

NOAA/MLML **Radiometric** Data **Acquisition and Processing**

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Abstract

This report is a detailed description of data processing of NOAA/MLML spectroradiometry data. It introduces the MLML_DBASE programs, describes the assembly of diverse data fues, and describes general algorithms and how individual routines are used. DefInitions of data structures are presented in Appendices.

Introduction

The purpose of this report is to document data processing of radiometric data obtained in support of the NOAA/MLML participation in the SeaWiFS project. This report presents details of program use, but a companion report (Broenkow, et al. 1993) presents a graphical, less-detailed summary of methods described here.

In 1978 the MLML oceanography group participated with the NOAA/NESDIS CZCS project by writing programs to acquire and process spectroradiometer data. Our current work provides similar support for the NOAA/NESDIS Marine Optical Characterization Experiment, MOCE, in preparation for the launch of the new ocean color satellite, SeaWiFS. During CZCS experiments, an 80 channel scanning radiometer measured in-water spectral upwelled radiance, $L_u(z,\lambda)$, and spectral downwelled irradiance, $E_d(z, \lambda)$, in the visible spectrum from 400 to 700 nm. A similar instrument onboard ship measured surface spectral irradiance, $E_s(\lambda)$. *In situ* observations of spectral L_u , E_d and E_s were required for calibration and validation of the spectral water-leaving radiance, $L_W(\lambda)$, as retrieved from the CZCS satellite sensor. These radiometric measurements were also used to develop and validate bio-optical algorithms and atmospheric correction algorithms.

The purpose of the data acquisition and reduction programs described here are similar to those of the CZCS program, and many of these programs directly parallel those used in the CZCS work. The purpose of the SeaWiFS project is to achieve remotely-sensed, water-leaving radiances within 5% accuracy, and chlorophyll *a* concentration within 35% over the range 0.05 to 50 mg m^{-3} . New instruments were developed by Dennis Clark of NOAA/NESDIS to surface-truth the SeaWiFS sensor. The dual spectrograph (340 to 900 nm) Marine Optical System, MOS, measures inwater $L_n(z,\lambda)$ and $E_d(z,\lambda)$. The instrument can be used either in the Marine Optical Buoy (MOBY) or by deployment from a ship. Data from either mode are processed identically. A 38-channel Surface Irradiance Spectrometer, SIS, measures $E_s(\lambda)$ at the sea surface. An older 80-channel PMT older 80-channel PMT spectroradiometer, Fastie (named after its optical designer), is also used from time to time. As part of the the NOAA/NESDIS project MLML, has provided C and FORTRAN software to acquire data from these instruments. Unlike their earlier CZCS counterparts, the new instruments are controlled by a VAXStation 4000.

Measurements of L_u , E_d , and E_s are used to calculate normalized water-leaving radiance, $L_{\text{WN}}(\lambda)$. These near nanometer resolution spectra are weighted by the eight SeaWiFS response functions to simulate SeaWiFS-determined radiances. The SeaWiFS $L_{\text{WN}}(\lambda)$ will be used for "vicarous" satellite sensor calibrations. In addition, these data observations will be used for development of bio-optical and atmospheric correction algorithms.

Measurements of $L_u(z,\lambda)$ and $E_d(z,\lambda)$, made over the upper few optical depths, are required to compute the diffuse attenuation coefficients $K_E(z, \lambda)$ and $K_L(z, \lambda)$. Optical depth, $\tau(z, \lambda)$, in the context of this report is the reciprocal integral of $K(z, \lambda)$ (for either radiance or irradiance), from the surface to a given depth z. Profiles of $E_d(z,\lambda)$ and $L_u(z,\lambda)$ must be normalized to account for changes during a cast in solar zenith angle, cloud cover and aerosols, hence the need for measuring $E_s(z, \lambda)$. $K_l(\lambda)$ is used to extrapolate profiles of $L_u(z,\lambda)$ to an infinitesimal depth below the surface $(z = 0)$, L_u(0⁻, λ), which, when propogated through the sea surface leads to the estimated value of $L_W(\lambda)$. To standardize SeaWiFS in-water algorithms, it is necessary to normalize $L_W(\lambda)$ for sun zenith angle, earth-sun distance, and atmospheric transmittance, yielding $L_{\text{WN}}(\lambda)$.

In conjunction with oceanographic and atmospheric measurements, MOS and SIS spectroradiometers are operated shipboard at a biooptics station. MOS is attached to a floating tetrahedron and lowered to typically three depths. The tetrahedron is floated away from the ship to avoid measuring a light field contaminated by the ship's shadow or reflection. Subsurface measurements are taken near local apparent noon, IAN, and require approximately two hours to complete. Typical sampling depths are 1, 6, and 11 m in eutrophic waters or 1, 11 and 21 m is oligotrophic waters. Simultaneous with underwater measurements, SIS measures $E_s(\lambda)$ to account for changing sky conditions. During the data acquisition session the workstation is used to control both instruments, to display the data in real-time and to save data in the MLML_DBASE format (Broenkow and Reaves, 1993).

MOS and SIS produce data sets, composed of several spectral scans and ancillary information. Each scan consists of two 500-channel spectra: the blue from about 340 to 640 nm; the red from about 600 to 900 nm. The spectral output from a diode array is measured as dark scans with the shutter closed and as radiance or irradiance scans when the shutter is open. MOS also contains light-emitting diodes for field calibration. The dark scans account for instrument background noise and are subtracted from radiance or irradiance measurements. Spectra are scanned in sets, typically composed of two dark scans followed by five radiance or irradiance scans, followed by another two dark scans. The number of scans is controlled by the operator depending on wave noise, sky variability and the achieved signal to noise ratio. MOS contains sensors to measure external water pressure and temperature, internal and array temperatures, instrument tilt and heading, heat exchanger cooler flow, and battery voltages. SIS measures internal temperature and reference voltages.

VAX FORTRAN and C software was developed at MLML to acquire and process large amounts of data gathered by the MOS and SIS spectroradiometers. Instrument control and acquisition software provides the interface between the person operating the instruments and the hardware. Control of MOS is done by sending FORTH commands to the instrument. MOS is controlled internally by a TattleTale 7 FORTH system described by Reaves and Broenkow (1993). The VAX MOS and SIS programs display menus, accept commands from the operators terminal, send commands to, and receive data from the instruments, and display spectral scans and ancillary data to the operator. Acquisition programs save all data in MLML DBASE or *PLOT* binary fIles. These scansets, are subsequently processed by converting raw data from digital counts to radiometric units, and by deriving required radiometric products. Broenkowand Reaves (1993), "Introduction to MLML DBASE Programs", discusses *PLOT* terminology which will be used throughout this report. Broenkow, *et al.* (1993), "ProcessingNOAASpectroradiometer Data", provides a terse, graphical outline of the NOAA/MLML data processing system.

A completely processed optical profile from a MOCE bio-optics station is contained in a single file and consists of:

- 1. L_u , E_d , and E_s spectral scans at three depths with ancillary information.
- 2. Radiance and irradiance diffuse attenuation coefficients, K_L and K_R .
- 3. Water-leaving radiances, L_{WN} normalized for time of day, location, atmospheric scattering functions, and L_{WN} weighted for the eight SeaWiFS bands.
- 4. Photosynthetically available radiation, PAR, solar zenith angle, airmass, and normalization factor, F_{N}

Each file contains a header identifying: cruise and ship name; station name, GMT date and time, geographic location; cloud cover and type, sky conditions; wind direction and velocity, sea state; Secchi depth and Munsell color; instrument identification; calibration date and file identification; processing methods and algorithms; and observer names.

Data Acquisition Outline

Data acquisition programs control MOS, SIS and Fastie spectroradiometers via four X-Windows displays:

1. The control window displays a menu of commands to start and stop scans, setup

Table 1. The NOAA Main Menu for Processing Radiometric Data. Note that some program names are the same as those used in the main MLML_DBASE program *PLOT.*

are selected via a mouse-driven menu bar and should be recorded and archived at 4 levels:executed by "clicking" on the highlightedcommand.

- 2. Program prompts are displayed in a text window.
- 3. A third window displays intrument analog
- displayed in a graphics window. for dark/zero offsets.

MOS Data Acquisition by measured $E_s(\lambda)$.

