Construction of topological defect networks with complex scalar elds

V J. A fonso<sup>a</sup>, D. Bazeia<sup>a</sup>, M. A. Gonzalez Leon<sup>b</sup>, L. Losano<sup>a</sup>, and J. Mateos Guilarte<sup>c</sup>

<sup>a</sup> Departamento de Fisica, Universidade Federal da Paraba, BRAZIL

<sup>b</sup> Departamento de Matematica Aplicada, Universidad de Salamanca, SPAIN

<sup>c</sup> Departamento de Fisica and IUFFyM, Universidad de Salamanca, SPAIN

This work deals with the construction of networks of topological defects in models described by a single complex scalar eld. We take advantage of the deformation procedure recently used to describe kinklike defects in order to build networks of topological defects, which appear from complex eld models with potentials that engender a nite number of isolated minima, both in the case where the minima present discrete symmetry, and in the non symmetric case. We show that the presence of symmetry guide us to the construction of regular networks, while the non symmetric case gives rise to irregular networks which spread throughout the complex eld space. We also discuss bifurcation, a phenomenon that appear in the non symmetric case, but is washed out by the deformation procedure used in the present work.

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#### I. INTRODUCTION

Defect structures have appeared in high energy physics almost fly years ago, with some of the pioneer results collected in Refs. [1, 2, 3, 4]. A long the years, the subject has grown in importance, accompanied with an increasing number of investigations on kinks in one spatial dimension, vortices in two dimensions and monopoles in three dimensions, among other topological defects { see, e.g. [5] for an extensive discussion of some of the most important results in the area.

The classical solutions which represent the defect structures can be of topological or non topological nature, and here we will deal with perhaps the sim plest topological structures, which appear in models of scalar elds. To be specic, we will consider models of the Wess-Zum ino type, described by a single complex scalar eld in the presence of discrete symmetry and in the more general case which engenders no specic symmetry. Some of the models have been studied before in [6, 7] { see also [8] for related issues { with particular attention to the presence and stability of kinklike defects and junctions, and in [9], where the kink orbits are written in terms of real algebraic curves and the equations of motion are shown to be fully expressed in terms of rst order dierential equations of the Bogom ol'nyi-Prasad-Sommer eld (BPS) type [4].

The kinklike structures have been used in many dierent contexts, in (1;1) and in higher space-time dimensions, in particular in the form of junctions and networks of defects [7, 10, 11, 12]. In (3;1) dimensions they are usually named domain walls, which can indeplications in several distinct scenarios, in particular as seeds for the form ation of structures in the early Universe. In this context, although the standard scenario seems to show that the presence of domain walls has little to contribute to the cosmic evolution, it has been suggested that domain walls may perhaps be used as a source for the dark energy necessary to feed the current cosmic acceleration [13].

A nother line of research has recently appeared in gravity in higher dimensions [14, 15], with the hope to solve the hierarchy and other problems in high energy physics. In (4;1) dimensions, the braneworld model with warped geometry involving a single extra dimension of in nite extent suggested in [15] has strongly in pacted the subject. In this braneworld scenario, the inclusion of scalar elds may contribute to smoothen the brane [16], to give rise to a diversity of situations of current interest, as one can see, for instance, in the recent investigations [17].

The present study is a continuation of a form erwork [18]. Here we will focus mainly on the deformation procedure introduced in [19], and extended to other scenarios in [20,21]. Those investigations have led us to india peculiar and very interesting feature of the deformation procedure there in plemented. The issue is that it is sometimes possible to deform a given model described by a potential containing some minima, to get to another model, with the potential giving rise to a different set of minima, which may increase periodically. This feature strongly suggests the possibility of using the deformation procedure to build lattices of minima in the two-dimensional eld space.

An interesting property of the deform ation procedure is that it also constructs the kink like solution of the deform ed model in terms of the kink solution of the original model. Thus, in the lattice of minima we can then nest a network of defects very naturally, that is, as internal feature of the deformation itself. This is the idea underlying this paper, in which we apply the deformation procedure to investigate the generation of networks of kinklike defects for the deformed models, which are expanded networks. Although it is possible to start with the more general case, considering models with an arbitrary set of minima, we shall rstly deal with the case involving N minima in a  $Z_N$  symmetric arrangement. We shall consider the symmetric N = 2, N = 3 and N = 4 cases explicitly, and later we relax

the constraint to deal with three and four minima in the non symmetric case. The main reason for this is that we want to keep the motivation set forward in our former work [18], where we have investigated the construction of regular networks. Moreover, as a pedagogical concern we believe that this route makes the problem easier to understand.

The idea of constructing networks of defects is not new, but the novelty here relies on the use of the deform ation procedure as a simple and natural way to generate networks. The mechanism is powerful and suggestive, and fully motivates the present work. To make it short, direct, we have decided to consider models of the Wess-Zum ino type, driven by a single complex scalar eld. These models are popular, of great importance and easy to manipulate, and so they very much help us to highlight the idea to be explored below.

We start the investigation in Section II, where we introduce the symmetric models and perform the deformation procedure on general grounds. In Section III we illustrate the procedure with some applications, considering two important cases, which engender three and four minima, forming an equilateral triangle and a square, respectively. There we show how the deformed models tile the plane replicating the sets of minima in the entire eld plane. We then consider other possibilities in Section IV, and there we deal with more general models which three and four minima, engendering no symmetry anymore. We use the deformation procedure to get to many distinct and interesting patterns. The more general case allows for a new phenomenon, bifurcation, and so in Section V we deal with bifurcation, which concerns the possibility of the system to allow for two or more distinct connections between two given minima. This is related to the marginal stability curve, and has to do with the energy balance involving distinct orbits in eld space, as already investigated in [7]. We end the paper in Section VI, where we introduce some comments and conclusions.

## II. DEFORMATION OF WESS-ZUMINO MODELS

In this Section we start with a brief sum mary of the bosonic sector of the standard W ess-Zum ino model engendering the D  $_{\rm N}$  sym metry. We then propose a simple but very interesting way to deform the model, to generate an in nite family of new Wess-Zum ino like models with their defect solutions.

## A. The general case

Let  $(x;t) = {}_{1}(x;t) + {}_{2}(x;t)$  be a complex scalar eld, written in terms of the two real partners  ${}_{1}(x;t)$  and  ${}_{2}(x;t)$  in (1;1) spacetimed in ensions. The dynamics of the bosonic sector of Wess-Zum in omodels is governed by the Lagrange density

$$L = \frac{1}{2} @ @ - V(;-)$$
 (1)

where the bar stands for com plex conjugation. We refer to these systems as Wess-Zumino, or Landau-Ginzburg, models if the potential energy density is determined from an holomorphic superpotential Wood such that the potential energy density of the scalar eld theory reads

$$V(\ ; \ ) = \frac{1}{2}W^{0}(\ )\overline{W^{0}(\ )}$$
 (2)

The interest of these models lies in the fact that all of them adm it a supersymmetric version with N=2 extended supersymmetry. It is also important to stress that there is a U (1) am biguity in the election of the superpotential:  $W = e^{i} + W = e^{i} +$ 

 $\underline{W}$  e shall rstly consider polynom ials of degree N + 1 in with real coe cients as superpotentials, in this case  $\underline{W}$  ( ) =  $\underline{W}$  ( ) is the same function of the conjugate complex eld. The vacua manifold, the set of zeros of V ( ;  $\overline{\phantom{W}}$ ), is given by the critical points of the superpotential, the N roots of the polynomial  $\underline{W}$  ( )

$$^{(j)}(x) = v^{(j)}; \quad W^{0}(v^{(j)}) = 0; \quad j = 1;2;$$
 ;N (3)

The BPS static kinks satisfy the system of rst-order ordinary di erential equations

$$\frac{d}{dx} = \overline{W^{0}()}; \quad \frac{d}{dx} = W^{0}()$$
 (4)

We use these expressions to show that the eld prolles and orbits should obey

$$dx = \frac{d}{\overline{W^{0}()}} = \frac{d}{\overline{W^{0}()}}; \quad W^{0}()d \quad \overline{W^{0}()}d = 0$$
 (5)

Equation (5) (right) m eans that  $d(W \overline{W}) = 0$  such that

$$Im W ((x)) = constant$$
 (6)

for the eld orbits. These orbits are kink orbits if they connect two vacua

$$\operatorname{Im} W_{(kj)}(v^{(k)}) = \operatorname{Im} W_{(kj)}(v^{(j)}) = \operatorname{constant}$$

$$\tag{7}$$

This criterion sets  $^{(k\,j)}$  by requiring that W  $(v^{(k)})$  W  $(v^{(j)}$  e  $^{i}$   $^{(k\,j)}$ ) be real, or

$${}^{(kj)} = \arctan \frac{ \frac{\mathbb{I} m \ \mathbb{W} \ (v^{(k)}) \ \mathbb{W} \ (v^{(j)})}{\mathbb{R} e \ \mathbb{W} \ (v^{(k)}) \ \mathbb{W} \ (v^{(j)})}} \text{ mod } ; \qquad {}^{(jk)} = {}^{(kj)} +$$
 (8)

Integration of (5) (left) gives

$$x \quad x_0 = \frac{1}{2} \quad \frac{d}{\overline{W^0}} + \frac{d}{\overline{W^0}}$$
 (9)

Interpretation of the meaning of this integral is reached from the identity between dierential one-forms

$$W^{0}()d + \overline{W^{0}()}d = 2 \dot{y} V^{0}() \dot{j} dx ) d(W + \overline{W}) = 2 \dot{y} V^{0}() \dot{j} dx ) ReW = \dot{y} V^{0}() \dot{j} dx = s (10)$$

The kink pro less are then obtained by inverting these relations between the real part of the superpotential and the \length" s on the kink orbits (6) { see, e.g., Ref. [9].

The energy of the static con gurations can be written a la Bogom ol'nyi in the form

$$E = \frac{1}{2}^{Z} dx \frac{d}{dx} \frac{W^{0}()^{2} + \frac{1}{2}^{Z} d(W + \overline{W})$$
 (11)

which shows that the solutions of the rst-order equations (4) for the kink phases (kj) (8) have energies given by

$$\texttt{M} \ (k \ j) = \ \texttt{R} \ \texttt{eW} \ \ _{(k \ j)} \ (v^{(k)}) \quad \ \ \, \\ \texttt{R} \ \texttt{eW} \ \ _{(k \ j)} \ (v^{(j)}) = \ \ \ \ \, \\ \texttt{W} \ (v^{(k)}) \quad \ \ \, \\ \texttt{W} \ (v^{(j)})$$

B. Sym m etric W ess-Zum ino m odels

The choice of the holom orphic superpotential in the form

W () = (x;t) 
$$\frac{N+1(x;t)}{N+1}$$
; N 2 N (12)

leads to the potential energy density

$$V(\ ; \overline{\ }) = \frac{1}{2} \ 1 \qquad {}^{N} (x;t) \ 1 \qquad {}^{-N} (x;t)$$
 (13)

$$(k)(x;t) = v^{(k)} = \exp(2 i(k 1) = N); \quad k = 1;2; \dots; N$$
 (14)

the D  $_{\rm N}$  sym m etry is spontaneously broken to the complex conjugation Z  $_{\rm 2}$  sub-group at every vacuum state. The BPS static kinks satisfy the system of rst-order ordinary di erential equations

$$\frac{d}{dx} = e^{i} (1 - (x)); \quad \frac{d}{dx} = e^{i} (1 - (x))$$
 (15)

These equations can also be written as

$$dx = \frac{e^{i} d}{1 N} = \frac{e^{i} d}{1 N}; e^{i} (1^{N}) d e^{i} (1^{N}) d = 0$$
 (16)

Thus, the real algebraic curves which solve (16) (right)

Im 
$$e^{i}$$
  $(x)$   $\frac{N+1}{N+1} = constant$  (17)

are the orbits of the solutions. Kink orbits pass through two minim a of the potential, so we have

Im 
$$e^{i} v^{(k)} = \frac{(v^{(k)})^{N+1}}{N+1} = \frac{N}{(N+1)} \sin(\frac{2}{N}(k-1)) = \text{constant}$$
 (18)

and

where  $^{(jk)} = ^{(kj)} +$ 

Integration of (16) (left) gives

$$x \quad x_0 = \frac{1}{2}^{Z} \quad \frac{d}{e^i (1 - x_0(x))} + \frac{d}{e^i (1 - x_0(x))}$$
(20)

This integral can be written in terms of the local parameter  $\frac{ds}{dx} = e^{-i} (1 - N(x))^2$  in order to implicitly obtain the kink proles

Re 
$$e^{i}$$
 (1  $N(x)$ ) =  $e^{i}$  (1  $N(x)$ )  $dx = s$  (21)

The kink energies are

$$M (kj) = \frac{2N}{N+1} \sin \frac{1}{N} (k j)$$
 (22)

In [6] it was shown that this superpotential in the N = 2 supersymm etric Landau-G inzburg action is an integrable deform ation of the N = 2 supersymm etric m in in al  $A_N$  series of conform almodels. It was also suggested in [6] the connection with the solitons of the a ne Toda  $A_N$  eld theories { see Ref. [8] for details { which can be directly envisaged in the above expression (22).

