four velocity of the $B$-meson. At tree level we get for both QCD and HQET the same wave function:

$$
\begin{equation*}
i f_{B} \phi_{+}^{(0)}(\omega)=N_{c} \delta(\omega-\tilde{\omega}) \bar{q} \not n_{+} \gamma_{5} \Psi \tag{C.7}
\end{equation*}
$$

where the spinor $\Psi$ fulfils the condition $\not \wp \Psi=\Psi$, and our convolution integral is

$$
\begin{equation*}
i f_{B} \int_{0}^{\infty} d \omega \frac{\phi_{+}^{(0)}(\omega)}{\omega}=\frac{1}{N_{c} \tilde{\omega}} \bar{q} \not n_{+} \gamma_{5} \Psi \tag{C.8}
\end{equation*}
$$

At NLO only the first diagram in fig. C. 1 needs to be considered as the other two are in leading power identical for QCD and HQET. The following expressions are given in the MS scheme, i.e. we redefine $\mu^{2} \rightarrow \mu^{2} \frac{e^{\gamma} \mathrm{E}}{4 \pi}$. For the diagram in fig. C.1(a) we get in QCD

$$
\begin{equation*}
\frac{\alpha_{s}}{4 \pi} C_{F} N_{c} \frac{1}{\tilde{\omega}} \bar{q} n_{+} \gamma_{5} \Psi\left(\frac{2+2 \ln \frac{\tilde{\omega}}{m_{b}}}{\epsilon}+4 \ln \frac{\mu}{m_{b}}+4-\frac{\pi^{2}}{6}-2 \ln ^{2} \frac{\tilde{\omega}}{m_{b}}+4 \ln \frac{\tilde{\omega}}{m_{b}} \ln \frac{\mu}{m_{b}}\right) \tag{C.9}
\end{equation*}
$$

and in HQET

$$
\begin{equation*}
\frac{\alpha_{s}}{4 \pi} C_{F} N_{c} \frac{1}{\tilde{\omega}} \bar{q} \not \varkappa_{+} \gamma_{5} \Psi\left(-\frac{1}{\epsilon^{2}}+\frac{2 \ln \frac{\tilde{\omega}}{\mu}}{\epsilon}-2 \ln ^{2} \frac{\tilde{\omega}}{\mu}-\frac{\pi^{2}}{4}\right) . \tag{C.10}
\end{equation*}
$$

The wave function renormalisation constants of the heavy quark field are given in the onshell scheme for the QCD $b$-field:

$$
\begin{equation*}
Z_{2 b}^{\frac{1}{2}}=1+\frac{\alpha_{s}}{4 \pi} C_{F}\left(-\frac{1}{2 \epsilon}-\frac{1}{\epsilon_{\mathrm{IR}}}-3 \ln \frac{\mu}{m_{b}}-2\right) \tag{C.11}
\end{equation*}
$$

and for the HQET field $h_{v}$ :

$$
\begin{equation*}
Z_{2 h_{v}}^{\frac{1}{2}}=1+\frac{\alpha_{s}}{4 \pi} C_{F}\left(\frac{1}{\epsilon}-\frac{1}{\epsilon_{\mathrm{IR}}}\right) \tag{C.12}
\end{equation*}
$$

The renormalisation of the $q$-field drops out in the matching. Diagrammatically the matching equation (C.5) reads:

$$
\begin{equation*}
\left.\left.Z_{2 b}^{\frac{1}{2}}\left(\llbracket^{---}\right\rceil+\rrbracket^{-\vartheta^{-}}\right\rceil\right)^{\mathrm{QCD}}=C_{\lambda_{B}} Z_{2 h_{v}}^{\frac{1}{2}}(\llbracket^{\overbrace{}^{---}\rceil+\overbrace{}^{-\vartheta^{-}}\rceil)^{\mathrm{HQET}} .} . \tag{C.13}
\end{equation*}
$$

Finally we obtain

$$
\begin{equation*}
C_{\lambda_{B}}(\mu)=1+\frac{\alpha_{s}}{4 \pi} C_{F}\left(2 \ln ^{2} \frac{\mu}{m_{b}}+\ln \frac{\mu}{m_{b}}+2+\frac{\pi^{2}}{12}\right) \tag{C.14}
\end{equation*}
$$

where we have renormalised the UV-divergences in the $\overline{\mathrm{MS}}$-scheme.

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[^0]:    ${ }^{1}$ At this point I want to apologise to the reader because of a mismatch in my notation: $A_{\text {spect. }}$ is used for the matrix elements of the effective Operators $\mathcal{O}_{i}$ between both free external quarks and hadronic meson states. It should become clear from the context what is actually meant.

[^1]:    ${ }^{2}$ Tensor integrals which contain expressions like $k^{\mu}, k^{\mu} k^{\nu}, \ldots$ in the numerator can be reduced to scalar integrals as described in [23]

[^2]:    ${ }^{1}$ Please note that the hard spectator scattering kernel starts at $\mathcal{O}\left(\alpha_{s}\right)$. So we call $T^{\mathrm{II}(1)}$ the LO and $T^{\mathrm{II}(2)}$ the NLO.

[^3]:    ${ }^{2}$ In order not to confuse the reader I stress that the symbols $\tilde{\otimes}, \otimes$ in this and the following equations are meant in terms of 3.2 and not of 3.46 .