This section is in preparation.

operating parameters, and save data. Commands has recomended that $E_c(\lambda)$, $E_d(z,\lambda)$, and $L_u(z,\lambda)$ data

- Level 0: Raw instrument digital ouput.
- All keyboard input is read from this window. Level 1: Instrument output in volts (or frequency) and depth.
- Level 2: Calibrated radiance/irradiance, ancillary measurements in appropriate geophysical 4. X-V line graphs of spectral scan information are or biological units, and depth, corrected
	- Level 3: Smoothed profiles of $K(z, \lambda)$ and associated $L_u(z, \lambda)$ or $E_d(z, \lambda)$ normalized
- This section is in preparation.
Level 4: Level 3 data normalized to clear-sky,
zenith sun at mean earth-sun distance, SIS Data Acquisition and spectrally adjusted to match SIS Data Acquisition SeaWiFS reference wavelengths and FWHM bandwidths.

NOAA/MLML radiometric data acquisition Fastie Data Acquisition software (the SIS, MOS and Fastie programs) produces level 0 files in output units of raw digital This section is in preparation. counts. Before numerical processing begins, the various files taken at a single optics station must be Data Processing Outline assembled and inspected. This is done using a set of routines for "merging" file names into a file called The SeaWiFS Prelaunch Science Working Group MERGE.LIS. These files are then "grouped" time-

wise. Following that, spectral data from these files are inspected by the NOAA LOOK program. After the files have been assembled into logical units constituting a single radiometer station, the programs described in Table 1 produce level 2 through level 4 files. The order in which those programs are applied to MOS and SIS data sets is usually top-down, i.e. Program 1 through Program 10, though this order can vary. Program 11 incorporates one or more of Programs 1 through 10 for batch-mode processing.

.Data Description

A level 0 scan-set is acquired and saved to an *Archive* file, which consists typically of the following 10 *Variables* (the order of *Variables* 2 through 10 corresponds to the temporal sequence in which they were acquired):

Variable #

1. Nominal Wavelength (nm) 2. Dark Scan $# 1$ (counts) 3. Dark Scan # 2 (counts) 4. L_u or E_d or E_s Scan # 1 (counts) 5. L_u or E_d or E_s Scan # 2 (counts) 6. L_u or E_d or E_s Scan # 3 (counts) 7. L_u or E_d or E_s Scan # 4 (counts) 8. L_u or E_d or E_s Scan # 5 (counts) 9. Dark Scan $# 3$ (counts) 10. Dark Scan # 4 (counts)

Each of the above *Variables* contains 1000 *Elements* for a MOS $L_n(z, \lambda)$ or $E_d(z, \lambda)$ scan, or 48 *Elements* (38 spectral channels + 10 reference channels) for a SIS $E_s(z,\lambda)$ scan. Each *Element* contains the 16-bit (0-65535) intensity level of a single diode array pixel. All raw scan data units are counts, and wavelengths are un-calibrated nominal values in nm. The *Variable Header* for each scan contains an *Extension* which holds the following information collected at the time of the individual scan (see Appendix 5 for definitions of MLML DBASE MOS, SIS and Fastie data structures):

Scan, Cruise, Station, Cast, Instrument Serial #, Instrument Configuration #, Multiplexer Serial # (MOS only), Scan Time, Navigation Type, Latitude, Longitude, Time of Last Satellite Overpass, Variable Type, Variable Subtype, Processing Code, Smoothing Code, Interpolation Code, Beginning Interpolation Wavelength, Interpolation Wavelength Interval, Number of Averages (SIS only), Multiplexer Position (MOS

only), Mirror Position (MOS only), Comments, Number of Points Edited, Number of Errors.

Ed and *Variable Header Extensions* contain the following for each of the two (blue and red) MOS spectrographs:

Cooler Intensity, Intensifier Setting, LED Level, Data Length, Gain, Integration Time.

Es *Variable Extensions* contain the following for the SIS spectrograph:

Shutter Setting, Number of Channels, Gain and Timeout for each channel.

Type, value, gain and average are saved for the following analog channels:

MOS

Internal Temperature, External Temperature, Pressure Transducer Temperature, External Pressure (depth), Battery Voltage, High and Low Voltage References, LED Temperature, Compass Direction, X-tilt, Y-tilt, Blue and Red Array Temperature, Heat Exchanger Cooler Flow, Compass Direction.

SIS

Internal Temperature, Battery Voltage, X-tilt, Y tilt, Array temperature.

Appendix 5 provides the defmition and ranges of values for data codes such as Variable Type, Processing Code, Cooler Intensity and Shutter Setting.

Level 1 files are not produced for the MOS and SIS instruments. This is because the diode array detectors with their integral analog to digital converters are calibrated as single units: their digital outputs are directly converted to radiometric units. This is also the case for all ancillary sensors, whose digital outputs are directly converted to physical units.

Assembling Data Files

The SIS, MOS and Fastie data acquisition programs produce independent fIles (one for each instrument) containing surface irradiances, upwelled radiance or downwelled irradiances at a number of depths. These instruments have different spectral resolution, and data acquisition from them is not done synchronously. This makes assembling and reducing the diverse data sets difficult. 3.

Perhaps the most difficult problem is to provide a flexible system to merge the various data sets with the goal of determining water-leaving radiances and diffuse attenuation coefficients. The data acquisition programs produce files with file names that uniquely determine their contents (Appendix 4). Data from a single cruise or calibration experiment are contained in a single directory. The data merging procedure is outlined as follows:

1. The first step is to create lists of file names which constitute data for a single optical profile or for a group of stations to be processed in one session. This is done with the command "ALIAS", which is runs the DCL command file \$MLMl.\$DIR:ALIAS.COM. ALIAS is executed twice, first to assemble the MOS (and perhaps
Fastie) files for a specific station and next to 4. assemble the SIS files for that same station. The examples below result in two fIles: MALIAS and SALlAS, which are located in the user's logical directory ALIAS\$DIR. The use of VMS wildcards (*) limits the search to selected files. Familiarity with the VMS file naming conventions is essential to accomplish this process. The syntax for this is \$ ALIAS input-files output-files, for example,

\$ ALIAS (DATA.MOCE 2.MOS.RAW]SIN2*.MLDAT M \$ ALIAS [DATA.MOCE_2.SIS.RAW]*.MLDAT S

Examples of the ALIAS files are shown in Table 2.

2. The next step is to fmd the number of *PLOT Variables* and the time-span for each file. This is done by passing the M.ALIAS and S.ALIAS files to the NOAA LIST HEADERS program, which can be executed using the VMS symbol NLH, which is equivalent to \$ RUN NOAA LIST HEADERS, from the NOAA\$BIN directory. That program is invoked twice, once for MOS (or Fastie) files and again for SIS files. The syntax for this is \$ NLH input-alias-file output-list-fue, for example,

Output from this procedure are the files, M.LIS and S.LIS, (Table 3) which contains header information from the selected MOS and SIS files.

The third step in merging files is to "group" the surface irradiance files (SIS) by time with the contemporaneous MOS files containing contemporaneous MOS files containing downwelled irradiance and upwelled radiance. This is done by use of the program NOAA GROUP FILES which can be run using the VMS symbol, NGF, equivalent to

\$ RUN \$NOAA\$BIN:NOAA_GROUP_FILES

The syntax to invoke the "grouping" program requires the ALIAS file names for the MOS and SIS fde names followed by the name of the "merge" file, for example,

\$ NGF M.LIS S.LIS MERGE.LIS;1

Output from this procedure is shown in Table 4.

Up to this point, data reduction can be done more-or-Iess blindly, with the emphasis on "less blindly", because it is generally impossible to trap for all errors or exceptional circumstances that may occur aboard ship during data acquisition. Examples of exceptions might be erroneous latitude and longitude, erroneous system clock time or empty data files. Steps 1-3 are done outside of *PLOT,* from the VMS command line. The next step requires running *PLOT* and the *NOAA* programs to examine each E_d or L_u and E_s scan-set to determine which are suitable. This requires plotting each spectral scan and noting which are to be retained. The procedure is done inside *PLOT.* The syntax for doing this from the main *PLOT* prompt is to run the *NOAA* program and its subprogram *LOOK*followed by the merge listing file name, for example,

ML> NOAA LOOK MERGE.LIS;1

While determining which files are to be combined, it is essential that a hardcopy of the MERGE.LIS;1 file is available on which to note the acceptable E_s , L_u , and E_d files. This procedure requires patience and practice.