## C. The deform ation procedure

We now turn attention to the deform ation procedure, which will allow us to obtain new Wess-Zum in olike models. A coording to Refs. [19, 20], we will express the deformed system in terms of a new complex eld (x;t) = 1(x;t) + 1(x;t); related to the original one by means of the (a priori) holomorphic function f(x;t) = 1(x;t) + 1(x;t) + 1(x;t)

$$= f() = f_1(_1;_2) + if_2(_1;_2)$$
 (23)

This function has to obey

$$\frac{\partial f_1}{\partial x_1} = \frac{\partial f_2}{\partial x_2}; \quad \frac{\partial f_1}{\partial x_2} = \frac{\partial f_2}{\partial x_1}$$
 (24)

The rst-order equations become

$$\frac{\mathrm{d}}{\mathrm{dx}} = e^{i} \frac{\overline{W^{0}(f(\cdot))}}{f^{0}(\cdot)}; \quad \frac{\mathrm{d}}{\mathrm{dx}} = e^{i} \frac{\overline{W}^{0}(f(\cdot))}{\overline{f^{0}(\cdot)}}$$
(25)

that we choose to understand as determ ining the absolute energy m inim a associated to the \deformed" Lagrange density

The dynam ics governed by L and  $L_D$  are dierent, but we can de ne V ( ; ) and W ( ) by

$$V(\ ;\ ) = \frac{V(f(\ );\overline{f(\ )})}{f(\ )} = \frac{1W^{0}(f(\ ))}{2\overline{f(\ )}} = \frac{1W^{0}(f(\ ))}{f^{0}(\ )} = \frac{1}{2}W^{0}(\ )\overline{W^{0}(\ )}$$
(27)

such that the \deform ed" rst-order equations (25) are

$$\frac{\mathrm{d}}{\mathrm{dx}} = e^{i} \overline{W^{0}()}; \quad \frac{\mathrm{d}}{\mathrm{dx}} = e^{i} W^{0}()$$
 (28)

The BPS kink solutions for this system are obtained from the solutions of (15) by simply taking the inverse of the deformation function:  $K(x) = f^{-1}(K(x))$ . Thus, we can make the following relation between the deformed and original equations: if K(x) is a kinklike solution of the original model, we have that

$$Im W (K(x)) = constant; ReW (K(x)) = s$$
 (29)

and so we get that  $K(x) = f^{-1}(K(x))$  is kinklike solution of the deform ed model, obeying

$$Im W (f^{-1}(K(x))) = constant; ReW (f^{-1}(K(x))) = (30)$$

where is de ned by

$$Z = \hat{W}^{0} f^{1}(K(x)) \hat{J} dx$$
 (31)

A liternatively, one could understand (25) as the rst-order equations of the original model written in the form

$$L = \frac{1}{2} f^{0}() \overline{f^{0}()} = V(f(); \overline{f()})$$
(32)

This interpretation means that the original system in the new variables appears as a nonlinear sigma model with target space a non-compact Riemannian manifold with metric

G (; ) = 0 = G--(; ); G -(; ) = 
$$f^0$$
() $\overline{f^0}$ () = G-(; )

The merit of our approach is that we infer the kink solutions of one complicated but interesting eld theoretical model from the well-known kinks of the associated simple system. The other point of view, in which one deforms the metric rather than the potential energy density, is sometimes also interesting. Despite dealing with the same model, the use of appropriate coordinates in eld space may lead to separation of variables in the rst-order equations, sometimes reducing its integration to quadratures; see eq., [22].

Inspired by form er investigations on the deform ation procedure, we select the deform ation function as being equal to the new superpotential, that is, we choose  $f(\ )=W(\ )$ . This choice constrains the function  $f(\ )$  to obey the equation

$$f^{0}()\overline{f^{0}()} = Q \frac{Q}{2V(f();\overline{f()})}$$
(34)

A function f satisfying this condition assures the relation (27) to be full led and presents the advantage of providing a potential for the new model which is well dened (nite) at the critical points of f (); i.e. the zeros of  $f^0$  (). As a bonus, the procedure leads to a very simple expression for the deformed superpotential.

## III. DEFORMATION OF SYMMETRIC WESS-ZUMINO MODELS

To clarify the general considerations, let us now illustrate the above results with explicit exam ples. We will consider the cases N=2, N=3, and N=4: The case N=2 is simpler, and it is very similar to the deformation used in the rst work in [20] to get to the sine-Gordon model. The cases N=3 and N=4 are harder. The deformation procedure leads to the formation of junctions of kink orbits from the original Wess-Zum ino kinks. Since the original non deformed models engender sets of minima which depict equilateral triangles and squares, respectively, the deformation will then naturally tile the plane, with networks of defect orbits which we name expanded kink networks.

We will solve (34) for the Wess-Zum ino model with solutions of  $f^0$  ()  $f^0$  (1); and for N = 3;4 these solutions are merom orphic functions. The issue here is that the deformation function  $f^0$  (1) fails to be holomorphic in a discrete (in nite) set of points, , and this induces the potential energy density to acquire a countably in nite set of poles (the metric in the target space in the second approach above acquires a countably in nite set of zeros). A nalogous physical systems are described by the elliptic Calogero-Mosermodels (the elliptic tops in the second framework); see e.g., Ref. [23]. The loss of holomorphicity can be avoided by restricting the new eld to take values away from the set , the lattice of poles of  $f^0$  (1). We shall then take  $C^0$  as the -eld space.

A. The case 
$$N = 2$$

In this case, we do not be eld—as a function of the new—eld—in the form—= f(). W ith this, we can use f() to rewrite

$$W () = \frac{1}{3}^{3}; V (; \overline{}) = \frac{1}{2}(1 {}^{2})(1 {}^{\overline{2}})$$
 (35)

as

W (f) = f() 
$$\frac{1}{3}f^3()$$
; V (f) =  $\frac{1}{2}(1 + f^2())(1 + f^2())$  (36)

The deform ation function f ( ) is the new superpotential if we im pose

$$f^{0}()\overline{f^{0}()} = q \frac{q}{(1 - f^{2}())(1 - \overline{f^{2}()})}$$
(37)

The particular choice f ( ) = sin ( ) complies with (37) and leads to the deformed system de ned by

$$\mathbb{W} () = \cos(); \quad \mathbb{V} (;) = \cos()\cos()$$

$$(38)$$

which is the complex sine-Gordon model. Here the rst-order equations are

$$\frac{d}{dx} = e^{i} \cos(\overline{(x)}); \qquad \frac{d}{dx} = e^{i} \cos(x)$$
(39)

The solutions for = 0; are given by

$$(x) = \operatorname{gd}(x) + 2n = \arcsin(\tanh(x)) + 2n \tag{40}$$

where gd stands for the Gudermannian function [24], and n is an integer. Here the kinks are analytic solutions and the superpotential is holomorphic.

B. The case N = 3

In the N = 3 case, we have

$$W () = \frac{1}{4}^{4}; V (; \overline{}) = \frac{1}{2}(1)^{3}(1)^{-3}$$

and putting = f(), we rewrite this form ula in the form

$$W(f) = f() \frac{1}{4}f^{4}(); \quad V(f) = \frac{1}{2}(1 f^{3}())(1 \overline{f^{3}()})$$
 (42)

As stated in (34), the deformation function must then satisfy

$$f^{0}()\overline{f^{0}()} = q \frac{q}{(1 + f^{3}())(1 + \overline{f^{3}()})}$$
(43)

W e choose in particular the holomorphic solution of (43) which satis es the separated equations

$$f^{0}()^{2} = f()^{3} + 1; \quad \overline{f^{0}()^{2}} = \overline{f()^{3}} + 1$$
 (44)

The solution of (44), henceforth a solution of (43), is the equianharm onic case of the W eierstrass P function

$$W () = f() = 4^{\frac{1}{3}} P (4^{\frac{1}{3}};0;1)$$
 (45)

Recall that the Weierstrass P function { see [24] { is dened as the solution of the ODE

$$(P^{0}(z))^{2} = 4P^{3}(z) \quad g_{2}P(z) \quad g_{3}$$
 (46)

The W eierstrass P  $(z;g_2;g_3)$  elliptic function and its derivative that solve the di erential equation above are doubly periodic functions de ned as the series

$$P(z) = \frac{1}{z^2} + \frac{X}{(z + 2m!_1 + 2n!_2)^2} = \frac{1}{(2m!_1 + 2n!_2)^2}$$
(47a)

$$P(z) = \frac{1}{z^{2}} + \frac{X}{m_{pn}} \frac{1}{(z + 2m!_{1} + 2n!_{2})^{2}} \frac{1}{(2m!_{1} + 2n!_{2})^{2}}$$

$$P^{0}(z) = \frac{2}{z^{3}} \frac{X}{m_{pn}} \frac{1}{(z + 2m!_{1} + 2n!_{2})^{3}}$$
(47a)

with m; n 2 Z and m<sup>2</sup> + n<sup>2</sup> 6 0: Therefore, the deformation function is, up to a factor, the Weierstrass P function with invariants  $g_2 = 0$  and  $g_3 = 1$ , and we denote it by  $P_{01}(z)$ . This function is merom orphic, with an in nite num ber of poles congruent to the irreducible pole of order two in the fundam ental period parallelogram (FPP). Thus, we suppose that the -eld takes values away from the set of points 3 in order to make the new superpotential holomorphic in the N = 3 case.

A brief rem inder of the essential properties of P  $_{01}$  and P  $_{01}^{0}$  is the following:

1.  $P_{01}(4^{-\frac{1}{3}})$  is a single-valued doubly periodic function with primitive periods:  $2!_1$  and  $2!_3$ . De ning  $!_2$  =  $4^{\frac{1}{3}}$  3 (1=3)=4, the primitive half-periods are

$$!_{1} = !_{2} \frac{1}{2} i \frac{p_{3}!}{2}; \quad !_{3} = !_{2} \frac{1}{2} + i \frac{p_{3}!}{2}$$
 (48)

These periods determ in the FPP. Note that only two of the half periods are irreducible:  $!_2 = !_1 + !_3$ .