After the LOOK program has been run, the MERGE.LIS;1 file is manually edited to produce the file, MERGE.LIS:2, which contains the file names of only those scans which have passed scrutiny. The others are deleted by a text editor. The MERGE.LIS:2 file should be kept as a permanent record of data processing procedure.

Table 2. Example of MOS and SIS ALIAS Files used to group all Station 2 files before numerical processing.

Table 3. Example of the M.LIS and S.LIS files formed by running the NOAA LIST HEADERS program.

Table 4. Example of a MERGE.LIS:1 file produced by the NOAA_GROUP_FILES program before running NOAA LOOK. Notice that the ALIAS file names (M20, *S797*) precede the complete file names.

MERGE.LIS;1 M1 DUA1: [DATA.NOAA.MOCE_2.MOS.RAW]STN02_11M_LU_05.MLDAT
S819 DUA1: [DATA.NOAA.MOCE_2.SIS.RAW]SIS_930331_195129.MLDAT S819 :OUA1:[DATA.NOAA.MOCE:2.SIS.RAWlSIS_930331_195129.MLDAT M20 DUA1: [DATA.NOAA.MOCE_2.MOS.RAW]STN02_1M_ED_01.MLDAT_
S797 DUA1: [DATA.NOAA.MOCE_2.SIS.RAW]SIS_930331_185546.ML S797 - DUA1: [DATA.NOAA.MOCE_2.SIS.RAWISIS_930331_185546.MLDAT
S798 - DUA1: [DATA.NOAA.MOCE_2.SIS.RAWISIS_930331_185646.MLDAT S798 -OUA1:£DATA.NOAA.MOCE-2.SIS.RAWlSIS-930331-185646.MLOAT S799 -OUA1:[OATA.NOAA.MOCE-2.SIS.RAWlSIS-930331-185747.MLOAT S800 -OUA1:[OATA.NOAA.MOCE-2.SIS.RAWlSIS-930331-185846.MLOAT S801 DUA1: [DATA.NOAA.MOCE 2.SIS.RAW]SIS 930331 185946.MLDAT M2 DUA1: [DATA.NOAA.MOCE 2.MOS.RAW] STN02 1M LU 02.MLDAT
S801 DUA1: [DATA.NOAA.MOCE 2.SIS.RAW] SIS 930331 185946.ML $\overline{}$ DUA1:[DATA.NOAA.MOCE $\overline{}$ 2.SIS.RAW]SIS_930331_185946.MLDAT M22 DUA1: [DATA.NOAA.MOCE_2.MOS.RAW]STN02_1M_LU_03.MLDAT
S801 DUA1: [DATA.NOAA.MOCE_2.SIS.RAW]SIS_930331_185946.ML \overline{D} DUA1:[DATA.NOAA.MOCE $\overline{2}.$ SIS.RAW]SIS_930331_185946.MLDAT M3 DUA1: [DATA.NOAA.MOCE_2.MOS.RAW]STN02_1M_LU_04.MLDAT_
S801 DUA1: [DATA.NOAA.MOCE_2.SIS.RAW]SIS_930331_185946.ML $_$ DUA1:[DATA.NOAA.MOCE $_2$.SIS.RAW]SIS $_9\overline{3}$ 0331 $_$ 185946.MLDAT M4 DUA1:[DATA.NOAA.MOCE 2.MOS.RAWlSTN02 ED 11M 08.MLOAT S863 -OUA1:[OATA.NOAA.MOCE-2.SIS.RAWlSIS 930331 203529.MLOAT S864 ⁻⁻DUA1:[DATA.NOAA.MOCE⁻⁻2.SIS.RAW]SIS⁻930331⁻⁻203629.MLDAT
S865 ---DUA1:[DATA.NOAA.MOCE⁻⁻2.SIS.RAW]SIS-930331⁻⁻203729.MLDAT 5865 -OUA1:[OATA.NOAA.MOCE-2.SIS.RAWlSIS-930331-203729.MLOAT S866 $\sqrt{2}$ DUA1:[DATA.NOAA.MOCE $\sqrt{2}$.SIS.RAW]SIS $\sqrt{2}$ 930331 $\sqrt{2}$ 03829.MLDAT S867 DUA1: [DATA.NOAA.MOCE_2.SIS.RAW]SIS_930331_203929.MLDAT #Var=10 31Mar93 19:45 to 31Mar93 19:50 #Var=10 31Mar93 19:51 to 31Mar93 19:51 #Var=14 31Mar93 18:55 to 31Mar93 19:02 #Var=10 31Mar93 18:55 to 31Mar93 18:56 #Var=10 31Mar93 18:56 to 31Mar93 18:57 #Var=10 31Mar93 18:57 to 31Mar93 18:58 #Var=10 31Mar93 18:58 to 31Mar93 18:59 #Var=10 31Mar93 18:59 to 31Mar93 19:00 #Var=10 31Mar93 19:07 to 31Mar93 19:12 #Var=10 31Mar93 18:59 to 31Mar93 19:00 #Var=19 31Mar93 19:07 to 31Mar93 19:18 #Var=10 31Mar93 18:59 to 31Mar93 19:00 #Var=28 31Mar93 19:07 to 31Mar93 19:27 #Var=10 31Mar93 18:59 to 31Mar93 19:00 #Var=10 31Mar93 20:35 to 31Mar93 20:39 #Var=10 31Mar93 20:35 to 31Mar93 20:35 #Var=10 31Mar93 20:36 to 31Mar93 20:36 #Var=10 31Mar93 20:37 to 31Mar93 20:37 #Var=10 31Mar93 20:38 to 31Mar93 20:38 #Var=10 31Mar93 20:39 to 31Mar93 20:39

Table 5. Example of the edited MERGE.LIS file containing those files that passed scrutiny in the NOAA LOOK program.

Table 6a. Schematic of MLML_DBASE Merged File Contents. Top, Mid and Bot refer to the sample depths.

Table 6b. Schematic of MLML_DBASE Derived File Contents. Attenuation coefficients between top and mid are designated K₁, between top and bot as K₂, and between mid and bot as K₂. Water-leaving radiances are designated by the surface-most upwelled radiance and the attenuation coefficient (1,2,3) used in the extrapolation.

The liberal use of comments, which are denoted by a leading exclamation point (I), is essential. An example of the edited MERGE.LIS:2 file is shown in Table 5.

5. The last step required to assemble scans before numerical processing can begin is to form *PLOT* data files containing the E_S , L_u , and E_d , data for the Top, Mid and Bot sampling depths. That is done within *PLOT* using the *NOAA LOOK* program which requires the name of the merge listing file, for example,

ML> NOAA MERGE MERGE.LIS;2

Notice that the second version number of the MERGE.LIS file is used.

Following preprocessing steps 1-5, the NOAA programs, NOAA_CONVERT, NOAA_DERIVE, NOAA CLIP, NOAA GRAPH, and NOAA LIST can be run to produce calibrated radiances, attenuation coefficients andwater-leaving radiances, as explained next. The second data processing phase involves numerical conversions and produces files containing a well-ordered set of radiometric values shown in Table 6.

- 1. Each raw MOS $L_u(z, \lambda)$, $E_d(z, \lambda)$ and SIS $E_s(z, \lambda)$ scan-set is converted from raw counts to radiometric units by applying the sensor's system response. Since E_d and L_u sensors are calibrated in air but are used to measure radiation underwater, an additional immersion correction factor, F_{w} , is applied.
- 2. For MOS data only, known "bad" pixels are replaced with a missing-data code, NaN.
- 3. For MOS data only, each scan is smoothed spectrally to reduce instrument noise.
- 4. Dark scans are averaged. Radiance or irradiance scans are averaged. Average dark count is subtracted from average radiance or irradiance. Percent errors are determined for each average, and Signal to Noise Ratio, SNR, is calculated.

These steps are repeated for each scan-set, with the resulting *Variables* recorded to *PLOTArchive* files in the following order (for example, using a SIS E_s scanset):

Variable #

- 1. Corrected SIS Wavelength (nm) 2. Dark Scan # $1 \, (\mu W/cm^2/nm)$ 3. Dark Scan $\# 2 \ (\mu W/cm^2/nm)$ 4. E_s Scan # 1 (μ W/cm²/nm) 5. E_s Scan # 2 (μ W/cm²/nm) 6. E_s Scan # 3 (μ W/cm²/nm) 7. E_s Scan # 4 (μ W/cm²/nm) 8. E_s Scan # 5 (μ W/cm²/nm) 9. Dark Scan $# 3 (\mu W/cm^2/nm)$ 10. Dark Scan $# 4 (\mu W/cm^2/nm)$ 11. Mean Dark $(\mu W/cm^2/nm)$ 12. Dark %RMSE 13. Mean E_c (μ W/cm²/nm) 14. E_s %RMSE 15. Mean E_s - Mean Dark (μ W/cm²/nm) 16. SNR
- 5. The dark-corrected radiance or irradiance and SNR (V*07#* 15 and 16 above) of several sets which represent a single depth or a single time are averaged. For example, several E_s sets may span the time interval of one L_u scan, and hence must be averaged together.

At this point, each set has been processed to a corrected and dark-adjusted radiance or irradiance spectra with SNR. Multiple scan-sets have been averaged if necessary. Each (average) E_d and L_u at a depth has been associated with an (average) E_s .

The final step is to combine SIS and MOS files, trim the scans to wavelengths having acceptable signal to noise, derive attenuation coefficients and waterleaving radiances and produce hard-copy reports.

- 6. Processed data files are merged into a single optical profile, containing E_d and L_u spectra at three depths along with corresponding E_s and SNR's. We use the terminology "Top" for the surface-most data; "Mid" for data obtained at the middle depth; "Bot" for data taken at the deepest or bottom depth. A merged profile contains 26 *Variables* in the following order:
	- 1. Corrected MOS Wavelength (nm)
	- 2. Corrected SIS Wavelength (nm)
	- 3. Top E_d (μ W/cm²/nm)
	- 4. SNR of Top E_d
	- 5. E_i for Top E_d (μ W/cm²/nm)
	- 6. SNR of E_s for Top E_d
- 7. Surface L_u (μ W/cm²/sr/nm) 8. SNR of Top L_u 9. E_i for Top L_u (μ W/cm²/nm) 10. SNR of E_s for Top L_u
- 11. Mid E_d (μ W/cm²/nm) 12. SNR of Mid E_d 13. E_i for Mid E_d $(\mu W/cm^2/nm)$ 14. SNR of E_s for Mid E_d 15. Mid L_u (μ W/cm²/sr/nm) 16. SNR of Mid L_u 17. E_i for Mid L_u (μ W/cm²/nm)
- 18. SNR of E_s for Mid L_u
- 19. Bot E_d (μ W/cm²/nm) 20. SNR of Bot E_d 21. E_i for deep $E_d (\mu W/cm^2/nm)$ 22. SNR of E_s for Bot E_d L_u(μ W/cm²/sr/nm) 23. Bot L_{μ} (μ W/cm²/sr/nm) 24. SNR of Bot L_u 25. E_i for Bot $L_u(\mu W/cm^2/nm)$ 26. SNR of E_s for Bot L_u
- 7. The SeaWiFS project group identifies a SNR of 100:1 as the minimum quality assurance level. Individual data *Elements* below this level can be removed, or clipped, from the merged *Variables.*
- 8. Diffuse attenuation coefficients and solarnormalized water-leaving radiances are calculated from *Variables* arranged in step 6 and clipped in step 7. SeaWiFS normalized responses are applied for the eight bands. PAR, FLH, solar altitude, and air mass are calculated and saved with processed spectral and analog data. This step derives the *Var#'s* 27 to 36. All 36 *Variables* are recorded in a single *Archive* file containing one optical profile. The file contents are summarized in Table 6.

27. Top to Mid K_F (m-¹) 28. Top to Mid K_L (m-¹) 29. Top to Bot K_E (m^{-1}) 30. Top to Bot K_L^{-} (m-¹) 31. Mid to Bot K_E (m-1) 32. Mid to Bot $K_L^{\sim}(m^{-1})$ 33. L_W Top/1 (μ W/cm²/sr/nm) 34. $L_W \text{Top}/2 \ (\mu\text{W}/\text{cm}^2/\text{sr}/\text{nm})$ 35. L_W Mid/1 $(\mu W/cm^2/sr/nm)$ 36. L_W Bot/2 (μ W/cm²/sr/nm)

- 37. L_{WN} Top/1 (μ W/cm²/sr/nm)
- 38. L_{WN} Top/2 (μ W/cm²/sr/nm)
- 39. L_{WN} Mid/1 $(\mu W/cm^2/sr/nm)$
- 40. L_{WN} Bot/2 (μ W/cm²/sr/nm)
- 41. F_N Solar normalizing factor
- 9. Two pages of X-Y line graphs formatted with three plots per page are produced (Figs. 1 and 2). These can be displayed to a terminal, PostScript printer or PostScript file. Plots are available in color or in black-and-white.
- 10. A data report listing is produced in tabular ASCII format and output to either a terminal, printer, or file (Figs 3 and 4). Graphical and tabular output are useful for archive and reference. The ASCII files generated here or more simple ASCII files generated by the *PLOT* program are used to distribute data to interested parties.

The above processing is implemented in 10 steps to allow flexibility in program sequencing, to accommodate exceptions, and to allow use of partially processed profiles. This is necessary, for example, while processing calibration data to generate system response, $R(\lambda)$. In this situation, *CONVERT* is not applied and *DERIVEd* parameters are not calculated. Instead, sets are *EDITed, SMOOTHed, ADJUSTed* and possibly *MERGEd.* System response is computed as the ratio of the calibrated lamp spectra and the processed MOS or SIS spectral counts.

The *CONVERT* step can be omitted if a valid system response is not available. This may be the case before post-cruise calibrations are performed. Scan-sets can be processed in raw units and later *CONVERTed* to radiometric units. Indeed, diffuse attenuation coefficients are identical if*DERIVEd* from *CONVERTed* or unconverted scan-sets.

Running Data Processing Programs

What follows illustrates use of the NOAA programs, which run under the MLML DBASE program, generally called *PLOT*. *PLOT* must be run from the VMS command line as

\$ PLOT 100 1000 4 4 2000

The four parameters passed to the *PLOT* program do the following: 100 indicates the number of *PLOT Variables;* 1000 indicates the number of data *Elements;*

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File: MOCE1: [MOS.PRC]STN07_1_PRC.MLDAT;1

MODIS Marine Optical Characterization Experiment - 1 NOAA/MlML

Figure 3. First page of the data report, observed spectra, produced by NOAA_LIST.

File: MOCE1: [MOS.PRC]STN07_1_PRC.MLDAT;1

MODIS Marine Optical Characterization Experiment - NOAA/MLML

Figure 4. Second page of the data report, attenuation coefficients and water-leaving radiances, produced by NOAA LIST.

which is sufficient for NOAA MOS or SIS spectra; 4 indicates use of MLML DBASE DATA TYPE 4, being REAL*4 values having about 7 significant digits; the second 4 indicates the DATA ELEMENT SIZE as 4 bytes; and 2000 sets the File Header Extension size to be 2000 bytes. The File Header Extension holds derived products,and Appendix 5 shows the details of the extension variables.

Two methods access the above programs from *PLOT:*

- 1. At the ML> prompt type "NOAA" to print a menu and display the NOAA> prompt.
- 2. At the ML> prompt type "NOAA" followed by the program name or number.

Typing EXIT or 0 at the NOAA> command line exits back to *PLOT* prompt, ML>, and entering an empty RETURN at NOAA> redisplays the menu shown below. As with the *PLOT* program, DCL commands can be run from the NOAA> prompt by preceding them with a $\$, or a DCL "shell" can be run by typing \$ RETURN.

To display the main NOAA processing menu:

ML> NOAA

Notice that the prompt changes to NOAA, and that NOAA program names or numbers rather than the PLOT program names are in use.