- 2.  $P_{01}(4^{-\frac{1}{3}})$  has only one pole of order two at = 0 in the FPP.
- 3. The values of  $P_{01}$  (4  $\frac{1}{3}$  ) at the half-periods are

$$P_{01}(4^{\frac{1}{3}}!_{1}) = 4^{\frac{1}{3}}; \quad P_{01}(4^{\frac{1}{3}}!_{2}) = 4^{\frac{1}{3}} \frac{1}{2} + i \frac{p_{3}^{-1}}{2}; \quad P_{01}(4^{\frac{1}{3}}!_{3}) = 4^{\frac{1}{3}} \frac{1}{2} i \frac{p_{3}^{-1}}{2}$$
(49)

4. The zeros of the derivative  $P_{01}^{0}$  (4  $^{\frac{1}{3}}$  ) in the FPP are at the half-periods of  $P_{01}$ 

$$P_{01}^{0}(4^{-\frac{1}{3}}!_{1}) = P_{01}^{0}(4^{-\frac{1}{3}}!_{2}) = P_{01}^{0}(4^{-\frac{1}{3}}!_{3}) = 0$$
(50)

and the origin is a third-order pole of  $P_{01}^{0}(z)$ .

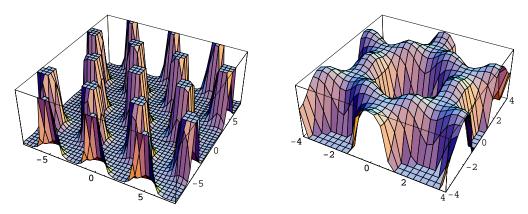


Figure 1: (Color online) The sym m etric case N = 3.3D graphics of the potential V(; ) (left panel), and of V(; ) near a point in the lattice  $_3$  (right panel). Note that in the right panel the zeros are now maxima.

W ith these ingredients we write the deformed potential

$$V(; ) = \frac{1}{2}^{q} \frac{1}{(1 - 4P_{01}^{3}(4^{\frac{1}{3}}))(1 - 4P_{01}^{3}(4^{\frac{1}{3}}))} = \frac{1}{2}P_{01}^{0}(4^{\frac{1}{3}})P_{01}^{0}(4^{\frac{1}{3}})$$
(51)

which is depicted in Fig. 1.

The new potential is doubly periodic with an structure inherited from the half-periods" of P. The set of zeros of the potential in the FPP (see Fig. 2) has three elements

$${}^{(1)} = !_1 = !_2 \quad \frac{1}{2} \quad \frac{p_{\overline{3}}!}{2}; \qquad {}^{(2)} = !_3 = !_2 \quad \frac{1}{2} + i \frac{p_{\overline{3}}!}{2}; \qquad {}^{(3)} = !_2 = 4^{\frac{1}{3}} \frac{{}^{3}(\frac{1}{3})}{4}$$
 (52)

The set of all the zeros of V form a lattice (see Fig. 3) which tile the entire con guration plane

$${\binom{1}{(m,n)}} = !_2 m + n + \frac{1}{2} + {\stackrel{p}{3}}i(m n \frac{1}{2})$$
 (53a)

$$\binom{(2)}{(m, n)} = !_2 m + n + \frac{1}{2} + \frac{p}{3} i(m n + \frac{1}{2})$$
 (53b)

$${}^{(3)}_{(m,p_1)} = !_2 m + n + 1 + {}^{p} \overline{3} i(m n)$$
 (53c)

The values of the superpotential at the m in im a are

$$W \quad \left( \begin{array}{c} (1) \\ (m,n) \end{array} \right) = e^{i} \quad ; \quad W \quad \left( \begin{array}{c} (2) \\ (m,n) \end{array} \right) = ie^{i\left( \begin{array}{c} \overline{6} \end{array} \right)} ; \quad W \quad \left( \begin{array}{c} (3) \\ (m,n) \end{array} \right) = ie^{i\left( \begin{array}{c} \overline{6} \end{array} \right)} \tag{54}$$

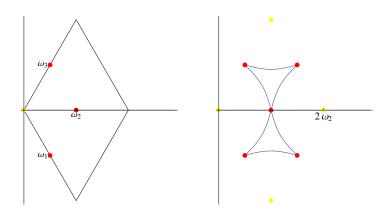


Figure 2: (Color online) The sym metric case N=3: Zeros (red) of V and points (yellow) of the set  $_3$ , and kink orbits (blue) connecting the zeros of the potential (right panel).

In sum, the potential obtained from the deform ation procedure has the same zeros in the FPP as the original model. Besides, one pole arises at the origin due to the merom orphic structure of  $P_{01}^0$ ; see Fig. 2 and 3. However, this structure is in nitely repeated in the deformed model, according to the two periods  $!_1$  and  $!_3$  determining the modular parameter  $= !_3 = !_1 = 1 = 2 + i = 3 = 2$  of the Riemann Surface of genus 1 associated with this P-W eierstrass function. As an aside, we note that the modular parameter = 1 = 2 + i = 3 = 2 gives the same Riemann surface because e = (a + b) = (c + d) where

is an element of the modular group SL(2;Z).

Contrarily to the deform ation function chosen in the case N=2, for N=3 the Weierstrass P function is not an entire function, that is, it is not holomorphic in the whole complex plane  $C=m_n=2m!_1+2n!_2$  (m; n 2 Z) is the lattice of points of P and P which are accordingly meromorphic functions. Thus, we suppose that the neweld take values in the space C=3 to avoid the loss of holomorphicity. This point of view is very close to consider the genus 1 R iem ann surface C=0 of modulus  $=!_3=!_2$  minus the origin (the FPP with the edges identified pairwise minus the origin) as the -eld space. Keeping, however, the in nite copies of this space contained in C=3 gives a richer kink structure.

We shall compare the P-kink orbits with the orbits of the original N = 3 polynomial Wess-Zumino model. If  $^{K}$  (x) is a N = 3 solution of (17) and (20) then  $^{K}$  (x) =  $4^{\frac{1}{3}}P_{01}^{-1}(4^{-\frac{1}{3}-K}(x))$  solves

Im 
$$e^{i} 4^{\frac{1}{3}} P_{01} (4^{\frac{1}{3}} K(x)) = constant; Ree^{i} 4^{\frac{1}{3}} P_{01} (4^{\frac{1}{3}} K(x)) =$$
 (55)

w here

$$= P_{01}^{0} \left(4^{\frac{1}{3}} K(x)\right)^{2} dx$$
 (56)

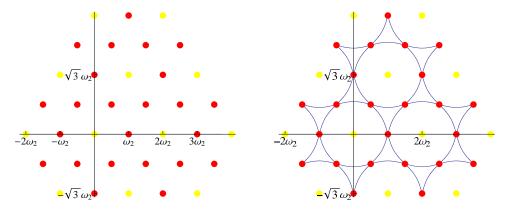


Figure 3: (Color online) The sym metric case N=3: Lattice of zeros (red) of V(;) and points (yellow) of the set  $_3$ , and the network of kink orbits (blue) in the lattice of minima (right panel).

Because of the relations

$$\mathbb{W} \ (^{(k)}) \ \mathbb{W} \ (^{(j)}) = \frac{3}{4} \ \mathbb{W} \ (^{(k)}) \ \mathbb{W} \ (^{(j)}) = \frac{3}{4} \ e^{i\frac{2}{3}(k-1)} \ e^{i\frac{2}{3}(j-1)} \tag{57}$$

the same values of as in the non deformed case,

$$\frac{(kj)}{\sin^{2}\frac{1}{3}(k-1) + \sin^{2}\frac{1}{3}(j-1)}{\cos^{2}\frac{1}{2}(k-1) + \cos^{2}\frac{1}{3}(j-1)}$$
(58)

give the deform ed kink orbits.

There are three types, which we show below.

Type (13), non deform ed

The condition Im W ( $^{(3)}$ ) = Im W ( $^{(1)}$ ) is satisfied only for  $^{(31)}$  = =6 (for antikinks) and  $^{(13)}$  = 7 =6 (for kinks). What is called kink and what is antikink is a matter of convention. Our convention is that kink/antikink orbits run clock/anticlockwise in the (W; W) plane. The orbits obey

$$\frac{3^{p}\overline{3}}{8} \quad \text{ReW}_{\frac{1}{6}}(K) \quad \frac{3^{p}\overline{3}}{8}; \quad \text{Im} W_{\frac{1}{6}}(K) = \frac{3}{8} \text{ with } \text{ReW}_{\frac{1}{6}}(M) = \frac{3^{p}\overline{3}}{8}; \quad \text{ReW}_{\frac{1}{6}}(M) = \frac{3^{p}\overline{3}}{8}$$

and

$$\frac{3^{\frac{p}{3}}}{8} \quad \text{ReW}_{\frac{7}{6}}(^{K}) \quad \frac{3^{\frac{p}{3}}}{8}; \quad \text{Im} \, \text{W}_{\frac{7}{6}}(^{K}) = \frac{3}{8} \quad \text{with} \quad \text{ReW}_{\frac{7}{6}}(^{(1)}) = \frac{3^{\frac{p}{3}}}{8}; \quad \text{ReW}_{\frac{7}{6}}(^{(3)}) = \frac{3^{\frac{p}{3}}}{8}$$

Type (13), deform ed

As mentioned above, the condition  $\operatorname{Im} W$  ( $\binom{(3)}{(\mathfrak{m},\mathfrak{m})}$ ) =  $\operatorname{Im} W$  ( $\binom{(1)}{(\mathfrak{m}^0,\mathfrak{m}^0)}$ ) is again satisfied only for  $\binom{(31)}{(31)} = -6$  and  $\binom{(13)}{(31)} = 7 = 6$ . The orbits obey

$$\frac{p_{\overline{3}}}{2} \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} \text{K} \end{array} \right) \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{Im W }_{\overline{6}} \left( \begin{array}{c} \text{K} \end{array} \right) = \quad \frac{1}{2} \quad \text{with} \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (3) \\ (m, p_1) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (m^0, p_1^0) \end{array} \right) = \quad \frac{p_{\overline{3}}}{2} \; ; \quad \text{ReW }_{\overline{6}} \left( \begin{array}{c} (1) \\ (1)$$

and

$$\frac{P_{\frac{1}{3}}}{2} \quad \text{ReW}_{\frac{7}{6}}(^{\text{K}}) \quad \frac{P_{\frac{3}{3}}}{2}; \quad \text{ImW}_{\frac{7}{6}}(^{\text{K}}) = \frac{1}{2} \quad \text{with} \quad \text{ReW}_{\frac{7}{6}}(^{(1)}_{(\mathfrak{m}^{0},\mathfrak{m}^{0})}) = \quad \frac{P_{\frac{3}{3}}}{2}; \quad \text{ReW}_{\frac{7}{6}}(^{(3)}_{(\mathfrak{m}^{m})}) = \frac{P_{\frac{3}{3}}}{2}$$

The nearest neighbor type (1) and type (3) m in im a are connected by the orbit following the sequence

$$\binom{(1)}{(m,n)}$$
 \$\(\begin{pmatrix} (3) \\ (m,n) \end{pmatrix} \\ \begin{pmatrix} (1) \\ (m+1,n) \end{pmatrix} \\ \begin{pmatrix} (3) \\ (m+1,n) \end{pmatrix} \\ \begin{pmatrix} (1) \\ (m+2,n) \end{pmatrix}

See Fig. 3.