> NOAA: NOAA/MLML Radiometric Data Processing rev: 24-May-1993

:/I Command Description

o EXIT-NOAA>-

To exit from NOAA>:

NOAA> EXIT

```
or 
    NOAA> 0
```
 $ML >$

To start the *SMOOTH* program, for example:

ML> NOAA SMOOTH

or ML> NOAA 3

NOAA programs will prompt the user for necessary input. Typing CTRL Z in response to a prompt exits that program without processing data. Many prompts provide default values which can be accepted, modified or re-entered. Program input requests can be satisfied by passing parameters to the program when it is started. When passing parameters to a program, default inputs can be accepted by typing the wildcard * in place of the specific *Variable* or *Element* ranges. It is best to run a program through the interactive mode the fust time to understand what input needs are required and what default values are provided. For example, *SMOOTH* prompts for the *Variable* list to be smoothed and the number of points to use for the smoothing window. *SMOOTH* provides 5 as the default window size. Input requests from *SMOOTH* can be answered in several ways:

1. Starting *SMOOTH* from the main NOAA> command line:

ML> NOAA (--- NOAA main menu printed here ---) NOAA> SMOOTII NOAA SMOOTII Enter Range of Variable # to Smooth: $2-10$ NOAA SMOOTH Enter # of Elmts/Window (odd): 5 Working... Done.

2. Starting *SMOOTH* from the ML> command line:

ML> NOAA SMOOTII NOAA_SMOOTH Enter Range of Variable # to Smooth: 2-10 $NOAA$ SMOOTH Enter $#$ of Elmts/Window (odd): 5 Working... Done.

3. Passing input to *SMOOTH:*

ML> NOAA (--- NOAA main menu printed here ---) NOAA> SMOOTH 2-10 7

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Working... Done.

or

ML> NOAA SMOOTH 2-107 Working... Done.

4. To use the default *SMOOTH* window size (5):

The following paragraphs define the user interface . and some details of processing algorithms and data handling for each of the *NOAA* programs.

MLML NOAA

The main NOAA program provides an interface to the ten radiometric data processing programs. The tasks of NOAA are four:

- 1. NOAA displays the processing menu and the NOAA> prompt (see example above), reads and parses user commands.
- *2.* NOAA handles any parameters passed from *PLOT.*
- *3.* NOAA handles any parameters passed to subprograms.
- 4. *NOAA* runs its subprograms.

The menu contents and physical location of subprogram code is known to NOAA via the ASCII file NOAA MAIN MENU.DAT (see Appendix 4). This file contains three lines for each subprogram: line $1 = program$ name; line $2 = program$ description; line 3 = complete VMS filename of program. NOAA reads this file, prints the name and description information to the user's terminal in a menu format, and uses the VMS filename to run subprograms. NOAA MAIN MENU.DAT can be updated to add, delete, or change the name, description, or location of a NOAA subprogram.

Programming Details

A temporary disk file is used to pass parameters between separate programs because FORTRAN does not allow directly for passing arguments to a main program. The existence of the temporary parameter file is known to separate program units via the VMS logical MLML_DBASE_PARAMETER in a VMS logical table (Appendix 3). Managing parameter files constitutes most of NOAA's work. If input is passed from *PLOT,* this information is extracted from the parameter file for interpretation by a subroutine. The name of the parameter file is retrieved from the VMS logical and saved, freeing the logical to be used, if necessary, to pass parameters to subprograms. Before exiting back to *PLOT,* the original parameter filename is replaced in the logical and returned to *PLOT.* If it is necessary for NOAA to pass parameters to a subprogram, it creates a temporary parameter file, writes information to the file, calls the subprogram, then deletes the file after the subprogram is finished.

If all the NOAA programs were coded as subroutines, NOAA would not have to pass parameters via external files. As separate programs, however, command line mode is enabled from both the ML> and NOAA> levels. This is convenient for an advanced user who may write batch mode command files from ML>, and for the novice user who runs the NOAA programs by passing parameters within the NOAA processing environment. As individual programs, the processing routines are also accessible to future high-level application programs, and the increased modularity facilitates program modification and maintenance.

NOAA CONVERT

CONVERT applies a system response and/or immersion correction factor to SIS and MOS data. *CONVERT* has three prompts:

The system response, $h(\lambda)$, is a spectral multiplier used to convert raw units of MOS or SIS radiance

scans $C(\lambda)$ and dark scans $D(\lambda)$ (both in counts) to radiometric units of radiance $L(\lambda)$ (μ W/cm²/sr/nm) or irradiance $E(\lambda)$ (μ W/cm²/nm) for integration times, At. Multipliers are

$$
L(\lambda) = \frac{C(\lambda) - D(\lambda)}{\Delta t} F_{w}(\lambda) h(\lambda)
$$
\n(1)

determined via laboratory calibration of L_u , E_d and E_s sensors. The immersion factor, $F_w(\lambda)$, is a spectral multiplier necessary when the calibration situation is different from that of the field-measurement. When measuring L_u and E_d underwater, the difference between the indices of refraction for air and water is accounted for by $F_w(\lambda)$. Immersion factors are only applied to MOS data. The air-water $F_W(\lambda)$ for radiance is calculated from values for indices of refraction for fused quartz (MOS radiance port), with seawater salinity of 35 PSU and temperature of 16°C (Austin, 1976). The immersion factor for irradiance must be empirically determined; we use data from Tyler and Smith (1970) for the Scripps Spectroradiometer. The radiance immersion factors are listed in Appendix 4.

NOAA EDIT

Erroneous data generated by "bad" pixels in a silicon diode array detector must be removed and replaced with a missing-data code, i.e. NaN "Not a Number". All other processing routines check for the presence of NaN and skip processing of such *Elements. EDIT* prompts the user for *Variable* and *Element* ranges to be replaced.

NOAA> EDIT

NOAA EDIT: Replace Bad Data rev. 22-Jul-1993

Currently, MOS has three unusable areas on the Blue array and one on the Red array for a total of 25 bad pixels out of 1024. These known *Elements* are the default response to the second prompt. There is currently no need to *EDIT* SIS data. In the *Variable Header Extension, NUM PTS EDITED* is updated with the number of bad points edited (see Appendix 5 for SIS and MOS *Extension* structures.)

NOAA SMOOTH

MOS's holographic diffraction grating disperses a bandwidth of approximately 300 nm onto a 512 diode array detector, giving a potential spectral resolution of 0.6 nm. However, to increase sensitivity, a 250 μ m wide spectrograph entrance slit leads to a 2.6 nm or 4 pixel bandwidth. This allows the option to *SMOOTHMOS* data to reduce standard error without loss of spectral information. *SMOOTH* applies an Npoint running-mean, even-weight smoothing function. The user is prompted to select which Variable to smooth and the number of points to use for window size.

NOAA SMOOTH: N-Point Running Mean Smooth 22-Jul-1993

A five point window is the default and is typically used. *SMOOTH* skips the first and last N/2 points, takes N *Elements* centered around each point, replaces any NAN's with the mean of surrounding elements, and sums the weighted N data to replace the middle *Element*. SIS E_s spectra are not smoothed.

$$
\overline{X} = \sum_{j=1}^{j=M} \frac{X(\lambda_j)}{M}
$$
 (2)

Applying a running mean smooth with a 5 point window does not change the frrst and last 2 data *Elements.* For example, a simple 5 point spectra smoothed with a 5 point window would only update the third point.

NOAA ADJUST

A typical SIS or MOS file or scan-set contains ten *Variables:* one wavelength, two dark scans, five radiance or irradiance scans, and two more dark scans. ADJUST averages and calculates the relative error of the four spectral dark *Variables,* averages and calculates relative error for radiance, subtracts average dark from average radiance, then calculates Signal to Noise Ratio, SNR. The user is prompted for *Var#'s* of wavelength, dark, and radiance data, and *Var#'s* to save the six output spectra (saving to *Var# 0* means do not save.)

NOAA> ADJUSf

Working... Done.

Relative error is calculated as the spectral quantity, root mean square error, RMSE:

$$
RMSE = \sqrt{\frac{\sum_{i=1}^{i=N} [X_i(\lambda) - X_m(\lambda)]^2}{N}}
$$

$$
X_m(\lambda) = \frac{\sum_{i=1}^{i=N} X(\lambda)}{N}
$$
(3)

where $X(\lambda)$ represents radiances from the N scans For SNR, the *signal* equals the radiance less the dark counts, while *environmental* noise is given by the error in radiance:

where $L_M(\lambda)$ is the mean radiance or irradiance,

$$
SNR(\lambda) = \frac{L_M(\lambda) - L_D(\lambda)}{RMSE_L(\lambda)}
$$
 (4)

 $L_D(\lambda)$ is the mean radiance or irradiance dark count and $RMSE_I(\lambda)$ is the RMSE from Eq. 3. For MOS *Variables,* data are optionally divided by the integration time. The SIS integration time is constant.

The average percent errors in the eight SeaWiFS bands are saved in the *Variable Header Description* of each %RMSE spectra. These eight mean errors are used as an indicator of "goodness" for the radiance or dark data. Also updated are the *Variable Header Extensions:* all analog values are averaged and saved. This is necessary, for example, to record with the Rad-Drk *Variable* the average depth of a MOS scanset or the average time of a SIS surface irradiance scan-set. Average depth is read from the *Extension* during the *DERIVE* step to calculate diffuse attenuation coefficients (below). *Variable Header* scale parameters are setup and saved to allow for easy plotting.

NOAA AVERAGE

When more than one MOS scan-sets at a particular depth are to be used, or several surface irradiance SIS sets taken at different times need combining, data can be spectrally averaged via *AVERAGE.* This program prompts the user to enter the *Var#s* to be averaged and the *Var#s* to save the mean and %RMSE. These results will not be saved if *Var#* 0 is selected.

Averages are calculated across rows to preserve spectral information. Analog data in *Variable Header Extensions* are also averaged and saved with the mean

100

Enter SNR cutoff:

Variable's Extension. Again, eight mean %RMSE's corresponding to the SeaWiFS bands are calculated as a quality index and saved in the %RMSE *Variable Header Description.*

NOAA MERGE

The *MERGE* program assists the user in combining processed *Archive* files containing L_u, E_d, and E_s data into *Workspace Variables* ready for the *DERIVE* step. The *MERGEd Variable* order is:

Variable #

- 1. MOS Corrected Wavelength
- 2. SIS Corrected Wavelength
- 3,4. MOS surface E_d and SNR
-
- 5,6. SIS E_s and SNR for Top E_d
7,8. MOS surface L, and SNR MOS surface L_u and SNR
-
- 9,10. SIS E_s and SNR for Top L_u
11-18. Mid data as $Var#$ 3-10 Mid data as *Var#* 3-10
- 19-26. Bot data as *Var# 3-10.*

MOS *Variables* hold 1000 *Elements* (500 pixels from each of the two diode arrays) and SIS contain only 48 (38 scan channels plus 10 references). This is why both MOS and SIS wavelength Vars are needed. Merge works on the manually edited MERGE.LIS;2 file that was assembled during preprocessing (page 6).

NOAA> MERGE

NOAA MERGE: Load Radiometric Archive rev: 22-Jul-1993

NOAA MERGE:

Enter Filename: MERGE.LIS;l

Loading...

NOAA CLIP

This program allows spectral data with a SNR less than 100:1 to be replaced with NaN. *CLIP* prompts the user for the *Variable* range to be clipped and the value of the SNR cutoff to be used.

NOAA> CLIP NOAA CLIP: Replace Data if SNR < 100:1 rev: 22-Jul-1993

NOAA DERIVE

Done.

From L_u, E_d, and E_s Variables 1-26, *DERIVE* calculates diffuse attenuation coefficients K_L and K_F , solar-normalized water-leaving radiance, L_{WN} , SeaWiFS-weighted solar-normalized water-leaving radiances in eight SeaWiFS bands, and other products such as PAR and solar inclination. K's are calculated between depths for a pair of L_u or E_d 's and are corrected for differences in associated E_s 's. K's are calculated between the Top and Mid depths, between Top and Bot, and between Mid and Bot. Waterleaving radiance is obtained by attenuating L_u to just below the sea surface using a K_L, then propagating $L_{u,0}$ across the water-air interface. Four such L_w 's $L_{u,0}$ across the water-an interface. Four such L_w s
are calculated: using the L_u _{Top} with K_L _{Top-Mid} and K_L _{Top-Bot}; using L_u _{Mid} with K_L _{Top-Mid}; and the L_{u} Bot with K_{L} Top-Bot.

Between observation depths, z_1 and z_2 irradiance is assumed to attenuate by a first-order absorption and scattering with a constant attenuation coefficient, K,

$$
\frac{E_{z_2}}{E_{z_1}} = e^{-K(z_2 - z_1)}
$$
 (5)

The diffuse attenuation coefficient is calculated by Eq. 6, where K denotes either K_L or K_E ; X denotes either L_u or E_d spectral data; Z are depths, and subscripts 2 and 1 indicate the deeper and shoaler depths, respectively. R_N (Eq. 7) is the incident

$$
K = -\frac{\ln\left(\frac{X_2 R_{N2}}{X_1 R_{N1}}\right)}{(z_2 - z_1)}
$$
(6)

irradiance normalization ratio determined from mean, spectrally- smoothed E_{s} , one (E_{s1}) for the shoaler E_{d} or L_u , the other (E_{s2}) for the deeper E_d or L_u . K cannot be derived if either of the two L_u or E_d scans or their corresponding surface irradiances E_s are missing. E_{sr} is the reference surface irradiance. This will normally be E_s taken for the surface-most (i.e. Top) scans.

$$
R_{N1} = \frac{E_{s}t}{E_{s1}}
$$

\n
$$
R_{N2} = \frac{E_{s}t}{E_{s2}}
$$
 (7)

The depths of L_u and E_d scans are retrieved from *Variable Header Extension* analog data, converted to pressure units (dbar), corrected for "deck" pressure offset, and corrected for pressure offset between the . sensor and pressure transducer. E_s spectra are used to normalize L_u and E_d scans taken under potentially different sky conditions. L_u and E_d are normalized to the time of the surface-most E_s scan. Normalization factors, R_N , can be calculated as a mean ratio or spectral ratio. For mean normalization, E_s spectra are summed across wavelengths, divided by the number of *Elements,* then ratioed. A spectral ratio is cubic spline interpolated to MOS wavelengths. If a spectral R_N ratio is used, however, MOS wavelengths outside the E_c bandwidth cannot be normalized and are replaced with NaN's.

Water-leaving radiance, L_W , is calculated by extrapolating a radiance spectrum from one Evening: (observation depth to the surface:

$$
L_{w} = 0.543 L_{u} e^{K_{L}(Z_{2} - Z_{1})}
$$

0.543 =
$$
\frac{(1 - R_{F})}{n_{v}^{2}}
$$
 (8)

where R_F is water-air Fresnel reflectance and n_w is the index of refraction for salinity of 35 PSU. L_w cannot be calculated if any one of E_s , L_u or K_l is missing. L_W 's can be solar normalized by the solar normalization factor, F_N , which is a function of

$$
L_{WN} = \frac{L_W}{F_N} \tag{9}
$$

atmospheric Rayleigh and ozone optical thicknesses, τ_R and τ_{oz} , solar zenith angle, θ_0 , and earth-sun distance, r_{es} (Eqs. 9,10,11).

The normalized SeaWiFS response functions, Wt1...Wt8, (Appendix 6) are applied by their spectral multiplication with the solar-normalized water-leaving radiances.

$$
F_N = \frac{t(\lambda, \theta_0) (1 - \rho(\theta_0)) \cos(\theta)}{r_{es}^2}
$$
 (10)

$$
t(\lambda, \theta_{0}) = \exp\left(-\frac{0.5 \tau_{R} + \tau_{oz}}{\cos(\theta_{0})}\right)
$$

The earth-sun distance r_{es} in astronomical units is given approximately by Eq. 11, where D is the Julian day of the year, and the solar zenith angle θ_0 is

$$
r_{es} = 1 + 0.0167 \cos\left(\frac{2\pi (D-3)}{365}\right)
$$
\n(11)

determined from time of the Top L_{WN} , station latitude and longitude from MLML DBASE library procedures based on the solar ephemeris of Von Flanderan and Pulkkinen (1979).

The eight SeaWiFS-weighted solar-normalized water-leaving radiances along with the Photosynthetically active radiation, PAR (Eq. 12), the solar zenith angle, and earth-sun distance are saved in the MOS *File Header Extensions.*

$$
PAR = \sum_{\lambda=400}^{\lambda=700} \frac{\lambda}{\hat{hc}} E(\lambda) d\lambda
$$
 (12)

All 41 *Variables* are setup with *Header* information and scale parameters tailored to the *Var* type and units. Default scale ranges are set for L_u , E_d , E_s and L_w 's while K_L and K_E scales are adjusted for their ranges. Time of each E_s set is retrieved from *Var Header Extension,* converted, formatted, and saved in the *Var's* description. Converted and corrected depth of each L_u and E_d scan is saved in their *Var* descriptions. If average incident ratios were used, the R_N value is recorded with the diffuse attenuation *Var* description.