The other two cases (23) and (12) follow similarly. Here we just add that for the case (23), (m;n) and  $(m^0;n^0)$  are restricted to link nearest neighbor type (3) and type (2) m in in a along the orbit. The sequence is

$$(2)$$
  $(3)$   $(3)$   $(2)$   $(m, n)$   $(m, n+1)$   $(m, n+1)$   $(m, n+2)$ 

A lso, for the case (12) we have that (m;n) and  $(m^0;n^0)$  m ust be chosen according to the following sequence

in order to connect nearest neighbor m in im a of type (1) and (2).

We end the case N=3 collecting the corresponding energies. We have that the defect energies for the original non deformed Wess-Zum ino model are given by M=3  $\overline{3}=4$ , for kinks and anti-kinks for all the three sectors, with (kj)=(12); (23); and (13). For the deformed model, the energies of the P-defects are given by  $M=\frac{7}{3}$ ; for the same cases.

C. The case N = 4

In the N = 4 case we deal with

$$W () = \frac{1}{5}^{5}; V (; \overline{}) = \frac{1}{2}(1 {}^{4})(1 {}^{-4})$$
 (59)

Thus, putting = f(), we have

W (f) = f() 
$$\frac{1}{5}f^5()$$
; V =  $\frac{1}{2}(1 f^4())(1 \overline{f^4()})$  (60)

The special deform ation function must satisfy

$$f^{0}()\overline{f^{0}()} = q \frac{q}{(1 + f^{4}())(1 + \overline{f^{4}()})}$$
(61)

A rguing like in (43), we choose the holom orphic solution that satis es the separated equations

$$f^{0}()^{2} = 1$$
  $f()^{4}; \overline{f^{0}()^{2}} = 1$   $\overline{f()^{4}}$  (62)

The solution of (62), henceforth a solution of (61), is the elliptic sine of parameter  $k^2 = 1$ , the Gauss's sinus lem niscaticus

$$W() = f() = sn(; 1)$$
 (63)

As the derivative of the Jacobi elliptic sine is  $\operatorname{sn} u^0 = \operatorname{cn} u \operatorname{dn} u$ , the identities

$$cn^{2}(; 1) = 1 sn^{2}(; 1); dn^{2}(; 1) = 1 + sn^{2}(; 1)$$
 (64)

show the solution of (61) very directly.

The deform ed potential reads

$$V(;) = \frac{1}{2} q \frac{1}{(1 + \sin^4(; 1))(1 + \sin^4(; 1))} = \frac{1}{2} j cn(; 1) j c j dn(; 1) j dn(; 1) dn($$

which is depicted in Fig. 4. The superpotential is given by W ( ) =  $e^{i}$  sn(; 1)

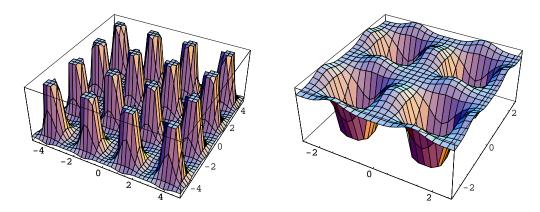


Figure 4: (Color online) The symmetric case N = 4:3D graphics of the potential V(; ) (left panel) and of V(; ) near four points of the set  $_4$  (right panel). Note that in the right panel the zeros are now maxima.

The new potential is doubly periodic with its structure inherited from the \quarter-periods" K (1) =  $!_1$ =4 and iK (2) =  $!_2$ =4 of the twelve Jacobi elliptic functions; see Fig. 5. Here K (1) 1:31103 is the complete elliptic integral of the rst type, a quarter of the length of the lemniscate curve in eld space:  $(\frac{1}{2} + \frac{2}{2})^2 = \frac{1}{2}$ : K (2) 1:31103 il:31103 is the complem entary complete elliptic integral of K ( 1).

The set of zeros of the potential in the FPP are  $^{(1)} = !_1 = 4;$   $^{(2)} = i!_1 = 4;$   $^{(3)} = !_1 = 4;$   $^{(4)} = i!_1 = 4;$  whereas the set of all the zeros of V form a quadrangular lattice in the whole con guration space. They are given by, explicitly

because

$$\operatorname{cn}(\binom{(k)}{(m,m)}; 1) \operatorname{dn}(\binom{(k)}{(m,m)}; 1) = 0; k = 1;2;3;4$$
 (67)

Thus,

$$W (^{(1)}) = e^{i}; W (^{(2)}) = ie^{i}; W (^{(3)}) = e^{i}; W (^{(4)}) = ie^{i}$$
 (68)

since  $\operatorname{sn}[K(1); 1] = 1$  and  $\operatorname{sn}[iK(1); 1] = i$ .

The modular parameter of the associated genus 1 Riemann surface is = iK [2]=K [1]=1+i. Identical Riemann surface is associated to the lemniscatic case,  $g_2 = 1$ ,  $g_3 = 0$ , of the Weierstrass P function. Like in the former case, however, the Jacobi elliptic sine is not an entire function, and so we restrict the new eld to live in C = 4 (where 4is the set of poles of f ( ) in this case) in order to make the superpotential holomorphic.

# C.1. Network of sn-kink orbits

We shall compare the sn-kink orbits with the orbits of the original N = 4 polynomial Wess-Zumino model. If K = 1is a solution of (29) then  $K(x) = \operatorname{sn}^{-1}(K(x); 1)$  solves

Im 
$$e^{i}$$
 sn( $K(x)$ ; 1) = constant; Ree  $i$  sn( $K(x)$ ; 1) = (69)

w here

The same values of as in the non deformed case give the kink orbits

$$W (^{(k)}) W (^{(j)}) = \frac{4}{5} W (^{(k)}) W (^{(j)}) = \frac{4}{5} e^{i\frac{\pi}{2}(k-1)} e^{i\frac{\pi}{2}(j-1)}$$

$$(71)$$

selects

$$\frac{\sin_{\frac{1}{2}}(k - 1) - \sin_{\frac{1}{2}}(j - 1)}{\cos_{\frac{1}{2}}(k - 1) - \cos_{\frac{1}{2}}(j - 1)}$$
; mod (72)

as the angles for both the original and deform ed kink orbits.

There are four types, which we show below.

Type (12)/(34), non deform ed

In this case we have  $\operatorname{Im} W$  (  $^{(1)}$ )=  $\operatorname{Im} W$  (  $^{(2)}$ ) and  $\operatorname{Im} W$  (  $^{(3)}$ )=  $\operatorname{Im} W$  (  $^{(4)}$ ) only for = 3 =4 (kinks) or = 7 =4 (antikinks). The kink orbits obey

$$\frac{2^{p}\overline{2}}{5} \quad \text{ReW}_{\frac{3}{4}}(^{K}) \quad \frac{2^{p}\overline{2}}{5}; \quad \text{Im} \, \text{W}_{\frac{3}{4}}(^{K}) = \frac{2^{p}\overline{2}}{5} \quad \text{with} \quad \text{ReW}_{\frac{3}{4}}(^{(1)}) = \frac{2^{p}\overline{2}}{5}; \quad \text{ReW}_{\frac{3}{4}}(^{(2)}) = \frac{2^{p}\overline{2}}{5} \quad \text{or} \quad \text{ReW}_{\frac{3}{4}}(^{(3)}) = \frac{2^{p}\overline{2}}{5}; \quad \text{ReW}_{\frac{3}{4}}(^{(4)}) = \frac{2^{p}\overline{2}}{5}$$

whereas for the antikinks

$$\frac{2^{p}\overline{2}}{5} \quad \text{ReW } \frac{7}{4} \text{ (}^{K} \text{ )} \quad \frac{2^{p}\overline{2}}{5} \text{ ; Im W } \frac{7}{4} \text{ (}^{K} \text{ )} = \frac{2^{p}\overline{2}}{5} \quad \text{with} \quad \text{ReW } \frac{7}{4} \text{ (}^{(1)} \text{ )} = \frac{2^{p}\overline{2}}{5} \text{ ; ReW } \frac{7}{4} \text{ (}^{(2)} \text{ )} = \frac{2^{p}\overline{2}}{5} \text{ }$$

Type (12)/(34), deform ed

Here we have  $\operatorname{Im} W = (\binom{1}{(\mathfrak{m} \, \mathfrak{m})}) = \operatorname{Im} W = (\binom{2}{(\mathfrak{m}^{\,0} \, \mathfrak{m}^{\,0})})$  and  $\operatorname{Im} W = (\binom{3}{(\mathfrak{m} \, \mathfrak{m})}) = \operatorname{Im} W = (\binom{4}{(\mathfrak{m}^{\,0} \, \mathfrak{m}^{\,0})})$  only for = 3 = 4 (kinks) or = 7 = 4 (antikinks). The kink/antikink orbits obey

$$\frac{p_{\frac{1}{2}}}{2} \quad \text{ReW}_{\frac{3}{4}}(\text{K}) \quad \frac{p_{\frac{1}{2}}}{2}; \quad \text{Im W}_{\frac{3}{4}}(\text{K}) = \frac{p_{\frac{1}{2}}}{2} \text{ with} \quad \text{ReW}_{\frac{3}{4}}(\text{Min}) = \frac{p_{\frac{1}{2}}}{2}; \quad \text{R$$

and

$$\frac{p_{\frac{1}{2}}}{2} \quad \text{ReW}_{\frac{7}{4}}(^{K}) \quad \frac{p_{\frac{1}{2}}}{2}; \quad \text{Im} \, \text{W}_{\frac{7}{4}}(^{K}) = \frac{p_{\frac{1}{2}}}{2} \quad \text{with} \quad \text{ReW}_{\frac{7}{4}}(^{(11)}_{(m_{p_{1}})}) = \frac{p_{\frac{1}{2}}}{2}; \quad \text{ReW}_{\frac{7}{4}}(^{(21)}_{(m_{p_{1}})}) = \frac{p_{\frac{1}{2}}}{2}; \quad \text{ReW}_{\frac{7}{4}}(^{(21)}_{($$

where (m;n) and  $(m^0;n^0)$  are restricted to link nearest neighbor type (1) with type (2) and type (3) with type (4) minima along the orbit. The sequences are

and

$$\binom{(3)}{(m,n)}$$
 \$  $\binom{(4)}{(m,n)}$  \$  $\binom{(3)}{(m-1,n+1)}$  \$  $\binom{(4)}{(m-1,n+1)}$  \$  $\binom{(3)}{(m-2,n+2)}$ 

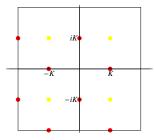
See Fig. 6.

The other three cases, (13), (14)=(23), and (24) follow similarly. Here we just add that in the case (13) we have that (m;n) and  $(m^0;n^0)$  are restricted to link nearest neighbor type (1) and type (3) m in in a along the orbit, and so they must be chosen according to the following sequence

$$\binom{(3)}{(m,m)}$$
 \$  $\binom{(1)}{(m,m)}$  \$  $\binom{(3)}{(m+1,m)}$  \$  $\binom{(1)}{(m+1,m)}$  \$  $\binom{(3)}{(m+2,m)}$ 

In the case (14)/(23) we have that (m;n) and  $(m^0;n^0)$  are restricted to link nearest neighbor type (1) and type (2) m in in a respectively with type (4) and type (3) m in in a along the orbit. Therefore, the sequences are

$$\binom{(1)}{(m,m)}$$
 \$  $\binom{(4)}{(m,m)}$  \$  $\binom{(1)}{(m,m+1)}$  \$  $\binom{(4)}{(m,m+1)}$  \$  $\binom{(1)}{(m,m+2)}$ 



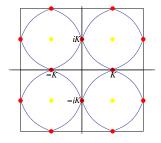
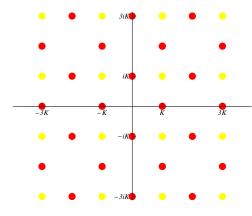


Figure 5: (Color online) The sym metric case N=4: Zeros (red) of V and points (yellow) of the set  $_4$ , and the two (blue and black) possible orbits connecting the zeros of the potential.