Because all variables have been formed by previous

processing the *DERIVE* program requires only two input arguments and operates on the currently loaded data:

NOAA> DERIVE

NOAA DERIVE: Derived Radiometric Parameters: rev: 22-Jul-1993

Var# 1-26: Lambda*2; Ed,SNR,Es,SNR,Lu,SNR,Es,SNR*3 Derives Var# 27-41: Ke,Kl *3; Lw *4; Lwn *4

NOAA DERIVE: Enter Data Units 1=Raw 2=Cnvrted: 1

NOAA DERIVE: Enter Ratio Type 1=Mean 2=Spctrl: 1

Working... Done.

NOAA GRAPH

Graphical output of spectral data is provided via *GRAPH* (Figs 1 and 2). Plots can be sent to a graphics terminal or PostScript ftle or printer. Color or black-and-white plots are available. Two pages are produced, each with three X-Y plots. Page one shows E_d , L_u and E_s (Fig. 1), and page two shows K_E , K_L and L_w (Fig. 2). X-axis wavelength range is selectable as either full-scale, 300 to 900 nm, or clipped, 400 to 700 nm. Y-axis units are selectable between raw units (counts/sec) and (counts), or converted units $(\mu W/cm^2/nm)$ and $(\mu W/cm^2/sr/nm)$. Cruise and station name, position and date (found in *File Header* strings# 1 to 3) are labeled on each page, and the user can enter an optional page title and footer to be printed. *Vars* 1 to 36 are assumed to be in the order previously described, however empty *Vars* are acceptable and do not plot.

Plotting data... Done.

NOAA LIST

Tabular output of spectral data is provided by *LIST.* The output device is selectable as either the user's terminal, a printer, or an ASCII file. *LIST* prints columns of user selectable E_d , L_u , E_s , K_E , K_L , L_{wsn} data and rows of wavelengths at user specified intervals. Two such tables are normally produced: the first showing observed radiance and irradiance spectra for the Top, Mid and Bot depths (Fig. 3); the other showing derived attenuation coefficient and waterleaving radiance spectra (Fig. 4). Cruise and station information is labeled at the top of each page along with Secchi disk depth, Munsell color, SeaWiFSweighted solar-normalized water-leaving radiances, PAR, solar zenith angle, and an optional user-entered title and page footer.

Working... Done.

Batch Mode Processing

NOAA/MLML data processing may be done in batch-mode to automate repetitive steps. A text file is created containing keystrokes used to invoke the NOAA programs and respond to program prompts. For example, a raw MOS scan-set is loaded and the first four processing programs run sequentially by the following batch file:

```
! Load the raw scan-set from disk Archive file 
load 
! 
! apply MOS system response to Var#'s 2-10 
noaa convert moo_response 2 2-10 
! 
1apply MOS immersion factor to Var#'s 2-10 
noaa convert mos_response 3 2-10
!
```

```
! throw out MOS bad data in Var#'s 2-10 
noaa edit * * 
! 
! smooth wi 5 pt running mean Var#'s 2-10 
noaa smooth 2-10 * 
! 
! subtract dark counts, get %Error 
noaa adjust * * * * * 
! 
! plot the results 
graph 
! 
! record processed set 
record
```
Other batch files can be created to process SIS sets, *A VERAGE* multiple sets, *MERGE* processed sets, and *DERIVE* radiometric quantities. If the above file's name is PROC1.MLCOM, it is run by typing:

ML> @PROC1.MLCOM

Any lines beginning with the exclamation mark in a batch file are comments and are not executed. When a *PLOT* or *NOAA* command is missing a parameter, or is supplied an improper parameter, the program will prompt the user for necessary input. In the above batch file, for example, the *LOAD* command will prompt the user for the filename and *Variable* range to load, and at which *Workspace Variable* to begin loading. Notice the use of asterisks (*) to use default values.

NOAA PROCESS

Batch fIle processing can be automated for a set of fue names by using the ALIAS command. If the ALIAS file SIS.ALIAS is created at the VMS command level,

\$ ALIAS *SIS*.MLDAT SIS

then all the files referenced by SISALIAS can be manipulated by NOAA_PROCESS program. The aliased files are referenced by an apostrophe followed by the alias file name, for example,

'SISALIAS

The NOAA PROCESS program can invoke all NOAA programs (Table 1), which derive much of their flexibility from their ability to operate on lists of files.

Acknowledgements

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Appendix 1. SeaWiFS standard radiometry symbols.

Appendix 1. (continued) SeaWiFS standard radiometry symbols.

Adapted from Mueller, et a1. 1991.

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Appendix 2. Glossary of acronyms used in text.

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Appendix 3a. The MLML_DBASE data structure. Variable types use FORTRAN notation. Pointers to variable are shown by *. Defined in [MLML_LIBRARY.MLCLIB]MLML_DBASE.H

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Appendix 3c. Data Structures used with SIS. Variable types use C notation. Pointers to variables are shown by *. The structure, SIS_VAR_INFO, △ ۞
Is an MLML_DBASE <u>File Header Extension</u>. The structure, SIS_DATA_BLOCK, is is an MLML_DBASE <u>file Header Extension</u>. The structure, SIS_DATA_BLOCK, is an MLML_DBASE Variable Extension and accompanies each SIS
spectral Variable.

Defined in NOAA\$SRC:SIS.C

Defined in NOAA\$INC:SIS_TYPES.H

(Each pointer in SIS_DATA_BLOCK references a data element of type long that accompanies each SIS spectrum)

g.

ம்

Defined in NOAA\$INC:FASTIE_TYPES.H

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Data structures used with reduced NOAA radiometer files. Variable types use C notation. Pointers to variables are indicated by *. Appendix 3e. The structure, RADIOM FILE INFO, is an MLML DBASE File Header Extension and accompanies each file. and accompanies each PLOT Variable which contains the spectral data.

(continued) Data structures used with reduced NOAA radiometer files. Variable types use C notation. Pointers to variables are Appendix 3e. indicated by *. The structure, The structure, RADIOM_VAR_INFO, is an MLML_DBASE Variable Extension and accompanies each PLOT Variable which contains the spectral data.

Defined in NOAA\$DIR:RADIOM TYPES.H

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Appendix 4. Spectral quantities used in radiometer algorithms. a) Air-water radiance immersion factors (Austin, 1976)

$$
n = n_0 + \frac{C}{\lambda - \lambda_0}
$$

$$
F_w = n_w \frac{(n_w + n_g)^2}{(1 + n_g)^2}
$$

Air-Water Radiance Factors

550.3 4.820E-2 550.9 4.854e-2 667.1 4.576E-2 667.8 4.594E-2 761.0 2.399E-2 761.6 2.398e-2 859.3 2.470E-2

Appendix 4. Spectral quantities used in radiometer algorithms.
b) SeaWiFS Normalized Response Functions as applie

385.8 O.OOOE+O 409.7 O.OOOE+O 462.3 O.OOOE+O 480.2 O.OOOE+O 525.2 O.OOOE+O 639.8 O.OOOE+O 734.9 O.OOOE+O 835.0 O.OOOE+O 386.4 5.984E-5 410.4 6

415.4 1

421.7

 433.0 433.6

434.2

436.1 3.450E-2 436.7 3.601E-2 488.2 4.412E-7 488.8 4.365E-8 505.9 4.141e-2 506.5 4.112E-2

387.0 1.438E-4 387.7 2.658E-4 388.3 4.000E-4

 388.9 2.151E-4 389.6 6.392E-5 390.2 1.510E-4

390.8 5.001E-4

392.7 3.024E-5 393.3 6.146E-4

394.6 2.685E-4 395.2 3.787E-4 395.9 8.087E-4 396.5 1.215E-3 397.1 1.690E-3 397.8 2.315E-3

404.7 3.219E-2 428.6 4

408.5 4.109E-2 409.1 4.376E-2 409.7 4.242E-2

410.4 4.260E-2

Appendix 4. Spectral quantities used in radiometer algorithms. b) (continued) SeaWiFS Normalized Response Functions as applied to the prototype MOS.