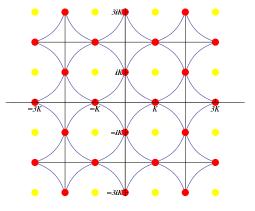


Figure 6: (Color online) The sym metric case N = 4: Lattice of zeros (red) of V(; ) and points (yellow) of the set 4, and the networks of kink orbits (blue and black) in the lattice (right panel).

and

In the case (24) we have that (m;n) and  $(m^0;n^0)$  are restricted to link nearest neighbor type (2) and type (4) m in in a along the orbit. Thus, they must be chosen according to the following sequence

$$\begin{pmatrix} 4 \\ (m, n) \end{pmatrix}$$
  $\begin{pmatrix} 2 \\ (m, n) \end{pmatrix}$   $\begin{pmatrix} 4 \\ (m, 2, n+1) \end{pmatrix}$   $\begin{pmatrix} 2 \\ (m, 2, n+1) \end{pmatrix}$ 

W e end the N = 4 case collecting the energies of the defect structures. For the non deform ed kinks we get

M (13) = M (24) = 
$$\frac{8}{5}$$
; M (12) = M (34) =  $\frac{4^{p}\overline{2}}{5}$ ; M (14) = M (23) =  $\frac{4^{p}\overline{2}}{5}$  (73)

whereas the energies of the deform ed sn-kinks read

$$M (13) = M (24) = 2;$$
  $M (12) = M (34) = \frac{p}{2};$   $M (14) = M (23) = \frac{p}{2}$  (74)

with M (kj) = M (jk) and M (kj) = M (jk).

## IV. DEFORMATION OF ABRAHAM -TOWNSEND MODELS

Let us now move on to the case where the original model engenders no specic symmetry. This study is inspired on Ref. [7], in which Abraham and Townsend consider some interesting situations, guided by more general superpotentials, which develop no specic symmetry. Similar potentials were also considered in [6], but there the investigation was mainly on the symmetric case. In [7], however, the focus was on the non symmetric case, as a basic model underlying the study of intersecting extended objects in supersymmetric eld theories { see also [25], which deals with the forces between soliton states in the same model.

The absence of sym m etry makes the investigation harder to follow, but it is still of current interest since it leads to more general possibilities, bringing somenew elects into the game and allowing for the presence of irregular network of defects. We follow as in the former section, and below we consider models which develop three and four minima with no specific symmetry anymore, using  $e_3$  and  $e_4$  as the set of poles of f () as before, but now in the asymmetric cases, with N = 3 and N = 4, respectively.

#### A. Irregular network of P-kink orbits

Let us start with the sim plest case, in which one considers a class of models that engenders three minima. Here we deal with the superpotential

$$W () = \frac{1}{2} {}^{2} \frac{1}{3} {}^{3} + \frac{1}{4} {}^{4}$$
 (75)

where  $= 1 + i_2$  is a complex coupling constant which parametrizes the family of models. In this case, the potential is given by

$$V() = \frac{1}{2}j$$
  $j$   $j$  (76)

This choice leads to the following set of m inima: two real m inima which are xed to be at  $v_1=1$  and  $v_2=1$ ; and a complex m inimum at  $v_3=1$ , which may move in the complex plane for dierent choices of the complex parameter: This is the most general case with three arbitrary m inima in the complex plane, since one can always choose the straight line joining two vacua as the abscissa axis, crossing the perpendicular ordinate axis through the middle point between the vacua, setting the distance between them to be 2 by an scale transformation. The values that the superpotential W ( ) =  $e^{-\frac{1}{2}}$  W ( ) takes now at the vacua are

W (1) = 
$$\frac{1}{4} \frac{2}{3}$$
 e <sup>i</sup>; W () =  $\frac{1}{12}$  <sup>2</sup>(6 <sup>2</sup>)e <sup>i</sup> (77)

From these expressions we obtain

$$W (1) W (1) = \frac{4}{3} e^{i} ; W (1) W (1) = \frac{1}{12} (3) (1)^{3} e^{i}$$
 (78)

Therefore, the angles of the kink orbits are (mod )

and now the kink energies are given by

$$M (12) = \frac{4}{3} \left( \begin{array}{ccc} 2 & 2 \\ 1 & 2 \end{array} \right)^{1-2}; & M (23) = \frac{1}{12} \left( \left( \begin{array}{ccc} 1 + 3 \right)^2 + \begin{array}{ccc} 2 \\ 2 \end{array} \right) \left( \left( \begin{array}{ccc} 1 & 1 \right)^2 + \begin{array}{ccc} 2 \\ 2 \end{array} \right)^3 & ^{1-2} \\ M (31) = \frac{1}{12} \left( \left( \begin{array}{ccc} 2 & 3 \right)^2 + \begin{array}{ccc} 2 \\ 2 \end{array} \right) \left( \left( \begin{array}{ccc} 2 + 1 \right)^2 + \begin{array}{ccc} 2 \\ 2 \end{array} \right)^3 & ^{1-2} \end{array}$$
 (79)

with M (kj) = M (jk). Note that the three masses M (12) = M (13) = M  $(23) = (4=3)^{p}$  for = i3, a value of the parameter for which the vacualie at the vertices of an equilateral triangle, leading us back to the sym metric case which engenders the  $Z_3$  sym metry.

We now go to the deform ation procedure, changing ! f(); w ith f() = W() : In the present case, we get that <math>f() : W() : W()

$$f^{0}()\overline{f^{0}()} = Q = QV(f();\overline{f()})$$
(80)

and this now gives, in the specially simple case which we have already considered in the former section,

$$f^{0}()^{2} = f f^{2} + f^{3}$$
 (81)

This is again the W eierstrass equation, and if we do no z =  $4^{-\frac{1}{3}}$  and f =  $4^{\frac{1}{3}}$  + =3 we not

$${}^{0}(z)^{2} = 4 {}^{3} q_{2} q_{3}$$
 (82)

where  $g_2 = 4^{\frac{1}{3}}(1 + {}^2=3)$  and  $g_3 = (2 = 3)(1 {}^2=9)$ . The solution for f ( ) is then given by

$$f() = \frac{1}{3} + 4^{\frac{1}{3}} P(4^{\frac{1}{3}}; g_2(); g_3())$$
(83)

The half-periods  $!_1$  and  $!_3$  of the W eierstrass P function are obtained from the invariants  $g_2$  and  $g_3$  by means of the equation

$$g_2 = 60$$
  ${}^{X}_{mn}$ ;  $g_3 = 140$   ${}^{K}_{mn}$  (84)

where  $m_n = 2m!_3 + 2n!_1$ ; with m; n 2 Z. We also have  $= g_2^3 - 27g_3^2 = 4(^2 - 1)^2$  and  $e_1 = 4^{\frac{1}{3}}(1 + =3)$ ;  $e_2 = 4^{\frac{1}{3}}(1 = 3)$ ; and  $e_3 = 4^{\frac{1}{3}}(2 = 3)$  are respectively the discriminant and roots of the cubic equation which appears from the right hand side of (82).

It is interesting to note that the modular parameter ( ) =  $\frac{1}{3}$  ( )= $\frac{1}{1}$  ( ) of the genus 1 R iem ann surface associated to this P-function depends on . Thus, variations of correspond to motions in the R iem ann surface moduli space. The deformed model is governed by the potential

$$V(\ ) = \frac{1}{2} \mathcal{P}^{0}(4^{-\frac{1}{3}} ; g_{2}(\ )); g_{3}(\ )^{2}$$
(85)

It has zeros in the FPP at  $P^0(!_1) = P^0(!_3) = P^0(!_1 + !_3) = 0$ . Thus, the vacua of the deformed model are the constant eld con gurations

$${}^{(1)}_{(m,n)} = 4^{\frac{1}{3}}(!_1 + m_n); \qquad {}^{(2)}_{(m,n)} = 4^{\frac{1}{3}}(!_1 + !_3 + m_n); \qquad {}^{(3)}_{(m,n)} = 4^{\frac{1}{3}}(!_3 + m_n)$$
(86)

The values of the superpotential

$$W () = \frac{1}{3} + 4^{\frac{1}{3}} P (4^{-\frac{1}{3}}; g_2(); g_3()) e^{-\frac{1}{3}}$$
(87)

at these vacua are

Therefore the angles of the deform ed kink orbits are (mod )

$$^{(12)} = 0;$$
  $^{(13)} = \arctan \frac{\text{Im}}{\text{Re}(+1)};$   $^{(23)} = \arctan \frac{\text{Im}}{\text{Re}(-1)}$  (89)

and the kink masses become

M (12) = 2; M (13) = 
$$\binom{1}{1} + \binom{1}{2} + \binom{2}{2} = \binom{1}{2}$$
; M (23) =  $\binom{1}{1} + \binom{1}{2} + \binom{2}{2} = \binom{1}{2}$  (90)

Note that for =  $\frac{p}{i}$   $\frac{1}{3}$  the three masses are equal, M (12) = M (13) = M (23) = 2, corresponding to a regular triangular lattice of m in in a. The sequences of m in in a connected by the kink orbits in these families are

In Fig. 7 and 8 we plot the potential, m in in a and network of kink orbits, respectively, for the speci c value of the complex parameter = 1 + i. Comparison of these gures with gures 2 and 3 shows to what extent the complex parameter induces irregularity when it diers from = 1 + i.

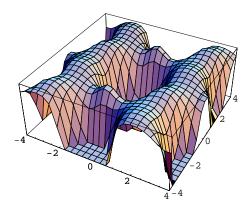


Figure 7: (Color online) The case of three asym metric minima. 3D plot of the deformed potential V(;) for = 1 + i near a point of the set  $\tilde{a}$ .

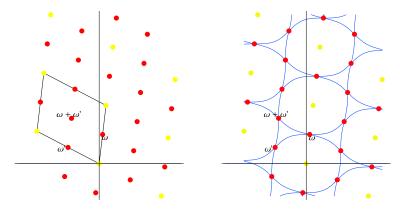


Figure 8: (Color online) The case of three asymmetric minima. Plots of the zeros (red) of the potential and points (yellow) of the set  $\tilde{a}$ , and the networks of kink orbits (blue) connecting the zeros for  $\tilde{a}$  = 1 + i (right panel).

## B. Irregular network of sn-kink orbits

W e consider now the superpotential

$$W () = +\frac{1}{2}(+)^{2} + \frac{1}{3}(-1)^{3} + \frac{1}{4}(+)^{4} + \frac{1}{5}^{5}$$
 (92)

where and are two complex coupling parameters which control the model. With this polynomial of fth-order Wwe get the potential

$$V(;) = \frac{1}{2}jl + j'j \qquad j'j \qquad j'j \qquad j'$$
 (93)

displaying the four m in im a  $v_1 = 1$ ;  $v_2 = v_3 = 1$ ; and  $v_4 = 0$ . We notice that two of the m in im a are again xed at the values 1, but this does not restrict generality of the procedure for the same reasons as before.