b) (continued) SeaWiFS Normalized Response Functions as applied to the prototype MOS. λ Band 1 λ Band 2 λ Band 3 λ Band 4 λ Band 5 λ Band 6 λ Band 7 λ 464.7 3.528E-4 465.4 2.790E-4 466.0 2.707E-4 466.6 1.785E-4 467.2 7.003E-5 467.8 5.464E-5 468.5 4.399E-5 469.1 8.728E-S 469.7 2.362E-5 470.3 O.OOOE+O 516.2 1.711E-3 516.8 517.4 518.0 1.087E-6 518.6 1.266E-7 519.2 9.967E-8 519.8 6.882E-9 520.4 O.OOOE+O 1.265E-4 1.426E-5 533.7 534.3 534.8 535.4 536.0 536.6 537.2 537.8 538.4 539.0 1.497E-5 539.6 3.778E-5 540.2 O.OOOE+O 1.438E-4 1.041E-4 7.188E-5 5.647E-5 3.891E-5 3.544E-5 3.812E-5 9.822E-6 1.447E-5 577.3 5.747E-4 577.9 578.5 579.0 3.664E-4 579.6 3.163E-4 580.2 2.253E-4 580.8 2.907E-4 581.3 2.400E-4 581.9 1.958E-4 582.5 1.710E-4 583.1 1.656E-4 583.7 1.688E-4 4.586E-4 4.319E-4 696.8 697.5 698.1 698.8 1.791E-4 699.4 1.256E-4 700.0 1.034E-4 700.7 O.OOOE+O 1.910E-4 2.004E-4 1.520E-4 789.1 789.7 790.3 4.738E-3 790.9 4.275E-3 791.5 3.878E-3 792.1 3.513E-3 792.7 3.202E-3 793.3 2.862e-3 793.9 2.629E-3 794.5 2.404E-3 795.1 1.874e-3 795.7 O.OOOE+O 5.890E-3 5.280E-3 884.8 885.3 885.8 1.676E-2 886.4 1.531E-2 886.9 1.393E-2 887.4 1.253E-2 888.0 1.108E-2 888.5 9.776E-3 889.0 8.573E-3

889.5 7.545E-3 890.1 6.629E-3 890.6 5.849E-3 891.1 5.163E-3 891.6 4.557E-3 892.2 4.077E-3 892.7 3.677E-3 893.2 3.270E-3 893.7 2.919E-3 894.2 2.587E-3 894.8 2.667E-3 895.3 1.364E-3 895.8 O.OOOE+O

Band 8

1.952E-2 1.816E-2

λ (nm)	$\boldsymbol{\tau}_\text{R}$	τ_{oz}		λ (nm)	τ _r	\mathfrak{r}_{oz}
340	0.7062	0.0188		625	0.0585	0.0352
345	0.6653	0.0070		630	0.0566	0.0326
350	0.6273	0.0032		635	0.0548	0.0301
355	0.5919	0.0007		640	0.0531	0.0279
360	0.5590	0.0000		645	0.0514	0.0254
365	0.5283	0.0000		650	0.0498	
370	0.4997	0.0000		655		0.0232
375	0.4730				0.0483	0.0214
380		0.0000		660	0.0468	0.0199
385	0.4480 0.4247	0.0000		665	0.0454	0.0185
390		0.0000		670	0.0440	0.0166
	0.4029	0.0000		675	0.0427	0.0152
395	0.3824	0.0000		680	0.0414	0.0137
400	0.3632	0.0000		685	0.0402	0.0123
405	0.3452	0.0000		690	0.0390	0.0108
410	0.3283	0.0000		695	0.0379	0.0090
415	0.3124	0.0000		700	0.0368	0.0079
420	0.2975	0.0000		705	0.0357	0.0075
425	0.2834	0.0000		710	0.0347	0.0064
430	0.2702	0.0003		715	0.0337	0.0057
435	0.2577	0.0007		720	0.0328	0.0052
440	0.2459	0.0010		725	0.0318	0.0047
445	0.2348	0.0010		730	0.0310	0.0042
450	0.2243	0.0010		735	0.0301	0.0038
455	0.2144	0.0018		740	0.0293	0.0034
460	0.2050	0.0029		745	0.0285	0.0031
465	0.1962	0.0029		750	0.0277	0.0028
470	0.1878	0.0025		755	0.0270	0.0025
475	0.1798	0.0043		760	0.0263	0.0022
480	0.1723	0.0065		765	0.0256	0.0020
485	0.1651	0.0072		770	0.0249	0.0018
490	0.1583	0.0065		775	0.0242	0.0016
495	0.1519	0.0079		780	0.0236	0.0015
500	0.1458	0.0101		785	0.0230	0.0013
505	0.1399	0.0137		790	0.0224	0.0012
510	0.1344	0.0145		795	0.0218	0.0011
515	0.1292	0.0137		800	0.0213	0.0009
520	0.1241	0.0163		805	0.0207	0.0008
525	0.1194	0.0195		810	0.0202	0.0008
530	0.1148	0.0228		815	0.0197	0.0007
535	0.1105	0.0254		820	0.0192	0.0006
540	0.1064	0.0261		825	0.0188	0.0005
545	0.1024	0.0279		830	0.0183	0.0005
550	0.0987	0.0304		835	0.0178	0.0004
555	0.0951	0.0323		840	0.0174	0.0004
560	0.0917	0.0348		845	0.0170	0.0003
565	0.0884	0.0388		850	0.0166	0.0003
570	0.0853	0.0421		855	0.0162	0.0003
575	0.0823	0.0431		860	0.0158	0.0002
580	0.0794	0.0424		865	0.0154	0.0002
585	0.0767	0.0402		870	0.0151	0.0002
590	0.0740	0.0392		875	0.0147	0.0002
595	0.0715	0.0417		880	0.0144	0.0001
600	0.0691	0.0450		885	0.0141	0.0001
605	0.0668	0.0450		890	0.0137	0.0001
610	0.0646	0.0428		895	0.0134	0.0001
615	0.0625	0.0402		900	0.0131	0.0001

Appendix 4. Spectral quantities used in radiometer algorithms. c) Rayleigh and ozone atmospheric optical thicknesses. τ_{oz} from Howard Gordon (personal
communication).

620 0.0604 0.0377

Appendix 5. File Naming Conventions and Directory Structures.

Device Names for Program and Data Directories:

Data Directories:

MLML DBASE Directories

Appendix 5. (continued) File Naming Conventions and Data Directory Structure.

Program Directories:

Appendix 5. (continued) Logical Name Table used in NOAA data processing programs.

VMS Logicals:

Setup by SYSSMANAGER:SETUP_NOAA_LOGICALS.COM:

Setup by SYSSMANAGER:SETUP_CTD_LOGICALS.COM:

Setup by USER\$DEV: [DATA.COM] SETUP.COM:

S DEFINE/EXEC/SYS DATA USERSDEV:[DATA.cruise.]

VMS Symbols:

Setup by USER\$DEV: [NOAA] LOGIN.COM:

- S MOS :== aNOAA\$COM:MOS
\$ SIS :== aNOAA\$COM:SIS
- \$ SIS :== @NOAA\$COM:SIS
\$ FASTIE :== @NOAA\$COM:FAS
- S FASTIE :== QNQAASCOM:fASTIE
- :== SUBMIT/LOG=DATA: [STATION] GPS.LOG/NAME=GPS_LOGGER_NOAA\$COM:START_GPS

Setup by USER\$DEV: [CTD] LOGIN.COM:

\$ CTD :== aCTD\$COM:START_CTD_ACQ
\$ GPS :== SUBMIT/LOG=DATA: ISTATI :== SUBMIT/LOG=DATA: [STATION]GPS.LOG/NAME=GPS_LOGGER NOAA\$COM:START_GPS

Setup by USERSOEV:[CTD]LOGIN.COM:

\$ CTD :== aCTD\$COM:START_CTD_ACQ
\$ GPS :== SUBMIT/LOG=DATA:[STATIO

:== SUBMIT/LOG=DATA: [STATION]GPS.LOG/NAME=GPS_LOGGER NOAA\$COM:START_GPS