Again, we choose the deform ation f() = W() determined from the specially sim ple case

$$f^{0}()^{2} = (f+1)(f-1)(f-1)(f-1)$$
 (94)

This equation can be written as the elliptic sine equation  $y^0 = \frac{p}{(1-y^2)(1-k^2y^2)}$  if we follow the usual procedure [26], in which we write the product  $(f^2-1)(f-1)(f-1)$  as the product of the two factors  $A_+(f-1)^2 + A_-(f-1)^2 + A_-(f-1)^2$ 

$$A = \frac{1}{2} \quad 1 \quad P \frac{1+}{(1-2)(1-2)} \quad ; \quad B = \frac{1}{2} \quad 1 \quad \frac{1}{2} P \frac{2^{-2} \quad 2}{(1-2)(1-2)} \quad ; \quad = \frac{1+}{(1-2)(1-2)} P \frac{(1-2)(1-2)}{(1-2)(1-2)}$$
(95)

We then make the homographic substitution p = (f) = (f) to obtain

$$\frac{dy}{d} = {}^{p} \overline{(1 + y^2)(1 + k^2 y^2)} \quad ) \quad y() = sn(;k^2)$$
 (96)

$$f() = \frac{p \overline{A = A_{+}} \operatorname{sn}(;k^{2})}{1 \overline{A = A_{+}} \operatorname{sn}(;k^{2})}$$

$$(97)$$

is the new superpotential from which we can construct the potential of the modi ed model in the form

$$V(;) = \frac{1}{2} \frac{(+)^{2^{p}} \overline{A B} cn(;k^{2}) dn(;k^{2})}{(P \overline{A} = A_{+} sn(;k^{2}) 1)^{2}}$$
(98)

The above investigation is very general, and the presence of the two complex parameters and make the illustrations awkward. For this reason, we shall restrict ourselves to the simpler case where = and only a single complex parameter is free. Thus,

$$W () = {}^{2} \frac{1}{3}(1 + {}^{2})^{3} + \frac{1}{5}^{5}; \quad V (;) = \frac{1}{2}J + {}^{2}J +$$

are respectively the superpotential and potential of the original model. The four minima of V are at the points  $v_1 = v_3 = 1$ ,  $v_2 = v_4 = v_4 = v_5$  in the -com plex plane.

From the values of the superpotential at the minima

$$W (1) = \frac{2}{3} (^{2} \frac{1}{5}) e^{\frac{1}{5}}, W () = \frac{2}{3} (1 \frac{2}{5}) e^{\frac{1}{5}}$$
 (100)

we obtain

W (1) W (1) = 
$$\frac{4}{3}$$
(2  $\frac{1}{5}$ )e i; W (1) W ( ) =  $\frac{2}{3}$   $\frac{5}{5}$  2( 1) e i (101a)

W () W () = 
$$\frac{4}{3}$$
 3 (1  $\frac{2}{5}$ ) e <sup>i</sup>; W (1) W () =  $\frac{2}{3}$   $\frac{5+1}{5}$  <sup>2</sup> (+1) e <sup>i</sup> (101b)

The angles of the kink orbits are (mod )

$$\frac{\text{Im } (1 \ 5^{2})}{\text{Re}(1 \ 5^{2})} ; \qquad \frac{\text{Im } (2 \ 1 \ 5^{2})}{\text{Re}(2 \ 5^{2})} ; \qquad \frac{\text{Im } (2 \ 1 \ 5^{2})}{\text{Re}(2 \ 1 \ 5^{2})}$$
 (102a)

$$\frac{\text{(24)}}{\text{Re(}^{3}(^{2} 5))} = \arctan \frac{\text{Im (}^{3}(^{2} 5))}{\text{Re(}^{3}(^{2} 5))}; \qquad \text{(23)} = \frac{\text{(14)}}{\text{arctan}} = \arctan \frac{\text{Im (}^{2} + 1 5^{2}( + 1))}{\text{Re(}^{2} + 1 5^{2}( + 1))} \qquad (102b)$$

The corresponding energies are given by

$$M(13) = \frac{4}{3}jl=5$$
  $^2j;$   $M(12) = M(34) = \frac{2}{3}jl=5+(1)^2$   $^5=5j$  (103a)

$$M(24) = \frac{4}{3}j(^2=5)^3j; \quad M(23) = M(41) = \frac{2}{3}j!=5 + (+1)^2 + ^5=5j$$
 (103b)

plus M (kj) = M (jk). Note that for = i; M (13) = M (24) =  $\frac{8}{5}$ , M (12) = M (34) = M (23) = M (14) =  $\frac{4}{5}$  $\frac{p}{2}$  and we recover the case of a regular square.

W e now deform the model, choosing f() = W(). As before, we consider the specially simple case, which gives

$$f^{0}()^{2} = (1+f)(1-f)(-f)(-f) = (1+f)(-f)^{2} + f()^{4}$$
 (104)

We compare this with the elliptic Jacobi sine equation to  $\inf() = \sin(; ^2)$  as the solution of equation (104). Thus, the deformed superpotential and potential are given by

$$W () = sn(;^{2})e^{i}; V(;) = \frac{1}{2}j \hat{j} jcn(;^{2})f jdn(;^{2})f$$
 (105)

The zeros of the potential form the lattice of vacua at the constant values of the eld

$$\binom{(3)}{(m,m)} = \frac{1}{4}(!_1 + !_2) + m_n; \qquad \binom{(4)}{(m,m)} = \frac{1}{4}!_1 + m_n$$
 (106b)

where the periodicity is determined by the quarter periods of the Jacobi elliptic sine

$$m_n = n!_1 + \frac{1}{2}m!_2; \quad !_1 = 4K (^2); \quad !_2 = 4iK (1 ^2)$$
 (107)

From the values of the superpotential at the minima

$$W (_{(m,n)}^{(3)}) = e^{i} = W (_{(m,n)}^{(1)}); W (_{(m,n)}^{(4)}) = e^{i} = W (_{(m,n)}^{(2)})$$

$$(108)$$

we derive

$$W ( ( _{(m ; n)}^{(3)} ) W ( ( _{(m ; n)}^{(1)} ) = 2e^{i} ; W ( _{(m ; n)}^{(4)} ) W ( _{(m ; n)}^{(2)} ) = 2 e^{i}$$
 (109a)

$$\mathbb{W} \quad \left( \begin{array}{c} (2) \\ (m \, \mu) \end{array} \right) \quad \mathbb{W} \quad \left( \begin{array}{c} (1) \\ (m \, \mu \, \mu) \end{array} \right) = (1 \qquad ) e^{\frac{i}{n}} = \mathbb{W} \quad \left( \begin{array}{c} (3) \\ (m \, \mu) \end{array} \right) \quad \mathbb{W} \quad \left( \begin{array}{c} (4) \\ (m \, \mu \, \mu) \end{array} \right)$$
 (109b)

$$W \ ( \ _{(m \ ; n)}^{(3)} ) \ W \ ( \ _{(m \ ; n)}^{(2)} ) \ = \ (1 + \ ) e^{i} \ = \ W \ ( \ _{(m \ ; n)}^{(4)} ) \ W \ ( \ _{(m \ ; n)}^{(n)} ) \ (109c)$$

The orbit angles and the energies of the deform ed kinks are given by

$$^{(13)} = 0$$
; M  $_{(13)} = 2$ ;  $^{(12)} = ^{(34)} = \arctan \frac{\text{Im} (1)}{\text{Re}(1)}$ ; M  $_{(12)} = \text{M} (34) = \text{jl}$  j  
 $^{(24)} = \arctan \frac{\text{Im}}{\text{Re}}$ ; M  $_{(24)} = 2\text{j}$  j;  $^{(23)} = ^{(14)} = \arctan \frac{\text{Im} (1+)}{\text{Re}(1+)}$ ; M  $_{(41)} = \text{M} (23) = \text{jl} + \text{j}$ 

with (kj) = (jk) + and M(kj) = M(jk): The vacua connected by the kink orbits are organized according to the following sequences

To illustrate the investigations, in Fig. 9 and 10 we plot the potential, m in im a, points in the set e, and orbits of the kinklike con gurations.

#### V. BIFURCATION

All the models which we have been studying so far present an interesting feature, which we now explore. It concerns the fact that they have three or four minima. Thus, if we choose two minima arbitrarily, it may be possible that they are connected with two or more distinct orbits. When this happens to be the case, we say that the system develops a bifurcation, since one can go from a given vacua to another one, following two or more distinct kink orbits. This possibility is directly related to the balance of kink energies providing an upper bound for the fusion of two of the kinks in a single kink of a third type. Such a process is energetically possible { and the outgoing kink stable { if, given e.g. three minimak; j; l, the energy of the (kj) kink is lower than the sum of the other two kink masses

$$M(kj) < M(kl) + M(lj)$$
 (112)

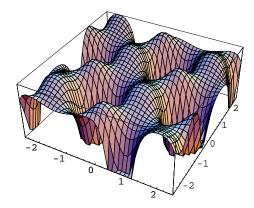


Figure 9: (Color online) The case of four asym metric minima. 3D plot of the deformed potential V(; ) near four points of the set  $\tilde{a}_4$  for = 1 + i.

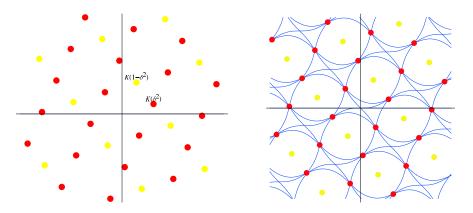


Figure 10: (Color online) The case of four asymmetric minima. Plots of the zeros (red) of the potential and points (yellow) of the set  $\tilde{a}$ , and networks of kink orbits (blue) connecting the zeros for  $\tilde{a}$  = 1 + i (right panel).

A llthe kink m asses depend on the N 2 com plex parameters, indicating the arbitrary positions of N 2 vacua, since the other two m inim a are xed at the 1 points in the -com plex plane. A bifurcation occurs when the inequality becomes equality, since we can go from k to j following the direct  $k \,!\, j$  path, or then visiting 1 through the path  $k \,!\, 1 \,!\, j$  with the same energetic cost. In the case N = 3 it is shown in [7], in the search for intersecting domain walls, that this possibility indeed happens  $\{$  see also Ref. [25]. In the generic case N; there is a sub-manifold of real dimension 2N - 5 of the parameter space characterized by the equation M(kj) = M(kl) + M(lj). We shall refer to this sub-manifold as the marginal stability variety because some irreducible component of it is a boundary between two regions of the (N - 2)-dimensional complex parameter space; one region where M(kj) < M(kl) + M(lj) (the (kj) kink is unstable and cannot decay to (kl) and (lj) kinks), and the other region where M(kj) > M(kl) + M(lj) (the (kj) kink is unstable decaying to the (kl)-(lj) kink combination). In fact, things are slightly more complicated and a bifurcation occurs at the marginal stability variety, with the kink orbit going to in nity beyond this point, with M(kj) become ing divergent, breaking the direct connectivity between the vacua k and j:

Since the energy in the topological sector is controlled by W; we immediately see that bifurcation appears if and only if at least three m in imagets aligned in W space, that is, i W (k) W (j)j= W (k) W (l)j+ W (l) W (j)j, with W (l) in between W (k) and W (j). This is the general condition for bifurcation, and below we use it to investigate the models introduced in the former section.

## A. The case of three m in im a

Here we consider the case of three m inima, with the model being described by the complex parameter: We notice that bifurcation cannot appear in models which engender the  $Z_3$  symmetry, since the symmetry implies that M(12) = M(23) = M(31), henceforth M(12) < M(23) + M(31), etc. In the Abraham-Townsend model, the marginal stability curve is characterized by the alignment of the three minima in the  $(W; \overline{W})$  plane, i.e., it is the curve in this

plane for which (12)() = (23)() = (13)() m od , or,

$$\frac{\text{Im}((3)(+1^{\frac{3}{2}}))}{\text{Re}((3)(+1^{\frac{3}{2}}))} = \frac{\text{Im}}{\text{Re}} = \frac{\text{Im}((+3)(-1^{\frac{3}{2}}))}{\text{Re}((+3)(-1^{\frac{3}{2}}))}$$
(113)

These identities hold if

$$( )(3 + (^2 + ^2 + 6)) = 0 (114)$$

which is the algebraic equation that characterizes the marginal stability curve. We note that the curve has two irreducible components.

One component is the real axis in the -plane:  $_2$  = 0: For real values of the parameter it happens that all the three minima are aligned in the real axis of the complex -plane, and the system cannot develop bifurcation. In the two very special cases with  $_1$  = 1; we have only two minima, and so only one kink orbit remains in the system. For  $_1$  < 1; there exist only two kink orbits, which we label (31) and 12, because the vacuum  $v_1$  sits on the real axis in between  $v_3$  and  $v_2$ . For 1 < 1; there exist the two kink orbits (13) and (32), and for  $_1$  > 1; there exist the two kink orbits (12) and (23) for similar reasons.

Things are more interesting for the other algebraically irreducible component. The quartic curve

which allow for each one of the three possibilities, M (12) = M (23) + M (31), or M (23) = M (31) + M (12), or yet M (31) = M (12) + M (23), depending on the speci c value of the complex parameter , is the true boundary between regions in the plane where there exist three or two kinks and the phenomenon of bifurcation takes place. This case is fully studied in [7], and below in Fig. 11 we plot the curves of marginal stability in the complex plane.

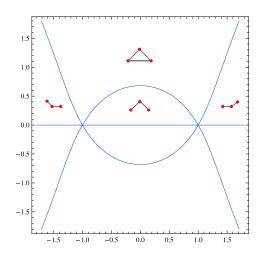


Figure 11: (Color online) The curves of marginal stability in the case of three minima. The set of (red) points illustrates the positions of the minima and the kink orbits for distinct values of the parameter in the related regions.

The condition for alignment in the deformed model in the W plane  $^{(12)} = ^{(13)} = ^{(23)}$  is

$$0 = \frac{\operatorname{Im}}{\operatorname{Re}(+1)} = \frac{\operatorname{Im}}{\operatorname{Re}(-1)} \tag{116}$$

There is only one component,  $_2=0$ ; but we have already seen that in this case there is no bifurcation anymore. In fact on the real axis the kink masses are M (12)=2, M  $(13)=j_1+1j$ , and M  $(13)=j_1-1j$ , such that: M (13)=M (12)+M (23) if  $_1>1$ , M (12)=M (13)+M (23) if  $_1<1$ , and M (23)=M (12)+M (13) if  $_1<1$ . If  $_2\not\in 0$  the mass of one of the deformed kinks is always lighter than the sum of the masses of the other two kinks because of the triangle inequality, the sum of the lengths of two sides of a triangle is greater than the length of the third side. Thus, we notice that the deformation procedure wash out the marginal stability curve, leaving no room for bifurcation in the deformed models which we are dealing with in this work. In Fig. 12 we illustrate the N = 3 case with several distinct possibilities for the complex parameter .

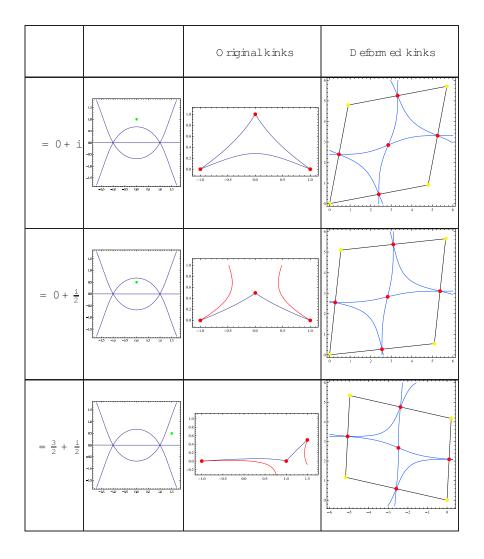


Figure 12: (Color online) The bifurcation curves and some related illustrations with (green) points in the rst column representing distinct values of the complex parameter in the case of three minima. The broken (red) lines in the second column indicate how the orbit goes to in nity, leading to divergent contributions which should be discarded.

## B. The case of four m in im a

Let us now consider the case of four m inim a, with the model being driven by the two complex parameters and . The general case is very complicate, so we move to the simpler case in which we use =; leading to a single complex parameter .

The condition for alignment of all the minima in the W plane now is  $^{(12)} = ^{(13)} = ^{(23)} = ^{(24)}$ ; mod (note that because the vacua in the pairs  $(v_1; v_3) = (1; 1)$  and  $(v_2; v_4) = ($ ; ) are always aligned in the W -plane, there cannot be alignments of three non vanishing vacua). The identities between kink angles hold if

$$\frac{\text{Im} \left[1 \quad 5^{2} + \ ^{3}\left(5 \quad ^{2}\right)\right]}{\text{Re}\left[1 \quad 5^{2} + \ ^{3}\left(5 \quad ^{2}\right)\right]} = \frac{\text{Im} \left[\ ^{3}\left(5 \quad \ ^{2}\right)\right]}{\text{Re}\left[\ ^{3}\left(5 \quad \ ^{2}\right)\right]} = \frac{\text{Im} \left[1 \quad 5^{2}\right]}{\text{Re}\left[1 \quad 5^{2}\right]} = \frac{\text{Im} \left[1 \quad 5^{2} \quad ^{3}\left(5 \quad \ ^{2}\right)\right]}{\text{Re}\left[1 \quad 5^{2} \quad ^{3}\left(5 \quad \ ^{2}\right)\right]} \tag{117}$$

which is satis ed if

( ) 
$$5(^2 + ^2)$$
  $^4$   $^4 + 5(^2 + ^2 + )$   $26 = 0$  (118)

One irreducible component is again the abscissa axis  $_2 = 0$ : The system does not support bifurcation since all the minima are now in the real axis of the complex -plane. There are three special values:  $_1 = 1$  and  $_1 = 0$ ; for  $_1 = 1$ ; there exist only two minima and one kink orbit (12) and for  $_1 = 0$  there exist three minima, with the two

distinct kink orbits (12) and (23):0 ther possible values are: for  $_1<1$ ; there exist four m inim a, but only the three kink orbits (21), (13), and (34); for  $_1<0$ ; there exist four m inim a, but only the three kink orbits (12), (24), and (41); for  $_1<1$ ; there exist four m inim a, but only the three kink orbits (14), (42), and (23); for  $_1>1$ ; there exist four m inim a, but only the three kink orbits (41), (13), and (32).

The other algebraically irreducible component is given by

$$15_{1}^{2} \quad 5_{2}^{2} \quad 30_{1}^{4} \quad 26_{2}^{4} \quad 40_{12}^{22} + 15_{1}^{6} \quad 5_{2}^{6} + 25_{12}^{42} + 5_{12}^{24} = 0$$
 (119)

In Fig. 13 we plot the marginal stability curves which follow from the above expression, and we illustrate how the minima behave in the complex plane. The two minima 1 and any other pair of asymmetrically positioned minima on the curves introduces two distinct kink orbit possibilities, and so it illustrates the bifurcation phenomenon once again.

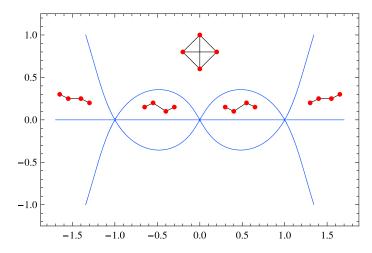


Figure 13: (Color online) The curves of marginal stability in the case of four minima, for = . The set of (red) points illustrates the positions of the minima and the kink orbits for distinct values of the parameter in the related regions.

It is interesting to notice that the condition for alignment in the deformed model in the W plane, leads to  $_2=0$  again. But we have already seen that in this case there is no bifurcation anymore. Thus, once again we notice that the deformation procedure washout the marginal stability curve, leaving no room for bifurcation also in this case. In Fig. 14 we illustrate the case N = 4 with several distinct possibilities, driven by the complex parameter:

#### VI. FINAL COMMENTS

In this work we have rstly dealt with the standard W ess-Zum ino model driven by a complex scalar eldengendering discrete  $Z_N$  symmetry. The model is defined in terms of a superpotential, a holomorphic function of the complex eld which contains N m in in a, the vacua manifold which represents a set of points with the very same  $Z_N$  symmetry of the model. The min in a determine several topological sectors, which can be represented by algebraic curves and solved by rst-order differential equations of the BPS type, as shown in Ref. [9].

The main idea of this work concerns the deform ation procedure developed in [19, 20], which was used to deform the model in a way such that the set of N m inima could be replicated in the entire con guration space, the plane described by the complex eld. In this way, the algebraic orbits in eld (target) space of the original Wess-Zum ino model are also replicated in the entire plane, naturally leading to a network of defects, a spread network of kinklike orbits in eld space. As we have seen, the idea was in plan ented very e ciently, and we have illustrated the procedure with three interesting cases, involving the  $Z_2$ ,  $Z_3$  and  $Z_4$  sym metries, with the sets of minima forming an equilateral triangle and a square in the last two cases, respectively.

We have then investigated the more general case, with models engendering N non symmetric minima, with N = 3 and N = 4. The case N = 3 is driven by a single complex parameter; and we have illustrated the results with two distinct values of; showing how it models the regular structure that we have obtained in the symmetric case. The case N = 4 is more complicated, since it is controlled by two complex parameters, and . We have then made an interesting simplication, reducing the model to a single complex parameter, with = . In this case, we have also illustrated the results with two distinct values of the parameter, to show how it changes the regular structure of the symmetric case. In the non symmetric case, we have also examined bifurcation and the marginal stability curve for

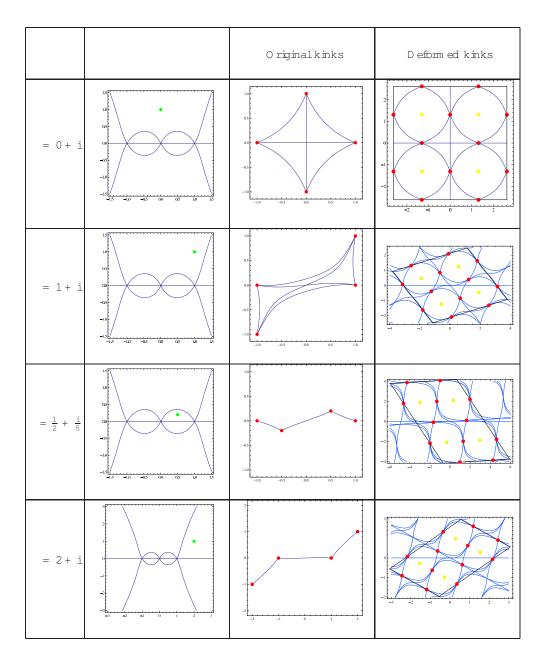


Figure 14: (Color online) The bifurcation curves and some related illustrations with (green) points in the rst column representing distinct values of the complex parameter in the case of four minima.

both n = 3; and n = 4; in the last case after introducing the sim pli cation which leads to models described by a single complex parameter. In both cases, the complex parameter gives rise to a diversity of very nice possibilities, but the deformation washes out bifurcation from the deformed models.

The present investigation poses several issues, one of them concerning the N=5 and N=6 cases, which follows the natural course of this work. In the asymmetric case of four minima, we could also consider other relations between the two parameters and . A nother issue concerns the tiling of the plane with regular polygons, which is constrained to appear with regular hexagons, squares and equilateral triangles, in this order of decreasing e ciency. It would be interesting to search for possible connection among these regular tilings, the respective basic polygons, and the deformation procedure. It is also interesting to consider the issue studied in Ref. [25], in which one deals with the problem of marginal stability in terms of forces between soliton states in the case of three minima. A natural extension to the case of four minima seems desirable, and could also include investigations on how the forces between solitons would behave under the deformation procedure used in the present work.

A nother route of interest is related to a topological change in the eld plane itself: if we let the elds to live in a

genus 0 or genus 1 R iem ann surface, that is, if we consider the target space to be the two-sphere  $S^2$  or the two-torus  $T^2 = S^1$   $S^1$ ; instead of the plane  $R^2$ ; we could be able to tile the surface  $S^2$  or  $T^2$  with other patterns, nesting distinct networks of defects. This last case m ay possibly lead us to an issue of current interest: the conict between geometry and topology, related to the geometric features of the tiling with regular polygons and the topological properties of the conguration space itself. It seems plausible that the modular parameter of the target space  $T^2$  is forced to be our  $S^2 = 1$  and the sphere  $S^2 = 1$  is restricted to a 2:1 embedding of this  $T^2$  (!) in  $T^2 = 1$  is forced to be now under consideration, and we hope to report on them in the near future.

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- [1] T. H. R. Skyrme, Nucl. Phys. 31,556 (1962); D. Finkelstein, J. M. ath. Phys. 7,1218 (1966); J. Rubinstein, J. M. ath. Phys. 11,258 (1970); R. F. Dashen, B. Hasslacher, and A. Neveu, Phys. Rev. D 10,4130 (1974).
- [2] H. B. Nielsen and P.O. lesen, Nucl. Phys. B 61, 45 (1973); H. J. de Vega and F. A. Schaposnik, Phys. Rev. D 14, 1100 (1976).
- [3] G. 't Hooft, Nucl. Phys. B 79, 276 (1974); A M. Polyakov, JETP Lett. 20, 194 (1974).
- [4] M. K. Prasad and C. M. Sommer eld, Phys. Rev. Lett. 35, 760 (1975); E.B. Bogom ol'nyi, Sov. J. Nucl. Phys. 24, 449 (1976).
- [5] R. Rajaram an, Solitons and Instantons (North-Holland, Amsterdam, 1982); S. Coleman, Aspects of Symmetry (Cambridge, Cambridge, UK, 1985); A. Vilenkin and E.P.S. Shellard, Cosmic Strings and Other Topological Defects (Cambridge, Cambridge, UK, 1994); N. Manton and P. Sutclie, Topological Solitons (Cambridge, Cambridge, UK, 2004).
- [6] P. Fendley, S.D. Mathur, C. Vafa, and N.P. Warmer, Phys. Lett. B 243, 257 (1990); S. Cecotti, and C. Vafa, Comm. Math. Phys. 158, 569 (1993).
- [7] E.R.C. Abraham and P.K. Townsend, Nucl. Phys. B 351, 313 (1991).
- [8] T. Hollowood, Nucl. Phys. B 384, 523 (1992).
- [9] A. Albonso Izquierdo, M. A. Gonzalez Leon, and J. M. ateos Guilarte, Phys. Lett. B. 480, 373 (2000); D. Bazeia, J. M. enezes, and M. M. Santos, Phys. Lett. B. 521, 418 (2001); Nucl. Phys. B. 636, 132 (2002).
- [10] G W .G ibbons and P K .Townsend, Phys.Rev.Lett.83,1727 (1999); P M .Sa n, Phys.Rev.Lett.83,4249 (1999); H .
  Oda, K .Ito, M .Naganum a, and N .Sakai, Phys.Lett.B 471,140 (1999); D .Bazeia and F A .Brito, Phys.Rev.Lett.84,
  1094 (2000); S M .Carroll, S.Hellerm an, and M .Trodden, Phys.Rev.D 61,065001 (2000).
- [11] D. Bazeia and F.A. Brito, Phys. Rev. D 61,105019 (2000); 62,101701(R) (2000); R. Hofmann, Phys. Rev. D 62,065012 (2000); S.K. Nam, JHEP 03,005 (2000); S.K. Man and K. Olsen, JHEP 08,001 (2000); D. Binosi and T. ter Veldhuis, Phys. Lett. B 476,124 (2000); D. Binosi, M. Shifman, and T. ter Veldhuis, Phys. Rev. D 63,025006 (2000); F.A. Brito and D. Bazeia, Phys. Rev. D 64,065022 (2001); M. Naganuma, M. Nitta, and N. Sakai, Phys. Rev. D 65,045016 (2002); D. Tong, Phys. Rev. D 66,025013 (2002); JHEP 04,031 (2002); P. Sutcli e, Phys. Rev. D 68,085004 (2003).
- [12] M. Bowick, A. De Felice, and M. Trodden, JHEP 0310, 067 (2003); L. Pogosian and T. Vachaspati, Phys. Rev. D 67, 065012 (2003); N. D. Antunes, E. J. Copeland, M. H. indm. arsh, and A. Lukas, Phys. Rev. D 69, 065016 (2004); N. D. Antunes and T. Vachaspati, Phys. Rev. D 70, 063516 (2004); A. de Souza Dutra, Phys. Lett. B 626, 249 (2005); A. A. lonso Izquierdo, M. A. Gonzalez Leon, M. de la Torre M. aiado, and J. M. ateous Guilarte, Physica D 200, 220 (2005); M. Eto, Y. Isozumi, M. N. itta, K. Ohashi, and N. Sakai, Phys. Rev. D 72, 085004 (2005); D. Bazeia, L. Losano, and R. M. enezes, Physica D 208, 236 (2005); S. Franco, A. Hanany, K. D. K. ennaway, D. Vegh, and B. W. echt, JHEP 0601, 096 (2006); A. A. lonso Izquierdo and J. M. ateous Guilarte, Physica D 220, 31 (2006); A. Prikas, J. M. ath. Phys. 47, 112503 (2006); D. Bazeia, F. A. Brito, and L. Losano, Europhys. Lett. 76, 374 (2006); M. Eto, Y. Isozumi, M. N. itta, K. Ohashi, and N. Sakai, J. Phys. A. 39, R. 315 (2007); M. Eto et. al., Phys. Rev. D. 75, 045010 (2007).
- [13] H. Kubotani, Prog. Theor. Phys. 87,387 (1992); M. Bucher and D. N. Spergel, Phys. Rev. D. 60,043505 (1999); L. Conversi, A. M. elchiorri, L. M. ersini-Houghton, and J. Silk, Astropart. Phys. 21,443 (2004); P. Avelino, C. M. artins, J. M. enezes, R. M. enezes, and J. O. liveira, Phys. Rev. D. 73,123519 (2006); D. 73,123520 (2006); Phys. Lett. B. 647,63 (2007); R. A. Battye, E. Chachoua, and A. M. oss, Phys. Rev. D. 73,123528 (2006); R. A. Battye and A. M. oss, Phys. Rev. D. 74,023528 (2006).
- [14] N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, Phys. Lett. B 429, 263 (1998); I. Antoniadis, N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, Phys. Lett. B 436, 257 (1998); L. Randalland R. Sundrum, Phys. Rev. Lett. 83, 3370 (1999).
- [15] L.Randalland R.Sundrum, Phys. Rev. Lett. 83, 4690 (1999).
- [16] W D.Goldberger and M B.W ise, Phys.Rev.Lett.83, 4922 (1999); O.DeW olfe, D Z.Freedman, S.S.Gubser, and A.
  Karch, Phys.Rev.D 62,046008 (2000); C.Csaki, J.Erlich, T.J.Hollowood, and Y.Shirman, Nucl. Phys.B 581,309
  (2000); C.Csaki, J.Erlich, C.Grogean, and T.J.Hollowood, Nucl. Phys.B 584,359 (2000); M.Gremm, Phys.Lett.B
  478,434 (2000).

- [17] A. Karch and L. Randall, JHEP 0105,008 (2001); M. Porrati, Phys. Lett. B 498,92 (2001); F. A. Brito, M. Cvetic, and S.-C. Yoon, Phys. Rev. D 64,064021 (2001); M. Cvetic and N. D. Lambert, Phys. Lett. B 540,301 (2002); A. Campos, Phys. Rev. Lett. 88,141602 (2002); A. M. elfo, N. Pantoja, and A. Skirzewski, Phys. Rev. D 67,105003 (2003); D. Bazeia, F. A. Brito, and J. R. Nascimento, Phys. Rev. D 68,085007 (2003); D. Bazeia, C. Furtado, and A. R. Gomes, JCAP 0402, 002 (2004); D. Z. Freedman, C. Nunez, M. Schnabl, and K. Skenderis, Phys. Rev. D 69,104027 (2004); D. Bazeia and A. R. Gomes, JHEP 0405,012 (2004); O. Castillo-Felisola, A. M. elfo, N. Pantoja, and A. R. amirez, Phys. Rev. D 70,104029 (2004); K. Takahashi and T. Shiromizu, Phys. Rev. D 70,103507 (2004); D. Bazeia, F. A. Brito, and A. R. Gomes, JHEP 0411,070 (2004); A. Celietal. Phys. Rev. D 71,045009 (2005); E. Anderson and R. Tavakol, JCAP 0510,017 (2005); R. Guerrero, R. Omar Rodrigues, and R. S. Torrealba, Phys. Rev. D 72,124012 (2005); D. Bazeia, F. A. Brito, and L. Losano, JHEP 0611,064 (2006); V. I. Afonso, D. Bazeia, and L. Losano, Phys. Lett. B 634,526 (2006); A. Celi, JHEP 0702,078 (2007); K. Farakos and P. Pasipoularides, Phys. Rev. D 75,024018 (2007); A. Ceresole and G. Dall'Agata, JHEP 0703, 110 (2007); M. Giovannini, Phys. Rev. D 75,064023 (2007).
- [18] V. J. A. fonso, D. Bazeia, M. A. G. onzalez Leon, L. Losano, and J. M. ateos G. uilarte, Phys. Lett. B. 662, 75 (2008).
- [19] D. Bazeia, L. Losano, and JM C. Malbouisson, Phys. Rev. D 66, 101701(R) (2002).
- [20] C.A.A. Imeida, D.Bazeia, L.Losano, and J.M. C.M. albouisson, Phys. Rev. D. 69,067702 (2004); D.Bazeia and L.Losano, Phys. Rev. D. 73,025016 (2006); V. J.A. fonso, D.Bazeia, and F.A. Brito, JHEP 0608,073 (2006); D.Bazeia, M.A. Gonzalez Leon, L.Losano, and J.M. ateos Guilarte, Phys. Rev. D. 73,105008 (2006); V.J.A. fonso, D.Bazeia, M.A. Gonzalez Leon, L.Losano, and J.M. ateos Guilarte, Phys. Rev. D. 76,025010 (2007).
- [21] A. de Sousa Dutra, Deform ed solitons: the case of two coupled scalar elds [arX iv:0705.3237].
- [22] A. A lonso Izquierdo, M. A. Gonzalez Leon, J. M. ateos Guilarte, J. Phys. A. 31, 209 (1998); Nonlinearity 13, 1137 (2000); S.R. W. oodford and I.V. Barashenkov, J. Phys. A. 41, 185203 (2008).
- [23] A. Perelom ov, Integrable systems of classical mechanics and Lie algebras (Birkhauser, Basel, 1992); M. Olshanetsky, J. Nonlin. Math. Phys. 12, 5522 (2005).
- [24] M. Abram ow itz and I. Stegun, Handbook of Mathematical Functions, (Dover, New York, 1970).
- [25] R. Portugues and P.K. Townsend, Phys. Lett. B 530, 227 (2002).
- [26] E.T.W hittaker and G.N.W atson, A. Course of Modern Analysis (Cambridge, Cambridge, UK, 1973).